

1 **Long-term Growth and Subsidence of Ascension Island:** 2 **Constraints on the Rheology of Young Oceanic Lithosphere**

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12 The dating of material from deep boreholes drilled in volcanic ocean islands allows
13 constraints to be placed on their growth and long-term subsidence rates. We dated lavas from
14 a 3 km geothermal borehole at Ascension Island by the laser-heating ⁴⁰Ar/³⁹Ar technique. The
15 samples yield ages of up to 3.4 Ma and volcanic growth rates of ~0.4 km/Myr. The transition
16 from submarine to subaerial eruption occurs at ~710 m below present sea level and 2.5 Ma.
17 Since 2.5 Ma, there has been ~520 m of subsidence over and above the expected ~190 m due
18 to lithospheric cooling. Plausible elastic thicknesses and growth histories would generate a
19 maximum elastic subsidence since 2.5 Ma of ~200 m. We infer that the subsidence includes
20 a component of viscous relaxation resulting from rapid loading prior to 2.5 Ma, and place
21 constraints on the timescale of this relaxation, and hence the viscosity of the underlying
22 lithosphere.

23

23 1. Introduction

24 The growth of volcanic ocean islands places a time-varying load on the lithosphere, and the
25 response of the lithosphere to that load provides information about lithospheric rheology.
26 Over long timescales, to a good approximation, the lithosphere behaves like a thin elastic
27 plate and the response may be characterized by a single parameter, the effective elastic
28 thickness [e.g., *Calmant et al.*, 1990; *Watts and Cochran*, 1974]. However, on shorter
29 timescales there is a viscous component to the response [e.g., *Walcott*, 1970]. This
30 component may be parameterized by considering the response of a thin visco-elastic sheet on
31 an inviscid substrate [e.g., *Lambeck and Nakiboglu*, 1981], or alternatively by a model in
32 which viscosity varies more smoothly with temperature and hence depth [e.g., *Courtney and*
33 *Beaumont*, 1983; *Watts and Zhong*, 2000]. If the timing of load emplacement is sufficiently
34 well known, subsidence rates of volcanic islands can provide constraints on this viscous
35 behavior [e.g., *Watts and Zhong*, 2000]. Although subsidence rates may be measured on
36 geological timescales through coral growth rates [e.g., *Moore et al.*, 1996] or dating of
37 submerged terraces, volcanic loading histories are usually poorly known. However, the
38 dating of material from deep boreholes in ocean islands can allow constraints to be placed not
39 only on the growth rate but also, through measurement of the depth and age of the submarine-
40 subaerial transition, the long-term subsidence rate [*Hyndman et al.*, 1979; *Sharp and Renne*,
41 2005]. Here we use such measurements from Ascension Island in the central Atlantic to
42 investigate the response of young oceanic lithosphere to volcanic loading.

43

44 2. Isotopic Dating of Borehole Samples

45 Ascension Island forms the summit of a ~4-km-high volcanic edifice lying on 7 Ma oceanic
46 lithosphere, 90 km west of the mid-Atlantic Ridge (Fig. 1). K-Ar dating of surface samples
47 has yielded ages of up to 1.5 Ma, while sparse Ar-Ar geochronology has yielded ages of 0.4-
48 1.0 Ma. [*Harris et al.*, 1982; *Kar et al.*, 1998; *Nielson and Sibbett*, 1996]. A tomographic
49 seismic experiment showed that the lower oceanic crust is thickened beneath the island but no
50 evidence of a velocity discontinuity between pre-existing crust and a magmatic “underplate”,
51 leading to the suggestion that the bulk of the volcanic edifice may have formed close to the
52 ridge axis [*Evangelidis et al.*, 2004; *Klingelhoefer et al.*, 2001]. The Ascension #1(ASC-1)
53 borehole was drilled to a depth of 3.1 km as part of a geothermal exploration project, which
54 also drilled several shallower boreholes [*Nielson and Stiger*, 1996]. The borehole was not
55 cored, but cuttings were collected for each 3-m depth interval and archived. Most of the
56 cutting material consisted of heavily altered volcanic dust, but some larger and fresh

57 fragments were also available. Basaltic to doleritic clasts of well-crystallized glass-free fresh
58 groundmass 1-2 in cm size were dated using the $^{40}\text{Ar}/^{39}\text{Ar}$ laser-heating technique. At each
59 depth only clasts of most dominant rock type which clearly represents the lithofacies at the
60 sampling depth were considered for analysis. Age determinations were conducted using
61 $^{40}\text{Ar}/^{39}\text{Ar}$ geochronology facility at the Geological Survey of Japan/AIST [Ishizuka *et al.*,
62 2003; Ishizuka *et al.*, 2009]. Details of analytical procedure are described in the electronic
63 supplement.

64

65 Samples from four different depths of the ASC-1 gave plateau ages. Two fragments from a
66 single sample from 834-837 m depth (all depths are relative to the Kelly bushing at 181 m
67 above sealevel) were analyzed separately, and returned identical plateau ages (2.43 ± 0.07 and
68 2.42 ± 0.09 Ma). This result demonstrates the high reproducibility of the analysis on small
69 amount of cuttings material. Two different samples from 894-897 m depth returned well-
70 defined plateau ages of 2.23 ± 0.07 and 2.48 ± 0.10 Ma, respectively. These plateau ages are
71 also identical within 2σ error. In between these depths, the shallowest hyaloclastite section
72 (which was erupted in a submarine environment) was recovered between 887 and 939 m
73 depth [Nielson and Stiger, 1996], and 887 m marks the transition from subaerial to submarine
74 eruption, which is therefore dated accurately. A sample from 1050-1053 m depth gave a
75 similar plateau age of 2.5 ± 0.3 Ma. A clast from 1233-1236 m depth yielded a partly disturbed
76 age spectrum with a significantly older plateau age of 3.41 ± 0.25 Ma. Inverse isochrons for
77 each plateau obtained here gave Ar initial ratios identical to atmospheric ratio within 2σ
78 error, indicating no presence of extraneous Ar. No samples suitable for analysis were found
79 outside the depth range 834-1236 m.

80

81 **3. Modeling the Growth and Subsidence of Ascension Island**

82 The radiometric dates show that the volcanic edifice is indeed much older than indicated by
83 surface samples, that substantial subsidence has occurred since 2.5 Ma, and that the mean
84 volcanic growth rate during that period has been ~ 0.4 km/My (Fig. 3). First, we explored
85 whether this subsidence could be adequately explained by an elastic response. We calculated
86 the flexural isostatic response to three possible loading scenarios, using the load volume
87 defined by Evangelidis *et al.* [2004] and load and infill densities of 2323 kg/m^3 based on the
88 seismic velocity structure of the volcanic edifice [Evangelidis *et al.*, 2004]. The mantle
89 density was 3330 kg/m^3 . In each scenario, it is assumed that any space created by flexural
90 subsidence is filled by new igneous material, and vertical distances are rounded to the nearest

91 10 m. In scenario 1, all the currently subaerial load has been added since 2.5 Ma. In scenario
92 2, the growth rate of the volcanic edifice is in proportion to its current height, such that 180 m
93 of load has been added at the borehole location and no new material has been added around
94 the perimeter of the volcanic edifice. In scenario 3, a 180 m layer of load has been added at
95 all locations, except where the current edifice is less than 180 m high, where the whole of the
96 edifice has been added since 2.5 Ma.

97

98 Analysis of both gravity and seismic data indicate that the lithosphere has significant flexural
99 strength, with an effective elastic thickness T_e of 2-4 km inferred from gravity data [*Minshull*
100 *and Brozena, 1997*] and a value of at least 6 km inferred from the Moho shape derived from
101 wide-angle seismic data [*Evangelidis et al., 2004*]. Both estimates involve assumptions about
102 the processes controlling the shape of the Moho depression beneath the island, but the latter
103 estimate involves fewer assumptions and is therefore more reliable. For a T_e of 6 km, the
104 maximum subsidence predicted since 2.5 Ma is 110 m (Fig. 3). Even if we assume a T_e of as
105 low as 1 km, the elastic subsidence predicted since 2.5 Ma is 130 m, 220 m and 280 m for
106 scenarios 1-3, respectively. Ignoring eustatic sea-level changes, which cannot be accounted
107 for given the age uncertainties for the subaerial to submarine transition, the observed
108 subsidence is 710 m, There are 600-700 m of additional subsidence to explain for a 6 km
109 elastic thickness, or 430-580 m for 1 km elastic thickness. Some of this additional
110 subsidence might be explained by thermal subsidence since 2.5 Ma. However, if the
111 lithosphere has subsided according to plate cooling models [*Stein and Stein, 1992*], only ~190
112 m of subsidence is predicted between ages of 4.5 and 7.0 My. There is no significant residual
113 depth anomaly in the surrounding lithosphere [*Minshull et al., 1998*], so there is no evidence
114 for anomalous thermal subsidence. Some vertical motion may arise from flexural bending
115 associated with the adjacent Ascension Fracture Zone, but the island is ~60 km from this
116 fracture zone, so such motion is likely to be small. Anomalous subsidence might be attributed
117 to removal of dynamic support associated with an “Ascension plume”, but the timescale
118 required is very short for such a large-scale event. Hence we must seek an alternative
119 explanation for the large subsidence observed.

120

121 The anomalous subsidence might be explained if a viscous component leads to a delay
122 between volcanic loading and the isostatic response of the lithosphere. *Watts and Zhong*
123 [2000] infer from observations of flexural rigidity D as a function of both plate age and load
124 age that under volcanic loading, the oceanic lithosphere responds as a medium with viscosity

125 decaying gradually with depth as temperature increases. The simplest possible model
 126 capturing the isostatic, elastic and viscous response of the lithosphere to vertical loads is the
 127 thin viscoelastic plate model, which incorporates the widely used flexural response plus a
 128 time-dependent viscous relaxation due to a vertically-averaged lithosphere viscosity. We
 129 extended our analysis to include a viscous component by using the finite-difference code
 130 TISC [*Garcia-Castellanos, 2002*], which calculates the vertical deflection $w(x,y)$ of a
 131 viscoelastic thin plate when submitted to a load distribution $q(x,y)$. The deflection has two
 132 components: an instantaneous elastic response and a subsequent deflection velocity dw/dt
 133 related to viscous stress relaxation. In the absence of horizontal forces and lateral rigidity
 134 variations the elastic component can be calculated with the following equation [*van Wees and*
 135 *Cloetingh, 1994*]:

$$136 \quad D \frac{\partial^4 w}{\partial x^4} + D \frac{\partial^4 w}{\partial y^4} + 2D \frac{\partial^4 w}{\partial x^2 \partial y^2} + \Delta\rho g w = q(x, y)$$

137 where ν is Poisson's ratio (assumed to be 0.25) and $\Delta\rho$ is the density difference between
 138 sublithospheric mantle and the infill. Load and infill densities were as for the elastic
 139 calculation above. D is calculated from T_e assuming a Young's modulus of 7×10^{10} Pa [e.g.,
 140 *Panteleyev and Diament, 1993*]. To calculate the viscous component of the deflection, the
 141 same equation is solved but w is replaced by dw/dt , ν is 0.5, q is replaced by $(q - \Delta\rho \cdot g \cdot w)/\tau$,
 142 where τ is the viscous relaxation time, and the total deflection is computed by integrating
 143 over time.

144

145 The viscous flexural equation predicts that the deflection tends towards local isostatic
 146 equilibrium (similar to a reduction of T_e through time for an elastic plate) at timescales
 147 similar to τ , which relates to viscosity μ through $\tau = 2(1 + \nu)\mu/E$. We solved both equations
 148 (the elastic and the viscous one) for a range of values of T_e and τ , and for two different
 149 loading histories (Fig. 4). In both, we assume loading according to scenario 2, such that 12%
 150 of the load has been added since 3.4 Ma. In the first, we assumed the remainder of the
 151 loading occurred through rapid volcano growth at 5 Ma, and in the second we assumed that
 152 this loading occurred at 4 Ma. For loading at 5 Ma, the maximum predicted subsidence since
 153 2.5 Ma is less than 500 m (Fig. 4a), so insufficient to account for the 520 m difference
 154 between the observed subsidence and the predicted thermal subsidence. For loading at 4 Ma,
 155 the predicted subsidence since 2.5 Ma exceeds 500 m for T_e values of ~ 3 -7 km and τ values
 156 of ~ 0.5 -1.0 My (Fig. 4b), corresponding to a viscosity of ~ 0.4 - 0.7×10^{24} Pa s. For lower
 157 values of τ , the subsidence is largely complete before 2.5 Ma, while for higher values the

158 subsidence is too slow. For higher values of T_e , there is too little subsidence to match the
159 observations, while for lower values, although the overall subsidence is greater, it becomes
160 focused at earlier times and the subsidence since 2.5 Ma is too small. These results also
161 suggest that the main loading event cannot have occurred much before 5 Ma (because then
162 subsidence would be largely complete by 2.5 Ma), and the borehole data suggest that it
163 cannot have occurred after 3.4 Ma.

164

165 **4. Discussion and Conclusions**

166 The observed growth rate of Ascension Island since 3.4 Ma of 0.4 km/My is very similar to
167 the post-shield growth rate of 0.9 km/My inferred from $^{40}\text{Ar}/^{39}\text{Ar}$ dating of borehole samples
168 from the Hawaii Scientific Drilling Project, and an order of magnitude less than the 8.6
169 km/My growth rate of Mauna Kea during its shield-building phase [Sharp and Renne, 2005].
170 The mean subsidence rate of 0.3 km/My since 2.5 Ma is much smaller than the 2.6 km/My
171 value inferred at Hawaii [Sharp and Renne, 2005]. Both observations place Ascension Island
172 well into its post-shield phase, consistent with the evolved composition of most outcropping
173 volcanic rocks [e.g., Kar *et al.*, 1998], and isotopic evidence that the shield-building phase
174 was fed by a different magma reservoir than the rocks that outcrop at the surface [Paulick *et*
175 *al.*, 2010]. During the shield-building phase, the volcano was only 30-40 km from the ridge
176 axis, in a similar location to a present-day subcircular caldera identified by Klingelhoefer *et*
177 *al.* [2001].

178

179 Both elastic and visco-elastic models would predict a bowl-shaped depression of the Moho
180 beneath the edifice. Whilst there is some evidence from seismic data that the Moho deepens
181 beneath Ascension, the depression is far from bowl-shaped and rather appears to be elongated
182 in an east-west direction [Evangelidis *et al.*, 2004]. However, the observation that substantial
183 subsidence occurs > 1 My after the main shield-building phase of volcano growth requires a
184 time constant on the order of 1 My for a visco-elastic model and therefore the inferred value
185 of lithospheric viscosity is robust (within a factor of 2-3) whatever the precise details of the
186 growth history and nature of magmatic addition. However, given that significant lithospheric
187 cooling likely occurred during volcano growth, the viscosity may also have changed
188 significantly, so that the inferred value will be a time-averaged value.

189

190 We conclude the following:

- 191 1. Ascension Island has a complex growth history that is not revealed by surface
192 sampling.
- 193 2. For plausible growth histories, the subsidence at the ASC#1 borehole cannot be
194 explained by elastic models.
- 195 3. The main volcanic edifice has been built during the period 5.0-3.4 Ma.
- 196 4. The viscoelastic component of subsidence is consistent with a relaxation time of ~0.5-
197 1.0 My and a lithosphere viscosity of $0.4-0.7 \times 10^{24}$ Pa s.

198

199 **Acknowledgments**

200 We thank Dennis Nielson and David Langton (University of Utah) for assisting us to sample
201 the borehole material. We thank two anonymous reviewers for constructive comments.

202

202 **Figure captions**

203

204 **Figure 1.** Topography of Ascension Island and the location of the ASC-1 borehole. White
205 lines mark the magnetic anomaly picks of *Brozena* [1986]. Inset shows the tectonic setting of
206 the island on the west flank of the Mid-Atlantic Ridge.

207

208 **Figure 2.** $^{40}\text{Ar}/^{39}\text{Ar}$ age spectra with Ca/K plot for groundmass samples of basaltic rocks
209 drilled from the ASC#1 borehole on the Ascension Island. Error for each step is given at the
210 1σ level. Plateau ages were calculated as weighted means of ages of plateau-forming steps,
211 where each age was weighted by the inverse of its variance.

212

213

214 **Figure 3.** a) Circles mark ages and present-day depths below sea-level of ASC1 borehole
215 samples, with uncertainties. Thick bar marks age range of surface samples. Dotted lines
216 mark present-day sea-level and estimated depth of subaerial to submarine transition. Dashed
217 line indicates approximate mean island growth rate since 3.4 Ma. b) Predicted elastic
218 subsidence since 2.5 Ma, at the borehole site, for the three island growth scenarios described
219 in the text and a range of effective elastic thicknesses (T_e). Solid, dotted and dashed lines
220 correspond to scenarios 1, 2 and 3, respectively. Vertical line marks the lower limit on T_e
221 inferred by [*Evangelidis et al.*, 2004]

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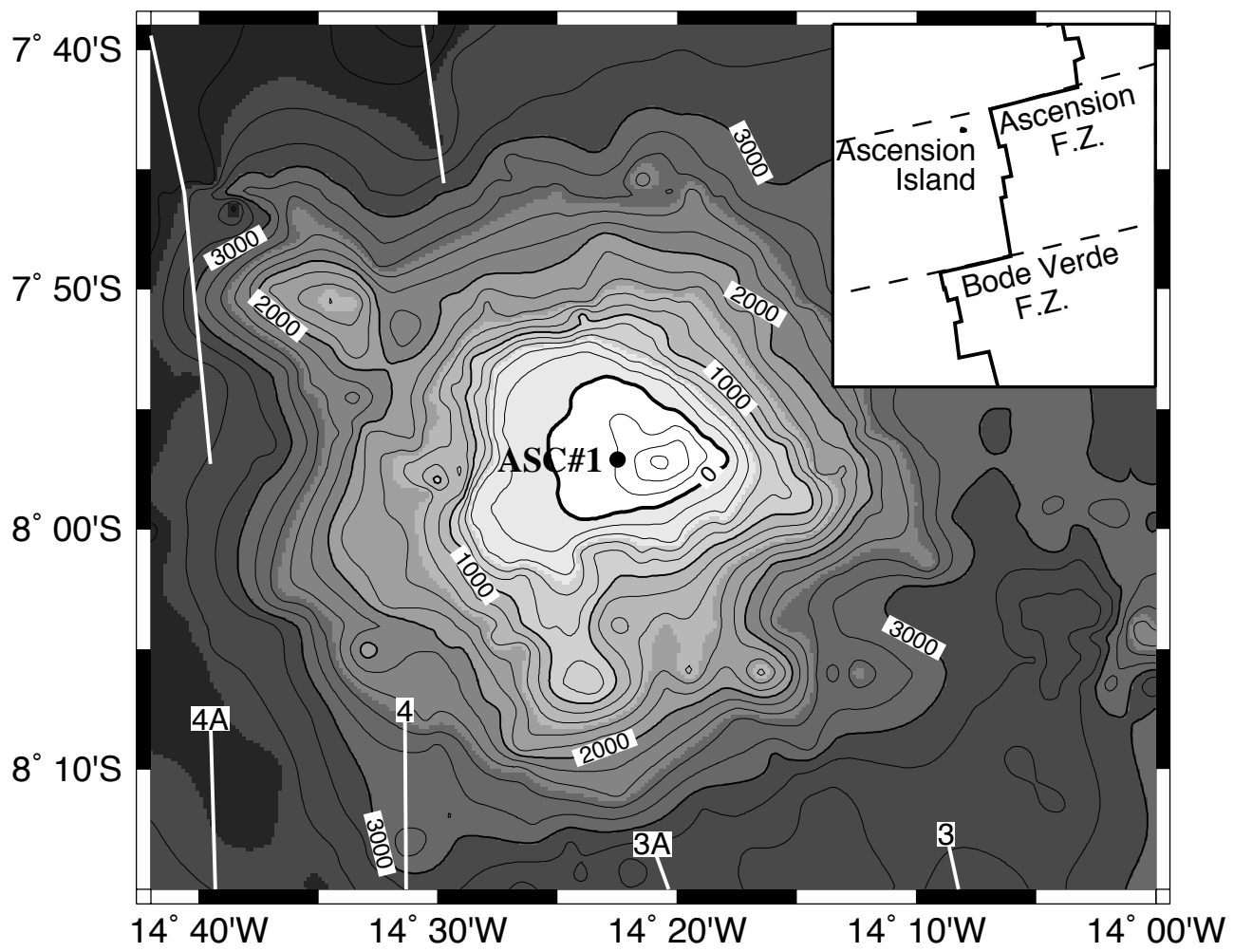
223 **Figure 4.** Predicted visco-elastic subsidence at the borehole site since 2.5 Ma for two
224 different loading histories. a) 88% of load emplaced at 5 Ma and 12% since 3.4 Ma. b) 88%
225 of load emplaced at 4 Ma and 12% since 3.4 Ma.

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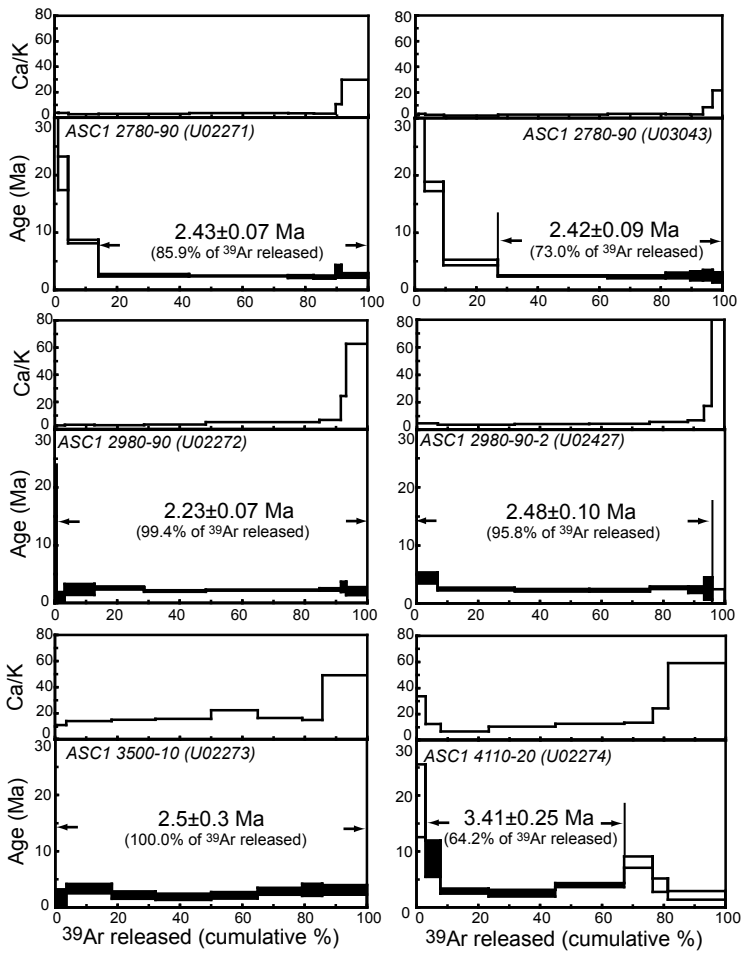
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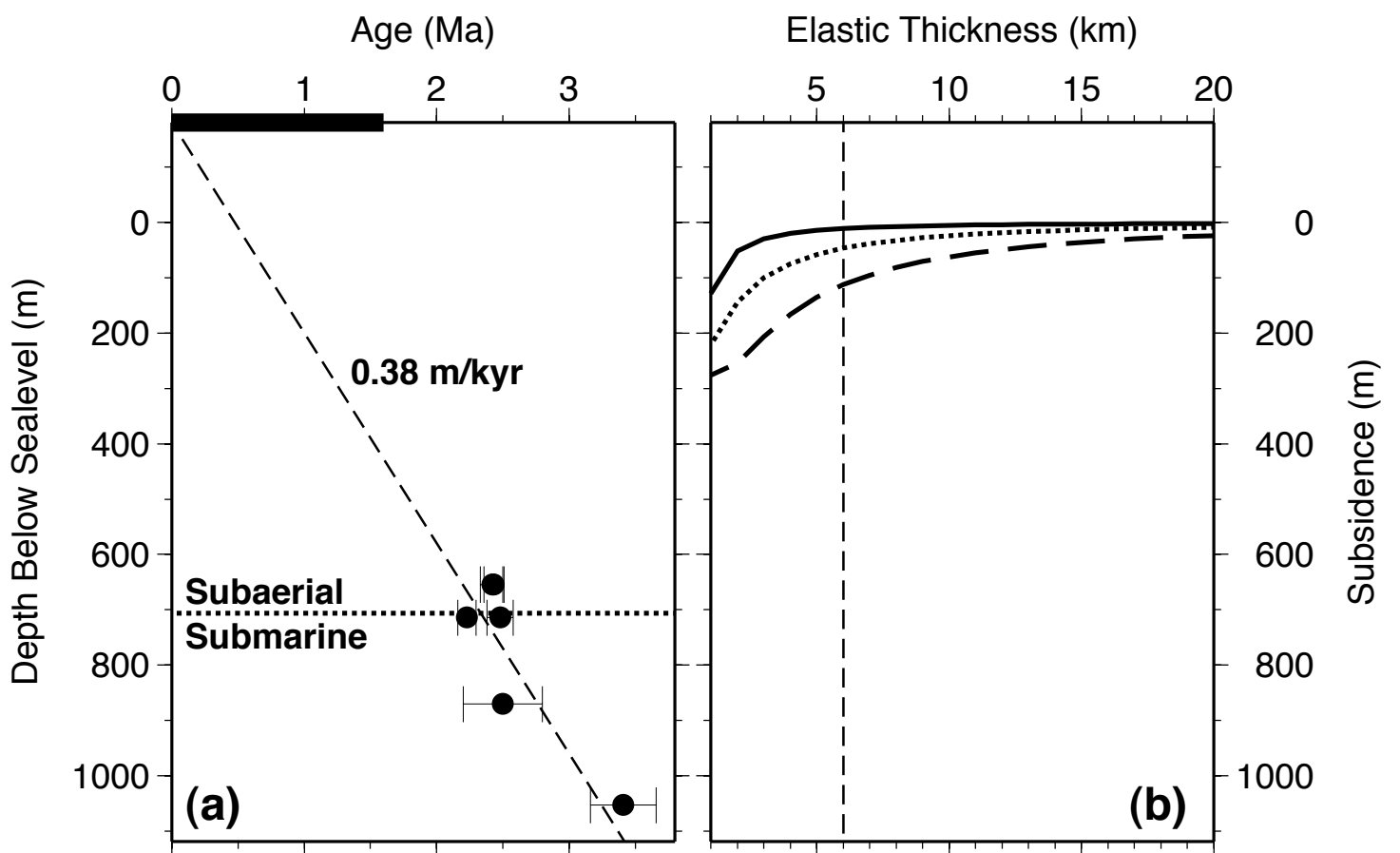
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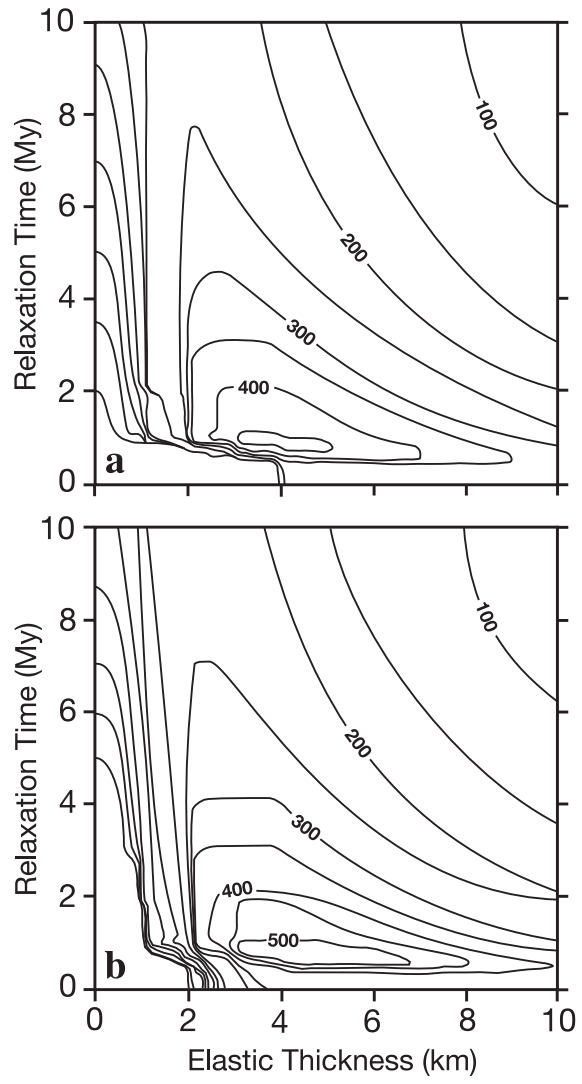
Minshull et al. Fig. 1



Minshull et al. Fig. 2



Minshull et al. Fig. 3



Minshull et al, fig. 4

Dating method

Age determinations were conducted using $^{40}\text{Ar}/^{39}\text{Ar}$ geochronology facility at the Geological Survey of Japan/AIST. Laser step-heating experiments were conducted on 6.6-10.3 mg groundmass samples. We analysed on a groundmass separate from a single chip of drill cuttings. Slabs 1 mm-thick were taken out from the freshest part of the samples with a water-cooled saw. This is partly because flat surfaces are required to precisely monitor the thermal energy distribution on the samples during laser heating. The slabs were gently crushed into small pieces of about 0.5-1 mg weight. The samples were treated ultrasonically in 3N HCl for 30 minutes and then 4N HNO₃ for 30 minutes to remove possible alteration products (clays and carbonates) prior to irradiation. After this acid treatment, groundmass separate was prepared under binocular microscope by handpicking. Sanidine separated from the Fish Canyon Tuff (FC3) was used for the flux monitor and assigned an age of 27.5 Ma [Lanphere and Baadsgaard, 2001].

Samples were baked at 250°C for 72 hours after being placed in an extraction line before analysis. A continuous Ar ion laser was used for sample heating. The groundmass samples were heated for 3 minutes in each step keeping laser power constant. Laser beam diameter was adjusted to 2 mm to ensure uniform heating of the sample. Extracted gas was purified for 10 minutes with three Zr-Al getters (SAES AP-10) and one Zr-Fe-V getter (SAES GP-50). Two Zr-Al getters were maintained at 400°C and other getters were at room temperature. Argon isotopes were measured on a VG Isotech VG3600 noble gas mass spectrometer fitted with a BALZERS electron multiplier. The sensitivity of the collector was about 5×10^{-10} ml STP/V. Mass discrimination was

monitored using diluted air.

Correction for interfering isotopes was achieved by analyses of $\text{CaFeSi}_2\text{O}_6$ and KFeSiO_4 glasses irradiated with the samples. The blank of the system including the mass spectrometer and the extraction line was 7.5×10^{-14} ml STP for ^{36}Ar , 2.5×10^{-13} ml STP for ^{37}Ar , 2.5×10^{-13} ml STP for ^{38}Ar , 1.0×10^{-12} ml STP for ^{39}Ar and 2.5×10^{-12} ml STP for ^{40}Ar . A blank analysis was done every 2 or 3 steps of the analyses.

All errors for $^{40}\text{Ar}/^{39}\text{Ar}$ results are reported at one standard deviation. Errors for ages include analytical uncertainties for Ar isotope analysis, correction for interfering isotopes and J value estimation. An error of 0.5 % was assigned to J values as a pooled estimate during the course of this study. The age plateaus were determined following the definition by *Fleck et al.* [1977]. Inverse isochrones (Fig. A1) were calculated using York's least-squares fit, which accommodates errors in both ratios and correlations of errors [*York*, 1969].

Results from the analyses are tabulated in Tables A1 and A2 and inverse isochron plots are shown in Figure A1.

Figure Caption

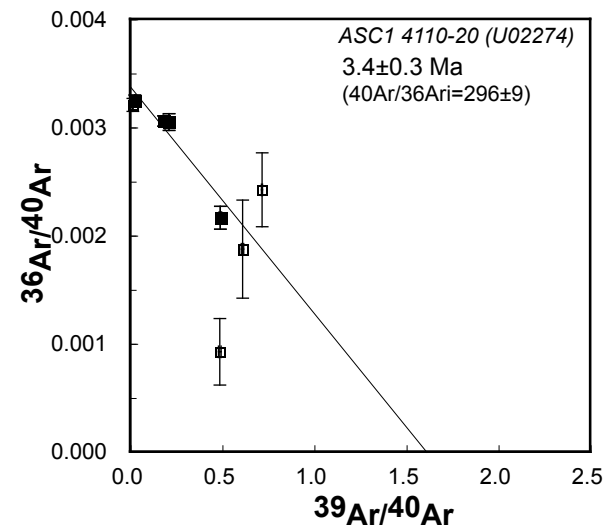
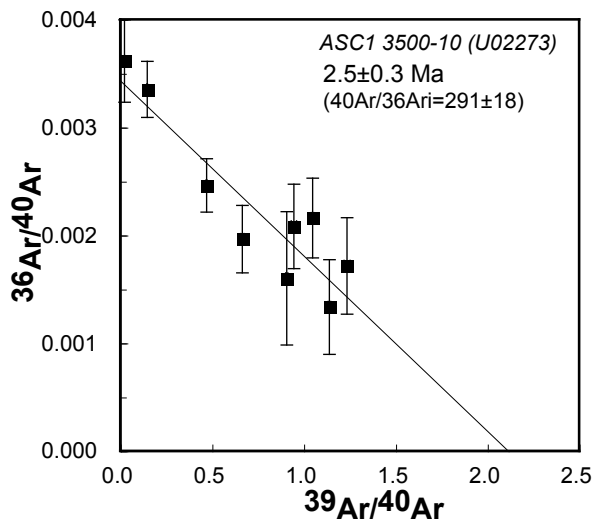
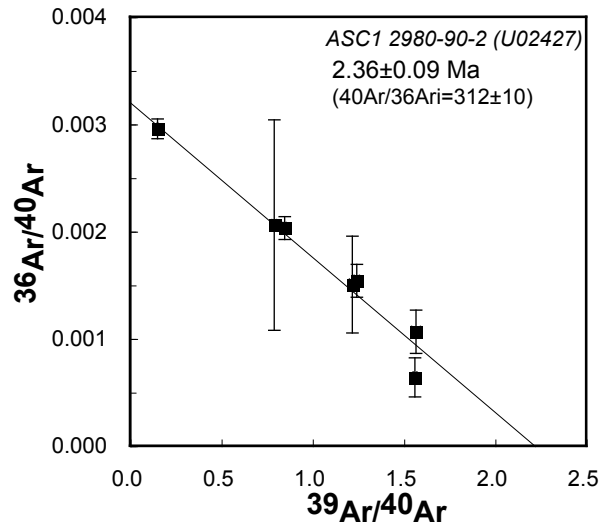
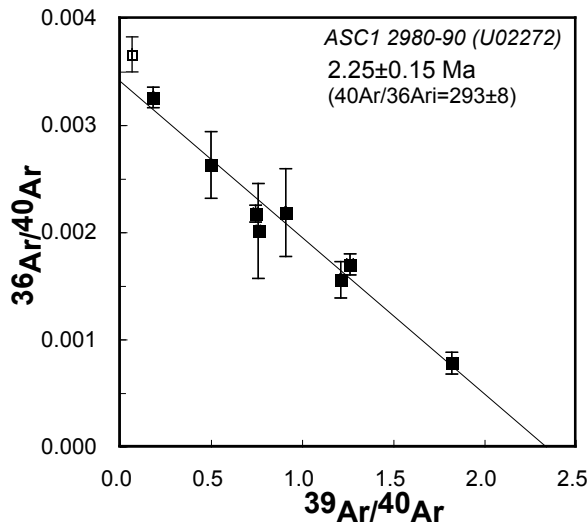
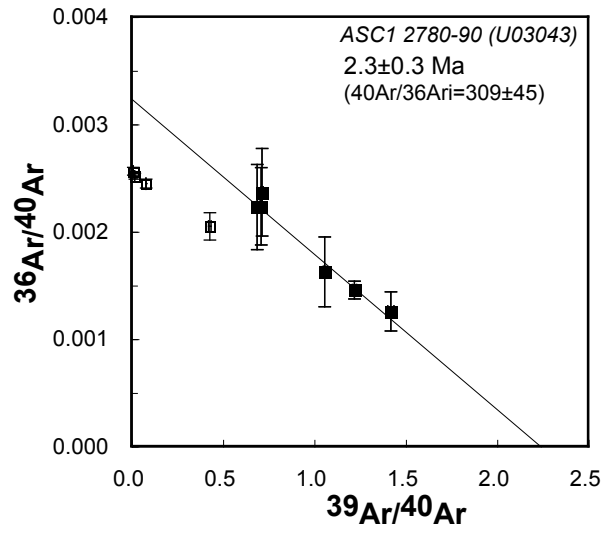
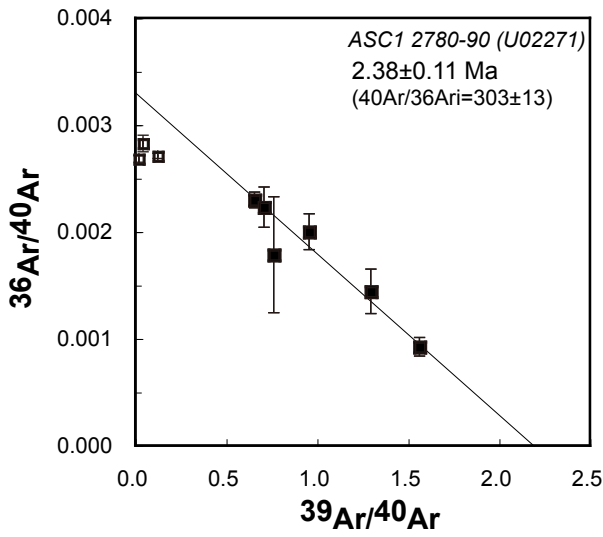
Fig. A1. Inverse isochron plots for groundmass samples of volcanic rocks from the ASC1 drilling site on the Ascension Island. Solid symbols in the inverse isochron plots are the steps used for isochron (plateau forming steps). Error bar for each step is given at the 1σ level.

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Minshull et al. Fig. A1

Table A1: Results of stepwise-heating analyses of volcanic rocks from the ASC1 drilling site on the Ascension Island.								
Analysis	Sample No.	depth	Total age ($\pm 1\sigma$)	Plateau age ($\pm 1\sigma$)				
No.		from surface	integrated age	weighted average	inv. isochron	$^{40}\text{Ar}/^{36}\text{Ar}$	MSWD	fraction of
		(m)	(Ma)	(Ma)	age (Ma)	intercept		^{39}Ar (%)
U02271	ASC1 2780-90	834-837	4.25 \pm 0.13	2.43\pm0.07	2.38 \pm 0.11	303 \pm 13	0.60	85.9
U03043	ASC1 2780-90	834-837	6.78 \pm 0.16	2.42\pm0.09	2.3 \pm 0.3	309 \pm 45	0.32	73.0
U02272	ASC1 2980-90	894-897	2.22 \pm 0.12	2.23\pm0.07	2.25 \pm 0.15	293 \pm 8	1.08	99.4
U02427	ASC1 2980-90-2	894-897	2.45 \pm 0.16	2.48\pm0.10	2.36 \pm 0.09	312 \pm 10	0.92	95.8
U02273	ASC1 3500-10	1050-1053	2.52 \pm 0.29	2.5\pm0.3	2.5 \pm 0.3	291 \pm 18	0.89	100.0
U02274	ASC1 4110-20	1233-1236	4.22 \pm 0.35	3.41\pm0.25	3.4 \pm 0.3	296 \pm 9	2.10	64.2

Table A2

Results of isotopic analysis for $^{40}\text{Ar}/^{39}\text{Ar}$ dating of samples from the ASC1 drilling site on Ascension Island																		
Laser output	$^{40}\text{Ar}/^{39}\text{Ar}$			$^{37}\text{Ar}/^{39}\text{Ar}$			$^{36}\text{Ar}/^{39}\text{Ar}$			K/Ca	$^{40}\text{Ar}^*$	$^{39}\text{Ar}_k$	$^{40}\text{Ar}^*/^{39}\text{Ar}_k$			Age($\pm 1\sigma$)		
											(%)	fraction (%)				(Ma)		
	<i>ASC1 2780-90</i>																	
	J= 0.002896																	
0.6W	50.05	\pm	0.51	2.254	\pm	0.025	135.1	\pm	1.7	0.26	20.8	1.2	10.41	\pm	0.53	54	\pm	3
0.92W	23.64	\pm	0.33	2.176	\pm	0.042	67.59	\pm	1.86	0.27	16.5	3.2	3.912	\pm	0.565	20	\pm	3
1.3W	8.074	\pm	0.038	1.597	\pm	0.015	22.45	\pm	0.20	0.37	20.0	9.6	1.614	\pm	0.059	8.4	\pm	0.3
1.75W	1.525	\pm	0.022	1.750	\pm	0.023	4.146	\pm	0.112	0.34	32.1	28.9	0.4901	\pm	0.0355	2.56	\pm	0.19
2.1W	0.6388	\pm	0.0025	2.119	\pm	0.009	1.368	\pm	0.050	0.28	72.7	31.6	0.4653	\pm	0.0164	2.43	\pm	0.09
2.48W	0.7709	\pm	0.0032	2.018	\pm	0.009	1.852	\pm	0.159	0.29	57.4	8.1	0.4433	\pm	0.0476	2.31	\pm	0.25
3.13W	1.045	\pm	0.004	1.844	\pm	0.015	2.767	\pm	0.173	0.32	40.9	6.9	0.4279	\pm	0.0517	2.2	\pm	0.3
4.05W	1.306	\pm	0.010	6.218	\pm	0.038	4.615	\pm	0.705	0.09	47.2	1.9	0.6203	\pm	0.2106	3.2	\pm	1.1
fusion	1.393	\pm	0.005	17.23	\pm	0.05	9.435	\pm	0.186	0.03	34.1	8.5	0.4821	\pm	0.0789	2.5	\pm	0.4
	<i>ASC1 2780-90</i>																	
	J= 0.002896																	
0.6W	106.9	\pm	1.0	1.403	\pm	0.199	274.1	\pm	4.3	0.42	24.4	1.0	26.07	\pm	1.24	131	\pm	6
0.92W	56.37	\pm	0.34	1.990	\pm	0.109	142.4	\pm	1.7	0.30	25.7	2.2	14.53	\pm	0.47	74	\pm	2
1.3W	12.51	\pm	0.06	1.505	\pm	0.034	31.16	\pm	0.53	0.39	27.7	6.1	3.475	\pm	0.157	18	\pm	1
1.53W	2.327	\pm	0.039	1.184	\pm	0.020	5.197	\pm	0.295	0.50	39.5	17.7	0.9203	\pm	0.0906	4.8	\pm	0.5
2.1W	0.8176	\pm	0.0079	1.622	\pm	0.015	1.784	\pm	0.066	0.36	57.1	35.7	0.4672	\pm	0.0211	2.44	\pm	0.11
2.48W	0.7031	\pm	0.0084	1.924	\pm	0.030	1.589	\pm	0.126	0.31	62.9	18.9	0.4431	\pm	0.0385	2.31	\pm	0.20
2.1W	0.9442	\pm	0.0056	1.797	\pm	0.024	2.193	\pm	0.306	0.33	52.0	8.0	0.4921	\pm	0.0910	2.6	\pm	0.5
2.45W	1.406	\pm	0.009	1.716	\pm	0.046	3.772	\pm	0.505	0.34	34.0	4.1	0.4783	\pm	0.1499	2.5	\pm	0.8
3W	1.446	\pm	0.012	4.951	\pm	0.099	5.037	\pm	0.570	0.12	34.2	3.1	0.4969	\pm	0.1704	2.6	\pm	0.9
fusion	1.383	\pm	0.010	12.54	\pm	0.07	7.872	\pm	0.548	0.05	30.1	3.2	0.4213	\pm	0.1689	2.2	\pm	0.9
	<i>ASC1 2980-90</i>																	
	J= 0.002913																	
0.6W	14.49	\pm	0.26	1.341	\pm	0.047	53.48	\pm	2.38	0.44	0.0	0.6	0.0002	\pm	0.6953	0.0	\pm	3.7
0.92W	5.403	\pm	0.063	1.591	\pm	0.022	18.17	\pm	0.52	0.37	3.8	2.6	0.2078	\pm	0.1528	1.1	\pm	0.8
1.3W	1.990	\pm	0.022	1.871	\pm	0.027	5.909	\pm	0.621	0.31	22.5	9.5	0.4481	\pm	0.1838	2.4	\pm	1.0
1.75W	1.337	\pm	0.022	1.674	\pm	0.024	3.513	\pm	0.105	0.35	35.9	15.9	0.4806	\pm	0.0338	2.52	\pm	0.18
2.1W	0.7896	\pm	0.0115	2.031	\pm	0.027	2.085	\pm	0.076	0.29	49.9	19.8	0.3945	\pm	0.0244	2.07	\pm	0.13
2.48W	0.5471	\pm	0.0071	3.004	\pm	0.036	1.527	\pm	0.046	0.20	77.1	36.3	0.4229	\pm	0.0176	2.22	\pm	0.09
3.15W	0.8211	\pm	0.0031	3.976	\pm	0.014	2.734	\pm	0.133	0.15	54.1	6.9	0.4461	\pm	0.0414	2.34	\pm	0.22

Table A2

3.9W	1.297	± 0.019	14.15	± 0.16	7.800	± 0.557	0.04	40.7	1.7	0.5341	± 0.1726	2.8	± 0.9
fusion	1.063	± 0.004	35.81	± 0.11	15.47	± 0.20	0.02	35.6	6.8	0.3904	± 0.1330	2.1	± 0.7
ASC1 2980-90-2	<i>U02427</i>												
	J= 0.002904												
0.6W	6.615	± 0.095	2.809	± 0.039	20.59	± 0.62	0.21	12.6	6.8	0.8385	± 0.1840	4.4	± 1.0
0.95W	1.178	± 0.013	2.214	± 0.022	3.205	± 0.124	0.27	40.0	25.0	0.4723	± 0.0384	2.47	± 0.20
1.25W	0.8032	± 0.0095	2.478	± 0.028	2.146	± 0.120	0.24	54.5	24.2	0.4389	± 0.0371	2.30	± 0.19
1.6W	0.6366	± 0.0100	2.603	± 0.032	1.632	± 0.126	0.23	68.6	19.6	0.4378	± 0.0391	2.29	± 0.20
2.1W	0.6386	± 0.0042	3.313	± 0.017	1.624	± 0.111	0.18	81.2	12.5	0.5199	± 0.0347	2.72	± 0.18
2.7W	0.8175	± 0.0112	4.035	± 0.046	2.711	± 0.366	0.15	55.6	5.2	0.4559	± 0.1098	2.4	± 0.6
4.1W	1.255	± 0.020	10.07	± 0.11	6.281	± 1.227	0.06	39.2	2.5	0.4963	± 0.3674	2.6	± 1.9
fusion	1.465	± 0.026	47.84	± 0.53	24.41	± 1.44	0.01	0.0	4.2	0.0001	± 0.4694	0.0	± 2.5
ASC1 3500-10	<i>U02273</i>												
	J= 0.002985												
0.6W	42.69	± 1.08	5.748	± 0.194	156.4	± 16.7	0.10	0.0	0.4	0.0011	± 4.8459	0.0	± 26.1
0.95W	6.772	± 0.077	6.416	± 0.084	25.04	± 1.77	0.09	1.0	3.2	0.0680	± 0.5232	0.4	± 2.8
1.4W	1.492	± 0.021	8.255	± 0.092	5.961	± 0.459	0.07	42.0	14.5	0.6312	± 0.1394	3.4	± 0.7
1.8W	1.049	± 0.016	8.853	± 0.105	5.435	± 0.401	0.07	38.5	14.0	0.4070	± 0.1227	2.2	± 0.7
2.3W	0.9470	± 0.0057	9.203	± 0.049	5.424	± 0.335	0.06	36.2	17.9	0.3458	± 0.1042	1.9	± 0.6
2.8W	0.8000	± 0.0103	12.98	± 0.13	6.138	± 0.331	0.04	49.3	14.9	0.3992	± 0.1073	2.1	± 0.6
3.4W	0.8703	± 0.0055	9.617	± 0.046	4.692	± 0.368	0.06	60.6	14.3	0.5319	± 0.1140	2.9	± 0.6
4.05W	1.094	± 0.007	8.723	± 0.041	4.953	± 0.670	0.07	52.7	6.4	0.5814	± 0.2015	3.1	± 1.1
fusion	2.075	± 0.009	28.21	± 0.10	15.47	± 0.41	0.02	27.3	14.4	0.5805	± 0.1556	3.1	± 0.8
ASC1 4110-20	<i>U02274</i>												
	J= 0.003012												
0.6W	65.64	± 1.03	19.45	± 0.29	217.6	± 4.5	0.03	5.3	2.9	3.526	± 1.209	19.1	± 6.5
0.95W	38.49	± 0.45	7.286	± 0.098	127.5	± 2.1	0.08	4.1	4.8	1.606	± 0.599	8.7	± 3.2
1.4W	5.508	± 0.051	3.912	± 0.039	18.27	± 0.26	0.15	9.7	15.6	0.5354	± 0.0788	2.9	± 0.4
1.8W	4.715	± 0.050	6.096	± 0.062	16.60	± 0.37	0.10	10.0	21.5	0.4720	± 0.1104	2.6	± 0.6
2.3W	2.026	± 0.028	7.427	± 0.096	7.105	± 0.203	0.08	36.1	22.3	0.7368	± 0.0657	4.0	± 0.4
2.8W	2.037	± 0.031	7.832	± 0.094	4.756	± 0.622	0.07	72.7	9.2	1.492	± 0.188	8.1	± 1.0
3.85W	1.616	± 0.018	14.23	± 0.12	8.252	± 0.719	0.04	44.7	4.9	0.7311	± 0.2199	4.0	± 1.2
fusion	1.354	± 0.005	33.71	± 0.13	15.66	± 0.29	0.02	28.5	18.7	0.3972	± 0.1408	2.2	± 0.8