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RESEARCH ARTICLE

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Kev Points:

- New proxy to measure average settling velocities using ²¹⁰Po-²¹⁰Pb profiles
- Suggest clear shallow and mesopelagic contribution of slow-sinking particles
- Depth increase in velocities suggests slow particles remineralization up to 400 m

Supporting Information:

- Readme
- Figure S1a
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- Figure S4
- Table S1Table S2

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Observations and modeling of slow-sinking particles in the twilight zone

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Abstract The biological carbon pump (BCP) transfers carbon from the surface ocean into the oceans' interior, mainly in the form of sinking particles with an organic component, and thereby keeps atmospheric CO₂ at significantly lower levels than if the oceans were abiotic. The depth at which these sinking particles are remineralized is a key control over atmospheric CO₂. Particle sinking speed is likely to be a critical parameter over remineralization depth. Carbon export is usually controlled by large, rapidly sinking particles (>150 m·d⁻¹); however, under some circumstances sinking velocity distributions are strongly bimodal with a significant fraction of total flux being carried by slowly (<10 m·d⁻¹) sinking particles. Therefore, there is an interest in determining sinking particle velocities and their variations with depth, as well as in understanding the interplay between sinking velocity distributions and carbon export. Here, we use profiles of total and particulate concentrations of the naturally occurring radionuclide pair ²¹⁰Po-²¹⁰Pb from the Porcupine Abyssal Plain (PAP) site (48°N, 16.5°W) to estimate depth variation in particle sinking speed using a one-box model and inverse techniques. Average sinking speeds increase from $60 \pm 30 \,\mathrm{m\cdot d^{-1}}$ at 50 m, to $75 \pm 25 \,\mathrm{m\cdot d^{-1}}$ and $90 \pm 20 \,\mathrm{m\cdot d^{-1}}$ at 150 and 500 m. Furthermore, a sensitivity analysis suggests that at the PAP site the measured ²¹⁰Po profiles are inconsistent with the usually assumed sinking velocities of 200 m·d⁻¹. We hypothesize that a trend of increasing velocity with depth might be caused by a gradual loss of slow-sinking material with depth, a factor with significant implications for regional carbon budgets.

1. Introduction

The downward flux of particles from the surface ocean consists of a complex mix of biogeochemical material including particulate organic carbon (POC) and particulate inorganic "ballast" material such as opal, calcite, and lithogenic material. POC is of particular importance for organisms living in the dark ocean, as it is their main source of food. More than 90% of the POC exported from surface waters is respired to CO_2 by the biota in the "twilight zone" (the region between the euphotic zone and 1000 m depth) [e.g., *Robinson et al.* [2010]]. This respiration limits the depth to which particles are being exported, with only 0.01–3% of the POC produced in the surface waters being stored in deep-sea sediments, e.g., *Falkowski et al.* [1998].

The depth at which POC is respired is, according to models, an important control over atmospheric CO₂ levels [Kwon et al., 2009; De La Rocha, 2003]. Yet, the mechanisms controlling this "remineralization depth" are unclear. Remineralization has been extensively studied in the last years (e.g., reviews by De La Rocha and Passow [2007], Burd et al. [2010], Sanders et al. [2014], and Turner [2014]) and is found to depend strongly on physical and geochemical factors as well as on the phytoplankton community in the site [Buesseler et al., 2007; Henson et al., 2012].

A factor likely involved in controlling remineralization depth is particle sinking speed. Slow-sinking particles $(<10\,\mathrm{m\cdot d^{-1}})$ are more likely to be recycled at shallower depths, whereas fast-sinking particles are likely to reach greater depths before the organic carbon they contain is respired. The size and composition of sinking particles can change significantly due to repackaging by organisms resident in the oceans interior (e.g., Lee et al. [2009] and Wilson et al. [2008]), potentially leading to changes in sinking speed with depth. However, particle sinking speed is a difficult parameter to measure, and few studies have measured it.

Direct measurements have been done either in situ by using under-water cameras [Sternberg et al., 1999] or after recovery in settling cylinders [Alldredge and Gotschalk, 1989; Riley et al., 2012]. McDonnell and Buesseler

Table 1. Station ID, Sampling Time, and Position Taken for ²¹⁰Po/²¹⁰Pb During the D341 Station ID Date Latitude Longitude 16477 13/07/2009 49.02°N 16.47°W 16497 16/07/2009 48.97°N 16.40°W 16523 20/07/2009 48.54°N 17.10°W 16544 23/07/2009 49.07°N 16.63°W 16583 29/07/2009 48.98°N 16.93°W 31/07/2009 16592 48.81°N 16.51°W 16619 03/08/2009 48.97°N 16.40°W 16640 06/08/2009 48.91°N 16.54°W 16659 08/08/2009 48.78°N 16.99°W

[2010] used in situ imaging of particles in viscous polyacrylamide gels to estimate the average sinking velocity of particles with equivalent spherical diameter between 45 µm and 10 mm as 10–150 m·d⁻¹. Xue and Armstrong [2009] developed a technique to determine settling velocities using the "benchmark" method, based on fitting Fourier series to time

series data from sediment traps, which estimated sinking speeds of $220\pm65\,\mathrm{m\cdot d}^{-1}$ in the northwestern Mediterranean Sea.

Another direct technique used to estimate particle sinking speeds is the Indented Rotating Sphere (IRS) sediment trap [*Peterson et al.*, 2005, 1993; *Trull et al.*, 2008]. A recent synthesis of IRS results from a variety of settings [*Alonso-Gonzalez et al.*, 2010] suggests that two scenarios occur: either most POC flux is carried by very fast settling particles (velocities >300 m·d⁻¹), or the sinking speed distribution is bimodal with a significant fraction (up to 62%) of POC flux being associated with very slowly sinking material (<10 m·d⁻¹). However, aggregation of particles on the sphere prior to rotation may alter particle characteristics and reduce their sinking speeds [*Peterson et al.*, 2009]. Sinking speeds diagnosed from IRS traps therefore need to be confirmed by alternative methodologies [*Trull et al.*, 2008], such as locating cameras above and below the IRS [*Peterson et al.*, 2009].

Here we use an alternative method to track particle flux and sinking velocity: the disequilibrium between polonium-210 and its parent lead-210 (the ²¹⁰Po-²¹⁰Pb pair) [*Le Moigne et al.*, 2013b]. This approach is analogous to the widely applied ²³⁴Th-method [*Stewart et al.*, 2010; *Verdeny et al.*, 2009] but has the advantage that the longer half-life of ²¹⁰Po (138 days vs 24 days for ²³⁴Th) allows the disequilibria to penetrate deeper into the twilight zone compared to the ²³⁴Th disequilibria [*Le Moigne et al.*, 2013b]. Furthermore, ²¹⁰Po is not only adsorbed to particles, but also assimilated by some phytoplankton and bacteria [*Fisher et al.*, 1983] and substitutes for S in proteins [*Stewart et al.*, 2007a]. Due to this strong biological affinity ²¹⁰Po is expected to trace POC export effectively.

In this paper, we use the observed profiles of ²¹⁰Po and ²¹⁰Pb to calculate particle-settling speeds throughout the water column. A one-box model with multiple vertical layers is built and fitted to ²¹⁰Po and ²¹⁰Pb profiles from the Porcupine Abyssal Plain (PAP) site (48°N, 16.5°W) in July 2009 [*Le Moigne et al.*, 2013b] to diagnose sinking velocity. Our calculations are validated by applying the same methodology to the ²³⁴Th-²³⁸U profiles measured in the PAP site in order to obtain alternative particle-settling speeds as well as by comparison to measurements made using the Marine Snow Catcher [*Riley et al.*, 2012]. Finally, the limitations of the approach are analyzed. We discuss the potential contribution of slow-sinking particles (<10 m·d⁻¹) to the downward flux as well as the dependence of this flux with depth and its penetration into the twilight zone.

2. Methods

2.1. Sample Collection

Samples for the analysis of ²¹⁰Po and ²¹⁰Pb were collected in July/August 2009 at 10 stations at the PAP site (Table 1) during cruise D341 onboard R.R.S *Discovery*. The PAP site is located on the boundary between the subpolar and subtropical gyres of the North Atlantic [*Henson et al.*, 2009]. Water was collected using a 24-bottle stainless steel sampling rosette equipped with 20-L Niskin® bottles, a CTD sensor, and a fluorimeter.

For total 210 Po and 210 Pb analysis, samples were taken at 10–15 depths from 5 to ~1000 m, with the highest resolution in the top 150 m of the water column. Five-liter samples were collected at each depth in acid-and MilliQ-cleaned carboys that were acid cleaned and pre-rinsed with the sample. Duplicate samples were taken at three different depths over the cruise, to check the reproducibility of the procedure.



Reproducibility was within the uncertainty of the measurements (~8% for ²¹⁰Pb and ~13% for ²¹⁰Po) with average values being 9% for ²¹⁰Pb and 15% for ²¹⁰Po.

Total ²¹⁰Po and ²¹⁰Pb analyses followed the GEOTRACES protocol [GEOTRACES Standards and Intercalibration Committee, 2010; Rigaud et al., 2013] which involves scavenging using Fe(OH)₃ followed by ²¹⁰Po selfdeposition onto a silver disk to give uncorrected ²¹⁰Po activity. ²⁰⁹Po and stable lead were added as yield tracers. Additional plating was performed to remove any traces of Po and to ensure that only ²¹⁰Pb remained in solution with less than 5% of ²⁰⁹Po being found on the second plating disk. A second ²⁰⁹Po spike was added at this point to monitor recovery after Po ingrowth. Since some of the initial ²⁰⁹Po tracer could have remained in the solution, this possibility was tested, and less than 5% of the initial tracer was found in the cleaning disk. Nevertheless, to eliminate overestimation of the radiochemical yield in the measurement of ²¹⁰Pb, when the solution was ²⁰⁹Po-spiked a second time the amount was an order of magnitude higher than the amount added on board. Samples were stored to allow ²¹⁰Po in-growth via ²¹⁰Pb decay to occur. After at least 6 months, ²¹⁰Po and ²⁰⁹Po were plated, and ²¹⁰Po activity was measured. ²¹⁰Pb activity was inferred from the second ²¹⁰Po analysis after the appropriate decay corrections were applied. Recoveries for Pb were obtained by ICP-MS measurement. Once ²¹⁰Pb had been determined, the initial measurement of ²¹⁰Po had to be corrected due to the decay of ²¹⁰Pb into additional ²¹⁰Po prior to ²¹⁰Po-²¹⁰Pb separations. ²¹⁰Po measurements were also decay corrected from the measurement to plating, as well as from plating to sampling. ²¹⁰Bi and ²¹⁰Po in-growth were both considered in the ²¹⁰Pb decay corrections. Contamination due to ²¹⁰Pb-²¹⁰Po present in the lead carrier was subtracted. The activity of the internal tracer ²⁰⁹Po was decay corrected from the reference date to the measurement date. Counting was undertaken using a PIPS-type alpha detector (Canberra) at CITIUS (Centro de Investigación, Tecnología e Innovación, Universidad de Sevilla), with a full background correction. Uncertainties were obtained by error propagation.

Particles were collected by filtering large volumes ($1000-2500\,L$) of seawater using in situ pumps (Stand Alone Pumping Systems—SAPS, Challenger Oceanic °) through 53 and 1 μ m pore size nylon mesh. The particles were rinsed off the meshes and split using a Folsom Splitter. Subsamples of the 53 (large) and 1 μ m (small) fractions were filtered onto glass fibre filters (GF/F 0.7 μ m pore size, Whatman) for subsequent 210 Po and 210 Pb analysis. Filters were fully digested in a mix of HNO₃, hydrochloric acid, and hydrofluoric acids. After evaporation, samples were redissolved in concentrated hydrochloric acid and then analyzed following the procedure documented above for water samples. Additional details of the radiochemical pre-treatment for water and particles can be found in *Le Moigne et al.* [2013b].

2.2. Model Development

A variety of approaches have been used to model, analyze, and interpret the observed activity of thorium in seawater [Savoye et al., 2006]. Explicit modeling of thorium geochemistry has been used to simulate thorium distributions and thus to diagnose fluxes, as well as to model, simulate, and test sampling strategies [Burd et al., 2000; Resplandy et al., 2012]. In contrast, inverse modeling techniques use the comparison between model results and existing data to tune model parameter values until model results give the best agreement with the data. Marchal and Lam [2012], Murnane and Cochran [1991], and Murnane et al. [1994, 1996] developed inverse techniques in a three-box export model to obtain rate constants describing thorium and particle cycling in the ocean.

Efforts to understand ²¹⁰Po particle dynamics via modeling are relatively rare. We focus here on a simple export model of ²¹⁰Po that combines particulate and dissolved polonium into a single box, based on the equations used to calculate ²³⁴Th or ²¹⁰Po fluxes [*Buesseler et al.*, 2001; *Verdeny et al.*, 2009]. The model explicitly incorporates particle-sinking speed as an unknown. Potential sinking speeds are derived by an inverse modeling approach that fits the observed profiles to the model whilst assuming steady state conditions. The systematic uncertainty of the latter assumption will be examined.

We assume that time variation in 210 Po activity (a_{210p_0}) at a given depth z is due entirely to the radioactive decay of 210 Po, the decay of 210 Pb into 210 Po, and the sinking of particles which carry 210 Po to the next depth,

$$\frac{da_{210p_o}(t)}{dt} = \lambda_{210p_o}(a_{210p_b} - a_{210p_o}) + \frac{dP_{210p_o}(z, t)}{dz}$$
(1)



where $a_{210p_b}(z)$ and $a_{210p_o}(z)$ are, respectively, ²¹⁰Pb and ²¹⁰Po total specific activities (e.g., in dpm·100 L⁻¹), λ_{210p_o} is the ²¹⁰Po decay constant (0.005 d⁻¹), and P_{210p_o} is downward ²¹⁰Po flux.

The simplest case of steady state conditions is assumed, where there is no time variation in activities or 210 Po flux over the half life of 210 Po (138 days). Thus, the total activity of 210 Pb and its daughter 210 Po can only differ from each other due to a downward flux of 210 Po, with the extent of this downward flux governing the size of the offset. The key point is that this assumption allows 210 Po flux to be written in terms of 210 Po activity in particles and their average bulk sinking speed v(z)

$$P_{210P_0}(z,t) = a_{210P_0}^{particle}(z) \cdot v(z) = a_{210P_0} \frac{a_{210P_0}^{particle}(z)}{a_{210P_0}} \cdot v(z) = a_{210P_0} \cdot f(z) \cdot v(z)$$
(2)

where f(z) is the fraction of ²¹⁰Po in particles:

$$f(z) = \frac{a_{210P_0}^{particle}(z)}{a_{210P_0}} \tag{3}$$

 $a_{210p_0}^{particle}$ is the ²¹⁰Po measured in particles, comprising the sum of the small (1–53 µm) and large (>53 µm) fractions. a_{210p_0} is the total ²¹⁰Po measured in seawater, including both particulate and dissolved ²¹⁰Po. Combining and rearranging equations (1) and (2) allows the size of the ²¹⁰Po-²¹⁰Pb disequilibrium (and hence the downward ²¹⁰Po flux) to be expressed in terms of sinking speed (equation (4)).

$$a_{210Pb} - a_{210Po} = -\frac{1}{\lambda_{210Po}} \frac{d(a_{210Po} \cdot (z) \cdot v(z) \cdot f(z))}{dz}$$
(4)

This equation is traditionally used in the literature to calculate ²³⁴Th and ²¹⁰Po fluxes [*Buesseler et al.*, 2001; *Cochran and Masque*, 2003; *Savoye et al.*, 2006; *Verdeny et al.*, 2009]; here it is written in terms of speed, allowing us to obtain particle sinking speed throughout the water column.

Rearranging equation (4) allows 210 Pb and 210 Po activities to be related to δ (z) which has dimensions of depth (in m) and which will be used to calculate sinking velocities:

$$\delta(z) = \frac{f(z) \cdot v(z)}{\lambda_{210_{P_0}}} \tag{5}$$

The traditional approach involves computing the downward Po flux using equation (1) and measured ^{210}Po and ^{210}Pb activities and $\delta(z)$ is not explicitly calculated. One of the aims of this work is to check the ability of equation (1) to describe real ^{210}Po profiles. Thus, inverse techniques are used here, and Po flux is not calculated. Instead, ^{210}Po activity is estimated, leaving $\delta(z)$ the only parameter to be tuned. To solve equation (4) ^{210}Po activity will be obtained from tuning the parameter $\delta(z)$ and from direct field measurements of $a_{210\text{ph}}(z)$.

2.3. Inverse Method

This section describes how ²¹⁰Po is estimated from observed ²¹⁰Pb specific activities using an inverse model. First, since our measurements are at discrete depths the finite difference is substituted for the derivative. Equation (4) is thus transformed into equation (6)

$$a_{210Pb} - a_{210Po} = -\frac{1}{h} \cdot \frac{(a_{210Po} \cdot (z+h) \cdot f(z+h) \cdot v(z+h)) - (a_{210Po} \cdot (z) \cdot f(z) \cdot v(z))}{\lambda_{210Po}}$$
(6)

where h is the depth between sampling points. Equation (6) allows us to calculate theoretical ²¹⁰Po specific activity from observed ²¹⁰Pb specific activity through the water column.

Equation (7) is obtained by writing equation (6) in terms of $\delta(x)$.

$$a_{210Po}^{MODELLED}(z) = a_{210Pb}^{MEASURED}(z) + \frac{\delta(z+h) \cdot a_{210Po}^{MODELLED}(z+h) - \delta(z) \cdot a_{210Po}^{MODELLED}(z)}{h}$$
(7)

Finally, modeled ²¹⁰Po at the next depth is calculated by solving the previous equation as follows:

$$a_{210p_0}^{MODELLED}(z+h) = \frac{1}{\delta(z+h)} \left(\delta(z) \cdot a_{210p_0}^{MODELLED}(z) - h \cdot \left(a_{210p_0}^{MEASURED}(z) - a_{210p_0}^{MODELLED}(z) \right) \right) \tag{8}$$

Equation (8) can be used to estimate 210 Po at any depth (z+h) from the measured 210 Pb activity at the previous depth, z, using modeled 210 Po activity calculated at z. The solution of the differential equation, either numerically or analytically, requires choosing an initial condition a_{210} Po(z_0) at some



depth. The initial value of 210 Po is, in most cases, chosen to be the value measured at the shallowest sampled depth, $a_{210Po}^{MEASURED}(z_0)$.

We tune the value of $\delta(z)$ and fit the model to measured ²¹⁰Po specific activity using a simplified least squares method to minimize the reduced chi squared, χ^2_{red} , defined as:

$$\chi_{red}^{2} = \frac{1}{n} \sum \left(\frac{a_{Po}^{MOD}(z) - a_{Po}^{EXP}(z)}{\Delta a_{Po}^{EXP}(z)} \right)^{2}$$
 (9)

where n is the number of degrees of freedom.

2.4. ²¹⁰Po Deficit as a Proxy to Calculate Sinking Speed

The tuning of $\delta(z)$ reveals potential changes in particle sinking speed with depth, which will be explored in subsequent sections. The average particle sinking speed at a given depth is obtained by evaluating v(z) using equations (3), (5), and (6);

$$v(z) = \frac{\delta(z) \cdot \lambda_{210_{P_0}} \cdot a_{210_{P_0}}(z)}{a_{210_{P_0}}^{particle}(z)} = \frac{\delta(z) \cdot \lambda_{210_{P_0}}}{f(z)}$$
(10)

where $\lambda_{210_{Po}}$ is the decay constant (0.005 d⁻¹). It follows that particle sinking speeds at a particular depth can be calculated if the fraction of ²¹⁰Po in particles (f(z)), the decay constant, and $\delta(z)$ are known. $\delta(z)$ was taken as the value that produced the best fit between model and observations (see above). f(z) was calculated from total ²¹⁰Po measured in the water samples and the particles collected from the in situ pumps (calculated as the sum of the two particle-size classes, 1–53 μ m and >53 μ m) (Table 3).

3. Results

3.1. Biogeochemical and Physical Conditions at PAP Site

At the Porcupine Abyssal Plain (PAP) site the annually integrated primary productivity is ~200 g C·m $^{-2}$ ·yr $^{-1}$ [Lampitt et al., 2010]. During cruise D341 (Figures S1a and S1b), chlorophyll-a concentrations derived from satellite data (MODIS) displayed two peaks prior to the cruise, one in mid-May (0.9 mg chl-a m $^{-3}$) followed by a second peak in mid-June (0.6 mg chl-a·m $^{-3}$). Primary productivity calculated from satellite data (PP) presented a typical pattern of increasing PP over the productive season reaching a peak in late June (2000 mg C·m $^{-2}$ ·d $^{-1}$) [Le Moigne et al., 2013b]. PIC concentrations showed a peak in mid May (6×10 $^{-4}$ mol·m $^{-3}$), which matched the Chl-a trend. The depth of the mixed layer (MLD) in winter is typically ~350 m at this site [Steinhoff et al., 2010]. That summer mixed layer depth (MLD) ranged between 31 (St. 16640) and 62 m (St. 16659), averaging 44±11 m (determined by temperature difference, see temperature and salinity plots, Figure S2).

The T-S diagrams (Figure S3) show that there is a distinct mixed layer. The water mass between MLD and ~600 m at all stations was East North Atlantic Central Water of polar origin (ENACWp), though some influence of the Mediterranean Overflow Water (MOW) could be seen below 600 m [Hartman et al., 2010].

The phytoplankton community was composed of a mixture of silicifying and calcifying phytoplankton [Le Moigne et al., 2013b]. During D341 cruise Riley et al. [2012] collected 459 particles containing organic matter using a marine snow catcher. Two categories were observed, Marine Snow Aggregates (MSA, 429 particles) and particles consisting of a distinct solid centre of biomineralizing protist enveloped in marine snow (APC, 30 particles). The sinking speed of the fast pool was calculated as a weighted average of the MSA ($180\pm20\,\mathrm{m\cdot d^{-1}}$) and the APC ($232\pm58\,\mathrm{m\cdot d^{-1}}$). The average sinking speed of the slow settling fraction was estimated by Riley et al. [2012] to be 9 m·d⁻¹. Previous cruises at the site revealed that the microplankton ($5-200\,\mu\mathrm{m}$) community in the area comprised diatoms, dinoflagellates, planktonic ciliates and nanoflagellates [Smythe-Wright et al., 2010], and radiolarians [Lampitt et al., 2009].

Deep sediment traps at the PAP site have shown strong regional and seasonal variations in downward particle flux at depths from 1000 to 4700 m [*Lampitt et al.*, 2009], and very large pulses of particulate organic matter intermittently sink into the deep waters of the open ocean in the Northeast Atlantic. These variations have been attributed to changes in upper-ocean biogeochemistry and phytoplankton community [*Lampitt et al.*, 2009] and to spatial-temporal variability in surface waters [*Smythe-Wright et al.*, 2010].



3.2. ²¹⁰Po-²¹⁰Pb Profiles

The minima in total 210 Po activity lay between 50 and 100 m depth and ranged from 2.7 to 6.6 dpm·100 L $^{-1}$ with an average value of 4.6 ± 2.0 dpm·100 L $^{-1}$ (Figure 1 and Table S1 in Supplementary material). The maxima in 210 Po specific activity were found at the surface (5–25 m) and in the mesopelagic zone (below 800 m) and ranged from 5.5 to 12.6 dpm·100 L $^{-1}$ with an average value of 7.6 ± 2.0 dpm·100 L $^{-1}$.

²¹⁰Po activities were lower than ²¹⁰Pb at every depth due to Po accumulation and adsorption onto sinking particles [*Le Moigne et al.*, 2013b]. Profiles of ²¹⁰Po and ²¹⁰Pb specific activities showed the greatest disequilibrium between 50 and 150 m, and indicated higher particle export rates in these regions. Below 500 m ²¹⁰Po activities slowly increased whereas ²¹⁰Pb activities decreased, with equilibrium being rarely found above 1000 m.

Table 3 shows measured particulate 210 Po activity at different depths and the associated values of f(z) calculated using equation (3). A weak dependence of both particulate 210 Po and the ratio of particulate to total 210 Po activity (f(z)) with depth is observed. Particulate 210 Po reached a minimum at 150 m in most stations with average f(z) being largest at 50 m, closest to the MLD.

3.3. Model Fit

 $\delta(z)$ was evaluated by tuning the model to fit the direct field measurements of ²¹⁰Po activities (a_{210} Po(z_0)), using Pearson's chi-squared test (χ^2_{rod}), defined in equation (9).

Initially, we assumed that $\delta(z)$ is constant through the water column (0–1000 m depth). The best solution for $\delta(z)$ (δ = 1100 m), described ²¹⁰Po specific activities at depth well, but failed to reproduce the ²¹⁰Po specific

activities at shallower depths (0 – ~50 m). This behavior indicates that $\frac{a_{210p_0}^{particle}}{a_{210p_0}} \cdot v(z)$ is depth dependent and allows us therefore to abandon the concept of a constant $\delta(z)$ throughout the entire water column.

 $\delta(z)$ is the product of the fraction f(z) of 210 Po activity in particles and the velocity v(z) of settling particles, a change of $\delta(z)$ with depth can therefore be caused by changes in either f or v with depth. Our data show a local extremum indicating a change in the slope of 210 Po and 210 Pb specific activity profiles at around 50 m (see Figure 1). In 210 Po this is a minimum in specific activity indicating where the maximum 210 Po export occurs. This maximum in turn signals the depth of maximum particle export and roughly coincides with the MLD in most cases (Figures 1 and S2), although the resolution of our profiles does not allow us to discern if this maximum export is above or below the MLD.

We therefore adopted a *two-stage* $\delta(z)$ approach, with δ_0 as the value of δ between the surface and some intermediate depth (h), and δ_h as the value of δ between h and greater depths. This intermediate depth, h, was chosen for each profile individually, at the point where the vertical ²¹⁰Po gradient changed (Figure 1) indicating the local minimum in ²¹⁰Po concentration. We then calculated ²¹⁰Po activity depth by depth using measured ²¹⁰Pb values and the measured value of ²¹⁰Po at the shallowest depth as a boundary condition. For each profile, values for δ_0 and δ_h values (shown in Table 2) were chosen to give the lowest value of reduced chi-square (χ^2_{red}).

Modeled 210 Po profiles (Figure 1 and Table S1) calculated from measured 210 Pb values matched measured 210 Po profiles, with higher values being seen at the surface and at 1000 m, and the highest disequilibrium being observed between 50 and 100 m (Figure 1). The Pearson's chi-squared test indicated a good fit for all profiles, with χ^2_{ted} ranging from 0.66 (n = 11) to 3.73 (n = 13), where n is the total number of depths (Table 2).

The lowest offsets between modeled and observed values are found between 200 and 800 m (9% averaged standard deviation, Table S1). The worst fits were found at 800–1000 m, where observations were lower than predicted by the model. Measured ²¹⁰Po and ²¹⁰Pb do not reach secular equilibrium; for instance, the difference at St. 16477 was as high as 35%. This contrasts with the predictions of the model, which suggests close activities for ²¹⁰Po and ²¹⁰Pb at 1000 m in all the profiles. We believe that this is caused by the model incorrectly assuming a constant value of δ below the euphotic zone. If we allow δ to increase by 25% at 800 m in the model, then modeled and measured ²¹⁰Po agree well. The reasons for the ²¹⁰Po-²¹⁰Pb disequilibrium found below 1000 m are not yet well determined [*Church et al.*, 2012]. According to the model presented here, secular equilibrium is broken below 1000 m when δ (x) increases. This increasing can be

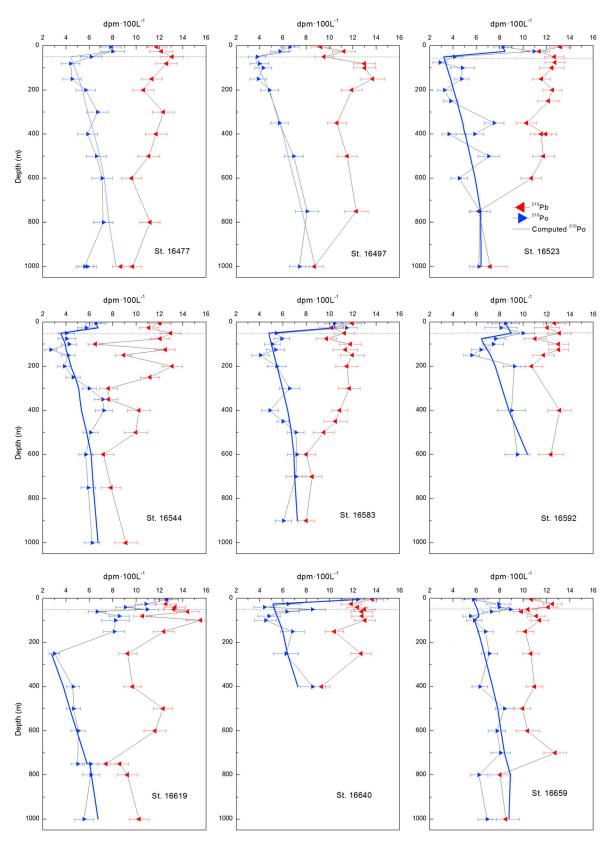


Figure 1. ²¹⁰Po and ²¹⁰Pb depth profiles measured in Porcupine Abyssal Plain (PAP) site during July 2009. Modeled data showing ²¹⁰Po from ²¹⁰Pb experimental data and using a one-box model. $\delta(z)$ is the parameter that needs to be tuned (see equations (5) and (6)) and is obtained by optimizing χ^2_{red} from calculated ²¹⁰Po and experimental ²¹⁰Po. Dash line corresponds to the mixed layer depth.

Table 2. $\delta(z)$	and χ^2_{red} Va	alues of the	²¹⁰ Po Fitting ^a						
Station	χ^2	z(m)	210 Po _{particles} (dpm·100 L ⁻¹)	±σ	f	±σ	δ (m)	$v(m\cdot d^{-1})$	±σ
16477	0.77	50	0.48	0.04	0.079	0.014	610	70	10
		150	0.25	0.03	0.057	0.012	1130	100	20
		400	0.53	0.06	0.091	0.017	1130	65	10
16497	0.37	50	0.56	0.08	0.149	0.035	595	40	9
		150	0.40	0.05	0.103	0.023	1135	55	15
		500	0.53	0.06	0.076	0.012	1135	75	15
16523	2.90	50	0.73	0.05	0.179	0.033	595	45	9
		150	0.39	0.04	0.082	0.013	1605	100	15
		400	0.53	0.06	0.091	0.015	1605	90	15
16544	1.87	50	0.73	0.05	0.183	0.034	595	30	6
		150	0.39	0.04	0.093	0.015	1170	65	10
		400	0.53	0.06	0.073	0.011	1170	80	15
16583	1.23	50	0.30	0.03	0.056	0.009	600	115	20
		150	0.40	0.05	0.099	0.023	1310	65	15
		300	0.47	0.08	0.071	0.015	1310	95	20
		900	0.31	0.05	0.051	0.009	1310	150	30
16592	1.96	50	0.42	0.05	0.042	0.007	410	50	8
		150	0.32	0.04	0.058	0.009	590	50	9
		400	0.32	0.05	0.036	0.007	590	80	15
		600	0.35	0.07	0.037	0.008	590	80	20
16619	1.04	50	0.35	0.04	0.032	0.004	600	95	15
		150	0.41	0.04	0.050	0.007	1000	120	20
		400	0.32	0.05	0.071	0.013	1180	85	15
16640	3.73	50	0.27	0.03	0.032	0.006	1130	180	35
		150	0.16	0.02	0.023	0.005	1130	250	50
		400	0.32	0.05	0.038	0.007	1130	150	30
16659	3.16	50	0.94	0.06	0.106	0.012	820	40	5
		150	0.58	0.05	0.086	0.011	820	50	7
		400	0.32	0.05	0.051	0.009	895	95	15
Average fron	า	50	0.60	0.20	0.10	0.06	1023	60	30
all stations	5	150	0.40	0.10	0.080	0.02	1127	75	25
		400-900	0.40	0.10	0.065	0.02	1067	90	20

 $^{^{}a}\delta(z)$ is the tuned parameter of the model that optimizes χ^{2}_{red} $^{210}Po_{particles}$ $(dpm\cdot100\,L^{-1})$ correspond to specific ^{210}Po activity measured in particles collected from in situ pumps filters (particles size >1 and $>1-53 \,\mu m$). f is the specific 210 Po activity in particulate fraction (particles size $>1 \,\mu m$) to total activity. v(z) is the average particle settling speed obtained using modeled δ and measured f.

ascribed either to an increase in the particle sinking velocities (which will be further discussed in section 4.2) or to a change in the ratio of particulate 210 Po to total 210 Po (i.e., a change in the f(x) parameter). An increase in the particulate to total ²¹⁰Po ratio would be associated to a constant scavenging of ²¹⁰Po by the particles while they sink, in contrast to the scavenging of ²³⁴Th where an equilibrium of ²³⁴Th in the particulate-dissolved phases is reached. This is supported by the fact that ²³⁴Th-²³⁸U reach equilibrium at shallow depth (200 m), whereas ²¹⁰Po-²¹⁰Pb never reach equilibrium before 500 m [*Bacon et al.*, 1976; *Hong* et al., 2013; Verdeny et al., 2008].

3.4. Sinking Speeds

Mean settling speeds diagnosed from the model at 50, 150, and 400-600 m (labeled as 500 m) using data from nine stations are shown in Figure 2 and Table 3. At 50 m, average sinking speeds show high scatter with a standard deviation of 50%. The scatter decreases to a standard deviation of 25% at 150 m and 20% at 500 m. Despite the spread in data, an increase in velocity with depth is observed. Average values increased from $60 \pm 30 \text{ m} \cdot \text{d}^{-1}$ at 50 m to $75 \pm 25 \text{ m} \cdot \text{d}^{-1}$ and $90 \pm 20 \text{ m} \cdot \text{d}^{-1}$, at 150 and 500 m, respectively, with a correlation between sinking speed and depth occurring (p = 0.002, R = 0.58, n = 27).

3.5. Application to ²³⁴Th and Sensitivity Analysis

Similar reasoning to that outlined above regarding the interpretation of ²¹⁰Po-²¹⁰Pb profiles can be applied to the analogous radioactive pair of ²³⁴Th-²³⁸U, which was also measured alongside our ²¹⁰Po-²¹⁰Pb

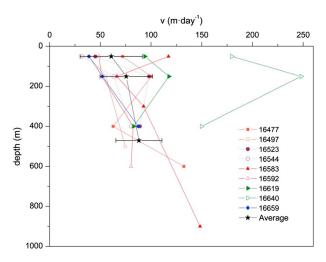


Figure 2. Average particle settling speeds $(m \cdot d^{-1})$ calculated at the PAP site during July 2009 for nine stations (uncertainties are given in Table 2). Average value for 50, 150, and 500 m depths are obtained. Uncertainty bars of the average values correspond to one standard deviation.

measurements. Details of collection, measurement, and results of ²³⁴Th specific activities at the PAP site are found in Le Moigne et al. [2013b], where POC-Th and POC-Po derived fluxes were presented. Sinking rates from ²³⁴Th results are shown in Table 3 (supplementary information in Table S2). ²³⁴Th and ²³⁸U modeled and measured profiles are presented in Figure S4. The differences between modeled and measured values range from 1 to 15%.

As the half-life of ²³⁴Th is only 24 days, equilibrium with ²³⁸U is found at around 150-200 m, which is shallower than the depth at which the ²¹⁰Po-²¹⁰Pb pair reaches equilibrium. Average velocities can therefore only be obtained at 50 and 150 m (Table 3). An increase in speed with depth is clearly observed at all stations with average velocities rising from

 $70 \pm 20 \,\mathrm{m}\cdot\mathrm{d}^{-1}$ at 50 m to $130 \pm 20 \,\mathrm{m}\cdot\mathrm{d}^{-1}$ at 150 m. These are slightly higher but similar to the velocities obtained from the Po-Pb profiles. An analogous correlation between sinking speed and depth is observed (p = 0.00004, R = 0.8, n = 22). However, low resolution of the ²³⁴Th profiles collected during D341 (Table S2) prevents the full potential ability of the inverse model to extract sinking velocities from ²³⁴Th measurements to be explored here. If this method is to be used systematically with ²³⁴Th data to derive sinking velocities we recommend that high resolution ²³⁴Th profiles be taken from surface to 150–200 m, the depth at which ²³⁸U and ²³⁴Th are likely to reach secular equilibrium.

Table 3. $\delta(z)$ and χ^2_{red} Value	s of the	²³⁴ Th Fitting ^a						
Station	z(m)	234 Th _{particles} (dpm·100 L ⁻¹)	±σ	f	±σ	δ (m)	$v(m\cdot d^{-1})$	±σ
16477	50	0.050	0.006	0.024	0.006	70	82	31
	150	0.038	0.004	0.016	0.003	70	127	33
16497	50	0.058	0.006	0.021	0.004	350	97	24
	150	0.035	0.004	0.013	0.002	350	153	36
16523	50	0.056	0.004	0.028	0.003	45	71	10
	150	0.041	0.004	0.014	0.002	45	140	23
16544	50	0.056	0.006	0.027	0.003	45	76	13
	150	0.041	0.004	0.017	0.002	45	118	22
16559	50	0.056	0.006	0.023	0.003	30	86	14
	150	0.041	0.004	0.017	0.002	30	121	19
16583	50	0.079	0.009	0.032	0.004	30	62	11
	150	0.046	0.005	0.018	0.002	30	112	19
16592	50	0.066	0.007	0.044	0.005	40	46	8
	150	0.031	0.003	0.013	0.002	40	153	25
16619	50	0.072	0.008	0.031	0.004	20	64	13
	150	0.063	0.007	0.025	0.003	20	80	14
16640	50	0.048	0.005	0.019	0.002	15	105	18
	150	0.036	0.004	0.014	0.002	15	140	22
16659	50	0.092	0.011	0.058	0.008	35	35	7
	150	0.041	0.027	0.016	0.002	35	122	23
Average from all stations	50	0.065	0.015	0.030	0.010	68	70	20
-	150	0.04	0.01	0.015	0.003	68	130	20
3		- 224			1		2	2.4

 $^{^{}a}\delta(z)$ is the tuned parameter of the model that optimizes χ^{2}_{red} . $^{234}Th_{particles}$ ($dpm\cdot100L^{-1}$) correspond to specific ^{234}Th activity measured in particles collected from in situ pumps filters (particles size >1 and >1–53 μ m). f is the specific ²³⁴Th activity in particulate fraction (particles size >1 μ m) to total activity. $\delta(z)$ is the fitting parameter of the model that optimizes χ^2_{red} . The term v(z) is the average particle settling speed obtained using modeled δ and measured f.

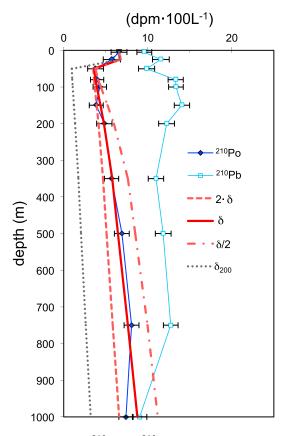


Figure 3. Specific ²¹⁰Po and ²¹⁰Pb activity profiles, Station 16977. 210 Po modeled profiles using: (i) optimized δ parameter, solid line; (ii) twice the optimized δ parameter, dashed line; (iii) half the optimized δ parameter, δ dotteddashed line; and (iv) δ parameter for sinking speed v = 200 m/d, dotted line.

We undertook a sensitivity analysis in which the speed is doubled or halved (and δ changed accordingly) in a profile (e.g., St 16977) and the shape of the modeled speed profile examined (Figure 3). No significant changes are observed at shallower depths; however, halving δ (lower velocities according to Equation (10) produces lower ²¹⁰Po-²¹⁰Pb deficits at depth, and the ²¹⁰Po-²¹⁰Pb equilibrium is reached faster. Conversely, according to our model doubling δ (higher velocities) leads to a higher deficit at depth and equilibrium is reached at depths greater than 1000 m. This indicates that both the extent of the ²¹⁰Po-²¹⁰Pb disequilibrium and the depth at which secular equilibrium is reached are very sensitive to sinking velocities and can therefore be used to qualitatively determine sinking velocities in that area.

Finally, since it was previously suggested that in the northwest Mediterranean the main contributors to the flux from the euphotic to the twilight zone are predominantly fast (200 m·d⁻¹) and large (>70 μm) particles [Abramson et al., 2010], we altered the sinking speed in our model to $200 \,\mathrm{m} \cdot \mathrm{d}^{-1}$ (δ_{fast}), to test this hypothesis at the PAP site. The results, shown in Figure 3, do not compare well with the observations and no equilibrium is found at depth.

4. Discussion

4.1. Limitations of the Approach

The most widespread application of ²¹⁰Po-²¹⁰Pb (and ²³⁴Th-²³⁸U) depth profiles is to quantitatively estimate particle export (e.g., Buesseler et al. [2008]; Le Moigne et al. [2013a]; Le Moigne et al. [2013b]; Verdeny et al. [2009] and Murray et al. [2005]) rather than to derive sinking speeds. 210Po (or 234Th) export is calculated using an export model, and subsequently the downward flux is converted to POC (or PIC, trace metal, etc.) flux via a known tracer/Th ratio. The simplest export model combines particulate and dissolved polonium (or thorium) into a single box [Savoye et al., 2006]. Two-box models differentiate between dissolved and particulate phases (e.g., Bacon and Anderson [1982] or Friedrich and van der Loeff [2002]), and more sophisticated models also discriminate between small and large fractions [e.g., Marchal and Lam, 2012].

The limitation of the one-box model is that it provides no information on removal processes such as adsorption, biological uptake, and colloidal interactions. In order to analyze such processes a more complex model is needed such as the three-box model described by Marchal and Lam [2012] which estimates the rate constants for ²³⁴Th sorption onto large and small particles using inverse techniques. In contrast, instead of the rate constants we can estimate depth variation in settling rate, since in our model the rate constant is written in terms of speed and a tuning parameter

In most models that extract export data from radiotracer profiles a steady state (SS) model is used to estimate flux, while the more difficult to treat non-steady state (NNS) models are undertaken in situations where rapidly changing conditions render the steady state assumption invalid, such as bloom events [Baena et al., 2008; Buesseler et al., 1992; Cochran et al., 2009]. The steady state assumption is less valid for ²¹⁰Po than for 234Th due to its longer half-life [Bacon et al., 1980; Friedrich and van der Loeff, 2002]. Nonetheless, most of the analyses undertaken for ²¹⁰Po assume steady state due to the inherent limitations of sampling strategies



during cruises (e.g., *Murray et al.* [2005]; *Stewart et al.* [2010] and *Verdeny et al.* [2008]): stations are normally sampled once while NSS equations can only be applied to ²¹⁰Po in stations sampled 2–3 months apart [*Stewart et al.*, 2007b].

Data from other authors obtained during the same cruise show that physical and biogeochemical conditions at the site were relatively constant during the course of this study. For instance, chlorophyll and PP data remained stable during the sampling period [*Le Moigne et al.*, 2013b]. In addition advection is believed to be moderate at the site [*Giering et al.*, 2014], T/S profiles show the presence of a single water mass (ENACWp) between the MLD and \sim 600 m at all stations (Figure S3), and current velocities were $<5 \, \text{m} \cdot \text{d}^{-1}$ [*Giering et al.*, 2014] with the MLD remaining relatively constant (range 31–62 m, Figure S2). Overall, the conditions during the cruise were relatively stationary, suggesting that steady state conditions prevailed over the 27 day sampling period and hence that the mean profile and the mean sinking speeds likely represent the average conditions at the site.

However, since D341 took place approximately 1 month after the bloom period [*Le Moigne et al.*, 2013b] there could be a temporal mismatch between the 210 Po- 210 Pb deficit in the water column (which might be partially driven by the spring bloom) and the particulate 210 Po sampled during the post-bloom period [*Buesseler et al.*, 2006], a factor which may affect the SS assumption [*Cochran et al.*, 2009]. Particles collected during the cruise represent a blend of recently synthesized material and any particles which might have been created prior to the sampling period and then sunk slowly out of the mixed layer, e.g., fast particles created during the bloom period (>100 m·d $^{-1}$) would not be present in the water column sampled in July. However, particles from the bloom period sinking at $10 \, \text{m·d}^{-1}$ would still be present below the 300 m when the sampling cruise took place. Thus, v at the shallowest depths is likely to be driven by the prevailing post-bloom situation during the cruise, with deeper values being more influenced by any remnant slowsinking particles which exited the bloom several weeks prior to sampling.

Any effect from the increase in production occurring during the bloom period will gradually decay with time until a steady state profile is reached, which will result in time-independent ^{210}Po fluxes over the course of the post-bloom period. Thus, in order to assess the validity of the steady state assumption one could simply compare $^{210}\text{Pb}-^{210}\text{Po}$ profiles sampled from the same location but at different times. Unfortunately, most of the locations where sampled only once during the cruise except for stations 16497 and 16619 which sampled the same location 18 days apart (Table 1). ^{210}Po fluxes at 16497 and 16619 stations were found to be significantly different from each other (65 ± 5 and 44 ± 5 dpm/m²/d respectively at 150 m) [*Le Moigne et al.*, 2013b]. Since advection was shown to be low at the PAP site, the difference between fluxes and ^{210}Po measured activities at the two stations could be due to non-steady state effects. With this assumption we estimate the time derivative in equation (1) as being $da_{210_{Po}}(t)/dt = 0.013\,\text{dpm}\cdot100\,\text{L}^{-1}\,\text{d}^{-1}$. When this value is included in equation (8) and the δ parameter is re-evaluated the changes in δ are <10%. We therefore conclude that the site was close enough to SS conditions for the purpose of this study and that the systematic error in estimating sinking speeds introduced by the steady state assumption is smaller than the error in sinking velocity quantified by the range observed from the various stations.

A final possible source of error is that large particles may be under-sampled due to the inlet design of the SAPS and to the methodology used to wash them off the $53\,\mu m$ mesh [Liu et al., 2009]. If particulate 210 Po were underestimated, then the sinking velocities from equation (6) would be overestimated, leading to a higher contribution of slow-sinking particles to flux.

One of the motivations of this study is to analyze the extent to which the simplest export model (one-box, steady state) can reproduce measured ²¹⁰Po profiles. The good agreement between modeled and observed specific ²¹⁰Po activities (Figure 1 and Table 2) and between ²³⁴Th and ²¹⁰Po derived results (Tables 2 and 3) suggests that the single box steady state model is an appropriate tool for interpreting data in this region during the study period, where the biogeochemical and physical conditions remained relatively constant and where ²¹⁰Po profiles sampled 18 days apart showed average relative variations of less than 10%.

4.2. Sinking Speeds

4.2.1. Contribution of Slow-Sinking Particles to Flux and Changes in Sinking Speed With Depth Minimum and maximum sinking velocities of 30 ± 6 (St. 16544) and $130 \pm 25 \text{ m} \cdot \text{d}^{-1}$ (St. 16583) were found at 50 and 400 m, respectively, with average values increasing from $60 \pm 30 \text{ m} \cdot \text{d}^{-1}$ at 50 m to $90 \pm 20 \text{ m} \cdot \text{d}^{-1}$ at 500 m. These values are significantly below $150-200 \text{ m} \cdot \text{d}^{-1}$, the reported average values for fast-sinking



particles [*Turner*, 2002]. 234 Th derived sinking velocities are $70\pm20\,\mathrm{m\cdot d^{-1}}$ at 50 m and $130\pm20\,\mathrm{m\cdot d^{-1}}$ at 150 m. Furthermore, our sensitivity analysis suggests that simulating the 210 Po profile using a sinking velocity of $200\,\mathrm{m\cdot d^{-1}}$ in the model produces results inconsistent with the experimental data (Figure 3). We interpret these results as suggesting that either most particles are sinking at speeds which are not seen in the IRS synthesis of *Alonso-Gonzalez et al.* [2010], or more likely, that a significant fraction of the flux at the PAP site is carried by very slow-sinking particles, consistent with the suggestion of *Alonso-Gonzalez et al.* [2010].

Our estimates of bulk particle sinking speeds are lower than directly measured aggregate (0.2–0.5 mm³) sinking speeds made at 50 m during the cruise [Riley et al., 2012] which averaged 180 m·d⁻¹. However, Riley et al. [2012] also observed that $63 \pm 12\%$ of the sinking POC was driven by slow settling (~9 m·d⁻¹) particles, with the remaining 40% made up of the fast settling aggregates. If we combine these relative fractions and their sinking rates (181 and 9 m·d⁻¹, respectively), we arrive at an estimated settling velocity of 70 m·d^{-1} at 50 m. This bulk estimate is remarkably similar to our estimate of $60 \pm 30 \text{ m·d}^{-1}$ at 50 m.

It is unclear how deep the slow-sinking particles penetrate into the mesopelagic, although *Riley et al.* [2012] hypothesize that most of these particles will be lost in the upper water column, with the deep flux being driven by remaining fast-sinking aggregates. Our average sinking speeds of 75 m·d⁻¹ at 150 m, and 90 m·d⁻¹ at 500 m are considerably slower than the sinking speeds of the aggregates measured by *Riley et al.* [2012] suggesting that either slow-sinking particles contribute to flux deep into the twilight zone or that the sinking speeds of aggregates decline significantly in the upper mesopelagic.

In equation (3), ²¹⁰Po flux can increase with depth due to either (i) an increase in the ²¹⁰Po transported by sinking particles (e.g., due to a creation of particles in the euphotic zone or to the accumulation of ²¹⁰Po in extant particles) or (ii) to an increase in sinking speed with depth. The increase in ²¹⁰Po flux between 0 and 1000 m is probably a combination of both factors.

We observe a moderate positive correlation between sinking velocities and depth (p = 0.002, R = 0.58, n = 27 using ^{210}Po and p > 0.001, R = 0.80, n = 22 using ^{234}Th) with average sinking velocities increasing from 50 m ($60 \text{ m} \cdot \text{d}^{-1}$) to 150 m ($75 \text{ m} \cdot \text{d}^{-1}$) and 500 m ($90 \text{ m} \cdot \text{d}^{-1}$). These data provide evidence for a moderate increase in bulk sinking velocity below the euphotic zone. We also found a positive correlation between depth and sinking speed (p < 0.001, p = 0.611, p = 0.611

4.2.2. Comparison With Previous Work

The presence of relatively slow-sinking material ($<10\,\mathrm{m\cdot d}^{-1}$) at depth was also documented in the VERTIGO campaigns at K2 and ALOHA sites by *Trull et al.* [2008] using IRS traps. IRS sediment traps at 260 m [*Alonso-Gonzalez et al.*, 2010] suggest that slowly sinking material makes substantial contributions to deep flux in the Canary Current (Central Atlantic) under some circumstances. During summer and fall 60% of total POC flux was in the slowest settling class (0.7–11 $\mathrm{m\cdot d}^{-1}$) with 25% of the POC being carried by fast-sinking particles sinking at $>320\,\mathrm{m\cdot d}^{-1}$. If these are weighted by their corresponding contributions to the flux then an average speed of 107 $\mathrm{m\cdot d}^{-1}$ is found, similar to our estimate of velocity at 500 m.

Larger contributions of the slow-sinking particles to flux have been identified in the subtropical gyre at ALOHA (subtropical) than in the subarctic at K2 [*Trull et al.*, 2008] due to differences in ballasting and zooplankton pellet size [*Buesseler et al.*, 2007]. In the North Atlantic the contribution of slow-sinking particles to flux at the PAP site is comparable to the contributions at ALOHA and at the Canary Islands despite the differences in the ecosystems involved, suggesting that the pattern we observe is rather widespread. It is worth noticing that the size and speed of sinking particles vary in space and time. Thus, subtropical oligotrophic ecosystems are dominated by picophytoplankton and export is often driven by aggregates of small cells and rarely by large, rapidly sinking individual cells [*Stewart et al.*, 2010]. This diverges from the community in a subarctic ecosystem such as K2, with high nutrient conditions and higher production and export and where a large fraction (15–45%) of the flux was contributed by fast-sinking classes (410–205 m·d⁻¹) [*Trull et al.*, 2008]. Also, the seasonality is a decisive parameter, since larger and faster organisms appear during the first stages of the bloom but when the productive season advances the community shifts to smaller and slower phytoplankton species.

Overall both the IRS trap data summarized by *Alonso-Gonzalez et al.* [2010] and the radiotracer data reported here indicate that slow-sinking particles on occasion make significant contributions toward flux beneath



the euphotic zone. Open questions include how significant this flux is, how deep it penetrates, and what regulates its presence. The interpretation of ²¹⁰Po-²¹⁰Pb profiles presented here would, if applied globally to a large enough dataset allow these questions to be addressed.

4.3. Implications for Our Understanding of Particle Dynamics

An increase in sinking speed with depth can occur either due to (i) a change in particle characteristics during export [e.g., *Fowler and Knauer* [1986]] or to (ii) the gradual elimination of the slow-sinking fraction, as suggested by *Riley et al.* [2012].

According to Stokes' Law, particle velocities can increase as they sink due to either an increase in size or to an increase in density or to both effects simultaneously [Armstrong et al., 2009]. Both changes might be associated with radiolarian populations at PAP site that contributed to the abnormally large POC flux observed during 2001 [Lampitt et al., 2009]. According to Riley et al. [2012] radiolaria may have acted as nucleation points for detrital material, incorporating it into fast-sinking aggregates termed Aggregate Protist Complexes. In addition the production of sticky transparent exopolymer particles (TEP) [De La Rocha and Passow, 2007] by diatoms might contribute to the formation of fast-sinking aggregates.

The combination of our results with those of *Riley et al.* [2012] and *Alonso-Gonzalez et al.* [2010] suggests that there may be a significant contribution of slow-sinking (<10 m·d⁻¹) particles to the flux at shallow (50 m) and mesopelagic (150–400 m) depths. The slow-sinking particles in the mesopelagic could originate from either (a) the slow-sinking material found at shallow depths sinking deeper into the water column or (b) to slowly sinking particles formed at depth by aggregation and disaggregation processes, by microbial activity or by the self-assembly of dissolved organic material [*Alonso-Gonzalez et al.*, 2010; *Maiti et al.*, 2010; *Trull et al.*, 2008].

Option (b) is supported by *Giering et al.* [2014] who analyzed carbon demand and remineralization processes at the PAP site during D341. Their major conclusion was that the key role of mesopelagic zooplankton is to degrade large (fast) sinking particles into smaller (slow) particles which are subsequently remineralized by microbes. The slow-sinking particles that we suggest occur below the euphotic zone are likely to be created this way.

Regardless of their origin, slow-sinking particles are likely to be preferentially remineralized compared to fast-sinking material [Abramson et al., 2010; Buesseler et al., 2007; Trull et al., 2008] due to their longer residence times at any particular depth. Their likely fate, rather than sinking, is ingestion and respiration by zooplankton and bacteria [Trull et al., 2008]. Furthermore, interior respiration is dominated by prokaryotic bacteria rather than zooplankton at the PAP site [Giering et al., 2014].

Boyd and Trull [2007] suggested that the power law expression that describes particle flux attenuation can be explained not only by the inherent decrease of particle degradation rates with depth but also by a downward increase of sinking rates. The balance of these two effects is hard to establish as most measurement methodologies probe fast-sinking particles preferentially [Berelson, 2002; Pilskaln et al., 1998].

Our work suggests that a downward increase in average speeds plays some part in creating the power law relationship observed, as slow-sinking particles are remineralized in the twilight zone with fast particles dominating the flux at 1000 m as suggested by *Armstrong et al.* [2009]. Alternatively or additionally the velocities of the large fast-sinking particles may themselves increase due to repackaging or remineralization [*Boyd and Trull*, 2007]. Our estimates of velocity at 500 m are lower than the fast-sinking particles observed by *Riley et al.* [2012], suggesting that this possibility cannot be excluded. Furthermore, we propose that at the PAP site the pool of slowly sinking particle is lost in the upper mesopelagic likely due to consumption by prokaryotes, as previously postulated by *Giering et al.* [2014]. A previous lack of accounting for slowly sinking particles in twilight zone carbon budgets may be a partial cause of the imbalance between supply and consumption of organic carbon in the mesopelagic, as discussed by *Burd et al.* [2010].

Our work cannot draw conclusions on the mechanisms of remineralization or repackaging which produce the sinking velocities diagnosed here. Further studies are therefore needed to determine the origin of the slow-sinking flux, its fate in the twilight zone in different biogeochemical settings, and its potential contribution toward understanding carbon budgets. *Bopp et al.* [2005] predicted future shifts from biomineralizing species such as diatoms toward smaller (more slowly sinking) species due to future changes in ocean stratification and associated reductions in nutrient supply. We therefore consider it of major



importance to ascertain the role of small particles throughout the water column in order to determine whether such a shift might change the efficiency of POC transport to depth.

5. Conclusions

The 210 Po- 210 Pb deficit in the water column is used to estimate the depth dependency of particle settling speed. Sensitivity tests show that the ²¹⁰Po deficit below the euphotic zone cannot be exclusively due to fast particles sinking with velocities close to 200 m s⁻¹. A moderate depth dependency in settling speed was found with average speeds of $60 \pm 30 \text{ m} \cdot \text{d}^{-1}$ at 50 m, $75 \pm 25 \text{ m} \cdot \text{d}^{-1}$ at 150 m and $90 \pm 20 \text{ m} \cdot \text{d}^{-1}$ at 500 m. Similar results have been found for ²³⁴Th-derived sinking activities. This approach yields sinking velocities similar to those obtained by Riley et al. [2012] from the same cruise and confirms that there might be a contribution of slow-sinking particles to the carbon flux below the euphotic zone. Our results may have implications for closing mesopelagic carbon budgets and suggest that ascertaining sinking velocities is significant for the prediction of biological carbon storage in an ocean whose population is predicted to shift toward small phytoplankton at the expense of diatoms.

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