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2 **The Development of Magmatism Along the**  
3 **Cameroon Volcanic Line: Evidence from Teleseismic**  
4 **Receiver Functions**

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5 **Abstract.** The Cameroon Volcanic Line (CVL) in West Africa is a chain  
6 of Cenozoic volcanism with no clear age progression. The reasons for its ex-  
7 istence are unclear and the nature of its magmatic plumbing system is poorly  
8 understood. Specifically, whether or not the CVL crust presently contains  
9 melt and/or mafic intrusions, as is often observed at hotspots and rifts else-  
10 where, is presently unknown. To address this issue, we present a receiver func-  
11 tion study of crustal structure using earthquakes recorded by the Cameroon  
12 Broadband Seismic Experiment. In regions of the CVL unaffected by Cre-  
13 taceous extension associated with the breakup of Gondwana (e.g., the Garoua  
14 rift),  $V_p/V_s$  ratios are markedly low (network average  $\sim 1.74$ ) compared to  
15 hotspots elsewhere, providing no evidence for either melt, or cooled mafic  
16 crustal intrusions due to CVL magmatism. The character of P-to-S conver-  
17 sions from beneath the CVL also indicates that lower-crustal intrusions (of-  
18 ten termed underplate) are not present beneath the region. Our observations  
19 thus corroborate earlier petrological studies that show CVL alkaline mag-  
20 mas fractionate in the mantle, not the crust, prior to eruption. Hypotheses  
21 for the formation of the CVL should not include markedly elevated upper-  
22 mantle potential temperatures, or large volumes of partial melt, both of which  
23 can explain observations at hotspots and rifts worldwide. The protracted,  
24 yet sporadic development of small-volume alkali melts beneath the CVL may  
25 instead be explained better by lower melt volume mechanisms such as shear  
26 zone reactivation or lithospheric delamination.

## 1. Introduction

### 1.1. Overview

27 The Cameroon Volcanic Line (CVL) is a region of intra-plate volcanism, with no ev-  
28 idence for age progression, that straddles the continent-ocean boundary in central West  
29 Africa [Figure 1: e.g., *Fitton*, 1980; *Halliday et al.*, 1990; *Nkouathio et al.*, 2008]. The  
30 reasons for the existence of the chain have long been debated, with end-member hypothe-  
31 ses including traditional mantle plumes [e.g., *Lee et al.*, 1994; *Burke*, 2001; *Ngako et al.*,  
32 2006], decompression melting beneath reactivated shear zones in the lithosphere [e.g.,  
33 *Freeth*, 1979; *Fairhead*, 1988; *Fairhead and Binks*, 1991; *Moreau et al.*, 1987], and small-  
34 scale upper-mantle convection [e.g., *King and Anderson*, 1995, 1998; *King and Ritsema*,  
35 2000] each having been proposed for the region. It has also been suggested that lat-  
36 eral flow of buoyant asthenosphere, beneath continental lithosphere thinned extensively  
37 during Mesozoic rifting, may now be contributing to the younger volcanism along the  
38 line [*Ebinger and Sleep*, 1998]. Analogies with other hotspot chains worldwide are thus  
39 not well established and the nature of the CVL's magmatic plumbing system remains  
40 relatively poorly understood.

41 One aspect of CVL magmatism that is now well established is the composition of its  
42 Oligocene-to-Recent lavas. Petrological studies show that the CVL's mostly basaltic volca-  
43 noes have erupted alkaline basalts in relatively low volume [e.g., *Fitton*, 1980; *Suh et al.*,  
44 2003]. These silica poor lavas have not undergone appreciable fractionation at depth,  
45 which raises the question of whether or not the crust beneath the CVL is characterized by  
46 cooled mafic intrusions and/or present-day melt, as is often observed at hotspots and rifts

47 worldwide [e.g., *Ebinger and Casey*, 2001; *Maguire et al.*, 2006; *Thybo and Nielsen*, 2009].

48 To help address this question, and to place fundamental new constraints on melt supply  
49 beneath the CVL, we perform a receiver function study of bulk crustal seismic structure  
50 (Moho depth,  $H$  and  $V_p/V_s$  ratio,  $\kappa$ ) using distant earthquakes recorded by the Cameroon  
51 Broadband Seismic Experiment [CBSE: e.g., *Reusch et al.*, 2010] between 2005 and 2007.  
52 The network samples crust that spans more than 3 billion years of the geological record  
53 in Cameroon, from Congo Craton basement formation during the Archean, to volcanism  
54 along the CVL during Holocene times.

55 If the crust beneath the CVL is host to appreciable volumes of melt and mafic crustal  
56 intrusion, this should be manifest as elevated bulk crustal  $V_p/V_s$  ratios, as is observed at  
57 hotspots and magmatic rifts such as Iceland [e.g., *Darbyshire et al.*, 1998] and Ethiopia  
58 [e.g., *Stuart et al.*, 2006], respectively. Lower crustal intrusions (with or without present-  
59 day melt) can also be identified using the receiver function technique via analysis of  
60 the amplitude and shape of P-to-S conversions from the Moho [e.g., *Stuart et al.*, 2006;  
61 *Zheng et al.*, 2008]. Our fundamental new constraints on Cameroon's bulk crustal seismic  
62 structure thus allow us to draw comparisons with rifts and hotspots worldwide, where  
63 the reasons for magmatism are relatively well established. Better constraints on the CVL  
64 magmatic plumbing system lead, in turn, to a better understanding of the mechanism of  
65 melt development in the mantle beneath the line.

## 1.2. Geological and Tectonic Setting

66 The oldest rocks in Cameroon are found within the 2.5–3.0 b.y. old Archean Congo  
67 craton in the SW of the country [Figure 1; e.g., *Cahen et al.*, 1984; *Nzenti et al.*, 1988;

68 *Tchameni et al.*, 2001]. The northernmost exposure of the craton in Cameroon is known  
69 as the Ntem Complex (Figure 1), which consists of Archean-age greenstone terranes,  
70 surrounded by tonalitic, trondhjemitic and granodioritic rocks [e.g., *Tchameni et al.*,  
71 2000]. Parts of the region were re-worked in Paleoproterozoic times with mafic doleritic  
72 intrusions modifying the crust [e.g., *Tchameni et al.*, 2001].

73 The Proterozoic Oubanguides Belt, comprising metamorphosed schists and gneisses,  
74 lies to the north of the Congo Craton. It forms part of the larger Neoproterozoic Pan  
75 African–Brazilian Belt, which underwent significant deformation during the Pan African  
76 Orogeny ca. 600 Ma when the Congo, São Francisco and West African Cratons collided  
77 during the formation of Gondwana [e.g., *Toteu et al.*, 1987; *Nzenti et al.*, 1988; *Toteu*  
78 *et al.*, 2001; *Oliveira et al.*, 2006]. During the collision, Proterozoic sediments were thrust  
79 on top of the edge of the Congo Craton ca. 565 Ma [e.g., *Toteu et al.*, 1987; *Nzenti et al.*,  
80 1988; *Ngako et al.*, 2006] such that its northern edge is now buried beneath the Yaoundé  
81 domain [Figure 1; e.g., *Toteu et al.*, 1987; *Oliveira et al.*, 2006].

82 Cross-cutting the Proterozoic Oubanguides Belt are several major shear zones, including  
83 the Central African Shear Zone (CASZ), a major tectonic feature that can be traced across  
84 the African continent from Sudan towards the Adamawa Plateau [e.g., *Guiraud et al.*,  
85 1992; *Binks and Fairhead*, 1992] and into SW Cameroon, where it is masked by Cenozoic  
86 volcanic rocks (Figure 1).

87 The breakup of Gondwana ca. 125 Ma resulted in the separation of Africa and South  
88 America, and the formation of the South Atlantic [e.g., *Burke and Dewey*, 1973]. At this  
89 time, SW Cameroon and the region that is now the Gulf of Guinea was characterized by

90 a triple junction, the third arm of which formed the Benue Trough, a continental rift that  
91 subsequently failed to develop [e.g., *Burke et al.*, 1971]. The Garoua rift (Figure 1) marks  
92 the eastward extension of the Benue Trough into northern Cameroon. The opening of  
93 the South Atlantic also resulted in reactivation of the CASZ [e.g., *Binks and Fairhead*,  
94 1992; *Browne and Fairhead*, 1983], which is known on the conjugate rifted margin in  
95 Brazil as the Pernambuco lineament [e.g., *Oliveira et al.*, 2006]. The CASZ consists of  
96 many individual shear zones that formed extensional basins during Cretaceous times to  
97 accommodate stress transferred to the African continental interior during the opening of  
98 the South Atlantic [e.g., *Browne and Fairhead*, 1983].

99 Since Oligocene times, Cameroon has experienced sporadic volcanism along the CVL,  
100 which trends SW–NE and extends  $\sim 1600$  km from offshore in the Gulf of Guinea to Lake  
101 Chad [Figure 1: *Fitton*, 1980; *Halliday et al.*, 1988; *Déruelle et al.*, 2007; *Nkouathio et al.*,  
102 2008], with an additional E-W trending arm at  $\sim 7^\circ$ N that extends across the Adamawa  
103 Plateau (Figure 1). The CVL is underlain by Pan African basement rocks consisting of  
104 schists and gneisses intruded by granites and diorites [e.g., *Fitton*, 1987; *Déruelle et al.*,  
105 2007]. Magmatism along the CVL began ca. 30 Ma at the Mandara Mountains and on the  
106 Island of Principe [e.g., *Halliday et al.*, 1988] but there is no evidence for age progression  
107 along the line [e.g., *Fitton and Dunlop*, 1985; *Halliday et al.*, 1988]. Table 1 shows a  
108 summary of the timing of CVL volcanism, compiled by *Halliday et al.* [1988].

109 *Fitton and Dunlop* [1985] showed using the elemental compositions of basaltic lava  
110 collected along the CVL, that magmas along the continental and oceanic sectors of the line  
111 are very similar. This led them to the conclusion that there is a single mantle source along

112 the chain, a hypothesis supported by several subsequent studies [e.g., *Halliday et al.*, 1988;  
113 *Rankenburg et al.*, 2005; *Déruelle et al.*, 2007]. Petrological data also show that hawaiiite  
114 and basanite alkali basaltic volcanism is commonplace along the CVL. These low-volume,  
115 high-pressure lavas suggest that little shallow (crustal) fractionation of magma is occurring  
116 beneath the line prior to eruption [e.g., *Fitton and Dunlop*, 1985; *Halliday et al.*, 1988;  
117 *Marzoli et al.*, 2000; *Suh et al.*, 2003; *Yokoyama et al.*, 2007; *Déruelle et al.*, 2007; *Njonfang*  
118 *et al.*, 2011; *Teitchou et al.*, 2011].

### 1.3. Previous Geophysical Work

119 The lithospheric seismic structure of Cameroon has been determined by continent-scale  
120 surface-wave studies [e.g., *Pasyanos and Nyblade*, 2007; *Priestley et al.*, 2008; *Fishwick*,  
121 2010]. They indicate that the CVL is underlain by mantle characterized by slow seismic  
122 wave-speeds and that the lithosphere-asthenosphere boundary beneath the Congo Craton  
123 is at a depth of  $\sim 250$  km. In contrast, the lithosphere is  $< 100$  km thick beneath the CVL  
124 [*Fishwick*, 2010].

125 *Pérez-Gussinyé et al.* [2009] constrain effective elastic plate thickness (a proxy for  
126 lithospheric strength) across Africa using coherence analysis of topography and Bouguer  
127 anomaly data. The weakest lithosphere is found in Ethiopia and is attributed both to  
128 the low wave-speed, hot mantle beneath the region, and the large degree of extension in  
129 that part of the East African rift system. Channels of relatively weak lithosphere extend  
130 across the African continent from this region to the CVL, where the effective elastic plate  
131 thickness is also relatively low compared to the surrounding areas. This was cited by  
132 *Pérez-Gussinyé et al.* [2009] as evidence in support of the model of *Ebinger and Sleep*

133 [1998], who proposed that flow of buoyant asthenosphere beneath continental lithosphere  
134 thinned extensively during Mesozoic rifting may now be contributing to volcanism along  
135 the CVL.

136 On a more local scale in Cameroon, *Dorbath et al.* [1986] analysed teleseismic P-wave  
137 travel-time residuals recorded by a 300 km-long profile of 40 short period seismograph  
138 stations across the Adamawa Plateau. Using these data they identified upper-mantle  
139 wave-speed contrasts of  $<2.5\%$  across the Plateau (Figure 1), an observation they cite as  
140 evidence for the presence of a mantle thermal anomaly. *Plomerova et al.* [1993] analysed  
141 data from the same seismic array and found evidence for  $\sim 70$  km lithospheric thinning  
142 beneath the Adamawa Plateau in the region where it is cross-cut by the CASZ. In support  
143 of these teleseismic studies, gravity **studies** by *Poudjom Djomani et al.* [1997] and *Nnange*  
144 *et al.* [2001] also found evidence for localized low density mantle beneath the region.

145 Most recently, *Reusch et al.* [2010] used body wave tomography to image the mantle  
146 seismic structure beneath Cameroon using data from the 2005–2007 CBSE experiment.  
147 They found that a continuous low velocity zone ( $\delta V_S = -2$  to  $-3\%$ ) underlies the entire  
148 CVL to a depth of at least 300 km and attributed this to a thermal anomaly of at least  
149 280 K.

150 Within Cameroon, gravity maps have often been used to delineate tectonic subdivisions  
151 inaccessible by traditional field geology due to younger sediment cover [e.g., *Fairhead and*  
152 *Okereke*, 1987; *Poudjom Djomani et al.*, 1992; *Djomani et al.*, 1995; *Poudjom Djomani*  
153 *et al.*, 1997; *Toteu et al.*, 2004; *Tadjou et al.*, 2009; *Shandini et al.*, 2010; *Basseka et al.*,  
154 2011]. A steep gradient in the Bouguer gravity field at  $\sim 4^\circ\text{N}$  and  $\sim 10^\circ\text{E}$  within the



155 Yaoundé domain (Figure 1), for example, has been interpreted by these **researchers** as  
156 the sediment-covered edge of the Congo Craton.

157 The earliest seismological constraints on crustal thickness in Cameroon came from *Stu-*  
158 *art et al.* [1985], who used direct as well as reflected and refracted waves from quarry blasts  
159 to determine a crustal thickness of  $\sim 33$  km beneath the Adamawa Plateau, which has an  
160 upper-mantle P-wave velocity of  $\sim 8$   $\text{kms}^{-1}$ . They also show that the crustal thickness in  
161 the Garoua Rift is  $\sim 23$  km with an upper-mantle P-wave velocity of  $\sim 7.8$   $\text{kms}^{-1}$ . More  
162 recently, *Tokam et al.* [2010] conducted a joint receiver function and surface-wave study  
163 of crustal structure across Cameroon using data from the CBSE network. Average shear  
164 wave velocity across Cameroon was found to be  $\sim 3.7$   $\text{kms}^{-1}$  and mean crustal thickness  
165  $\sim 36$  km. Beneath the Garoua Rift and towards the coast, the crust was found to be sig-  
166 nificantly thinner at 26-31 km. In contrast, the Congo Craton in Cameroon has thicker  
167 crust (43–48 km) and elevated mean crustal seismic velocities (3.9  $\text{kms}^{-1}$ ). The CVL and  
168 Oubanguides belt are characterized by crustal thicknesses of 35-39 km. Thickened crust  
169 and a  $\sim 25$  km-thick high-velocity lower-crustal layer beneath the northern margin of the  
170 Congo Craton was attributed to the collisional tectonic processes that characterized the  
171 formation of Gondwana. *Tokam et al.* [2010] also suggested that thin bodies of mafic  
172 material exist in the top 10–15 km of the crust throughout Cameroon.

173 Cameroon is relatively aseismic compared to magmatically active regions such as  
174 Ethiopia [*Keir et al.*, 2009; *Ebinger et al.*, 2010] in East Africa. However, studies of  
175 seismicity around Mount Cameroon, site of the most recent eruption along the CVL [e.g.,  
176 *Ateba et al.*, 2009], show that earthquakes occur at depths as great as 55–60 km in the

177 sub-continental lithospheric mantle [Tabod *et al.*, 1992; Ateba and Ntepe, 1997]. For a  
178 review of recent seismological work in Cameroon, see *Fishwick and Bastow* [2011].

## 2. Method

179 Data were obtained from 32 broadband stations throughout the CBSE network, span-  
180 ning the continental sector of the Cameroon Volcanic Line and the boundary with the  
181 Congo Craton (Figure 1). The CBSE was run between January 2005 and February 2007  
182 and the stations were positioned  $\sim 50$ – $150$  km from each other [Reusch *et al.*, 2010]. A  
183 second order Butterworth filter with corner frequencies of 0.04 and 3 Hz was applied to  
184 the data prior to analysis. Quality control was subsequently performed on the data such  
185 that only traces with a high signal-to-noise ratio and P arrivals were used. This resulted  
186 in a data set of 158 earthquakes which range between 3 and 18 per station (Figure 2).

187 **Receiver functions capture P-to-S wave conversions at velocity contrasts in**  
188 **the receiver crust and mantle that are recorded in the P-wave coda from**  
189 **teleseismic earthquakes** [Langston, 1977]. The receiver function method used in this  
190 study is the Extended-Time Multitaper Frequency-Domain Cross-Correlation Receiver-  
191 Function (ETMTRF) approach of *Helffrich* [2006]. This uses short multiple tapers which  
192 window the full length of the time series. This adds all of the Fourier transforms in the  
193 frequency domain such that the phase lags of each sub-window are preserved. The travel  
194 times of P-to-S conversions at the Moho can be studied to provide information on the  
195 crustal thickness (H) and  $V_p/V_s$  ratio ( $\kappa$ ). This is achieved with the stacking procedure  
196 outlined by *Zhu and Kanamori* [2000]. Both a linear [Zhu and Kanamori, 2000] and a  
197 phase weighted stacking [PWS: *Schimmel and Paulssen*, 1997] approach were used to

198 verify the results. The linear approach uses Equation 1 for stacking, in conjunction with  
 199 the equations for the predicted travel times (Equations 2-4). This results in a grid which  
 200 gives the plausible range of values for  $H$  and  $\kappa$  between, 20–60 km and 1.6–2.2 respectively:

$$s(H, \kappa) = \sum_{j=1}^N w_1 r_j(t_1) + w_2 r_j(t_2) - w_3 r_j(t_3), \quad (1)$$

201 where  $\omega_1$ ,  $\omega_2$  and  $\omega_3$  are weights,  $r_j(t_i)$  are the amplitudes at the arrival times for each  
 202 of the ray paths evaluated.  $N$  is the number of receiver functions:

$$t_1 = H \left[ \sqrt{\frac{1}{V_S^2} - p^2} - \sqrt{\frac{1}{V_P^2} - p^2} \right], \quad (2)$$

$$t_2 = H \left[ \sqrt{\frac{1}{V_S^2} - p^2} + \sqrt{\frac{1}{V_P^2} - p^2} \right], \quad (3)$$

$$t_3 = 2H \sqrt{\frac{1}{V_S^2} - p^2}, \quad (4)$$

203 where  $p$  is the ray parameter. For phase weighted stacking, the linear approach is altered  
 204 by adding the factor  $c$ , which is a measure of the correlation of the data between 0 and 1:

$$c(H, \kappa) = \frac{1}{N} \sum_{j=1}^N \frac{|\sum_{\kappa=1}^3 e^{i\Phi(t_\kappa)}|}{3}, \quad (5)$$

205 where  $\Phi$  is the instantaneous phase at time  $t$ . This is then combined with Equation 1  
 206 to give:

$$s(H, \kappa) = c^v \sum_{j=1}^N w_1 r_j(t_1) + w_2 r_j(t_2) - w_3 r_j(t_3), \quad (6)$$

207 where the value of  $v$  controls the sharpness of the filtering by  $c$ , between poorly cor-  
 208 related and well correlated data. The value for  $v$  was taken to be 2 throughout the  
 209 analysis as this provided sharper output than a value of 1, consistent with the studies of  
 210 *Schimmel and Paulssen* [1997] and *Thompson et al.* [2010]. The weights were chosen as  
 211  $\omega_1=0.5$ ,  $\omega_2=0.4$  and  $\omega_3=0.1$ , as the P-to-S and PpPs transitions are much clearer than the  
 212 PpSs+PsPs transition. The errors in the results were calculated by taking the maximum  
 213 axes of the 95% confidence interval of the grid plotted using the *Zhu and Kanamori* [2000]  
 214 stacking method, where the H- $\kappa$  grid consists of 10000 points (Figure 3).

### 3. Results

215 Of the 32 CBSE stations, 24 yielded receiver functions with a coherent P-to-S arrival. Of  
 216 these, **18** displayed PpPs and PpSs+PsPs arrivals, thus enabling H- $\kappa$  analysis. The phase  
 217 weighted stacking and linear methods yielded similar results with differences of no more  
 218 than 0.8 km for H and 0.06 for  $V_p/V_s$  (Table 2). Errors in the linear method were slightly  
 219 higher than those for the phase weighted stacking method, with differences of <0.7 km  
 220 for  $\delta H$  and 0.03 for  $\delta V_p/V_s$ . Examples of stacked receiver functions and H- $\kappa$  plots for  
 221 stations CM22 and CM29 are shown in Figure 3. For the **6** stations at which no clear  
 222 reverberant phases were found, receiver functions were combined to form single station  
 223 stacks. Crustal thicknesses were then determined using Equation 2 assuming  $V_p/V_s$  ratio  
 224 from the nearest station constrained by H- $\kappa$  analysis.

225 Six stations in the south of Cameroon (CM02–CM04, CM06, CM07, CM11) have a  
 226 crustal thickness of  $>40$  km (Figure 4). Other nearby stations (CM01, CM05, CM10,  
 227 CM12, CM17) have markedly thinner crust ( $H \approx 37$  km: Figure 4). Most of the southern  
 228 stations (e.g., CM02, CM03, CM07, CM10, CM12, CM17) have extremely low  $V_p/V_s$   
 229 ratios of 1.65–1.73 (Figure 5). Exceptions to this are stations CM04 and CM06 in the  
 230 Ntem Complex ( $V_p/V_s = 1.78$  and  $1.82$ , respectively). Transects A–A' and B–B' (Figure  
 231 6) show the abrupt change in crustal thickness of  $\sim 5$  km and the lack of variation in  $V_p/V_s$   
 232 ratios between the CVL and neighbouring Congo Craton. Throughout the CVL and  
 233 central Cameroon (CM13, CM16, CM20–CM27), crustal thickness is  $\sim 34$ – $39$  km (Figure  
 234 4);  $V_p/V_s$  ratios are generally low at 1.67–1.76 (Figure 5). Transect C–C' show the  
 235 consistently low  $V_p/V_s$  ratios along the CVL (Figure 6).

236 While most of the study area is characterized by relatively low  $V_p/V_s$  ratios (network  
 237 mean  $V_p/V_s = 1.74$ ) compared to the global mean of 1.768 [Christensen, 1996], localized  
 238 exceptions, in addition to CM04 and CM06, include CM29, CM30 and CM32 near the  
 239 Garoua Rift (Figure 1) in the northern part of the study area; these have  $V_p/V_s$  ratios  
 240 of 1.75–1.84; the crust is also thinner in this region ( $H = 25$ – $33$  km: Figure 4 & 6).

## 4. Discussion

### 4.1. Comparison with Previous Studies and Implications for Tectonic Subdivisions

241 Many of the tectonic subdivisions in Cameroon are based entirely on analysis of poten-  
 242 tial field data such as gravity studies [e.g., Fairhead and Okereke, 1987; Djomani et al.,  
 243 1995]; mapping by traditional field geology is often not possible because the putative

244 terrane boundaries are covered by younger rocks [e.g., *Basseka et al.*, 2011]. For exam-  
245 ple, the northern edge of the Congo Craton lies beneath thick Proterozoic sediments in  
246 the Yaoundé domain (Figure 1) that were thrust on top of it ca. 565 Ma during the  
247 formation of Gondwana [e.g., *Toteu et al.*, 1987; *Nzenti et al.*, 1988; *Ngako et al.*, 2006].  
248 A sharp crustal boundary identified as a steep gradient in the Bouguer gravity field at  
249  $\sim 4^\circ\text{N}$  appears to constrain the northern edge of the Congo craton at depth [e.g., *Fairhead*  
250 *and Okereke*, 1987; *Djomani et al.*, 1995; *Poudjom Djomani et al.*, 1997; *Toteu et al.*,  
251 2004; *Tadjou et al.*, 2009; *Shandini et al.*, 2010; *Basseka et al.*, 2011]. Consistent with  
252 these studies, we observe an abrupt change in crustal thickness at the same latitude. The  
253 average crustal thickness at stations CM02–CM04, CM06–CM07, and CM11 is  $\sim 44$  km,  
254 but north of  $\sim 4^\circ\text{N}$  it is consistently  $\leq 38$  km (Table 2, Figures 4 and 6). When reviewed  
255 in light of other geophysical constraints, including the receiver function study of *Tokam*  
256 *et al.* [2010], our results thus confirm that the northern limit of the Congo Craton is likely  
257  $\sim 4^\circ\text{N}$ , and its eastern limit in Cameroon  $\sim 10^\circ\text{E}$ .

258 In cratonic southern Cameroon,  $V_p/V_s$  ratios at stations CM04 and CM06 (Figure  
259 5) within the Archean Ntem complex (Figure 1) are elevated ( $V_p/V_s = 1.78$  and  $1.82$ ,  
260 respectively) compared to the values usually encountered in cratonic regions [e.g.,  $\sim 1.725$   
261 for Paleoproterozoic Rae domain of northern Canada: *Thompson et al.*, 2010]. The Ntem  
262 complex has been unaffected by volcanism associated with the Cretaceous breakup of  
263 Gondwana, and the subsequent development of the CVL; thus neither of these tectonic  
264 processes can explain the observations. Instead, mafic intrusion during the Neoproterozoic  
265 emplacement of greenstone terranes [e.g., *Tchameni et al.*, 2000] may have elevated  $V_p/V_s$

266 ratios compared to the surrounding relatively felsic cratonic crust at stations such as  
267 CM02, CM03 and CM07 where  $V_p/V_s = 1.65\text{--}1.73$  (Figure 5). A similar hypothesis was  
268 proposed by *Thompson et al.* [2010] to explain observations in the greenstone terranes of  
269 the  $\sim 2.7$  Ga Hearne domain in Canada, where  $V_p/V_s$  ratios are slightly higher at  $\sim 1.76$   
270 than in the neighboring Paleoproterozoic Rae domain ( $V_p/V_s = 1.725$ ). Additionally, mafic  
271 material may have been assimilated into the crust during collisional tectonics during the  
272 formation of Gondwana. This latter hypothesis was also proposed by *Tokam et al.* [2010],  
273 who cited high-velocity lower-crust in this region as evidence for mafic magmatic addition  
274 during the Pan African orogeny.

275 Previous studies of Precambrian terranes worldwide indicate that average Archean crust  
276 is thinner than average Proterozoic crust [ $\sim 35$  km and  $\sim 45$  km respectively; *Durrheim*  
277 *and Mooney*, 1991], with thicker Archean terranes generally the site of ancient collisional  
278 boundaries. On the other hand *Tedla et al.* [2011] used surface and satellite gravity data to  
279 produce a crustal model for Africa that shows average crustal thickness for the continent is  
280 almost identical to that of the Archean and Proterozoic domains within it (each  $\sim 39$  km).  
281 Since our constraints on crustal thickness of the Congo Craton are on its rifted margins,  
282 not its deep interior, we cannot determine with confidence the extent to which collisional  
283 processes during the Pan African Orogeny have served to thicken the crust compared to  
284 the neighbouring Proterozoic crust.

285 In the northern part of the study area, the Garoua Rift marks the eastward extension of  
286 the Benue Trough into northern Cameroon. Rifting here developed in Cretaceous times  
287 during the breakup of Gondwana. Several gravity and seismic field experiments have

288 been conducted both here and across the nearby Adamawa Plateau (Figure 1). *Stuart*  
289 *et al.* [1985], for example, constrained the average crustal thickness for the Plateau (Fig-  
290 ure 1) to be  $\sim 33$  km, which closely matches our results of  $H = 38.2$  km and  $H = 34.9$  km  
291 for stations CM24 and CM26 respectively (Table 2, Figures 4 and 6b). The same study  
292 constrained the crustal thickness beneath the Garoua Rift (Figure 1) to be  $\sim 23$  km, a  
293 value closely matched by our result of  $H = 25.3$  km at station CM29 (Table 2, Figures  
294 4 and 6b,c). Elevated  $V_p/V_s$  ratios here (e.g., 1.84 at CM29) may indicate addition of  
295 mafic material to the crust during Cretaceous extension, as is observed in both present-  
296 day magma rich [e.g., Ethiopia: *Maguire et al.*, 2006] and non-volcanic rifts [e.g., Baikal:  
297 *Thybo and Nielsen*, 2009] worldwide. Additionally, mafic addition to the crust may have  
298 occurred more recently at the Adamawa Plateau on the basis of thinned lithosphere [e.g.,  
299 *Plomerova et al.*, 1993], modestly low mantle seismic wave-speeds [e.g., *Stuart et al.*, 1985;  
300 *Dorbath et al.*, 1986; *Reusch et al.*, 2010], and low mantle densities [*Poudjom Djomani*  
301 *et al.*, 1992, 1997; *Nnange et al.*, 2001]. The Afar hotspot is a possible source of low  
302 density material, with buoyant asthenosphere flowing laterally beneath continental litho-  
303 sphere thinned extensively during Mesozoic rifting since the onset of East African hotspot  
304 volcanism  $\sim 40$  Ma [e.g., *Ebinger and Sleep*, 1998; *Pérez-Gussinyé et al.*, 2009; *Rooney*  
305 *et al.*, 2012a].

306 **Figure 7 shows a summary comparison of crustal thickness constrained in**  
307 **this study for CBSE stations with those of *Tokam et al.* [2010]. Our results are**  
308 **in close agreement with the earlier study, except in instances where the  $H$ - $\kappa$**   
309 **method was not applicable due to low signal-to-noise ratio P-to-S conversions**



310 (CM01, CM05, CM11, CM13, CM23, CM27: Table 2). At these stations,  
311 discrepancies between the results may be due to uncertainties in the  $V_p/V_s$   
312 ratio. Alternatively, the data at these stations may not be significantly high  
313 quality for use in the detailed 1D modeling employed by *Tokam et al.* [2010].

314 Towards the continent-ocean transition, in southern Cameroon the crustal  
315 thickness of  $\sim 37$  km found at stations CM01 and CM05 (Figures 4 and 6c) is  
316 somewhat thicker than the 29.9 km value constrained by *Tokam et al.* [2010]  
317 and the 26.8 km constrained by the gravity study of *Tadjou et al.* [2009].  
318 The easterly backazimuth of earthquakes used in our study to constrain H at  
319 CM01 and CM05 likely means that our study may be sensitive principally to  
320 the Congo craton, which lies immediately to the east of these stations (Figure  
321 1).

#### 4.2. Implications for the Crustal Magmatic Plumbing System Beneath the Cameroon Volcanic Line

322 Present-day melt within the crust would be expected to markedly raise  $V_p/V_s$  ratios  
323 [ $>1.9$ ; e.g., *Watanabe*, 1993]; mafic intrusions emplaced during the CVLs de-  
324 velopment would, even after cooling, also be expected to raise  $V_p/V_s$  ratios  
325 compared to the global average value of 1.768, which is dominated by the  
326 inherently felsic (low  $V_p/V_s$  ratio) shields. To this end, a striking observation in  
327 Figure 5 is that  $V_p/V_s$  ratios along portions of the CVL unaffected by Cretaceous rift-  
328 ing (e.g., the Garoua rift) and older magmatic events are markedly low compared to the  
329 global mean.  $V_p/V_s$  values are, in-fact, more akin to those found in cratons worldwide

330 [e.g., 1.72–1.76 in Archean northern Hudson Bay: *Thompson et al.*, 2010], despite the  
331 CVL being a region of active volcanism for  $\sim 30$  Ma.

332 Recent volcanism [ca. 1 Ma; *Halliday et al.*, 1988] at Mount Manengouba (Figure 1)  
333 might be expected to have resulted in elevated  $V_p/V_s$  ratios at stations CM10 and CM16,  
334 but our observations of  $V_p/V_s \leq 1.68$  indicate that this is not the case. It must be ac-  
335 knowledged that none of the CBSE stations used in this study were located on one of the  
336 CVL’s active volcanoes. Mount Manengouba, for example, is  $\sim 65$  km away from station  
337 CM13 (Figure 1). Therefore, we cannot preclude the possibility that melt and mafic mag-  
338 matic intrusions exist directly beneath the volcanic centres. However, the mean  $V_p/V_s$   
339 ratio across the entire study area is 1.74, implying little modification of Precambrian  
340 basement has occurred due to the development of the CVL, either by plutonism (e.g.,  
341 gabbroic intrusions) or by the emplacement of mafic cumulates due to fractionation of  
342 magmas in the crust prior to eruption. This is in slight contrast to *Tokam et al.* [2010],  
343 who cite seismological data as evidence for thin, shallow crustal intrusions beneath the  
344 CBSE network. Our low  $V_p/V_s$  ratios show that if these bodies exist, they are not vol-  
345 umetrically significant enough to affect our bulk crustal results. Our results thus agree  
346 instead with petrological studies that **attribute low volume, high pressure mag-**  
347 **mas to melting of sub-continental lithospheric mantle that has experienced**  
348 **only small amounts of crustal fractionation** [e.g., *Fitton and Dunlop*, 1985; *Halli-*  
349 *day et al.*, 1988; *Marzoli et al.*, 2000; *Suh et al.*, 2003; *Déruelle et al.*, 2007; *Yokoyama*  
350 *et al.*, 2007]. Such low-volume, high-pressure magmas are expected to form within the

351 sub-continental lithospheric mantle and exhibit relatively little fractionation within the  
352 crust [e.g., *Suh et al.*, 2003].

353 The presence of melt beneath the Moho can sometimes lead to magmatic addition to  
354 the base of the crust (often termed underplate). Underplate has anomalously fast seis-  
355 mic velocities when compared to unaltered crust, but anomalously slow seismic velocities  
356 when compared to the mantle [e.g., *Maguire et al.*, 2006]. Receiver function H- $\kappa$  anal-  
357 ysis is sometimes sensitive to the top, not the bottom, of such layers [e.g., *Stuart et al.*,  
358 2006], which can nevertheless still be detected via crustal receiver function analysis. The  
359 amplitude of the Ps arrival and, in particular, subsequent reverberent phases (PpPs and  
360 PpSs+PsPs) are highly sensitive to the sharpness of the velocity discontinuity that pro-  
361 duces them [e.g., *Zheng et al.*, 2008]. This is illustrated in Figure 8: note the lower  
362 amplitude arrivals in the underplate example, where the transition from  $6.5 \text{ kms}^{-1}$  crust  
363 to  $8 \text{ kms}^{-1}$  mantle occurs gradually, not abruptly.

364 Figure 9 shows a comparison of single station stacks for station CM04 on the Congo  
365 Craton, CM26 on the CVL, and CM29 in the Garoua Rift. In each case, the Ps arrival and  
366 subsequent reverberent phases are sharp, indicating an abrupt transition from relatively  
367 felsic lower crust to peridotitic mantle. Nowhere in the study area do we find evidence  
368 for thick magmatic underplate/lower-crustal intrusions.

369 When considered in light of the petrological literature, our low  $V_p/V_s$  ratio observa-  
370 tions and sharp Moho P-to-S conversions suggest strongly that the protracted magmatic  
371 activity along the CVL has resulted in very little compositional modification of the felsic  
372 Precambrian basement rocks. The CVL's alkali basaltic melts, sourced from sub-crustal

373 depths, are probably therefore transported quickly to the surface such that little shallow  
374 fractionation can occur. High strain rates associated with this rapid movement of magma  
375 are capable of generating seismicity in the lithospheric mantle [e.g., Hawaii: *Okubo and*  
376 *Wolfe*, 2008], where temperatures should be too high for brittle failure to occur [*McKenzie*  
377 *et al.*, 2005]. Accordingly, sub-continental lithospheric mantle ( $\sim 55\text{-}60$  km depth) earth-  
378 quakes have been observed beneath Mount Cameroon prior to its recent eruptions [*Tabod*  
379 *et al.*, 1992; *Ateba and Ntepe*, 1997; *Ateba et al.*, 2009], which may be the result of the  
380 aforementioned high strain rates.

### 4.3. Comparisons with Other Hotspots and Implications for the Upper Mantle Beneath the Cameroon Volcanic Line

381 Comparisons with rifts and other hotspots highlights the uniqueness of bulk crustal  
382 properties of the CVL (Table 3). The average  $V_p/V_s$  ratios of the other hotspots and rifts  
383 are all above the global average of 1.768 [*Christensen*, 1996]. Thus, explanations for CVL  
384 development based on observations at other hotspots worldwide are not straightforward.

385 The case for present-day melt within the crust is well established at many hotspots  
386 and rifts. In Iceland and Hawaii, for example, ongoing volcanic eruptions provide clear  
387 evidence for the melt migration required to maintain melt within the crust for periods of  
388 more than a few thousand years [e.g., *McKenzie*, 1984; *Rose and Brenan*, 2001]. The  
389 Ethiopian rift in East Africa, magmatically active for around the same period of time as  
390 the CVL ( $\sim 30$  Ma), is an excellent venue to compare and contrast styles of magmatism  
391 with those we observe Cameroon. Magnetotelluric studies image present-day melt beneath  
392 the Ethiopian rift [*Whaler and Hautot*, 2006], while geodetic studies provide evidence in

393 support of ongoing crustal melt migration, even in the absence of historical volcanism  
394 [*Biggs et al.*, 2011; *Keir et al.*, 2011]. These crustal properties result in bulk crustal  
395  $V_p/V_s$  ratios in Ethiopia of up to  $\sim 2$ , which is difficult to explain without invoking a  
396 hypothesis of melt within the crust [e.g., *Stuart et al.*, 2006].

397 In addition to present-day melt within the crust in Ethiopia, recent geophysical studies  
398 have found evidence for seismically fast [e.g., *Maguire et al.*, 2006; *Daly et al.*, 2008], dense  
399 [*Cornwell et al.*, 2006; *Mickus et al.*, 2007], cooled gabbroic intrusions that are believed to  
400 have have accommodated  $\sim 80\%$  of extension in the Ethiopian rift since Quaternary times  
401 [e.g., *Ebinger and Casey*, 2001]. Petrological data also indicate extensive fractionation  
402 and modification of the crust and mantle lithosphere beneath the Ethiopian rift [e.g.,  
403 *Rooney et al.*, 2005, 2011]. Further south in the EAR, seismic and gravity data from  
404 Kenya provide evidence for magma intrusion into the Pan African Mozambique belt [e.g.,  
405 *Simiyu and Keller*, 2001]. **Ongoing dike intrusion has been identified geodetically**  
406 **via InSAR study** [*Biggs et al.*, 2009]. **Underplating/magmatic addition to**  
407 **the lower crust has been detected seismically beneath the uplifted Ethiopia**  
408 **Plateau** [e.g., *Maguire et al.*, 2006], as well as in Hawaii [*Darbyshire et al.*,  
409 **2000**] and Iceland [*Leahy et al.*, 2010]. In receiver function data such layers  
410 are manifest as gradational velocity profiles at Moho depths, as illustrated in  
411 **Figure 8, but distinct from that which we observe beneath the CVL (Figure**  
412 **9).**

413 In contrast to the aforementioned hotspots and rifts, the  $V_p/V_s$  ratios observed along  
414 the CVL are low compared to the global average, indicating a bulk crustal composition

415 of mainly granodiorite and granite-gneiss [e.g., *Christensen, 1996*], with no requirement  
416 for present-day melt or significant volumes of intruded mafic material. In support of this,  
417 geophysical evidence for magmatic intrusion along the CVL is limited. While *Poudjom*  
418 *Djomani et al.* [1997], suggest tentatively that gravity data across the Adamawa plateau  
419 can be explained in part by magma crustal intrusions, residual Bouguer gravity maps for  
420 the CVL as a whole [e.g., *Fairhead and Okereke, 1987*], present no evidence for dense  
421 mafic intrusions beneath the CVL, consistent with our hypothesis of a relatively unmodi-  
422 fied CVL crust. **Instead, the principal reason for the plateau uplift is thus more**  
423 **likely a modestly buoyant mantle. Evidence in support of this view includes**  
424 **observations of lower mantle seismic wave-speeds compared to the surround-**  
425 **ing area [e.g., *Dorbath et al., 1986; Reusch et al., 2010*]. Gravity studies**  
426 **also provide evidence for a low density mantle at 70–90 km depth [*Poudjom***  
427 ***Djomani et al., 1992*].**

428 When reviewed in light of the low bulk crustal  $V_p/V_s$  ratios and relatively small volumes  
429 of eruptive products, the lack of lower-crustal intrusions in Cameroon indicate that melt  
430 generation is lower beneath the CVL than other hotspots and rifts worldwide. This may  
431 have the implication that mantle potential temperatures in the region are well below those  
432 at rifts and other hotspots. This is consistent with uppermost mantle P-wave velocities  
433 from previous studies of  $8 \text{ km s}^{-1}$  [*Stuart et al., 1985*], which are high for a volcanic region  
434 when compared to values of  $7.4\text{--}7.5 \text{ km s}^{-1}$  from the Ethiopian Rift [*Bastow and Keir,*  
435 *2011*], where mantle potential temperature is known to be elevated [e.g., *Rooney et al.,*  
436 *2012b*]. Alternatively, the volumes of melt due to adiabatic decompression beneath the

437 CVL [for example, due to shear zone reactivation: e.g., *Fairhead and Binks*, 1991], may  
438 simply be negligible compared to regions such as Ethiopia or Iceland, where extensional  
439 processes likely contribute significantly to melt production by decompression melting [e.g.,  
440 *Maguire et al.*, 2006; *Bastow et al.*, 2010]. Either way, hypotheses for the formation and  
441 development of the CVL should not likely incorporate large volumes of uppermost mantle  
442 melt. Figure 10 illustrates schematically how the melt supply and crustal plumbing system  
443 beneath the CVL compares to that in melt-rich settings such as the Ethiopian rift.

## 5. Conclusions

444 We have conducted a teleseismic receiver function study of crustal structure in  
445 Cameroon to place new constraints on the tectonic subdivisions of the region, and to  
446 determine the impact of Cenozoic volcanism on crustal structure along the Cameroon  
447 volcanic line.

448 Within cratonic Cameroon the edge of the Congo Craton, is characterized by an abrupt  
449 change in crustal thickness of  $\sim 5$  km, which constrains the northern and western edges  
450 of the craton to be  $\sim 4^\circ\text{N}$  and  $\sim 10^\circ\text{E}$  respectively, consistent with studies of the Bouguer  
451 gravity field.

452 Consistently low  $V_p/V_s$  ratios along the CVL (network average is  $\sim 1.74$ ) indicate  
453 strongly that the crust does not presently contain melt; earlier Cenozoic volcanism along  
454 the line has also not resulted in the additional of appreciable volumes of mafic material  
455 to the lower crust, either as gabbroic intrusions or as the fractionated by-product of alkali  
456 volcanism. These observations support previous petrological studies in Cameroon that

457 suggest alkali basaltic magma throughout the CVL has a very short residence time in the  
458 crust prior to eruption.

459 The similarity of P-to-S conversions from beneath the CVL and neighbouring cratonic  
460 regions indicates that lower-crustal intrusions (often termed underplate) are not likely  
461 present beneath the region. When reviewed in light of the low  $V_p/V_s$  ratios across  
462 Cameroon, this indicates that hypotheses for the formation of the CVL should not include  
463 markedly elevated upper-mantle potential temperatures or large volumes of crustal melt,  
464 both of which can explain observations at hotspots and rifts elsewhere. The protracted,  
465 yet sporadic development of small-volume alkali melts beneath the CVL may instead be  
466 explained better by low melt producing mechanisms such as shear zone reactivation or  
467 lithospheric delamination (Figure 10).

## 6. Figure Captions

468 **Table 1:** Age of Volcanism Along the Cameroon Volcanic Line. The timing of the  
469 various periods of volcanism north to south along the CVL as described by *Halliday et al.*  
470 [1988]. All centers of volcanism show morphological evidence for recent cinder cones with  
471 the exception of Principe, Etinde, Bambouto and Mandara Mountains (Figure1).

472 **Table 2:** Bulk crustal properties across the CBSE broadband network.

473 **Table 3:** A global comparison of global hotspot and rift bulk crustal properties.

474 **Figure 1:** Location map of the CBSE seismograph stations (triangles) superimposed  
475 on regional topography. Numbers are station codes. The Ntem Complex boundary and  
476 the Yaoundé domain boundary are after *Toteu et al.* [2004] respectively. Stars are selected  
477 CVL volcanoes. C.A.R.: Central African Republic; CASZ: Central African Shear Zone;



478 E.G.: Equatorial Guinea. The red areas are regions of Cenozoic volcanism along the CVL  
479 taken from *Tokam et al.* [2010].

480 **Figure 2:** The global distribution of all the earthquakes used in the study plotted with  
481 an azimuthal equidistant map projection. Concentric circles indicate  $30^\circ$  intervals from  
482 center of the CBSE network, marked by the triangle. Plate boundaries are black.

483 **Figure 3:** Examples of crustal thickness (H) versus  $V_p/V_s$  plots from the method of  
484 Zhu and Kanamori (2000) and receiver functions for stations CM22 and CM29. On each  
485 plot the arrival time of the Moho phase  $P_s$  ( $t_1$ ) and subsequent reverberent phases  $P_pP_s$   
486 ( $t_2$ ) and  $P_sP_s+P_pP_s$  ( $t_3$ ) (determined by Equations 2-4) are marked based on the crustal  
487 thickness and  $V_p/V_s$  ratios shown in Table 2. The back azimuth for each trace is shown in  
488 the top right. The horizontal and vertical lines on the contour plots mark the maximum  
489 of the stack defined by Equation 6. The white line is the  $2\sigma$  error contour.

490 **Figure 4:** Variations in crustal thickness (Table 2) across Cameroon determined from  
491 receiver function analysis. The thick black lines A-A', B-B', and C-C' show the orientation  
492 of transects in Figure 6.

493 **Figure 5:** Variations in  $V_p/V_s$  ratio across Cameroon determined from receiver func-  
494 tion analysis (Table 2). The thick black lines A-A', B-B', and C-C' show the orientation  
495 of transects in Figure 6.

496 **Figure 6:** Variations in elevation, crustal thickness, and  $V_p/V_s$  ratio (a) NW-SE  
497 across the southern portion of the CVL and into the Congo Craton; (b) NW-SE across  
498 the northern portion of the CVL and the southern tip of the Garoua rift; (c) NE-SW along

499 the CVL. The dashed line is the average crustal  $V_p/V_s$  ratio (1.768). The orientation of  
500 the transects is shown on Figures 4 and 5.

501 **Figure 7:** A comparison of crustal thickness constrained in this study with that of  
502 *Tokam et al.* [2010].

503 **Figure 8:** Forward models of receiver functions with (top) and without (bottom) the  
504 presence of a high velocity lower-crustal layer. Note the sharpness of the Ps phase and  
505 subsequent reverberations when this layer is absent.

506 **Figure 9:** Receiver function stacks for stations (a) CM04 on the Congo Craton, (b)  
507 CM26 in the CVL, and (c) CM29 in the Garoua Rift, (Figure 1). Note the similarity of  
508 shape and amplitude of the Ps and subsequent reverberation phases across the study area.  
509 These observations argue against the presence of high-velocity lower-crustal intrusions  
510 beneath the CVL (Figure 8).

511 **Figure 10:** A summary conceptual cartoon for the magma plumbing system in  
512 Cameroon as compared to that in the magmatically active Ethiopian rift. SCLM: sub-  
513 continental lithospheric mantle. The Ethiopian sketch is modified after *Rooney et al.*  
514 [2011].

515 **Acknowledgments.** The seismograms from the CBSE network were sourced from the  
516 IRIS DMC. Pennsylvania State University deployed and maintained the CBSE network.  
517 We thank Raphael De Plaen, Derek Keir, Jenny Di Leo, Ken Bailey, Luca Caricchi,  
518 Tyrone Rooney, Kathy Cashman and Steve Sparks for helpful discussions. Cindy Ebinger  
519 and Stewart Fishwick provided helpful reviews that improved markedly the focus of the  
520 contribution. IB is funded by the Leverhulme Trust.

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Sample	Age (Ma)
Biu Plateau	<5
Manadara Mountains	30
Adamawa Plateau	8
Oku	22
Bambouto	18
Manengouba	1
Mount Cameroon	0
Etinde	<0.5
Bioko	0
Principe	3.5, 6, 21, 31
Sao Tome	0, 3, 6.4, 7.4
Pagalu (Annobon)	5

**Table 1.** Age of Volcanism Along the Cameroon Volcanic Line. The timing of the various periods of volcanism north to south along the CVL as described by *Halliday et al.* [1988]. All centers of volcanism show morphological evidence for recent cinder cones with the exception of Principe, Etinde, Bambouto and Mandara Mountains (Figure 1)

Stations (Figure 1)	N	$H_{PWS}$ (km) (Figure 4)	$\delta H_{PWS}$ (km)	$\kappa_{PWS}$	$\delta \kappa_{PWS}$	$H_{lin}$ (km)	$\delta H_{lin}$ (km)	$\kappa_{lin}$	$\delta \kappa_{lin}$
CM02	3	44.6	0.6	1.71	0.02	44.2	1.3	1.70	0.04
CM03	5	44.6	0.7	1.65	0.03	44.6	1.0	1.66	0.03
CM04	4	47.5	1.0	1.78	0.05	47.5	1.2	1.76	0.05
CM06	7	41.8	0.7	1.82	0.04	41.0	1.0	1.87	0.05
CM07	12	41.8	0.7	1.73	0.03	41.4	1.1	1.74	0.04
CM10	6	35.8	0.7	1.68	0.04	35.4	1.1	1.72	0.06
CM12	5	37.8	0.7	1.73	0.03	37.4	1.2	1.73	0.05
CM16	4	34.1	0.6	1.67	0.03	34.1	1.2	1.68	0.05
CM17	5	37.8	0.9	1.66	0.04	37.0	1.3	1.69	0.05
CM20	3	33.7	0.6	1.84	0.04	33.7	1.0	1.86	0.06
CM21	6	34.1	0.7	1.76	0.04	34.1	1.0	1.76	0.05
CM22	5	35.8	0.6	1.71	0.03	35.4	1.0	1.76	0.06
CM24	13	38.2	0.7	1.68	0.04	37.8	1.0	1.70	0.05
CM25	4	38.2	0.6	1.74	0.03	37.8	1.1	1.75	0.05
CM26	18	34.9	0.7	1.72	0.04	34.9	1.1	1.71	0.06
CM29	6	25.3	0.8	1.84	0.07	25.3	1.2	1.82	0.09
CM30	4	26.5	0.9	1.75	0.06	26.5	1.2	1.73	0.06
CM32	4	33.3	1.1	1.80	0.08	33.7	1.0	1.74	0.06

Ps picked method

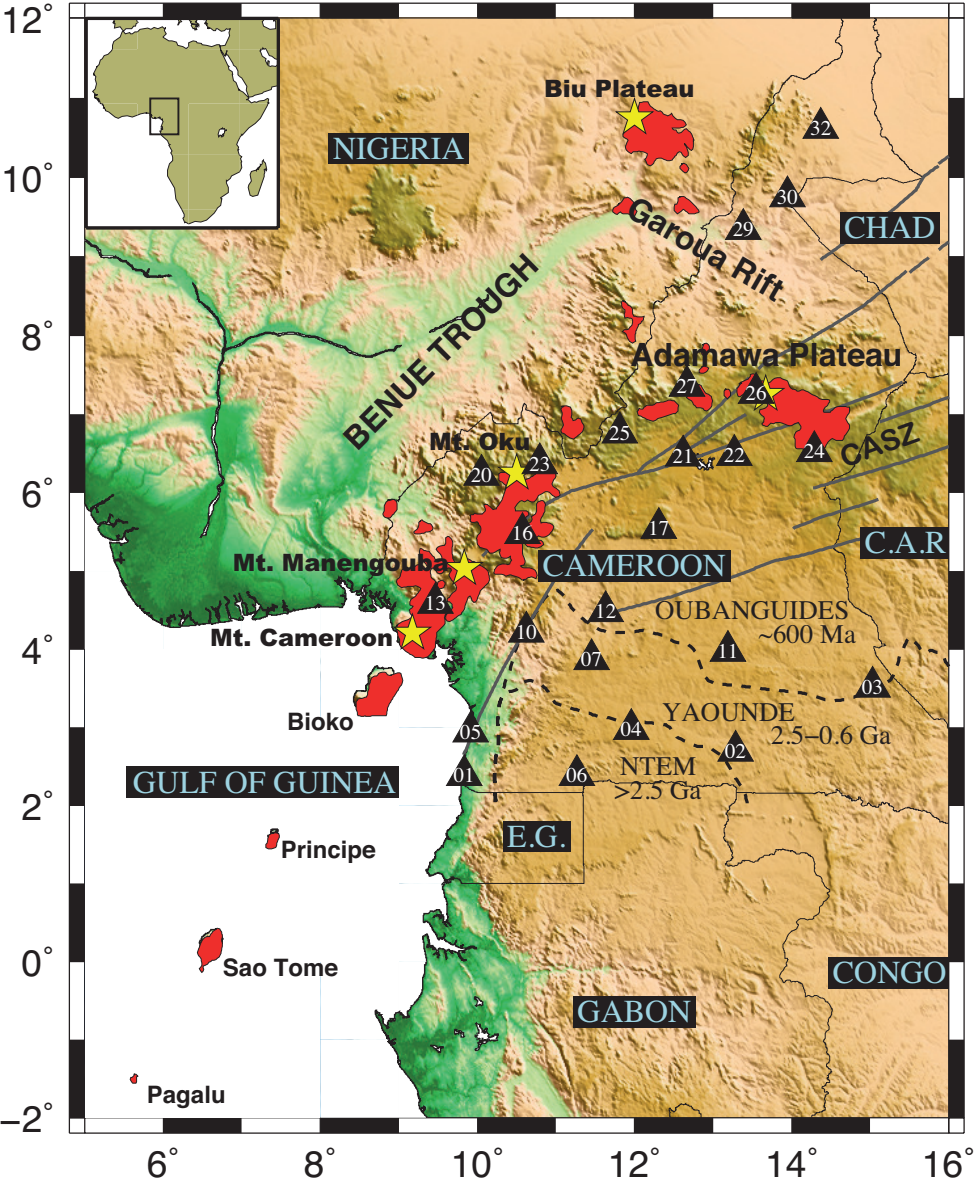
Station	N	H (km)	$\kappa$
CM01	6	<b>37.6</b>	<b>1.82</b>
<b>CM05</b>	<b>4</b>	<b>37.4</b>	<b>1.82</b>
CM11	7	43.3	1.71
CM13	3	36.5	1.68
CM23	10	38.3	1.84
CM27	4	39.0	1.72

**Table 2.** Bulk crustal properties across the CBSE broadband network.

Hotspot	Vp/Vs	H (km)	Reference
CVL	1.74	~35	This study.
Hawaii	1.80	10–20	<i>Leahy and Park</i> [2005].
Iceland	1.78	20–30	<i>Darbyshire et al.</i> [1998]; <i>Bjarnason and Schmeling</i> [2009].
Ethiopian Rift	2.00	30–40	<i>Stuart et al.</i> [2006].
Azores	1.80–1.90	20–30	<i>Silveira et al.</i> [2010].
Yellowstone	1.76–1.87	38–54	<i>Schutt et al.</i> [2008]; <i>Stachnik et al.</i> [2008].
S. EAR	1.84	38–42	<i>Dugda et al.</i> [2005].
Baikal	1.82	45	<i>Chen</i> [2002].
Rio Grande	1.78	35	<i>Wilson et al.</i> [2005].

**Table 3.** A global comparison of global hotspot and rift bulk crustal properties.

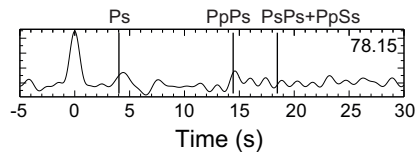
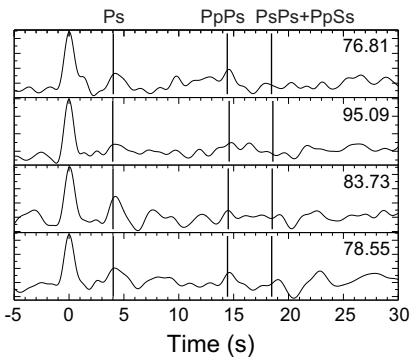
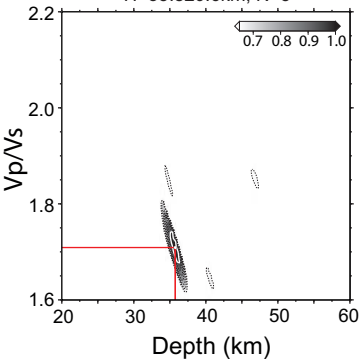






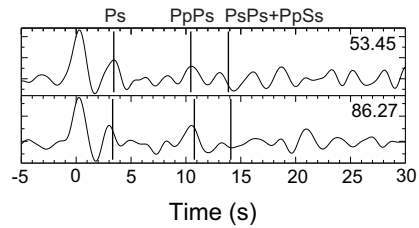
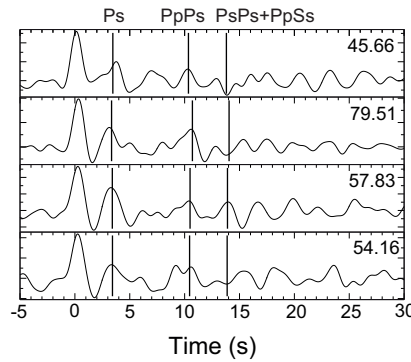
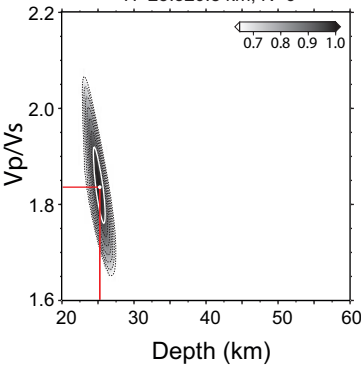
CM22,  $V_p=6.5\text{km/s}$ ,  $V_p/V_s=1.71\pm 0.03$

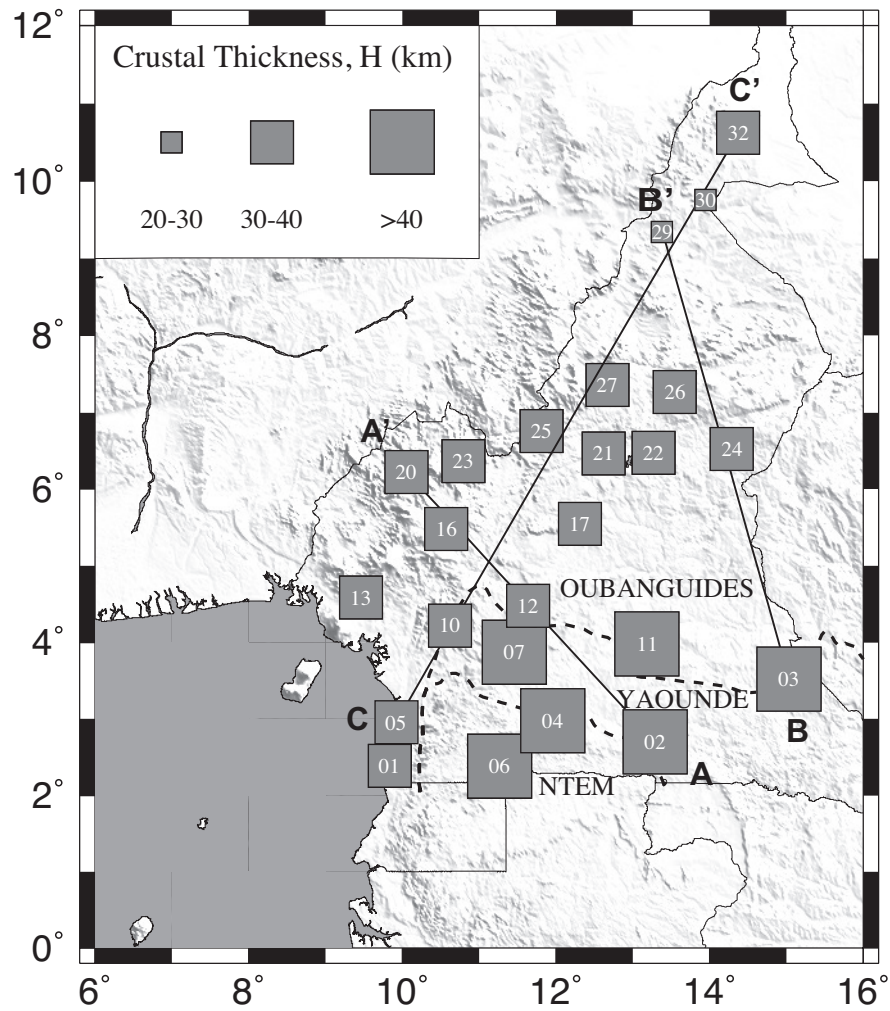
$H=35.8\pm 0.6\text{km}$ ,  $N=5$

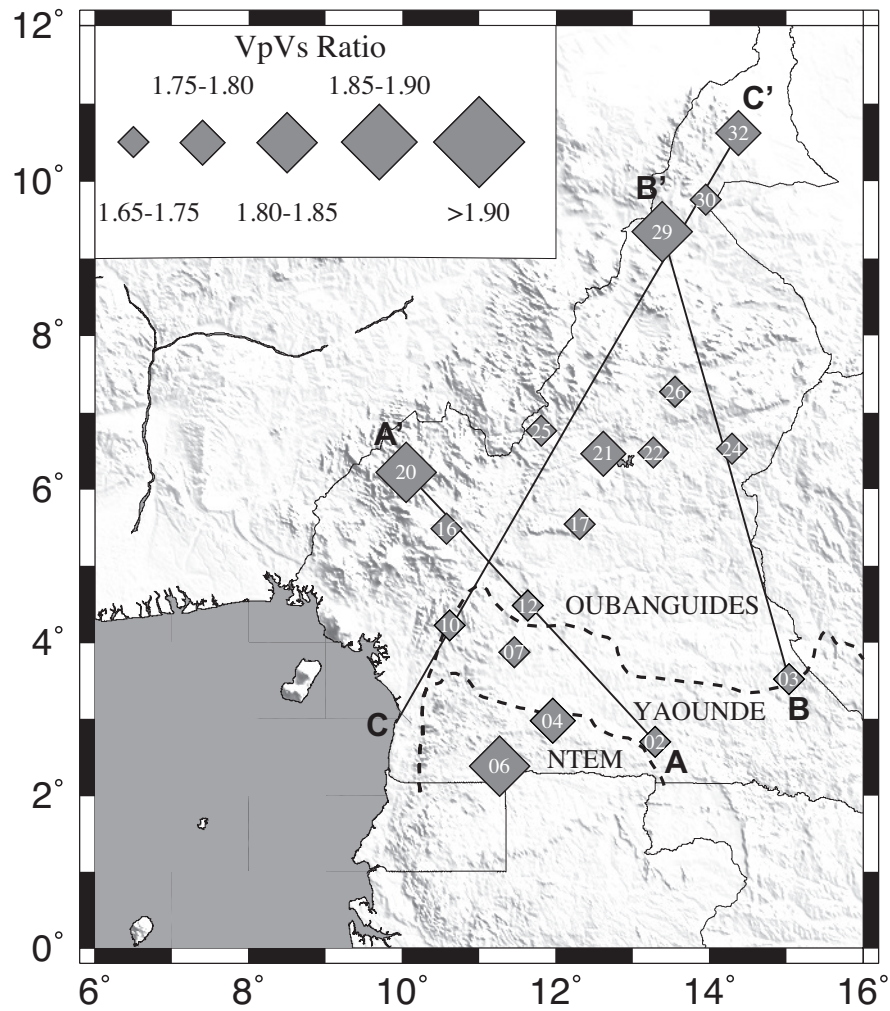


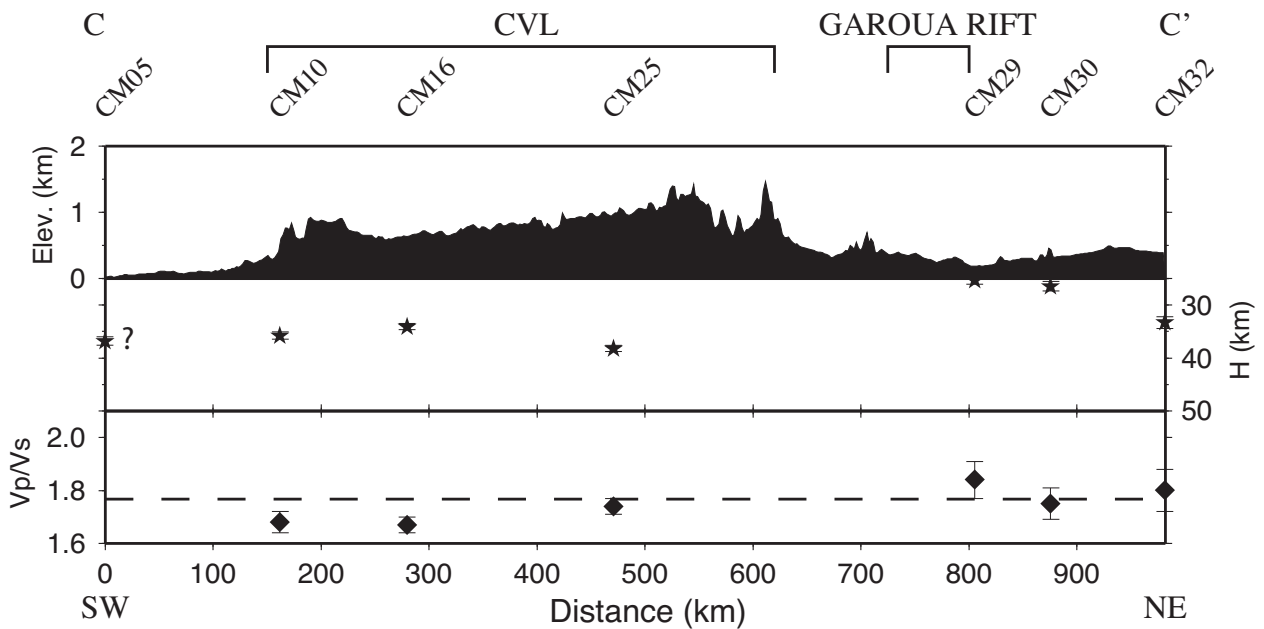
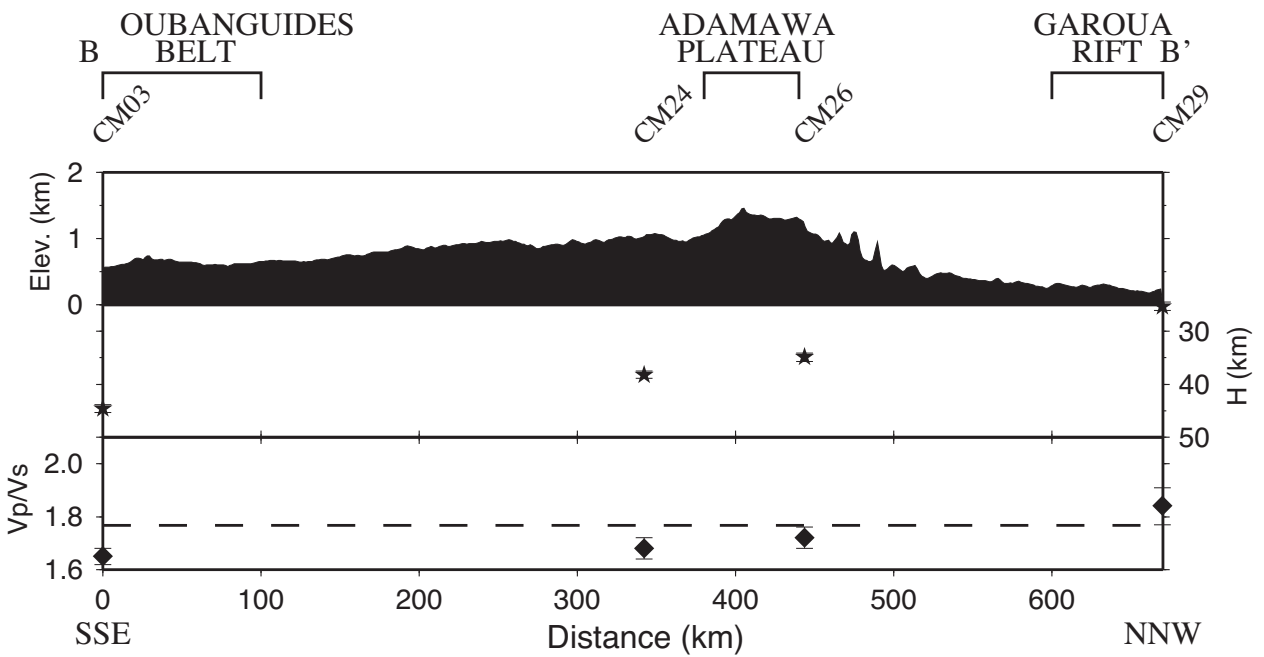
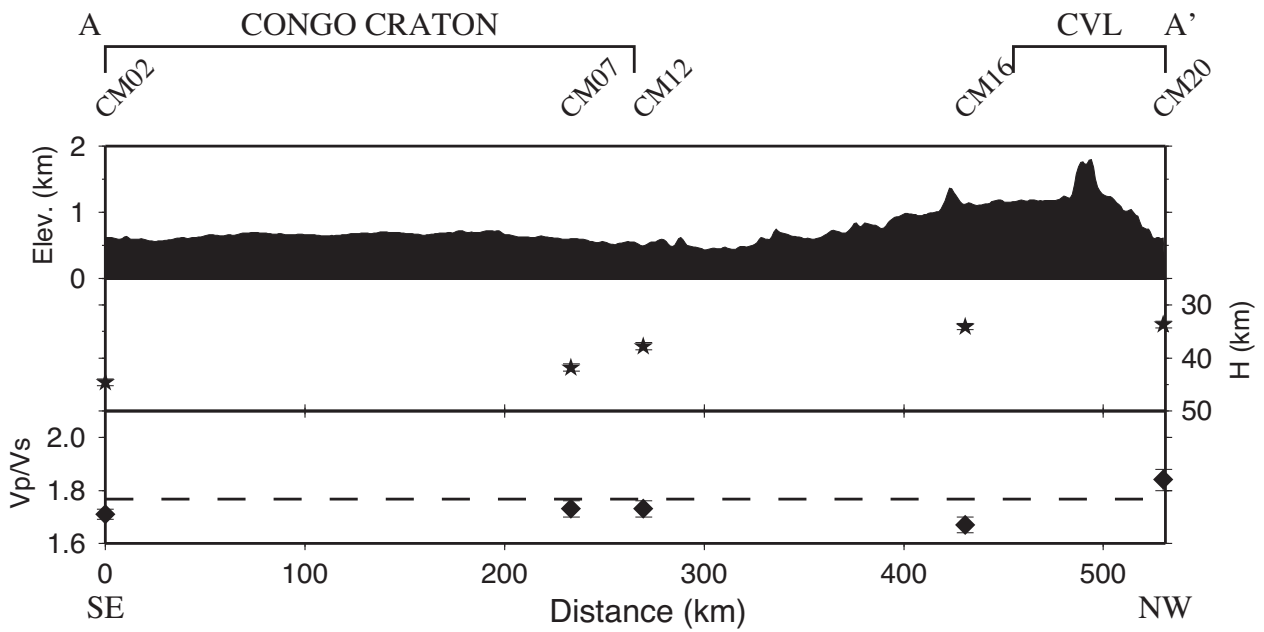
CM29,  $V_p=6.5\text{ km/s}$ ,  $V_p/V_s=1.84\pm 0.08$

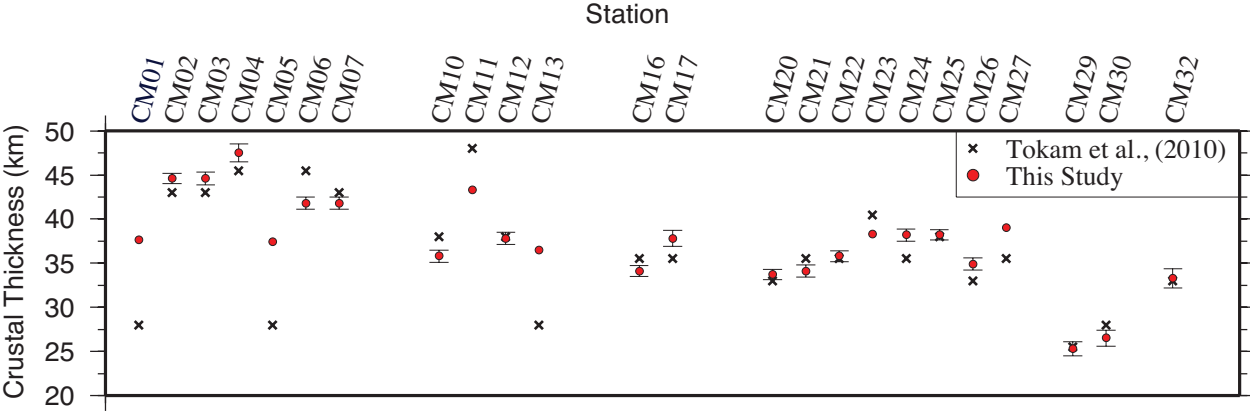
$H=25.3\pm 0.8\text{ km}$ ,  $N=6$

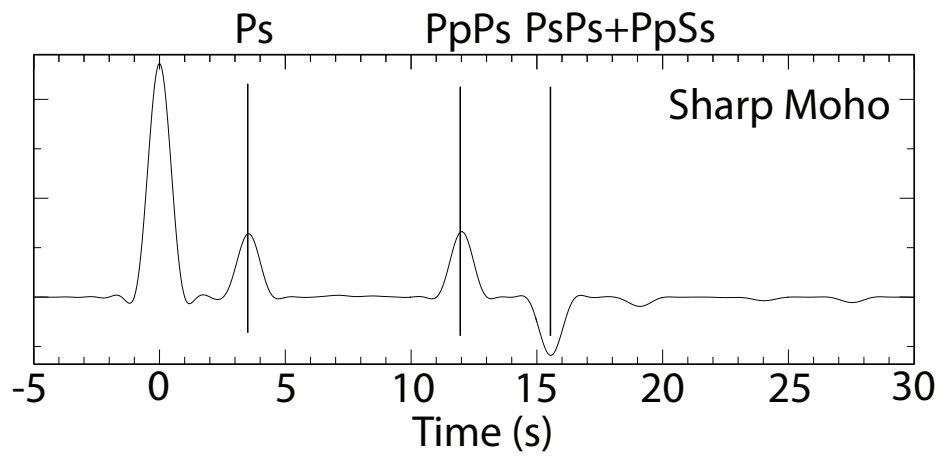
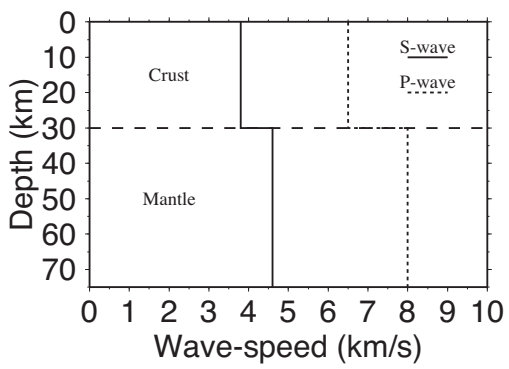
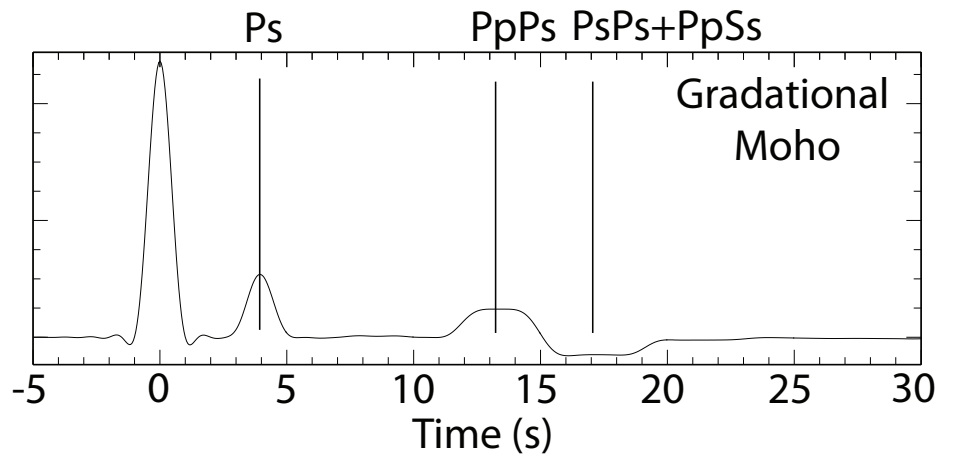
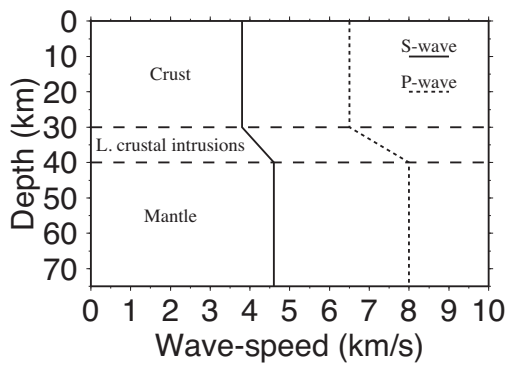




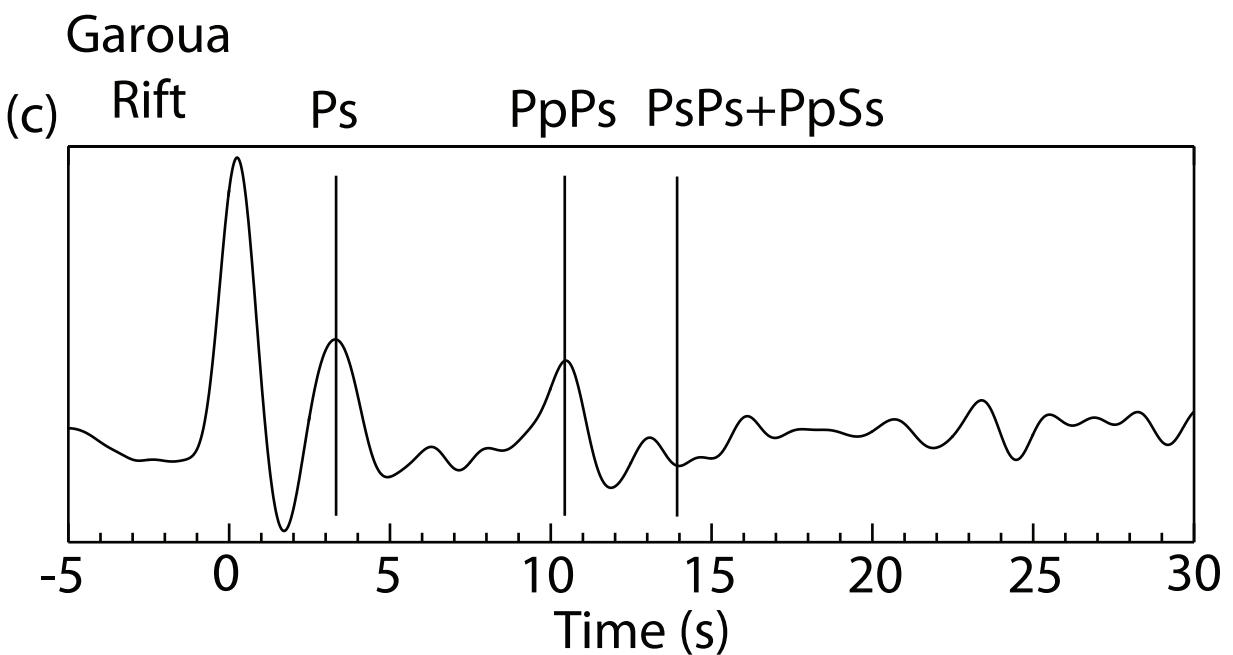
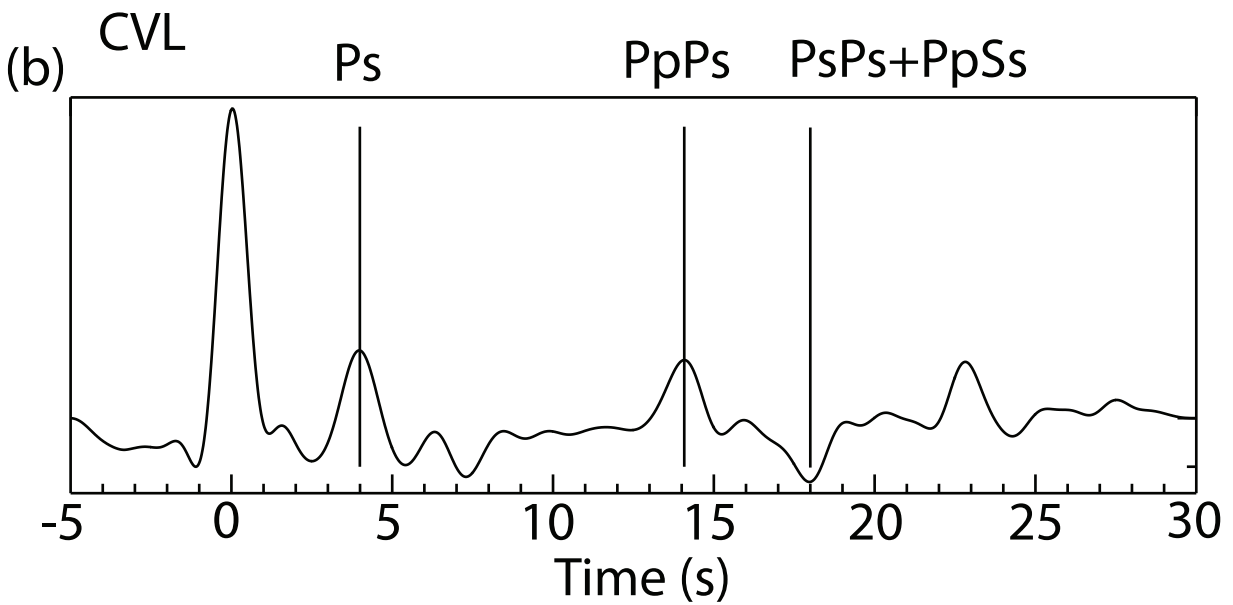
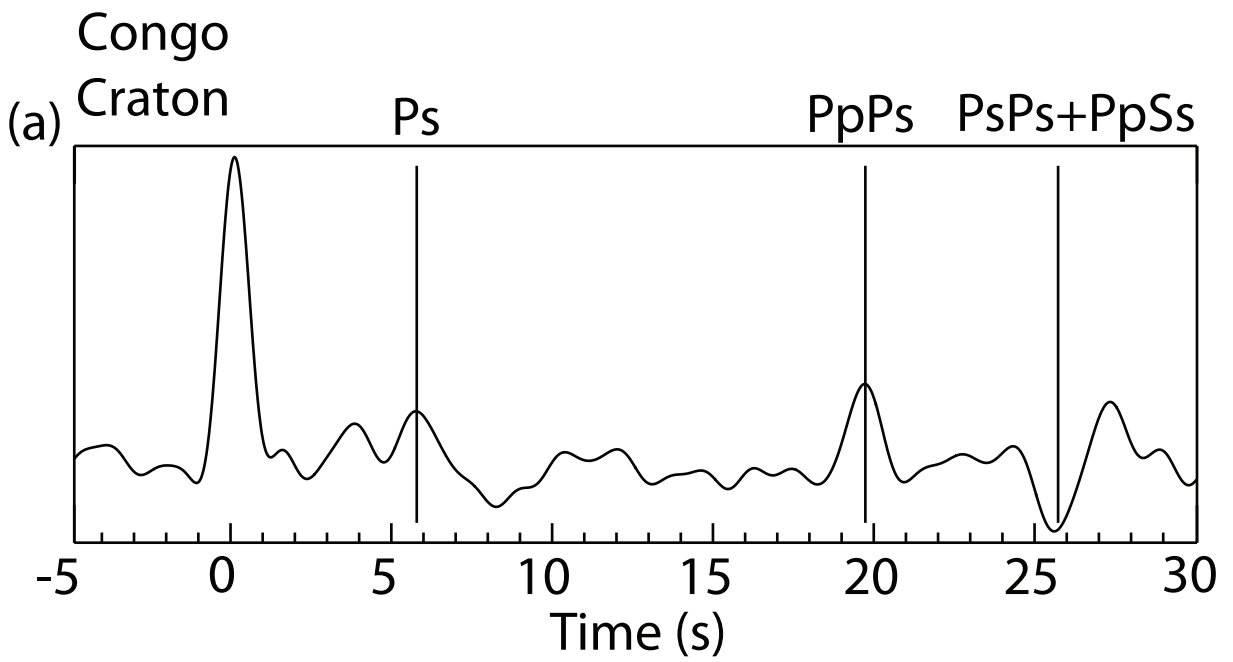


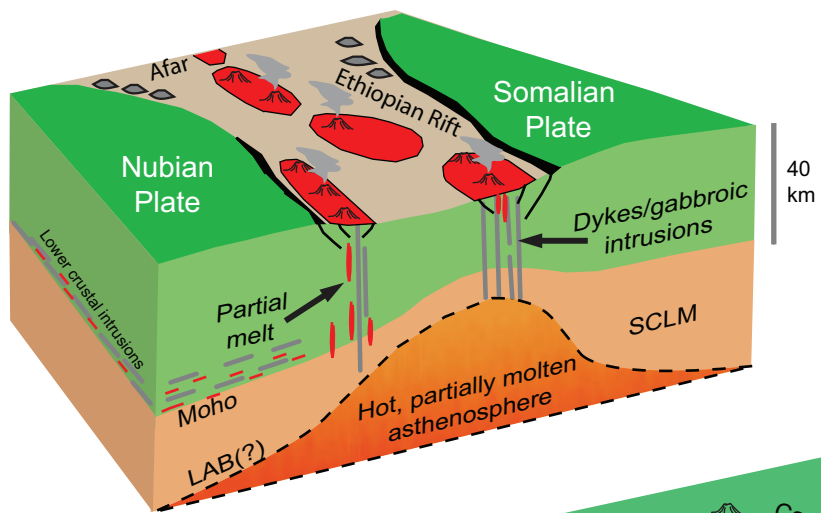












Ethiopia

Cameroon

