

# An improved quantitative measure of the tendency for volcanic ash plumes to form in water: implications for the deposition of marine ash beds

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## Abstract

Laboratory experiments and numerical simulations have shown that volcanic ash particles immersed in water can either settle slowly and individually, or rapidly and collectively as particle-laden plumes. The ratio of timescales for individual and collective settling, in the form of analytical expressions, provides a dimensionless quantitative measure of the tendency for such plumes to grow and persist which has important implications for determining particle residence times and deposition rates. However, existing measures in the literature assume that collective settling obeys Stokes' law and is therefore controlled by the balance between gravitational forces and viscous drag, de-

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spite plume development actually being controlled by the balance between gravitational forces and inertial drag even in the absence of turbulence during early times. This paper presents a new measure for plume onset which takes into account the inertial drag-controlled (rather than viscous drag-controlled) nature of plume growth and descent. A parameter study comprising a set of numerical simulations of small-scale volcanic ash particle settling experiments highlights the effectiveness of the new measure and, by comparison with an existing measure in the literature, also demonstrates that the timescale of collective settling is grossly under-estimated when assuming that plume development is slowed by viscous drag. Furthermore, the formulation of the new measure means that the tendency for plumes to form can be estimated from the thickness and concentration of the final deposit; the magnitude and duration of particle flux across the water's surface do not need to be known. The measure therefore permits the residence times of particles in a large body of water to be more accurately and practically determined, and allows the improved interpretation of layers of volcanoclastic material deposited at the seabed.

*Keywords:* Ash plumes, Settling rates, Volcanoclastic deposits, Computational modelling, Numerical simulations, Vertical density currents

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## 1. Introduction

Explosive volcanism generates vast quantities of small ash particles which can be transported over great distances, eventually depositing both on land and on the seabed to form particle layers (Carey and Schneider, 2011). These layers are a text-book example of isochroneity and have been used for strati-

6 graphic correlation of past eruption events (e.g. Ver Straeten (2004, 2008)),  
7 allowing a wealth of information regarding their duration and frequency to  
8 be determined. Furthermore, ash deposits can potentially preserve informa-  
9 tion about the environmental conditions at the time of an event (Manville  
10 and Wilson, 2004). However, the process behind the settling of ash and the  
11 resulting formation of the particle layers is far from simple.

12 It was once assumed that the settling of ash in the deep sea occurred  
13 passively such that particles always descend slowly and individually under  
14 Stokes' law (Ledbetter and Sparks, 1979; Carey and Schneider, 2011), but  
15 several field-based observations have provided contradictory evidence. For  
16 example, following the 1991 eruption of Mount Pinatubo, ash fallout in the  
17 South China Sea settled at speeds of over  $2 \text{ cms}^{-1}$  which is two to three orders  
18 of magnitude greater than the calculated Stokes' law velocities of individual  
19 particles (Wiesner et al., 1995). Through analogous laboratory experiments,  
20 Carey (1997) set out to explore this apparent contradiction in timescales and  
21 revealed the important role of vertical density currents in the rapid, collective  
22 transportation of material to the seabed.

23 The generation of vertical density currents is a complex multiphase pro-  
24 cess. Particles entering a body of water, either as fallout from ash clouds in  
25 the atmosphere or from a pyroclastic density current, undergo abrupt decel-  
26 eration as they cross the air-water interface. Initially, slow and individual  
27 settling under Stokes' law ensues, allowing the particle concentration near  
28 the surface to rapidly increase and form a layer of particle-rich water over  
29 time. However, if the particle concentration in the layer is large enough  
30 for the particles to affect each other's settling through drag reduction and

31 drifting such that the layer becomes gravitationally unstable, then finger-like  
32 Rayleigh-Taylor instabilities eventually form along the interface between the  
33 layer and the particle-free water below it. These instabilities grow exponen-  
34 tially to form plumes — clouds of particles that settle rapidly and collectively  
35 as vertical density currents.

36     Knowing whether plumes are likely to form, if at all, is important if one  
37 wishes to better determine the timescale of settling from the surface to the  
38 seabed. This can reveal information about the residence time of particles  
39 in the water and therefore the extent to which ambient ocean currents re-  
40 distribute volcanoclastic material as it settles (Carey and Schneider, 2011).  
41 Similarly, knowing the rate of deposition can help determine the degree of  
42 bioturbation of the growing particle layer by marine organisms (Bramlette  
43 and Bradley, 1941). Plume formation also has implications for fossil preser-  
44 vation and stratigraphy. Rapid sedimentation has long been recognised as a  
45 means of increasing the likelihood that an organism could be preserved as a  
46 fossil (Seilacher et al., 1985) and so ash plume formation can impact upon the  
47 completeness of the fossil record. Perhaps one of the most celebrated and ge-  
48 ologically significant examples of exceptional preservation beneath a marine  
49 ash deposit is that of the Neoproterozoic Ediacaran biota in Newfoundland  
50 which preserves some of the earliest metazoan fossils on Earth (Narbonne,  
51 2005).

### 52 *1.1. Theoretical Considerations*

53     Quantitatively describing the tendency for plumes of particles to form in  
54 an ambient fluid has been achieved in previous works (Marsh, 1988; Goldin,  
55 2008; Carazzo and Jellinek, 2012) through a dimensionless number  $B$ . This is

56 defined in such a way that values of  $B$  less than or equal to unity imply that  
 57 plumes do not form, whereas a value greater than unity implies favourable  
 58 conditions for plume growth and persistence. In particular, existing dimen-  
 59 sionless numbers have been defined by the ratio of timescales for individual  
 60 particle settling under Stokes' law and collective settling as a gravitationally  
 61 unstable plume, such that

$$B = \frac{\tau_{\text{individual}}}{\tau_{\text{collective}}}. \quad (1)$$

62 That is, given information about the current state of Rayleigh-Taylor insta-  
 63 bilities, the time required for particles to reach that state through individual  
 64 and collective settling modes can be approximated using analytical expres-  
 65 sions. Clearly a value of  $B \gg 1$  implies favourable conditions for plume  
 66 formation and persistence since collective settling happens over a shorter  
 67 timescale (e.g. days or weeks in the ocean) than individual settling (e.g.  
 68 months), whereas a value of  $B \approx 1$  implies that plumes cannot form since  
 69 the timescales of individual and collective settling are of the same order of  
 70 magnitude. Note that a value of  $B < 1$  also implies that plumes cannot form,  
 71 but when  $B$  is defined by the ratio of timescales this value has no physical  
 72 meaning except for the case of hindered settling (Kuenen, 1968) which is  
 73 not considered here. The parameters needed to compute these expressions  
 74 include the particle concentration and the thickness of the particle-rich layer  
 75 which often have to be estimated in practice. Alternatively, the measure  
 76 can be re-formulated in terms of a critical layer thickness that must be at-  
 77 tained in order for pluming to take place (discussed later). This only requires  
 78 knowledge of the mass influx across the water's surface and particle diameter

79 which is often readily available during or after an eruption event.

80 One such formulation of  $B$  is the one derived by Marsh (1988) for the  
81 study of crystal settling in magma, denoted  $B_{vv}$  in this paper. This formu-  
82 lation is based on the assumption that both individual particles and plumes  
83 obey Stokes' law and are therefore controlled by the balance between gravita-  
84 tional forces (weight and buoyancy) and the viscous drag force (i.e. the drag  
85 arising from the friction between the descending particles/plumes and the  
86 ambient fluid), hence the use of the subscript  $vv$  to denote 'viscous-viscous'.  
87 The time taken for an individual (spherical) particle to settle through a layer  
88 of thickness  $h$  is therefore given by

$$\tau_{\text{individual}} = \frac{18h\mu_f}{(\rho_p - \rho_f)gd_p^2}, \quad (2)$$

89 where  $d_p$  is the particle diameter,  $g$  is the acceleration due to gravity,  $\mu_f$  is  
90 the viscosity of the fluid phase, and  $\rho_f$  and  $\rho_p$  are the density of the fluid and  
91 particle phase, respectively (Stokes, 1851). The assumption that all particles  
92 have a perfect spherical shape is implicitly built-in to the timescale above  
93 through the Stokes drag coefficient. Furthermore, it has been shown (see for  
94 example Whitehead and Luther (1975); Goldin (2008)) that the timescale of  
95 collective settling is given by:

$$\tau_{\text{collective}} = \frac{18\mu_f}{\alpha_p(\rho_p - \rho_f)gh}, \quad (3)$$

96 where  $\alpha_p$  is the volume fraction of particles in the layer. Taking the ratio of  
97 these two timescales yields the dimensionless number  $B_{vv}$ :

$$B_{vv} = \frac{\alpha_p h^2}{d_p^2}. \quad (4)$$

98 Further work by Carazzo and Jellinek (2012) derived similar non-dimensional  
99 numbers for the scenario of volcanic ash settling through the atmosphere.  
100 Coarse-grained ash and lapilli can settle individually with a particle Reynolds  
101 number several orders of magnitude greater than that of fine ash (Bonadonna  
102 et al., 1998), so three forms of  $B$  were derived using different expressions for  
103  $\tau_{\text{individual}}$  to cover a wide range of individual particle settling regimes. How-  
104 ever, none of these measures address the fact that plume growth and descent  
105 are controlled by the balance between gravitational forces and the inertial  
106 drag force (Dalziel et al., 2008; Bergantz and Ni, 1999). This inertial drag  
107 force arises from the need for the plumes to accelerate and displace the sur-  
108 rounding fluid, even in the absence of fluid viscosity, and dominates the  
109 viscous drag force as shown by plume Reynolds numbers<sup>1</sup> much greater than  
110 unity (Jacobs et al., 2013). At this point Stokes' law no longer holds even  
111 if no turbulent effects are observed until the plumes are fully developed and  
112 begin to mix, which has a significant impact on entrainment and settling  
113 rates (Manville and Wilson, 2004). A measure which assumes that collective  
114 particle settling is slowed by inertial drag (rather than viscous drag) may  
115 therefore be more appropriate.

116 This paper presents a new measure of the tendency for particles to form  
117 plumes and settle collectively which accounts for the fact that collective par-  
118 ticle settling is slowed by inertial drag. The new non-dimensional number,  
119 denoted  $B_{vi}$ , is derived by applying Stokes' law and a well-founded expression  
120 for the growth rate of Rayleigh-Taylor instabilities (Youngs, 1984). The va-

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<sup>1</sup>The Reynolds number is a dimensionless quantity defined as the ratio of inertial to viscous drag force.

121 lidity of the measure for predicting the formation of plumes as particles settle  
122 in water is then evaluated and compared against  $B_{vv}$ . This is accomplished  
123 by (a) using data from the experiments by Carey (1997) which consider ash  
124 particles settling through a water tank, and (b) performing a parameter study  
125 through analogous numerical simulations with the multiphase computational  
126 fluid dynamics (CFD) code Fluidity (Piggott et al., 2008; Davies et al., 2011;  
127 Jacobs et al., 2013). The paper finishes with a discussion of the implications  
128 and applications of the new measure, other geophysical scenarios where the  
129 new measure could also be valid, and some concluding remarks. A list of  
130 notation used throughout the paper is provided in Appendix A.

## 131 2. Derivation of the New Measure

132 To derive a measure of the tendency of plumes to form which takes into  
133 account the fact that collective settling is slowed by inertial (rather than  
134 viscous) drag, consider the growth of wave-like instabilities with maximum  
135 amplitude  $\delta$  at the interface between a particle-water layer of thickness  $h$   
136 and the particle-free water beneath it, as illustrated in Figure 1. The water  
137 is treated as an incompressible fluid, and the particles have an idealised  
138 spherical shape.

139 From Stokes' law, the timescale required for an individual spherical par-  
140 ticle to settle through the layer of thickness  $h$  is given by (2) previously. A  
141 timescale for the settling of a cloud of particles with a growing amplitude  
142  $\delta$  can be derived from an ordinary differential equation describing the late-  
143 time growth rate of Rayleigh-Taylor instabilities (Ristorcelli and Clark, 2004;  
144 Youngs, 1984),



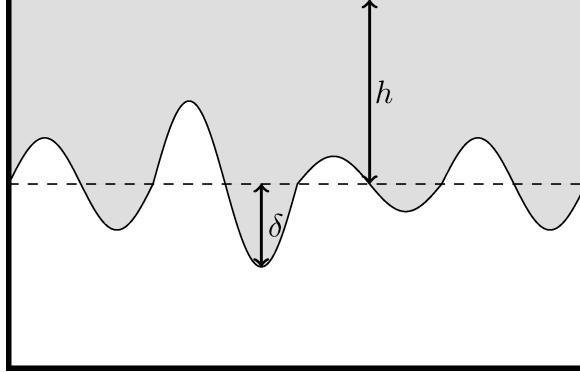


Figure 1: Illustration of particle plumes forming in a tank of water. The height of the particle-rich layer is denoted by  $h$ . The amplitude of the longest growing instability is denoted by  $\delta$ .

$$\frac{d\delta}{dt} = 2\sqrt{\beta Atg\delta}, \quad (5)$$

145 where  $\beta$  is a dimensionless constant growth parameter,  $At = \frac{\rho - \rho_f}{\rho + \rho_f}$  is the  
 146 Atwood number, and  $\rho$  is the bulk density of the plume defined as  $\rho =$   
 147  $\alpha_f \rho_f + \alpha_p \rho_p$  where  $\alpha_f$  is the volume fraction of the fluid. In this work,  
 148  $\beta = 0.03$  which is within the range of values estimated by experimental  
 149 and numerical techniques (Dimonte and Schneider, 2000; Dimonte et al.,  
 150 2004). This expression can be readily integrated to provide an expression for  
 151  $\tau_{\text{collective}}$ , given by (Youngs, 1984)

$$\tau_{\text{collective}} = \sqrt{\frac{\delta}{\beta Atg}}. \quad (6)$$

152 Note that the initial condition  $\delta(t = 0) = 0$  has been applied here. Although  
 153  $t = 0$  is supposed to be the point at which the flow reaches self-similarity

154 (that is, when the flow behaviour appears the same on any scale) such that  
 155 the initial condition becomes  $\delta(t = 0) = \delta_0$  for some  $\delta_0 > 0$ , this work chooses  
 156  $t = 0$  to correspond to the very start of the numerical simulation. This choice  
 157 was shown *a posteriori* to still provide a consistently close approximation to  
 158 the growth rate of the plumes across all simulations, even during very early  
 159 times. Furthermore, this choice was made in order to be consistent with the  
 160 expression for  $\tau_{\text{individual}}$  and to avoid any ambiguity in deciding exactly when  
 161 the flow becomes self-similar.

162 Taking the ratio of (2) and (6) yields the new dimensionless number

$$B_{vi} = \frac{18h\mu_f}{d_p^2} \sqrt{\frac{\alpha_p\beta}{(\rho + \rho_f)(\rho_p - \rho_f)\delta g}}. \quad (7)$$

163 It should be emphasised that this dimensionless quantity assumes that the  
 164 ambient fluid is incompressible, and that individual particle settling is con-  
 165 trolled by the balance between gravitational forces and viscous drag, whereas  
 166 plume growth and descent (i.e. collective particle settling) is controlled by  
 167 the balance between gravitational forces and inertial drag. Additional mea-  
 168 sures can be derived for a compressible ambient fluid (Goldin, 2008), which  
 169 is important for scenarios in which particles with a high initial momentum  
 170 move through the atmosphere, and for different regimes of individual and col-  
 171 lective particle settling. For completeness, the  $B_{ii}$  measure appropriate for  
 172 very coarse-grained particles that settle individually at Reynolds numbers  
 173 much greater than unity, implying that the inertial drag force dominates vis-  
 174 cous drag, is presented in Appendix B. However, this measure is not tested  
 175 here.

### 176 3. Numerical Simulations

177 To determine the ability of  $B_{vi}$  and  $B_{vv}$  to predict plume onset, a suite of  
178 two-phase numerical simulations of particle settling in water was performed  
179 using a multiphase computational fluid dynamics code called Fluidity, vary-  
180 ing the particle diameter and constant particle mass flux (into the water from  
181 above) over a range that encompassed the laboratory particle settling exper-  
182 iments of Carey (1997). The size of the water tank in the simulations was  
183  $0.3 \text{ m} \times 0.3 \text{ m} \times 0.7 \text{ m}$ , replicating the geometry of Carey’s experiments.

184 Initially, no particles were present in the domain, except along the surface  
185 where random perturbations in the particle volume fraction were introduced  
186 such that  $10^{-7} \leq \alpha_p \leq 10^{-5}$ . This essentially ‘seeded’ instabilities in the  
187 growing particle-water layer so that plumes could form. For numerical rea-  
188 sons,  $\alpha_p$  was bounded below by a value of  $10^{-7}$  instead of zero to avoid  
189 singularities in the system of linear equations. The velocity of both phases,  
190 denoted  $\mathbf{u}_f$  and  $\mathbf{u}_p$  respectively, was set to  $\mathbf{0} \text{ ms}^{-1}$  (where  $\mathbf{0}$  is the zero vector)  
191 at  $t = 0 \text{ s}$ . Throughout the simulations, no-normal flow conditions  $\mathbf{u}_f \cdot \mathbf{n} = 0$   
192 and  $\mathbf{u}_p \cdot \mathbf{n} = 0$  (where  $\mathbf{n}$  is the normal vector) were enforced along each bound-  
193 ary of the domain to prevent the fluid and particles from exiting. Particles  
194 entered the domain through the top boundary at a constant user-specified  
195 mass flux rate (defined later).

196 The following physical parameters were used and remained constant through-  
197 out all simulations:  $\rho_p = 2,340 \text{ kgm}^{-3}$ ,  $\rho_f = 1,000 \text{ kgm}^{-3}$ ,  $\mu_f = 0.001 \text{ Pas}$ ,  
198 and  $g = 9.8 \text{ ms}^{-2}$ . The particle phase was assumed to be inviscid such that  
199  $\mu_p = 0 \text{ Pas}$ . The range of mass flux was  $2.50 \times 10^{-4} - 6.11 \times 10^{-4} \text{ kgm}^{-2}\text{s}^{-1}$   
200 (the range determined for the eruption of Mount St Helens on 18 May 1980

201 (Sarna-Wojcicki et al., 1981; Scheidegger et al., 1982; Carey, 1997)), and  $d_p$   
202 ranged between 20 and 64  $\mu\text{m}$  as per the experiments by Carey (1997). In  
203 total, four different mass fluxes and six different particle diameters within  
204 these ranges were chosen, detailed in Table 1.

205 The domain was discretised using an unstructured mesh of solution nodes,  
206 composed of triangular and tetrahedral elements in two and three dimensions  
207 respectively, produced by Gmsh (Geuzaine and Remacle, 2009). The char-  
208 acteristic element length was fixed at 0.0025 m, except in the preliminary  
209 three-dimensional simulation mentioned in the next paragraph which used  
210 mesh adaptivity (Piggott et al., 2008) to optimise the mesh throughout the  
211 simulation and place high resolution only where necessary in order to reduce  
212 computational costs; in this case, the upper and lower bounds on the ele-  
213 ment length were set to 0.1 m and  $10^{-5}$  m respectively (Jacobs, 2013). The  
214 spatial discretisation of the model equations was performed using a Galerkin  
215 finite element method for the continuity and momentum equations, and a  
216 control volume method for the volume fraction fields (Jacobs et al., 2013; Ja-  
217 cobs, 2013). The implicit backward Euler method was used for the temporal  
218 discretisation, in conjunction with an adaptive time-stepping scheme which  
219 maximised the time-step subject to a Courant number of 0.5. All simulations  
220 were performed until  $t = 600$  s, which was enough time for plumes to form  
221 for all combinations of particle diameters and mass fluxes.

222 To establish any possible effect of problem geometry on plume formation,  
223 both 2D and 3D simulations were first performed using  $d_p = 48 \mu\text{m}$  and  
224 a mass flux of  $4.72 \times 10^{-4} \text{ kgm}^{-2}\text{s}^{-1}$  (see Figure 2). In both cases, initial  
225 particle settling happened individually at the appropriate Stokes' law veloc-

Reference	Mass flux ( $\text{kgm}^{-2}\text{s}^{-1}$ )	$d_p$ ( $\mu\text{m}$ )
A1 – A6	$2.50 \times 10^{-4}$	26, 32, 40, 48, 56, 64
B1 – B6	$3.61 \times 10^{-4}$	26, 32, 40, 48, 56, 64
C1 – C6	$4.72 \times 10^{-4}$	26, 32, 40, 48, 56, 64
D1 – D6	$6.11 \times 10^{-4}$	26, 32, 40, 48, 56, 64
E1 – E2	$4.72 \times 10^{-4}$	26, 48

Table 1: Reference table for the 24 simulations in the numerical parameter study (A1 – A6, B1 – B6, C1 – C6 and D1 – D6), and for the experimental data points (E1 – E2).

ity, forming a uniform layer of thickness  $h$ . Eventually, instabilities at the base of this layer grew into plumes that settled to the base of the tank much more rapidly than the initial, individual particle settling speed. The layer thickness, particle volume fraction and time at the onset of plume formation differed by less than 10% between the 2D and 3D simulations. Therefore, for computational expedience, only 2D simulations were performed for the remaining particle diameters and mass fluxes. Note that for some simulations the nominal  $0.3 \text{ m} \times 0.7 \text{ m}$  domain was extended in the vertical direction to accommodate plumes that grew longer than 0.7 m.

To quantify the conditions at the onset of plume formation and hence evaluate the accuracy of the dimensionless quantities for predicting plume onset ( $B$  values), the values of  $h$  and  $\delta$  needed to be extracted from the simulation results. By assuming that particles in the layer settle under Stokes' law (at least until plumes have formed), the layer thickness  $h$  was consistently found using the Stokes' law settling velocity multiplied by the time at the onset of pluming. This assumption was tested *a posteriori* and shown to be

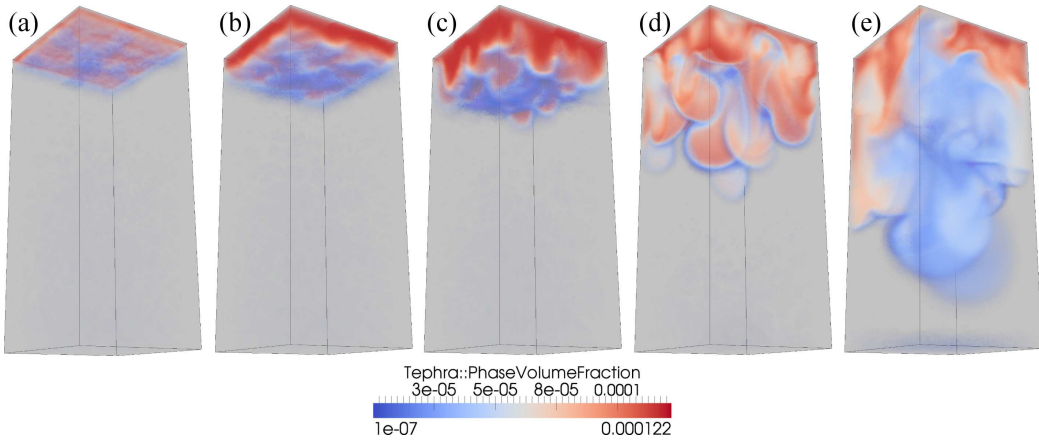


Figure 2: Three-dimensional simulation of particles settling through a tank of water at  $t =$  (a) 10, (b) 30, (c) 50, (d) 80, and (e) 120 s. All visualisations show the whole  $0.3 \text{ m} \times 0.3 \text{ m} \times 0.7 \text{ m}$  domain.

242 valid across all simulations. At a given time, the head of the growing plume  
 243 of greatest amplitude  $\delta$  was calculated by finding the lowermost position of  
 244 the  $10^{-5}$  particle volume fraction contour. This contour was chosen *a pos-*  
 245 *teriori* as a sensible lower bound on the volume fraction of particles in the  
 246 layer. The amplitude  $\delta$  was then computed by taking the difference between  
 247 the depth of the layer and the position of the plume head.

248 As one might expect, there is a certain amount of ambiguity involved  
 249 when deciding when an instability is developed enough to be defined as a  
 250 plume. Since the amplitude of a growing instability is known to be a function  
 251 of the layer thickness (Manville and Wilson, 2004), this work defined the onset  
 252 of pluming as the moment when  $\delta = h$ . The validity of this choice is discussed  
 253 in Section 5. At this time, the quantities  $h$  and  $\alpha_p$  were determined, and the  
 254 dimensionless numbers  $B_{vv}$  and  $B_{vi}$  were calculated.

255 **4. Experimental Data**

256 The experiments performed by Carey (1997) used ultrasound imaging to  
 257 track particle positions, which did not permit the accurate measurement of  
 258 the parameters  $h$  and  $\alpha_p$ . Some assumptions were therefore made in order  
 259 to calculate estimates for experimental values of  $B_{vv}$  and  $B_{vi}$  for compari-  
 260 son with the numerical simulations. Assuming that particles in the growing  
 261 particle-laden layer settled at Stokes' law velocity,  $\mathbf{u}_p = \mathbf{u}_{\text{stokes}}$ , the distance  
 262 the particles had travelled at the time of plume onset,  $t_{\text{onset}}$ , provided an  
 263 approximation for the layer thickness:

$$h \approx |\mathbf{u}_{\text{stokes}}| t_{\text{onset}}. \quad (8)$$

264 Furthermore, assuming the volume fraction of particles in the layer was uni-  
 265 form (because of the constant mass flux), and the total volume of the layer  
 266 (including the water) was given by

$$V_{\text{layer}} = hA, \quad (9)$$

267 where  $A$  is the area through which particles fluxed in ( $A = 0.9 \text{ m}^2$  for these  
 268 particular experiments), then

$$\alpha_p = \frac{V_p}{V_{\text{layer}}}, \quad (10)$$

269 where  $V_p$  is the volume occupied by the particles. The mass flux of particles  
 270 per unit area,  $\dot{M}_p$ , was used to calculate the volumetric flux per unit area  $\dot{V}_p$   
 271 using

$$\dot{V}_p = \frac{\dot{M}_p}{\rho_p}. \quad (11)$$

272 From this, the volume of the particle phase in the layer was calculated as

$$V_p = A\dot{V}_p t_{\text{onset}}, \quad (12)$$

273 and the volume fraction followed from

$$\alpha_p = \frac{V_p}{V_{\text{layer}}}. \quad (13)$$

274 Carey (1997) noted that plumes had formed after approximately 30 s in  
 275 experiment 96-5 which used 20–32  $\mu\text{m}$  diameter particles, and after approx-  
 276 imately 60 s in experiment 96-1 which used 32–64  $\mu\text{m}$  diameter particles.  
 277 These times were used as approximations to  $t_{\text{onset}}$  for the purpose of estimat-  
 278 ing  $B_{vv}$  and  $B_{vi}$ , giving two data points for each measure, denoted E1 and  
 279 E2 (see Table 1).

## 280 5. Evaluation of the Measures

281 The results from the parameter study reinforced the expected relation-  
 282 ship between the particle diameter, mass flux and layer instability. Smaller  
 283 particle sizes decrease the time required for plume onset because the slower  
 284 Stokes' law settling results in a higher average particle concentration in the  
 285 near-surface layer. This behaviour was also witnessed in the experiments  
 286 performed by Carey (1997) where, for two ranges of particle diameter (20–  
 287 32  $\mu\text{m}$  and 32–64  $\mu\text{m}$ ), there was a difference of approximately 30 s in the  
 288 onset time. Similarly, a higher particle flux also causes a denser build-up of  
 289 particles in the growing layer, further encouraging plume formation.



290 As expected, the calculated values of  $B_{vv}$  and  $B_{vi}$ , shown in Figures 3a  
 291 and 3b respectively, are all greater than unity since the parameters  $h$  and  
 292  $\alpha_p$  were measured at the point where plumes formed. Most importantly, the  
 293 values from the measure  $B_{vi}$  (which assumes that collective settling is slowed  
 294 by inertial drag) lay consistently on a particular contour ( $\sim 1.2$ ), whereas the  
 295 values from the measure  $B_{vv}$  (which assumes that collective settling obeys  
 296 Stokes' law and is therefore slowed by viscous drag) did not. In theory,  
 297 one would expect plume onset to occur at a constant  $B$  value because the  
 298 definition of when a plume has formed does not change between simulations.  
 299 By correctly describing the drag on the plumes, the  $B_{vi}$  measure robustly  
 300 estimated the timescale of collective particle settling, even when the system  
 301 became more and more unstable and non-linear as a result of increasing  
 302 particle diameter and flux rate. In contrast, the  $B_{vv}$  measure grossly under-  
 303 estimated the timescale of collective settling.

304 Plume formation in every numerical simulation was robustly predicted  
 305 by a  $B_{vi}$  value of  $\approx 1.2$ . This threshold value for  $B_{vi}$  was derived by defining  
 306  $\delta = h$  as the condition for the onset of pluming. While the coefficient of  $h$  in  
 307 this expression was chosen arbitrarily, other coefficients close to unity would  
 308 still result in a consistent plume-onset  $B_{vi}$  value, but the exact threshold  
 309 value would differ from 1.2. This is because for any  $\delta$  proportional to  $h$  the  
 310 ratio of timescales between individual and collective particle settling is the  
 311 same to within a constant factor for a given plume scenario.

312 Although the estimated experimental data points do not follow an exact  
 313 contour for either measure, the two experimental  $B_{vi}$  values are much more  
 314 consistent than the two  $B_{vv}$  values. The small discrepancy in the  $B_{vi}$  values

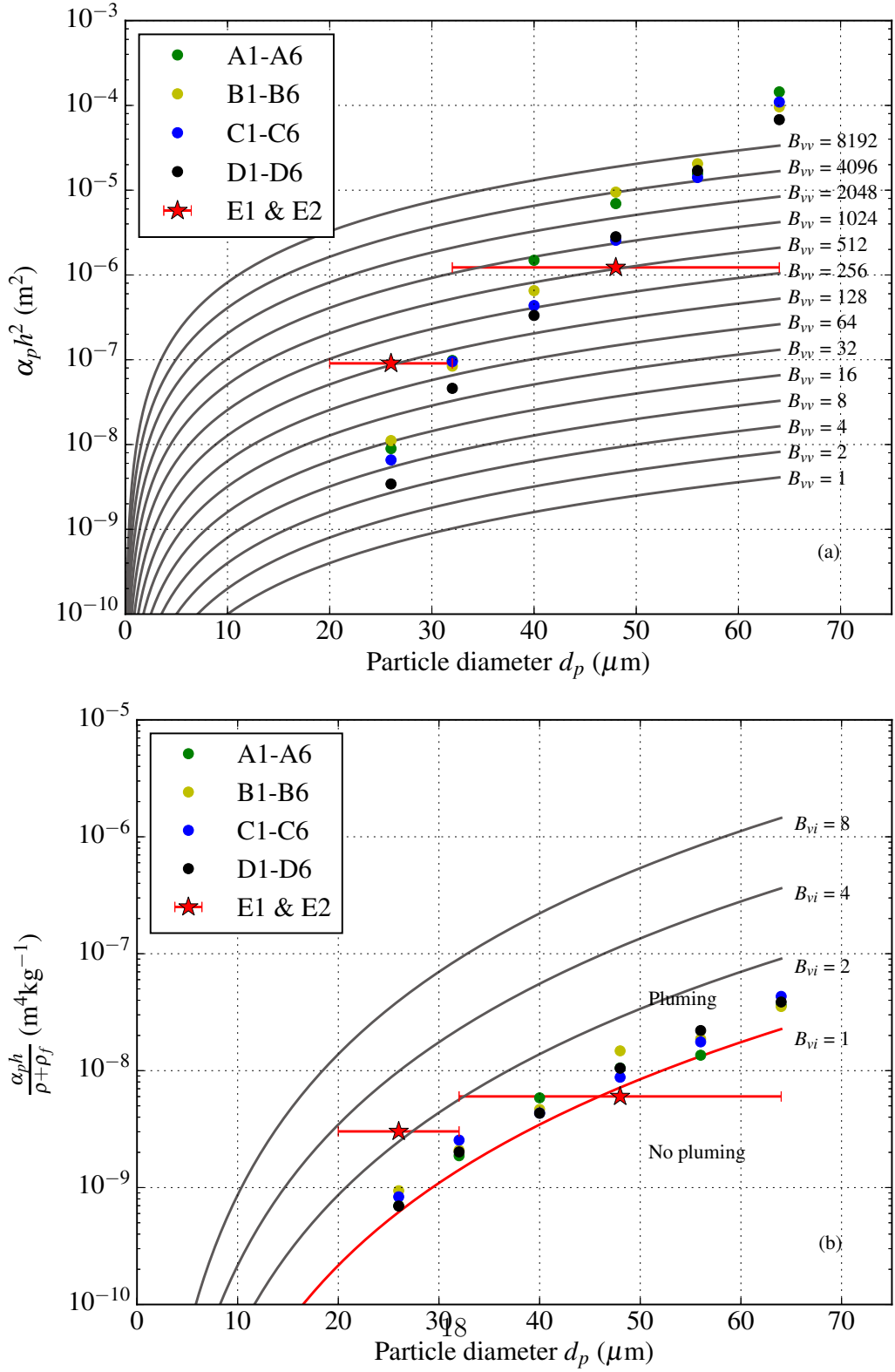


Figure 3:  $B_{vv}$  and  $B_{vi}$  results (presented in (a) and (b), respectively) using  $\delta = h$ . Several contours of  $B_{vv}$  and  $B_{vi}$  are given by solid lines. Due to the differences in the formulations of the measures, different quantities were considered along the  $y$ -axis. These quantities are related to the input mass flux of particles across the water surface, and are functions of the volume fraction of particles  $\alpha_p$ , the layer thickness  $h$ , the fluid density  $\rho_f$  and the bulk density  $\rho$ .

315 is easily explained by the ambiguity in plume onset time, which could not be  
316 accurately determined from the ultrasound images. Moreover, the approxi-  
317 mate time of plume onset in the experiments does not necessarily correspond  
318 to the point at which  $\delta = h$ , thus potentially introducing further uncertainty  
319 in the experimental estimates.

320 At earlier times when plumes had not formed (i.e. before the point at  
321 which  $\delta = h$ ), the values of  $B_{vv}$  and  $B_{vi}$  were also calculated to show that  
322  $B_{vi}$  is less than unity, while  $B_{vv}$  is much greater than unity, demonstrating  
323 the inaccuracy of the measure that assumes collective settling is slowed by  
324 viscous drag. Simulation C4 is considered here for demonstration purposes  
325 because the relatively low mass flux and large particle diameter favoured the  
326 stability of the growing layer. Figure 2a shows the particle volume fraction  
327 at  $t = 10$  s. Clearly plumes had not formed at this point, and only very small  
328 initial perturbations (with  $\delta \ll h$ ) are present along the base of the layer. It  
329 was found that all particles were still travelling at their Stokes' law velocity  
330 at this point in time. The  $B_{vi}$  measure yielded a value less than unity ( $\sim 0.3$ ),  
331 correctly implying that individual particle settling dominated the dynamics.  
332 This also agrees with an estimated  $B_{vi}$  value of  $\sim 0.37$  (see the contour plot  
333 in Figure 4), computed using estimates for the volume fraction and layer  
334 thickness as described in Section 4. However, a  $B_{vv}$  value of  $\sim 15$  implied  
335 that plumes were already well into the growth stage. This demonstrates  
336 that the measure which assumes collective settling is slowed by viscous drag  
337 grossly under-estimates the timescale of plume growth and descent. On the  
338 other hand, the new measure which assumes collective settling is slowed by  
339 inertial drag is able to more accurately measure the tendency for plumes to

340 form at early times.

### 341 *5.1. Alternative Formulation*

342 The measures in their current form require knowledge about the state of  
343 the layer, in particular the layer thickness, the volume fraction of particles  
344 within it, and (in the case of  $B_{vi}$ ) the amplitude of the growing instabilities.  
345 Given this information, the non-dimensional number can be used to deter-  
346 mine whether plumes will form. These quantities have to be estimated in  
347 practice since measuring them after or during an eruption event would be  
348 infeasible or impossible. However, as an alternative to calculating  $B_{vv}$  and  
349  $B_{vi}$  directly from the state of the system, the measures can be re-formulated  
350 in terms of a critical layer thickness, denoted  $h_{\text{crit}}$ . For pluming to occur,  
351 the value of  $h$  must satisfy  $h_{\text{crit}} < h < H$ , where  $H$  is the height of the  
352 water column. The thickness of the layer  $h$  can be estimated throughout  
353 time using Stokes' law since the particles within the layer settle individually.  
354 Furthermore, the critical value is expressed only in terms of the volumetric  
355 influx of particles and the particle diameter, such that the measures can be  
356 useful regardless of whether the exact values for  $h$ ,  $\delta$  and  $\alpha_p$  are known.

357 By using a similar technique to that used when estimating the values  
358 of  $B_{vv}$  and  $B_{vi}$  from the experiments of Carey (1997), an expression for  $\alpha_p$   
359 (assumed to be constant and uniform in the layer) was formulated:

$$\alpha_p = \frac{\dot{V}_p}{|\mathbf{u}_{\text{stokes}}|}, \quad (14)$$

360 where  $\dot{V}_p$  is the volumetric flux (per unit area) and  $\mathbf{u}_{\text{stokes}}$  is the Stokes' law  
361 velocity. This was then used to re-arrange both measures in terms of  $h$ , and

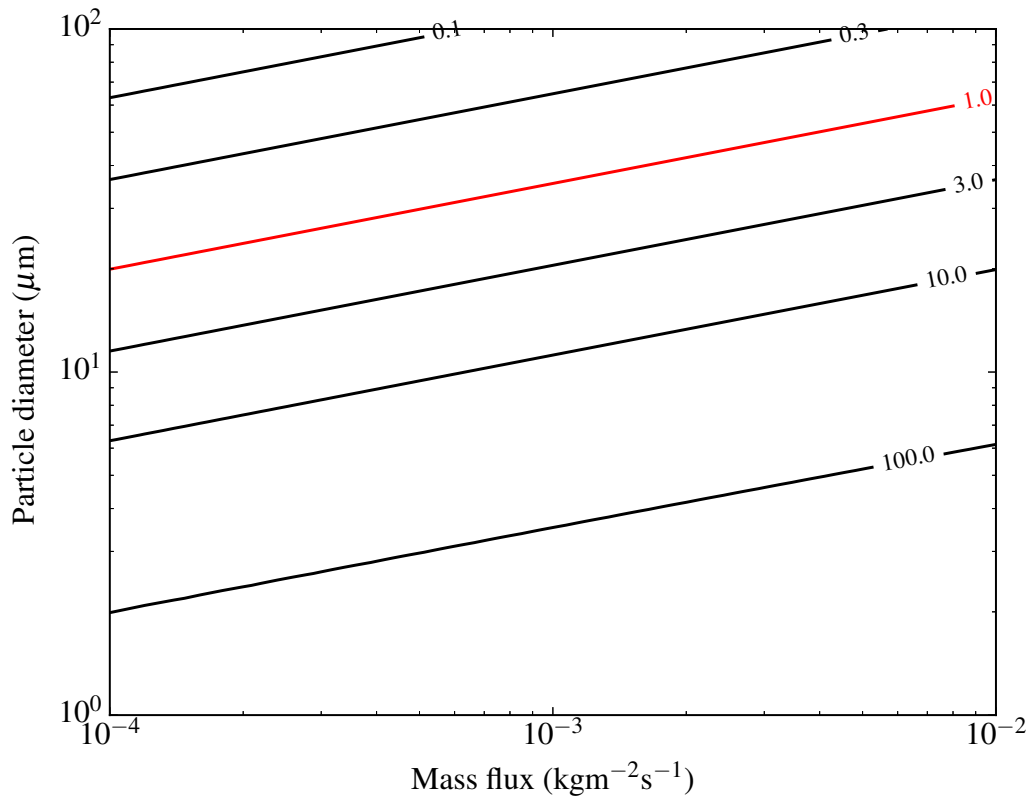


Figure 4: A contour plot of  $B_{vi}$ , computed using estimates for the particle volume fraction and layer thickness, at  $t = 10$  s for various mass fluxes and particle diameters. The line  $B_{vi} = 1$  is highlighted in red. The plot further reinforces the finding that higher mass flux and/or smaller particle diameter encourages plume formation.

362 by setting  $B$  equal to unity,  $h_{\text{crit}}$  was derived. For  $B_{vv}$ , the value of  $h_{\text{crit}}$  is  
 363 given by

$$h_{\text{crit}} = \sqrt{\frac{d_p^4 g (\rho_p - \rho_f)}{18 \mu_f \dot{V}_p}}. \quad (15)$$

364 On the assumption that  $\rho \approx \rho_f$  in (7) because  $\alpha_f \approx 1$ , and that plumes have  
 365 formed when  $\delta = h$ , the value of  $h_{\text{crit}}$  for  $B_{vi}$  is given by

$$h_{\text{crit}} = \left( \frac{2\rho_f}{\beta} \right) \left( \frac{(\rho_p - \rho_f)^2 g^2 d_p^6}{5832 \mu_f^3 \dot{V}_p} \right). \quad (16)$$

366 Figures 5a and 5b illustrate the relationship between the particle diameter  
 367 and the critical value  $h_{\text{crit}}$  for both measures, for all volumetric flux rates  
 368 considered in this paper. While all the values of  $h_{\text{crit}}$  and  $h$  were such that  
 369  $h_{\text{crit}} < h < H$  was satisfied, a measure could only be considered meaningful  
 370 and useful if the expected  $h_{\text{crit}}$  values *consistently* agree with the actual values  
 371 of  $h$  at the time of plume formation (i.e. if the values of  $h_{\text{crit}}$  run parallel  
 372 to all the layer thicknesses determined from the numerical simulations). As  
 373 demonstrated in Figure 5a, this is clearly not the case for the  $B_{vv}$  measure  
 374 whose values for  $h_{\text{crit}}$  start to diverge from the theoretical prediction. In  
 375 contrast, the values of  $h_{\text{crit}}$  obtained from the  $B_{vi}$  measure, which takes into  
 376 account the inertial drag acting on the particles, run parallel to all the data  
 377 points as shown in Figure 5b. This further demonstrates the robustness and  
 378 applicability of the  $B_{vi}$  measure when the exact values of  $h$ ,  $\delta$  and  $\alpha_p$  are not  
 379 readily available.

380 Since the volumetric flux and particle diameter are two quantities that  
 381 are often known during or after an eruption event, a plot of  $h_{\text{crit}}$  (for the  $B_{vi}$   
 382 measure) against the volumetric flux for various particle diameters is given

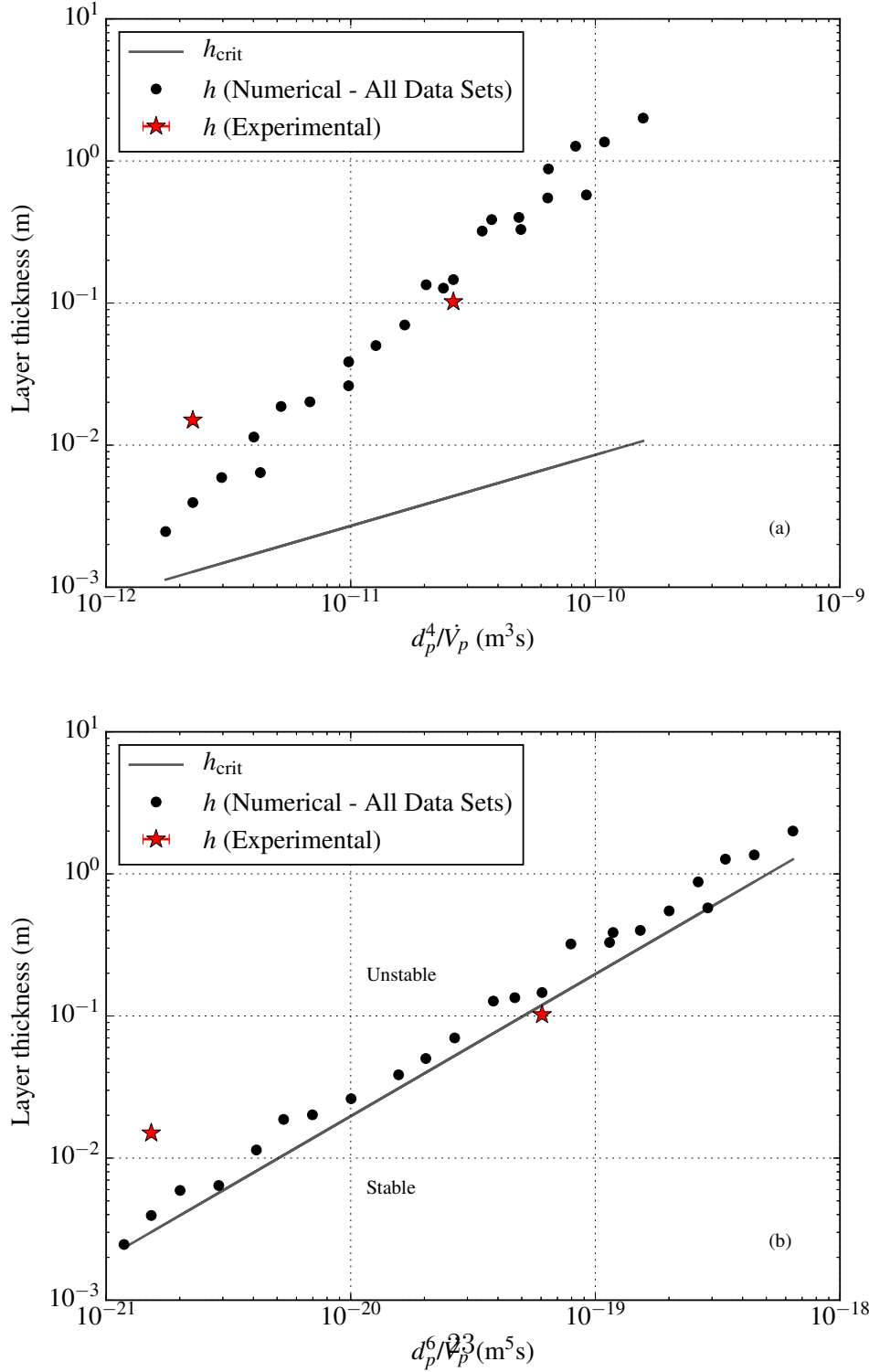


Figure 5: Plots of layer thickness  $h$  and the critical layer thickness  $h_{crit}$  for (a) the  $B_{vv}$  measure, and (b) the  $B_{vi}$  measure. The quantities  $d_p^4 / \dot{V}_p$  and  $d_p^6 / \dot{V}_p^3$  (which are functions of the particle diameter  $d_p$  and the volumetric flux  $\dot{V}_p$ ) were considered along the  $x$ -axis to allow all data points to be plotted against a single  $h_{crit}$  line (rather than having a separate line for each volumetric flux or particle diameter).

383 in Figure 6 for reference. This also helps to demonstrate once again how  
384 increasing the volumetric flux rate and/or decreasing the particle diameter  
385 makes the system increasingly unstable, as shown by the smaller  $h_{\text{crit}}$  values.

386 In the context of subaqueous explosive volcanic eruptions, in which the  
387 near-surface layer is formed from particles being forced upwards, high inertia  
388 and buoyancy are necessary to sustain particle ascent before the eruption  
389 column spreads out laterally along the water’s surface (White, 2000; White  
390 et al. (2003), pp. 9–12). If the mass flux of particles at the surface is greater  
391 than that typically achieved by atmospheric ash fallout, then Figure 6 im-  
392 plies that a much thinner layer will be required to initiate plume onset (for  
393 a given particle diameter). It is also important to note that, since plume  
394 size is related to  $h$  and therefore  $h_{\text{crit}}$ , any eruption column that is unable  
395 to sustain its upward motion and is thicker than  $h_{\text{crit}}$  will collapse as a den-  
396 sity current/plume, regardless of whether the ash particles reach the water’s  
397 surface.

## 398 *5.2. Including Additional Particle Sizes*

399 All the simulations presented thus far have considered multiphase flows  
400 comprising ash particles of the same diameter, known as monodisperse flows.  
401 Such flows are certainly an idealisation since real volcanic ash particles can  
402 vary greatly in diameter (Rose and Durant, 2009). The inclusion of addi-  
403 tional particle phases each defined by a different particle diameter, forming  
404 a so-called polydisperse flow (Crowe et al. (1998), p. 37), can therefore sig-  
405 nificantly alter the behaviour and enhance the realism of the results. To  
406 investigate the effect of multiple particle diameters on the transport of ash in  
407 water, and to determine how the theoretical measures defined earlier should



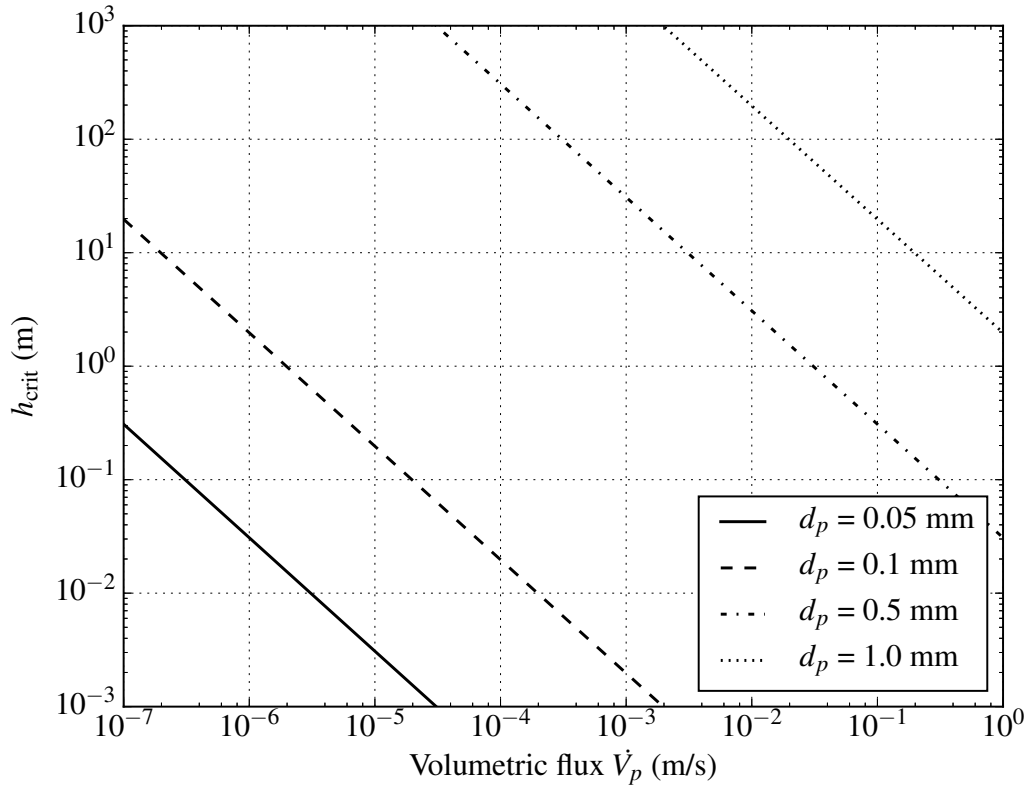


Figure 6: Plot of critical layer thickness  $h_{\text{crit}}$  (from the  $B_{vi}$  measure) against volumetric flux  $\dot{V}_p$ , for various particle diameters. The volumetric fluxes (per unit area) used in the experiments by Carey (1997) are of  $O(10^{-7})$   $\text{ms}^{-1}$ .

408 be modified to support polydisperse flows, a three-phase simulation was set  
409 up in Fluidity which extended the earlier two-phase simulations.

410 Two particle diameters  $d_{p_1} = 26 \mu\text{m}$  and  $d_{p_2} = 48 \mu\text{m}$  in the range of  
411 those considered by Carey (1997) were employed. Both particle phases had  
412 the same density of  $2,340 \text{ kgm}^{-3}$ . A previously used (total) mass flux of  
413  $4.72 \times 10^{-4} \text{ kgm}^{-2}\text{s}^{-1}$  was chosen and remained constant, but was divided  
414 equally between the two particle phases such that each one fluxed in at  
415  $2.36 \times 10^{-4} \text{ kgm}^{-2}\text{s}^{-1}$ . All other aspects of the set-up remained the same as  
416 the earlier two-phase simulations.

417 After performing the simulation, it was found that at early times the 26  
418  $\mu\text{m}$  particles and 48  $\mu\text{m}$  particles behaved just like their monodisperse ver-  
419 sions. That is, Stokes' law settling ensued once the particles first entered  
420 the water tank, as shown by the good agreement with the Stokes' law veloc-  
421 ities of  $0.00049 \text{ ms}^{-1}$  and  $0.00168 \text{ ms}^{-1}$  (for  $d_{p_1} = 26\mu\text{m}$  and  $d_{p_2} = 48\mu\text{m}$ ,  
422 respectively) in Figure 7. The near-surface layer of particles that formed  
423 was essentially divided up into two parts as a result of the different settling  
424 velocities; the smaller 26  $\mu\text{m}$  particles formed their own relatively thin and  
425 more concentrated 'sub-layer', while the larger 48  $\mu\text{m}$  particles were able to  
426 overtake the 26  $\mu\text{m}$  particles and form a thicker layer as shown in Figures  
427 8a and 8f. After the initial growth of the layer (as a whole), plumes formed  
428 from the thinner sub-layer layer of 26  $\mu\text{m}$  particles while the layer of 48  $\mu\text{m}$   
429 particles remained almost uniform in shape, as shown in Figures 8b and 8g.  
430 This occurred at approximately the same time as the monodisperse 26  $\mu\text{m}$   
431 simulation, but the plumes grew at a slightly slower rate which may have  
432 been the result of the presence of larger particles that typically increase the

433 stability of the system. Despite this small difference, the dynamics of the each  
434 particle phase were qualitatively similar to the monodisperse simulations of  
435 26  $\mu\text{m}$  and 48  $\mu\text{m}$  particles up until this point.

436 The plumes of 26  $\mu\text{m}$  particles that grew from the thin sub-layer eventu-  
437 ally started to influence the dynamics of the other part of the layer composed  
438 solely of 48  $\mu\text{m}$  particles, which were still settling at near-Stokes' law veloc-  
439 ity, by entraining them. The growth of any small instabilities in the 48  $\mu\text{m}$   
440 particle sub-layer was essentially over-ridden by the presence of the plumes  
441 of smaller particles. Therefore, while the two particle phases behaved almost  
442 independently at early times, in a similar manner to the separate monodis-  
443 perse versions, it was the smaller particles in the system that influenced the  
444 dynamics of the whole polydisperse system at later times.

445 As the plumes continued to grow and entrain material the two particle  
446 phases became strongly coupled to one another (as shown by the similar  
447 velocity profiles in Figure 7 at late times). This resulted in their volume  
448 fraction fields becoming almost identical in shape (see Figures 8c–e and 8h–  
449 j). The plumes were of a comparable length to those composed solely of  
450 26  $\mu\text{m}$  particles, although they appeared to be a few millimetres thicker as  
451 a result of the larger particles. Furthermore, as the plumes descended, the  
452 smaller particles tended to move a small distance away from the surface of  
453 the plumes and instead drift behind a thin outer layer of larger particles  
454 because of drag reduction effects. This suggests that a degree of sorting by  
455 settling velocity takes place during collective particle descent and deposition,  
456 which is commonly seen in the real world (Carey, 1997; Manville and Wilson,  
457 2004).

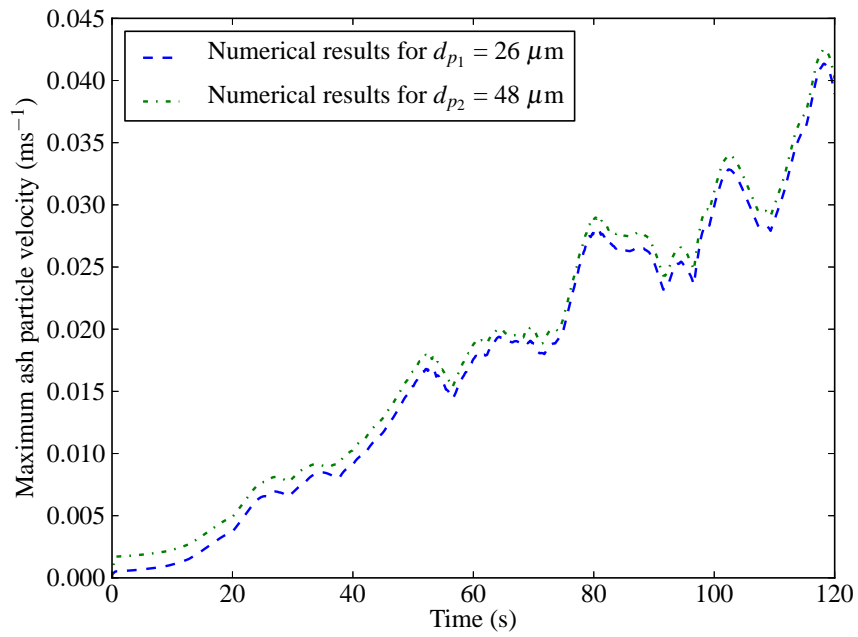


Figure 7: Maximum velocity of ash particles in each particle phase against time, with  $d_{p1} = 26 \mu\text{m}$  and  $d_{p2} = 48 \mu\text{m}$ , in a two-dimensional polydisperse simulation of the experiments by Carey (1997).

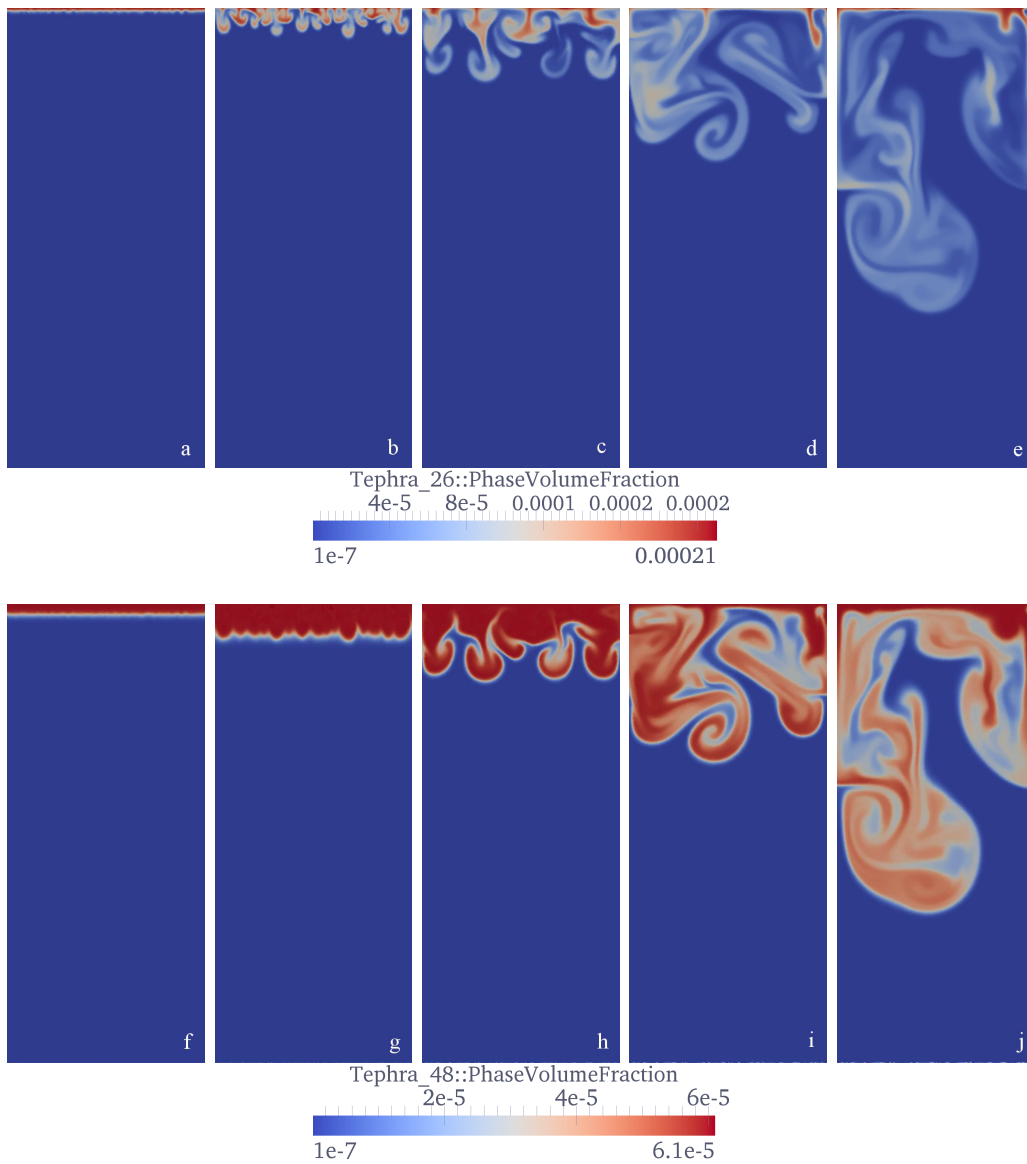


Figure 8: Visualisation of a three-phase, polydisperse ash settling simulation in Fluidity, with  $d_{p_1} = 26 \mu\text{m}$  (top row) and  $d_{p_2} = 48 \mu\text{m}$  (bottom row), at  $t = 10, 30, 50, 80$  and  $120$  s (from left to right). The volume fraction of the particle phase ( $\alpha_p$ ) is shown; warmer colours represent a higher volume fraction. All visualisations show the whole  $0.3 \text{ m} \times 0.7 \text{ m}$  domain.

Reference	Mass flux ( $\text{kgm}^{-2}\text{s}^{-1}$ )	$d_{p1}$ ( $\mu\text{m}$ )	$d_{p2}$ ( $\mu\text{m}$ )
P1	$4.72 \times 10^{-4}$	20	26
P2	$4.72 \times 10^{-4}$	26	32
P3	$4.72 \times 10^{-4}$	32	48
P4	$4.72 \times 10^{-4}$	48	64

Table 2: Reference table for the polydisperse simulations in the numerical parameter study.

458 Since the theoretical measures of the tendency for plume formation de-  
459 pend on the particle diameter, it is worth considering how the measures  
460 should be modified to support multiple particle diameters. To this end, four  
461 additional polydisperse simulations were performed. The particle diameters  
462 chosen covered the range used by Carey (1997) and are detailed in Table 2.

463 For the purpose of computing the dimensionless quantities  $B_{vv}$  and  $B_{vi}$ ,  
464 plumes were once again said to have formed when  $\delta = h$ . However, the  
465 calculation of the layer thickness through Stokes' law (and also the calculation  
466 of  $\tau_{\text{individual}}$ ) needs to be considered carefully. It has already been shown here  
467 that the dynamics of ash settling in water can be affected heavily by the  
468 end members of the particle size range, so simply using an average for  $d_p$   
469 when computing both the layer thickness and  $\tau_{\text{individual}}$  may not be accurate  
470 in general. It is also not appropriate to define the layer thickness as the  
471 maximum of the thicknesses of the two 'sub-layers' that form within the  
472 whole near-surface layer, because the thicker sub-layer (comprising larger  
473 particles) will eventually become entrained within the plumes growing from  
474 the shallow sub-layer (comprising smaller particles). It is because of this  
475 reason that using the Stokes' law settling velocity of the smaller particles

476 instead of the larger particles gave a good estimation of the layer thickness.  
477 Therefore, when computing  $h$  and  $\tau_{\text{individual}}$ ,  $d_p$  was chosen to be equal to  $d_{p1}$ .

478 The results from the parameter study of the polydisperse simulations are  
479 plotted in Figure 9 (for the  $B_{vi}$  measure only). Once again, the values for  $B_{vv}$   
480 did not lie consistently on a particular contour, whereas the measure that  
481 took into account the balance between gravitational forces and inertial drag  
482 ( $B_{vi}$ ) did. Moreover, this particular contour was approximately the same as  
483 the one from the monodisperse simulations, suggesting that the measures are  
484 robust even when multiple particle sizes are considered. Note also that only  
485 the definition of the layer thickness and  $d_p$  (in  $\tau_{\text{individual}}$ ) needed to be treated  
486 carefully; the formulation of the dimensionless quantity itself did not need to  
487 be changed.

## 488 6. Discussion

489 By once again assuming that  $\rho \approx \rho_f$  in (7) because  $\alpha_f \approx 1$ , a useful  
490 property of (7) is that  $B_{vi}$  is a function of the product of  $h$  and  $\alpha_p$ , which is  
491 the volume of particles per unit area in the particle-laden water layer at the  
492 onset of pluming. Assuming that material reaching the sea or lake floor by  
493 plumes spreads laterally as it is deposited to form a semi-continuous layer  
494 of approximately uniform thickness, mass conservation implies that the final  
495 deposit should contain the same volume of particles per unit area as the  
496 original particle-water layer. Hence, (7) provides a measure of the tendency  
497 for plumes to form which can be calculated from the properties of the final  
498 deposit: the product of the volume fraction of particles in the deposit  $\alpha_{p,\text{deposit}}$   
499 and the deposit thickness  $h_{\text{deposit}}$ . Knowledge of the mass flux and duration

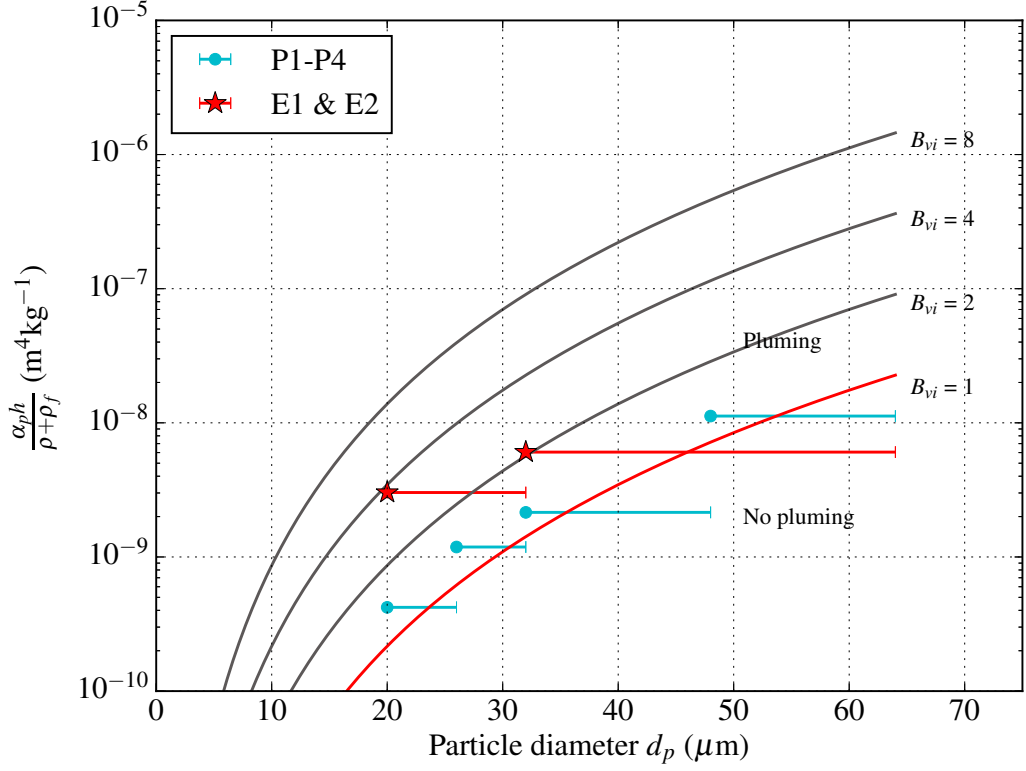


Figure 9:  $B_{vi}$  results from the four polydisperse simulations (points P1–P4, detailed in Table 2), using  $\delta = h$ . Several contours of  $B_{vi}$  are given by solid lines. Note that the  $x$ -coordinate of each numerical data point corresponds to the smallest particle diameter used in each polydisperse simulation, since this value is used to compute  $B_{vi}$ . As before, points E1 and E2 correspond to the experimental data; however, in light of the findings from the polydisperse simulations, the smallest particle diameters from the original experiments by Carey (1997) ( $d_p = 20 \mu\text{m}$  and  $d_p = 32 \mu\text{m}$ ) were used instead of the averages given in Table 1.



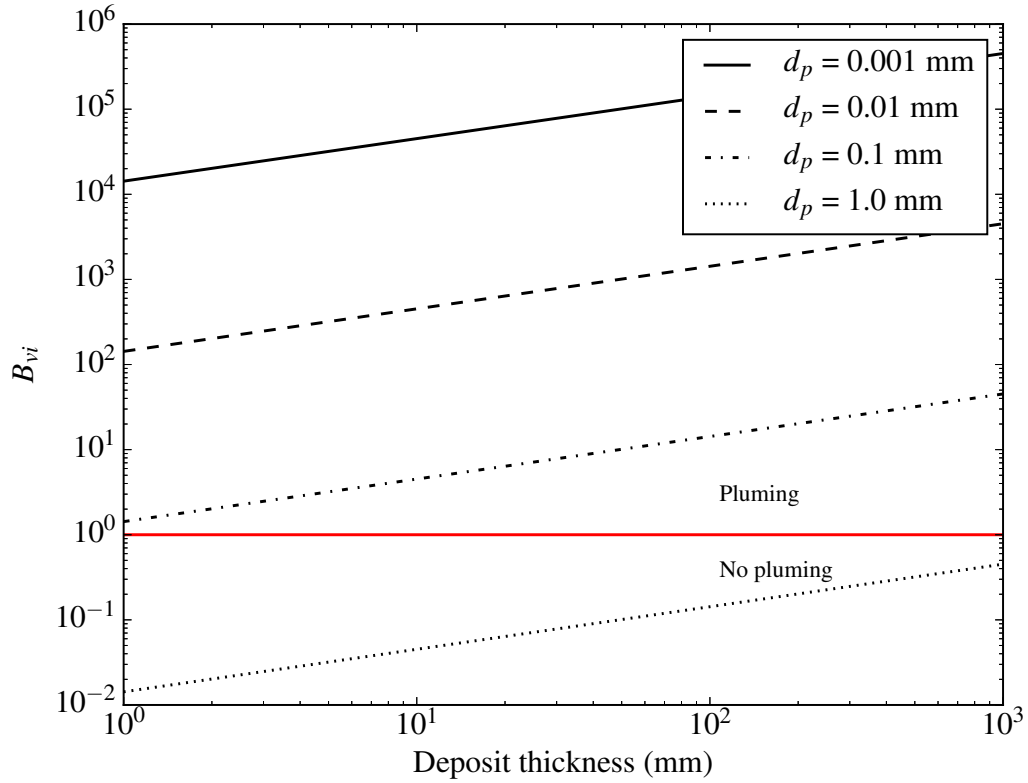


Figure 10:  $B_{vi}$  as a function of deposit thickness for various particle diameters. The horizontal red line represents  $B_{vi} = 1$  and indicates the threshold for stability; values of  $B_{vi} > 1$  imply that plume formation is likely.

500 are not required. Adopting this approach, Figure 10 shows how  $B_{vi}$  depends  
 501 on the deposit thickness for various particle diameters. A particle volume  
 502 fraction of 0.55 has been assumed for the final deposit, based on typical bulk  
 503 densities of compacted wet ash (Macedonio and Costa, 2012).

504 The plot shows that for particles smaller than 0.1 mm in diameter,  $B_{vi}$   
 505 is greater than unity for final deposits thicker than 1 mm, suggesting that  
 506 plume formation is expected in the formation of most benthic ash deposits,

507 particularly thick (single) deposits or those comprised of fine particles. On  
 508 the other hand, a larger particle diameter helps to stabilise the system and  
 509 prevent pluming. For  $d_p \geq 1$  mm,  $B_{vi}$  values are less than unity for the  
 510 range of deposit thicknesses considered, suggesting that in such cases the  
 511 Stokes' law settling velocity is high enough to prevent a concentrated particle-  
 512 water layer from building up near the surface and causing Rayleigh-Taylor  
 513 instabilities to develop.

514 The Reynolds number is a useful dimensionless quantity for determining  
 515 whether viscous or inertial drag effects dominate the dynamics. By defining  
 516 separate Reynolds numbers for individual particles and plumes, one can de-  
 517 cide which  $B$  measure is appropriate for a given geophysical scenario. These  
 518 Reynolds numbers are respectively defined as

$$\text{Re}_{\text{particle}} = \frac{\rho_f |\mathbf{u}_p| d_p}{\mu_f}, \quad (17)$$

519 and

$$\text{Re}_{\text{plume}} = \frac{\rho_f |\mathbf{u}_{\text{plume}}| d_{\text{plume}}}{\mu_f}, \quad (18)$$

520 where  $|\mathbf{u}_{\text{plume}}|$  and  $d_{\text{plume}}$  are the velocity and length scale of the plume.

521 The  $B_{vi}$  measure is appropriate in cases where  $\text{Re}_{\text{particle}}$  is small and  
 522  $\text{Re}_{\text{plume}}$  is typically much greater than unity, implying that individual particle  
 523 and plume settling are dominated by viscous and inertial drag, respectively.  
 524 For micrometre-sized ash particles settling in water with a velocity that obeys  
 525 Stokes' law, this is clearly the case for  $\text{Re}_{\text{particle}}$  (e.g.  $O(10^{-2})$  for the 48  $\mu\text{m}$   
 526 particles considered here). In contrast,  $\text{Re}_{\text{plume}} \gg 1$  as the plume diameter  
 527 and velocity is typically several times (or even several orders of magnitude)  
 528 larger than those of the individual particles, as shown by the numerical simu-

529 lations presented in this paper and the original experiments by Carey (1997).  
530 The measure  $B_{vi}$  is therefore appropriate in this case.

531 In addition to ash particles settling through bodies of water, the new  
532 measure  $B_{vi}$  may also be applicable to other geophysical processes which  
533 have the potential to form plumes. One example is the settling of volcanic  
534 ash through the atmosphere following an explosive volcanic eruption event.  
535 For small pyroclasts,  $Re_{\text{particle}}$  will still be less than unity (typically between  
536  $O(1)$  and  $O(10^{-5})$  for fine ash of the same size and a similar density to that  
537 considered here (Bonadonna et al., 1998)) if Stokes' law continues to hold,  
538 while the sheer diameter (tens to hundreds of metres) and settling velocity of  
539 the growing plumes results in  $Re_{\text{plume}}$  becoming large enough to imply that  
540 inertial drag forces dominate the plume's dynamics. However, it is important  
541 to note that the individual descent of larger pyroclasts will be controlled by  
542 inertial rather than viscous drag as a result of their size. Furthermore, unlike  
543 the particles settling in water, individual particles may begin their descent  
544 with a high inertia. This is certainly the case for impact ejecta re-entering  
545 the atmosphere, for example. In these cases, Stokes' law will no longer hold  
546 and the  $B_{ii}$  measure given in Appendix B may be more appropriate.

547 The process of crystals settling in a magma chamber is another example  
548 of where a different measure is necessary (Marsh, 1988). Here, the dynamics  
549 of the particles will obey Stokes' law regardless of whether they settle in-  
550 dividually or collectively due to the high viscosity of the ambient fluid. In  
551 this scenario, the  $B_{vv}$  measure would be more appropriate. However, unlike  
552 water or air, any significant variation in the viscosity of the magma would  
553 need to be taken into account.

## 554 7. Conclusion

555 This paper presented a new measure of the tendency for volcanic ash  
556 particle plumes to form in water which, unlike existing measures, takes into  
557 account the fact that plume growth and descent are controlled by the bal-  
558 ance between gravitational forces and inertial (rather than viscous) drag.  
559 The measure was evaluated, along with a measure by Marsh (1988) that  
560 assumes Stokes' law-based (i.e. viscous drag-controlled) collective settling,  
561 using results from a suite of particle settling simulations and previous ana-  
562 logue experiments by Carey (1997). The measure that assumes collective  
563 settling is slowed by viscous drag ( $B_{vv}$ ) did not consistently predict the on-  
564 set of pluming and in some cases grossly under-estimated the timescale of  
565 collective particle settling. In contrast, the new measure that assumes collec-  
566 tive settling is slowed by inertial drag ( $B_{vi}$ ) correctly predicted plume onset  
567 conditions for all numerical simulations, and was much more consistent with  
568 experimental data, highlighting the need to take the inertial drag force into  
569 account.

570 The robustness of the new measure became even more apparent when it  
571 was re-arranged in terms of a critical layer thickness  $h_{\text{crit}}$ , such that the layer  
572 thickness must satisfy  $h_{\text{crit}} < h < H$  (where  $H$  is the height of the water  
573 column) for pluming to occur. This quantity requires only the volumetric  
574 flux of particles and the particle diameter to be known, and is therefore more  
575 suitable in field studies. The values of  $h_{\text{crit}}$  for the  $B_{vv}$  measure did not  
576 consistently agree with the layer thicknesses determined from the numerical  
577 simulations, and in fact diverged away from them. This means that  $B_{vv}$  can-  
578 not be used to robustly predict the tendency for plumes to form, since the

579 corresponding values of  $h_{\text{crit}}$  imply that plumes may form much sooner than  
580 they actually do. In contrast, the layer thicknesses all ran parallel to the  $h_{\text{crit}}$   
581 line for the new measure as expected, further reinforcing its validity. The  
582 ability of the new measure to predict plume onset accurately and consistently  
583 allows the residence times and deposition rates of particles in a large body  
584 of water to be determined more reliably. The measure therefore has signif-  
585 icant implications for geological field studies since it permits the improved  
586 interpretation of the layers of volcanoclastic material along the seabed.

587 The formulation of the new measure itself brought an additional benefit;  
588 the value of  $B_{vi}$  could be estimated from the properties of the final deposit,  
589 such that knowledge of the particle mass flux and duration are not required.  
590 It was found that for typical fine-grained ash deposits greater than 1 mm in  
591 thickness, it is likely that particles would have settled collectively as plumes.  
592 However, care must be taken when using this estimation since it introduces  
593 assumptions about the layer itself (e.g. uniform in thickness) which may not  
594 always be justifiable in practice.

595 Despite the study focussing mainly on monodisperse systems with just  
596 one particle size, it was demonstrated that the measure can also correctly  
597 predict plume onset conditions for a polydisperse flow. Plume onset was  
598 found to be governed by the smaller particles in such flows, so the value of  
599  $d_p$  in  $B_{vi}$  should be chosen to be the diameter of the smallest particle in the  
600 system. Furthermore, it is worth noting that while the new measure was  
601 only applied to situations involving volcanic ash, it is likely that it will also  
602 be valid for other geophysical events involving small particles in water, such  
603 as impact ejecta fallout.

## 604 8. Acknowledgments

605 CTJ and SCK were funded by the Institute of Shock Physics at Im-  
606 perial College London and the Atomic Weapons Establishment; GSC was  
607 funded by the Natural Environment Research Council, Fellowship Grant  
608 NE/E013589/1. Support from the Imperial College High Performance Com-  
609 puting Service was also gratefully received. The interested reader is referred  
610 to the thesis by Jacobs (2013), from which most of the content in this paper is  
611 based. The authors would like to thank Vernon Manville and an anonymous  
612 reviewer for their constructive feedback which greatly improved the quality  
613 of this paper. AWE © Crown Owned Copyright (2012).

## 614 Appendix A. Notation

615 A list of notation used throughout this paper is given in Table A.3.

## 616 Appendix B. Derivation of the $B_{ii}$ measure

617 At high Reynolds numbers the terminal velocity of an individual particle  
618 can be approximated by balancing the inertial drag force with the buoyancy  
619 force and the particle's weight:

$$\frac{1}{2}C_D A_p \rho_f |\mathbf{u}|^2 = \frac{1}{6}(\rho_p - \rho_f) g \pi d_p^3, \quad (\text{B.1})$$

620 where  $C_D$  and  $A_p$  are the drag coefficient and cross-sectional area of a spher-  
621 ical particle, respectively (Batchelor (1973), pp. 233–234). Using the expres-  
622 sion  $A_p = \frac{1}{4}\pi d_p^2$  and re-arranging for the particle speed  $|\mathbf{u}|$  gives

$$|\mathbf{u}| = \sqrt{\frac{4(\rho_p - \rho_f) g d_p}{3C_D \rho_f}}, \quad (\text{B.2})$$

Notation	Units	Description
$t$	s	Time
$t_{\text{onset}}$	s	Time of plume onset
$\tau_{\text{individual}}$	s	Timescale of individual particle settling
$\tau_{\text{collective}}$	s	Timescale of collective particle settling
$\alpha_p$	Dimensionless	Volume fraction of the particles
$\alpha_f$	Dimensionless	Volume fraction of the fluid
$\rho_p$	$\text{kgm}^{-3}$	Density of the particles
$\rho_f$	$\text{kgm}^{-3}$	Density of the fluid
$\rho$	$\text{kgm}^{-3}$	Bulk density ( $\rho = \alpha_f \rho_f + \alpha_p \rho_p$ )
$\mathbf{u}_p$	$\text{ms}^{-1}$	Velocity of the particles
$\mathbf{u}_f$	$\text{ms}^{-1}$	Velocity of the fluid
$\mu_f$	Pa s	Viscosity of the fluid
$g$	$\text{ms}^{-2}$	Acceleration due to gravity
$d_p$	m	Diameter of the particles
Re	Dimensionless	Reynolds number
At	Dimensionless	Atwood number
$h$	m	Thickness of the near-surface layer
$h_{\text{crit}}$	m	Critical layer thickness
$H$	m	Height of the water column
$\delta$	m	Maximum amplitude of the growing plumes
$B_{vv}$	Dimensionless	The measure by Marsh (1988)
$B_{vi}$	Dimensionless	The new measure presented in this paper
$\beta$	Dimensionless	Constant plume growth parameter
$\dot{M}_p$	$\text{kgm}^{-2}\text{s}^{-1}$	Mass flux (per unit area) of particles
$\dot{V}_p$	$\text{ms}^{-1}$	Volumetric flux (per unit area) of particles
$A$	$\text{m}^2$	Area through which particles enter the water
$V_p$	$\text{m}^3$	Volume of the near-surface layer occupied by particles
$V_{\text{layer}}$	$\text{m}^3$	Total volume of the near-surface layer (including the water)

Table A.3: The notation used throughout this paper.

623 which is similar to the expression used by Bonadonna et al. (1998) for Re  
 624 > 500. It follows that the timescale of individual particle settling through a  
 625 layer of thickness  $h$  is

$$\tau_{\text{individual}} = \frac{h}{\sqrt{\frac{4(\rho_p - \rho_f)gd_p}{3C_D\rho_f}}}. \quad (\text{B.3})$$

626 Finally, dividing (B.3) by the timescale for inertial drag-based collective  
 627 settling:

$$\tau_{\text{collective}} = 2\sqrt{\frac{\rho_f\delta}{(\rho_p - \rho_f)\alpha_p g}}, \quad (\text{B.4})$$

628 and simplifying produces the non-dimensional number  $B_{ii}$ :

$$B_{ii} = \frac{h}{2}\sqrt{\frac{3C_D\alpha_p}{4\delta d_p}}. \quad (\text{B.5})$$

## 629 **References**

- 630 Batchelor, G. K., 1973. An Introduction to Fluid Dynamics. Cambridge Uni-  
 631 versity Press, Cambridge, UK, 615 pages.
- 632 Bergantz, G. W., Ni, J., 1999. A numerical study of sedimentation by drip-  
 633 ping instabilities in viscous fluids. International Journal of Multiphase  
 634 Flow 25 (2), 307–320.
- 635 Bonadonna, C., Ernst, G. G. J., Sparks, R. S. J., 1998. Thickness variations  
 636 and volume estimates of tephra fall deposits: the importance of particle  
 637 Reynolds number. Journal of Volcanology and Geothermal Research 81 (3-  
 638 4), 173–187.



- 639 Bramlette, M. N., Bradley, W. H., 1941. Geology and biology of North At-  
640 lantic deep-sea cores between Newfoundland and Ireland: lithology and  
641 geological interpretation. U.S. Geological Survey Professional Paper 196-  
642 A, 1.
- 643 Carazzo, G., Jellinek, A. M., 2012. A new view of the dynamics, stability and  
644 longevity of volcanic clouds. *Earth and Planetary Science Letters* 325-326,  
645 39–51.
- 646 Carey, S., 1997. Influence of convective sedimentation on the formation of  
647 widespread tephra fall layers in the deep sea. *Geology* 25 (9), 839–842.
- 648 Carey, S., Schneider, J.-L., 2011. Volcaniclastic Processes and Deposits in the  
649 Deep-Sea. In: Hüneke, H., Mulder, T. (Eds.), *Deep-Sea Sediments*. Vol. 63  
650 of *Developments in Sedimentology*. Elsevier, Ch. 7, pp. 457–515.
- 651 Crowe, C. T., Sommerfeld, M., Tsuji, Y., 1998. *Multiphase Flows with*  
652 *Droplets and Particles*. CRC Press, Boca Raton, USA, 471 pages.
- 653 Dalziel, S. B., Patterson, M. D., Caulfield, C. P., Coomaraswamy, I. A., 2008.  
654 A numerical study of sedimentation by dripping instabilities in viscous  
655 fluids. *Physics of Fluids* 20 (6), 065106.
- 656 Davies, D. R., Wilson, C. R., Kramer, S. C., 2011. Fluidity: A fully unstruc-  
657 tured anisotropic adaptive mesh computational modeling framework for  
658 geodynamics. *Geochemistry Geophysics Geosystems* 12 (6).
- 659 Dimonte, G., Schneider, M., 2000. Density ratio dependence of Rayleigh-  
660 Taylor mixing for sustained and impulsive acceleration histories. *Physics*  
661 *of Fluids* 12 (2), 304–321.

- 662 Dimonte, G., Youngs, D. L., Dimits, A., S, W., Marinak, M., Wunsch, S., C.,  
663 G., Robinson, A., Andrews, M. J., Ramaprabhu, P., Calder, A. C., Fryxell,  
664 B., Biello, J., Dursi, L., MacNeice, P., Olson, K., Ricker, P., Rosner, R.,  
665 Timmes, F., Tufo, H., Young, Y.-N., Zingale, M., 2004. A comparative  
666 study of the turbulent Rayleigh-Taylor instability using high-resolution  
667 three-dimensional numerical simulations: The Alpha-Group collaboration.  
668 *Physics of Fluids* 16 (5), 1668–1693.
- 669 Geuzaine, C., Remacle, J.-F., 2009. Gmsh: A 3-D finite element mesh gener-  
670 ator with built-in pre- and post-processing facilities. *International Journal*  
671 *for Numerical Methods in Engineering* 79 (11), 1309–1331.
- 672 Goldin, T., 2008. Atmospheric Interactions During Global Deposition of  
673 Chicxulub Impact Ejecta. Ph.D. thesis, University of Arizona.
- 674 Jacobs, C. T., 2013. Modelling of Multiphase Flows on Adaptive Unstruc-  
675 tured Meshes with Applications to the Dynamics of Volcanic Ash Plumes.  
676 Ph.D. thesis, Imperial College London.
- 677 Jacobs, C. T., Collins, G. S., Piggott, M. D., Kramer, S. C., Wilson, C.  
678 R. G., 2013. Multiphase flow modelling of volcanic ash particle settling in  
679 water using adaptive unstructured meshes. *Geophysical Journal Interna-*  
680 *tional* 192 (2), 647–665.
- 681 Kuenen, P. H., 1968. Settling convection and grain-size analysis. *Journal of*  
682 *Sedimentary Research* 38 (3), 817–831.
- 683 Ledbetter, M. T., Sparks, R. S. J., 1979. Duration of large-magnitude explo-

- 684 sive eruptions deduced from graded bedding in deep-sea ash layers. *Geology*  
685 7 (5), 240–244.
- 686 Macedonio, G., Costa, A., 2012. Brief Communication "Rain effect on the  
687 load of tephra deposits". *Natural Hazards and Earth System Sciences*  
688 12 (4), 1229–1233.
- 689 Manville, V., Wilson, C., 2004. Vertical density currents: a review of their  
690 potential role in the deposition and interpretation of deep-sea ash layers.  
691 *Journal of the Geological Society, London* 161, 947–958.
- 692 Marsh, B., 1988. Crystal Capture, Sorting, and Retention in Convecting  
693 Magma. *Geological Society of America Bulletin* 100, 1720–1737.
- 694 Narbonne, G. M., 2005. The Ediacara Biota: Neoproterozoic Origin of An-  
695 imals and Their Ecosystems. *Annual Review of Earth and Planetary Sci-*  
696 *ences* 33, 421–442.
- 697 Piggott, M. D., Gorman, G. J., Pain, C. C., Allison, P. A., Candy, A. S.,  
698 Martin, B. T., Wells, M. R., 2008. A new computational framework for  
699 multi-scale ocean modelling based on adapting unstructured meshes. *In-*  
700 *ternational Journal for Numerical Methods in Fluids* 56 (8), 1003–1015.
- 701 Ristorcelli, J. R., Clark, T. T., 2004. Rayleigh-Taylor turbulence: self-similar  
702 analysis and direct numerical simulations. *Journal of Fluid Mechanics* 507,  
703 213–253.
- 704 Rose, W. I., Durant, A. J., 2009. Fine ash content of explosive eruptions.  
705 *Journal of Volcanology and Geothermal Research* 186 (1-2), 32–39.

- 706 Sarna-Wojcicki, A., Shipley, S., Waitt, R., Dzurisin, D., Wood, S., 1981.  
707 Areal distribution, thickness, mass, volume, and grain size of air-fall ash  
708 from the six major eruptions of 1980. In: Lipman, P., Mullineaux, D.  
709 (Eds.), *The 1980 Eruptions of Mount St. Helens*, Washington. U.S. Geo-  
710 logical Survey Professional Paper 1250. pp. 577–600.
- 711 Scheidegger, K., Federman, A., Tallman, A., 1982. Compositional hetero-  
712 geneity of tephra from the 1980 eruptions of Mount St. Helens. *Journal*  
713 *of Geophysical Research* 87 (B13), 10,861–10,881.
- 714 Seilacher, A., Reif, W.-E., Westphal, F., 1985. Sedimentological, Ecological  
715 and Temporal Patterns of Fossil Lagerstätten. *Philosophical Transactions*  
716 *of the Royal Society of London. Series B, Biological Sciences* 311 (1148),  
717 5–23.
- 718 Stokes, G. G., 1851. On the Effect of the Internal Friction of Fluids on  
719 the Motion of Pendulums. *Transactions of the Cambridge Philosophical*  
720 *Society* 9, 8.
- 721 Ver Straeten, C. A., 2004. K-bentonites, volcanic ash preservation, and im-  
722 plications for Early to Middle Devonian volcanism in the Acadian orogen,  
723 eastern North America. *Geological Society of America Bulletin* 116 (3-4),  
724 474–489.
- 725 Ver Straeten, C. A., 2008. Volcanic Tephra Bed Formation and Condensa-  
726 tion Processes: A Review and Examination from Devonian Stratigraphic  
727 Sequences. *The Journal of Geology* 116 (6), 545–557.

- 728 White, J. D. L., 2000. Subaqueous eruption-fed density currents and their  
729 deposits. *Precambrian Research* 101 (2–4), 87–109.
- 730 White, J. D. L., Smellie, J. L., Clague, D. A., 2003. Explosive Subaque-  
731 ous Volcanism. American Geophysical Union, Washington, DC, USA, 379  
732 pages.
- 733 Whitehead, J., Luther, D., 1975. Dynamics of Laboratory Diapir and Plume  
734 Models. *Journal of Geophysical Research* 80 (5), 705–717.
- 735 Wiesner, M., Wang, Y., Zheng, L., 1995. Fallout of volcanic ash to the deep  
736 South China Sea induced by the 1991 eruption of Mount Pinatubo (Philip-  
737 pines). *Geology* 23 (10), 885–888.
- 738 Youngs, D. L., 1984. Numerical simulation of turbulent mixing by Rayleigh-  
739 Taylor instability. *Physica D: Nonlinear Phenomena* 12 (1-3), 32–44.