

1                   **Hydrothermal contributions to global biogeochemical cycles:**

2                   **Insights from the Macquarie Island ophiolite**

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4                   *Submitted to Lithos*

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26 **Highlights:**

27 • Macquarie Island comprises a unique complete section of altered ocean crust.

28 • The Macquarie crust was a net source of Si, Ti, Al, and Ca to the oceans.

29 • The Macquarie crust was a net sink for H<sub>2</sub>O, Mg, Na, K, and S.

30 • Veins make important contributions to hydrothermal chemical exchange budgets.

31

32

33 **Abstract**

34 Hydrothermal circulation is a fundamental process in the formation and aging of the  
35 ocean crust, with the resultant chemical exchange between the crust and oceans  
36 comprising a key component of global biogeochemical cycles. Sections of  
37 hydrothermally altered ocean crust provide time-integrated records of this chemical  
38 exchange. Unfortunately, our knowledge of the nature and extent of hydrothermal  
39 exchange is limited by the absence of complete oceanic crustal sections from either  
40 submarine exposures or drill core. Sub-Antarctic Macquarie Island comprises ~10 Ma  
41 ocean crust formed at a slow spreading ridge, and is the only sub-aerial exposure of a  
42 complete section of ocean crust in the ocean basin in which it formed. Hydrothermally  
43 altered rocks from Macquarie Island therefore provide a unique opportunity to evaluate  
44 the chemical changes due to fluid-rock exchange through a complete section of ocean  
45 crust. Here we exploit the immobile behavior of some elements during hydrothermal  
46 alteration to determine the precursor compositions to altered Macquarie whole rock  
47 samples, and evaluate the changes in bulk rock chemistry due to fluid-rock interaction  
48 throughout the Macquarie crust. The extent to which elements are enriched or depleted  
49 in each sample depends upon the secondary mineral assemblage developed, and hence  
50 the modal abundances of the primary minerals in the rocks and the alteration conditions,  
51 such as temperature, fluid composition, and water: rock ratios. Consequently the  
52 chemical changes vary with depth, most notably within the lava-dike transition zone  
53 where enrichments in K, S, Rb, Ba, and Zn are observed. Our results indicate that  
54 hydrothermal alteration of the Macquarie crust resulted in a net flux of Si, Ti, Al, and  
55 Ca to the oceans, whereas the crust was a net sink for H<sub>2</sub>O, Mg, Na, K, and S. Our  
56 results also demonstrate the importance of including the contribution of elemental

57 uptake by veins for some elements (e.g., Si, Fe, Mg, S). Extrapolation of our results,  
58 assuming a crustal production rate of  $3 \text{ km}^2/\text{yr}$ , yields estimates of the hydrothermal  
59 contribution to global geochemical cycles. For example, the Mg flux to the crust is  
60 estimated to be  $3.3 \pm 1.1 \times 10^{12} \text{ mol/yr}$ , sufficient to balance the riverine Mg input to the  
61 oceans given the uncertainties involved. However, the relationship between spreading  
62 rate and hydrothermal chemical exchange fluxes remains poorly understood, and the  
63 approach described here should be applied to crust produced at a range of spreading  
64 rates to refine the global hydrothermal flux estimates.

65

66 **Keywords:**

67 Ocean crust, hydrothermal alteration, biogeochemical cycles, Macquarie Island,  
68 ophiolite.

69

70 **Abbreviations:**

71 LDTZ = lava-dike transition zone; DGTZ = dike-gabbro transition zone; Mg# =  $100 \times$   
72  $\text{Mg}/[\text{Mg} + \text{Fe}^{2+}]$ , calculated assuming that 90% of the total iron is  $\text{Fe}^{2+}$ ; LOI = loss on  
73 ignition; MORB = mid-ocean ridge basalt.

74

## 75 1. Introduction

76 Hydrothermal circulation is an important component of global biogeochemical cycles.  
77 Chemical exchange between seawater and the ocean crust influences the composition of  
78 the oceans, the ocean crust, and via subduction the composition and heterogeneity of the  
79 mantle. Despite nearly 50 years of scientific ocean drilling, the ultimate goal of drilling  
80 a continuous in situ section through the entire ocean crust has not yet been achieved  
81 (Teagle and Ildefonse, 2011). The absence of complete oceanic crustal sections makes  
82 full quantification of the ocean crust's primary compositions and its hydrothermal  
83 contributions to global geochemical cycles difficult. In particular, our knowledge of the  
84 nature and extent of fluid-rock interaction in the lower crust is limited by the absence of  
85 accessible submarine exposures or drill core. Ophiolites, sub-aerially exposed sections  
86 of crust formed via seafloor spreading and subsequently emplaced on continental  
87 margins, have been used to investigate chemical exchange between seawater and the  
88 ocean crust (e.g., Bednarz and Schmincke, 1989; Gregory and Taylor, 1981). However,  
89 most intact ophiolites formed in supra-subduction settings and whether these outcrops  
90 are representative of normal mid-ocean ridge crust remains controversial (Miyashiro,  
91 1973; Rautenschlein et al., 1985). These ophiolites are more intensely recrystallized,  
92 and consistently record a greater extent of fluid-rock exchange than crust from mid-  
93 ocean ridges (Alt and Teagle, 2000; Bickle and Teagle, 1992).

94

95 Sub-Antarctic Macquarie Island is the only sub-aerial exposure of a complete section of  
96 ocean crust in the open-ocean basin in which it formed (Varne et al., 2000; Varne et al.,  
97 1969). Macquarie Island provides a unique opportunity to investigate how the ocean

98 crust is accreted, including the role of hydrothermal circulation in cooling the upper and  
99 lower crust and the resultant fluid-rock chemical exchange.

100

101 Chemical exchange between hydrothermal fluids and the crust is facilitated by mineral  
102 dissolution and the formation of secondary minerals, which replace igneous minerals  
103 and mesostasis, fill primary porosity (vugs and interstices) and form veins.

104 Hydrothermally altered rocks therefore provide a time-integrated record of the effects of  
105 fluid-rock exchange on the composition of the crust. Here we evaluate the hydrothermal  
106 changes in bulk rock chemistry of Macquarie Island rocks, and compare our results with  
107 other independent estimates of seawater-ocean crust hydrothermal exchange.

108

109 2. Macquarie Island

110 Sub-Antarctic Macquarie Island (54°30'S, 158°56'E) in the Southern Ocean comprises  
111 ~10 Ma ocean crust formed during slow spreading (10 mm/yr half rate) along the  
112 Australian-Pacific plate boundary (Armstrong et al., 2004; Duncan and Varne, 1988;  
113 Quilty et al., 2008; Varne et al., 2000). Spreading initiated along this boundary during  
114 the Eocene when the Tasman Sea rifted from the Campbell Plateau along a series of  
115 short ridge segments separated by fracture zones (Sutherland, 1995). Subsequent  
116 migration of the pole of Australian-Pacific relative plate motion caused progressive re-  
117 orientation of the spreading segments, producing the distinctive regional fracture zone  
118 curvature, with the system ultimately entering a dextral transpressional regime when  
119 extension was sub-parallel to the plate boundary (Cande et al., 2000; Lamarche et al.,  
120 1997; Massell et al., 2000; Sutherland, 1995). Shortening along the boundary uplifted  
121 the Macquarie Ridge Complex, exposed at Macquarie Island since ~0.6 Ma (Adamson

122 et al., 1996). The island comprises crust formed as the magmatism waned, with rocks  
123 from all crustal levels and the uppermost mantle exposed (Fig. 1) (Varne et al., 2000;  
124 Wertz et al., 2003).

125  
126 Basaltic lavas crop out across the southern three quarters of the island and on North  
127 Head (Fig. 1). To the northwest is an uplifted section of intrusive ocean crust  
128 comprising a sheeted dike complex, massive gabbros, layered gabbros, and a cumulate  
129 sequence of ultramafic rocks (Fig. 1), with the underlying residual mantle harzburgites  
130 exposed at Eagle Point. The extrusive portion of this block is not exposed, but a  
131 sequence of steeply dipping sheeted dikes that grade upwards into gently dipping lavas  
132 occurs within a 5 km wide fault-bounded block on the west coast near Double Point and  
133 Mount Waite, within which the original seafloor relationships are preserved and the  
134 lava-dike boundary is the original extrusive-intrusive crustal transition zone (Davidson  
135 et al., 2004).

136  
137 Three major faults juxtapose rocks from different crustal levels or metamorphic grade  
138 (Selkirk et al., 1990; Varne et al., 2000) (Fig. 1). The Finch-Langdon Fault separates the  
139 intrusive crustal section from lavas to the south. It is cemented by hydrothermal  
140 minerals, has talus breccias with clasts of basalt, dolerite and gabbro and is overlain by  
141 lava flows, consistent with it being a relict axial seafloor spreading fault (Wertz et al.,  
142 2003). The Isthmus fault emplaces greenschist facies dikes against the lower grade  
143 North Head lavas and the Major Lake Fault Zone emplaces sheeted dikes and lavas  
144 against extrusives of a lower metamorphic grade (Daczko et al., 2005; Lewis, 2007;

145 Portner et al., 2010; Rivizzigno and Karson, 2004). There are numerous other minor  
146 faults throughout the island (Goscombe and Everard, 2001).

147

148 All levels of the Macquarie crust have been affected by fluid-rock interaction. O and C  
149 isotopic analyses reveal that the crust interacted with seawater-derived fluids, rather  
150 than meteoric water. The lavas interacted with cold seawater at high water: rock ratios,  
151 whereas the lower crust interacted with hotter (300-600 °C) hydrothermal fluids at low  
152 water: rock ratios (Cocker et al., 1982).

153

## 154 2.1 Macquarie Island sample transects and stratigraphic reconstruction

155 Using the published geological maps (Goscombe and Everard, 1998a), supplemented by  
156 our own field observations, we selected three transects through the Macquarie crust to  
157 develop a complete ocean crustal section: a parallel pair of transects through the upper  
158 crust near Mount Waite and Double Point (A and B, Fig. 1); and the coastal Isthmus to  
159 Eagle Point transect through the intrusive crust and uppermost mantle (C, Fig. 1).

160

### 161 2.1.1 Mount Waite and Double Point transects

162 The Mount Waite transect traverses a narrow (<50 m) zone of steeply dipping (~75° W)  
163 sheeted dikes and the overlying extrusive rocks (Fig. 2a). The extrusive sequence  
164 consists of variably (<5 % to >30 %) plagioclase-phyric pillow lavas, volcaniclastic  
165 breccias and turbiditic sediments, dipping 20 - 30° E. Pillow units are typically 50-100  
166 m thick, and can be traced laterally for several kilometers. The lavas are pervasively  
167 altered. The upper lavas contain smectite, iron-oxides and carbonate, typical of low  
168 temperature ‘ocean floor weathering’ (Alt et al., 1986), whereas the underlying lavas

169 were partially altered under zeolite or lower greenschist facies conditions and contain  
170 chlorite ± zeolites ± smectite ± calcite ± albite ± prehnite ± epidote. The inter-lava  
171 sediments include 1-10 m thick graded conglomerate, sandstone and red shale beds.  
172 Assuming the paleo-vertical is perpendicular to the lava flows, the extrusive section is  
173 ~850 m thick. A small volcanic cone of radially outward dipping elongate pillow lavas  
174 near Pyramid Peak, to the east of Mount Waite, indicates that the top of the Mount  
175 Waite extrusive sequence was at or near the paleo- seafloor (Griffin and Varne, 1980).  
176 The lava-dike transition (LDTZ) is therefore reconstructed to 850 m below seafloor,  
177 with sample depths assigned from this datum.

178

179 The Double Point transect is similar to the Mount Waite transect, but with a greater  
180 expanse of sheeted dikes exposed (Fig. 2b). The dikes are 1.5 to 3 m wide, with  
181 narrower cross-cutting dikes increasingly abundant up section (Davidson et al., 2004),  
182 and are partially altered under greenschist facies conditions (Table S1). The LDTZ is  
183 cut by minor faults with disseminated pyrite and chlorite halos (Davidson et al., 2004).  
184 Similar sulfide anomalies have been identified at the LDTZ in ocean crust from ODP  
185 Holes 504B and 1256D, and the Troodos ophiolite, Cyprus (Alt, 1994, 1995b; Alt et al.,  
186 2010). A 40 m wide fault zone with epidote-actinolite-zoisite-quartz-chalcopyrite-pyrite  
187 veins is interpreted as a hydrothermal upwelling site (Davidson et al., 2004). The lavas  
188 at the western tip of Double Point were altered under greenschist to amphibolite facies  
189 conditions and were intruded by the dikes during spreading (Davidson et al., 2004),  
190 consequently the transect is treated as a continuous section. Sample depths are assigned  
191 assuming the paleo-vertical is perpendicular to the bedding, and the LDTZ was 850 m  
192 below seafloor.

193

## 194 2.1.2 The Isthmus-Eagle Point transect

195 Sheeted dikes are exposed in Hasselborough Bay. The dikes are typically 0.5-3 m wide  
196 and can be traced laterally <30 m along strike. They are sub-parallel, and their chilled  
197 margins dip ~60° SW on average (Fig. 3). Some dikes enclose and are chilled against  
198 gabbro screens that make up <5% of the outcrops. The dikes are aphyric to highly  
199 plagioclase phryic, with increasing phenocryst abundance and grainsize of groundmass  
200 and phenocrysts away from the well-defined chilled margins. Both porphyritic and  
201 aphyric dikes occur throughout the sheeted dike section, cross-cut by later stage narrow  
202 (~10 cm wide) aphyric dikes. The dike-gabbro transition (DGTZ) is characterized by an  
203 increase in size and abundance of gabbro screens and a decrease in the abundance of  
204 dolerite dikes. The sheeted dikes were altered under greenschist facies conditions, with  
205 amphibolite grade dikes in the DGTZ. The most intense alteration typically occurs in  
206 'halos' that flank veins.

207

208 The lower crust is predominantly massive gabbro with enclaves of anorthosite and  
209 olivine gabbro. Compositional layering is restricted to Handspike Point where felsic and  
210 mafic layering is oriented ~126/48°SW, sub-parallel to sheeted dike margins indicating  
211 it was originally sub-vertical. The gabbros are typically fresher than the overlying  
212 sheeted dikes and were altered under greenschist to amphibolite facies conditions (Table  
213 S1). The alteration is highly heterogeneous and most intense adjacent to veins. Sub-  
214 parallel amphibole + chlorite veinlets impart a weak fabric and are cut by wider (<5  
215 mm) epidote + prehnite veins and later-stage cataclastic prehnite + chlorite veins.  
216 Dolerite dikes similar to those of the sheeted dike complex occur throughout the gabbro,

217 but were altered under similar conditions to the host gabbro. Narrower (<40 cm)  
218 microgabbro dikes and veins are common, with diffuse margins suggesting  
219 emplacement before the host gabbro had completely cooled. The transition between the  
220 gabbro and the underlying harzburgite comprises a complex association of highly to  
221 completely altered ultramafic rocks including dunite, plagioclase dunite, wehrlite,  
222 plagioclase wehrlite, olivine gabbro, troctolite and harzburgite. The contacts between  
223 the rock units are poorly constrained by limited exposure. Troctolite exposed at  
224 Elizabeth and Mary Point contains alternating mafic and felsic layers that dip ~30° east-  
225 northeast. The layers are sub-parallel and typically a few centimeters to tens of  
226 centimeters thick. Increasing plagioclase abundance upwards within layers indicates  
227 they are the correct way up (Goscombe and Everard, 2001). Dolerite sills intrude the  
228 troctolite at an oblique angle to the layering. The harzburgite exposed at Eagle Point is  
229 massive, serpentinized, and intruded by pegmatitic gabbro and rare <1 m wide  
230 plagioclase phryic dolerite dikes that are pervasively prehnitized or rodingitized. The  
231 paleo-vertical of this section is constrained to lie within the average plane of the dolerite  
232 dike chilled margins, which dips ~ 60° SW. The troctolite compositional layering is  
233 orthogonal to the sheeted dikes' chilled margins, with the pseudo sedimentary structures  
234 indicating they are paleo-horizontal layers. Its pole is therefore taken as the paleo-  
235 vertical for this transect (Fig. 3b).

236

## 237 2.2 The proto-Macquarie Island ocean crust

238 Although slow-spread ocean crust is architecturally complex (e.g., Cannat et al., 2008;  
239 MacLeod et al., 2009), layered crust similar to that formed at fast spreading rates  
240 (Penrose Conference Participants, 1972) may be produced at the middle of relatively

241 robust slow-spreading accretionary segments (Dick et al., 2003; Dick et al., 2006; Sinha  
242 et al., 1998). The proto-Macquarie Island crust includes all the components of normal  
243 layered ocean crust. Our sample transects are combined to give a complete section  
244 through this crust, with the LDTZ of the Mount Waite and Double Point transects  
245 aligned at 850 meters below seafloor (mbsf), and the alteration assemblages of the  
246 sheeted dikes of the Eagle Point-Isthmus and Double Point transects correlated (Fig. 4).

247 Our composite section is similar to previous Macquarie stratigraphies (Dijkstra and  
248 Cawood, 2004; Goscombe and Everard, 1998b), but with significantly different layer  
249 thicknesses. This in part reflects considerable lateral variability in crustal layer  
250 thicknesses across Macquarie Island, for example the maximum exposed thickness of  
251 lavas is 1.4 km. Goscombe and Everard (1998b) use the average thickness of gabbro in  
252 other ophiolites, which is approximately double the maximum gabbro thickness  
253 observed on Macquarie Island. The reconstructed thickness of the proto-Macquarie  
254 crust (3-4 km; Fig. 4) is consistent with its formation on a short segment of a slow  
255 spreading ridge in a waning magmatic system (Cannat, 1996; Chen, 1992; Dick et al.,  
256 2003).

257

258 The syn-volcanic tectonism along the Finch-Langdon Fault (Rivizzigno and Karson,  
259 2004; Wertz et al., 2003) and the intrusion of dolerite dikes into relatively cool gabbros  
260 and harzburgite indicate that the Macquarie crust was generated during multiple  
261 magmatic episodes alternating with periods of tectonic extension, analogous to the  
262 ‘tectono-magmatic cycles’ observed at slow-spreading ridges (Sinha et al., 1998). The  
263 timing of the different magmatic events remains poorly constrained, and the co-genetic  
264 relationships between the different units unknown. For example, the majority of

265 Macquarie basalts have highly enriched mid-ocean ridge basalt (MORB) compositions  
266 indicative of small-degrees of fractional melting (Kamenetsky et al., 2000) that are not  
267 consistent with their formation during the melting event that depleted the Macquarie  
268 upper mantle (Wertz, 2003). Despite its formation during multiple magmatic events,  
269 Macquarie Island represents the best available continuous section through the lower  
270 ocean crust produced at a slow-spreading ridge and complements lower crustal sections  
271 drilled at slow spreading ridges into gabbro (for example ODP Hole 735B and IODP  
272 Hole 1309D; (Blackman et al., 2006; Dick et al., 1999; Robinson et al., 1989)).

273

### 274 3. Analytical Techniques

275 Representative samples (> 1 kg) were taken every 20 m along transects orthogonal to  
276 bedding at Double Point and Mount Waite, with additional samples taken from fracture  
277 zones. Along the steep cliff sections of these transects the outcrop is nearly continuous,  
278 with soft sediment interbeds preserved in the walls of transecting gullies. Along the  
279 Isthmus – Eagle Point coastal transect outcrops are discontinuous, separated by cobble  
280 beaches and areas of marshland. Large samples (1-5 kg) representative of the rock types  
281 in each outcrop were taken.

282

283 Thin sections were prepared to include the variations in igneous, metamorphic, or  
284 tectonic features. Samples for whole rock geochemical analyses were cut, using a  
285 diamond saw, avoiding these heterogeneities with sub-samples of more intensely altered  
286 zones prepared for comparison. Exterior weathered surfaces were removed, and samples  
287 were ultrasonicated in deionized water, dried for 12 hours at 70°C, fragmented to <1 cm  
288 chips between sheets of clean paper in a hardened pure-iron fly press, and powdered

289 using a hardened pure-iron tema. Lithologies were determined from the modal  
290 mineralogy. The key petrographic features of each rock type are summarized in Table  
291 S1.

292

293 Geochemical analyses of 236 Macquarie Island whole rock samples are presented in  
294 Table S2. Major and trace elements were analyzed by X-ray fluorescence (XRF),  
295 following the methods of Brewer et al., (1998), using a PW1400 X-Ray Spectrometer at  
296 the University of Leicester. Geochemical reference materials were used to construct  
297 calibrations for individual elements and to evaluate precision and accuracy (Table S3).

298 A subset of samples were analyzed at the University of Tasmania School of Earth  
299 Sciences using a PW1410XRF, following the methods of Norrish and Hutton (1969).

300 Carbon and Sulfur concentrations were analyzed using a LECO Carbon/Sulfur Analyser  
301 by high-temperature combustion, at the University of Leicester. The lower limit of  
302 detection was 10 ppm, with a precision of  $\pm 5\%$  and  $\pm 8\%$  for C and S, respectively.

303

304 Trace elements concentrations were determined by Inductively Coupled Plasma Mass  
305 spectrometry (ICP-MS) at the University of Southampton using a VG PlasmaQuad  
306 PQ2+. Precision and accuracy were better than  $\pm 5\%$  RSD and  $\pm 8\%$  RMSD,  
307 respectively, for the majority of elements (Table S4).

308

309 4. Primary magmatic diversity of Macquarie Island igneous rocks  
310 Published analyses of fresh glasses from Macquarie Island hyaloclastites and pillow  
311 margins reveal they are basaltic in composition, with 47.4-51.1 wt% SiO<sub>2</sub> and 5.65-8.75  
312 wt % MgO, but range from depleted N-MORB to compositions more enriched than

313 typical E-MORB (Kamenetsky et al., 2000; Wertz, 2003). They span a wider range in  
314 K<sub>2</sub>O content (0.1-0.8 wt%) than usual for a single MORB suite (Kamenetsky et al.,  
315 2000) and span most of the range in Zr/Y ratios of other MORB suites combined (Fig.  
316 5a). This diversity reflects a complex melting history. Glass compositions have been  
317 divided into two groups on the basis of their Mg-numbers (Mg# = 100 x Mg/[Mg +  
318 Fe<sup>2+</sup>]) and K<sub>2</sub>O contents, Group 1 being those with the highest Mg-number at a given  
319 K<sub>2</sub>O content (Kamenetsky et al., 2000). Group 1 glasses, which have the lowest Y  
320 concentrations at a given Zr concentration, are interpreted to be near-primitive  
321 'parental' melts from which Group 2 glasses fractionated (Kamenetsky et al., (2000);  
322 Fig. 5a). The seriate variation in parental melt compositions is attributed to an  
323 increasing degree of partial melting of a mantle source that is homogeneous on a large  
324 scale (Kamenetsky and Maas, 2002). This interpretation is supported by a progressive  
325 decrease in the degree of incompatible element enrichment up through several extrusive  
326 sequences on the island (Wertz, 2003). Aphyric lavas and dikes have Zr/Y ratios  
327 consistent with fractionation from parental melts with Group 1 glass compositions (Fig.  
328 5b). The anorthosites and gabbros have similar Zr/Y ratios to the dikes and lavas  
329 produced by higher degrees of partial melting.

330  
331 All whole rock samples are altered to some extent and alteration effects are  
332 superimposed on and may obscure primary magmatic compositional variations.  
333 However, there are variations in whole rock chemistry that correspond to changes in  
334 rock type, reflecting differences in primary mineral modal proportions, and hence are of  
335 magmatic origin. Consequently there are broad trends with depth; for example

336 decreasing concentrations of Ti, Fe, Na, P, K, Sr, Zr, and Nb and increasing Cr and Ca  
337 contents and magnesium number (Fig. 6).

338

339 Since Zr, Y, and the REEs are relatively immobile during hydrothermal alteration the  
340 Zr/Y and La/Sm ratios should record primary magmatic variations through the crust. In  
341 general the La/Sm ratio decreases with depth, despite significant variability at a given  
342 depth. Gabbros, dikes and lavas with similar Zr/Y and La/Sm ratios may be co-  
343 magmatic. Systematic variations in melt composition through the gabbro section due to  
344 cyclic magma chamber processes, such as fractionation and magma replenishment,  
345 cannot be distinguished given the sampling scale.

346

347 5. Calculating the chemical changes due to hydrothermal alteration  
348 The impact of hydrothermal circulation on ocean chemistry depends on the changes in  
349 bulk rock chemistry due to seawater-rock interaction. As hydrothermal alteration of the  
350 ocean crust is pervasive the fresh igneous precursors to altered rocks are rarely  
351 preserved, and their compositions are difficult to determine if the altered rocks comprise  
352 a fractionated suite. However, the behavior of elements that remain immobile during  
353 hydrothermal alteration can be used to account for: (i) passive changes in elemental  
354 concentrations in response to net changes in mass; and (ii) the primary magmatic  
355 variation in precursor suites (Grant, 1982; Gresens, 1967; MacLean, 1990; MacLean  
356 and Barrett, 1993). During alteration, immobile element concentrations change only in  
357 passive response to net mass changes, being concentrated by bulk rock mass loss or  
358 diluted by mass gains (Fig. 7a). If one immobile element is incompatible (e.g. Zr, Y or  
359 Nb) and another is compatible, their fractionation trend is distinct to the alteration

360 trends, which radiate from the origin. Their concentrations in the precursor to a given  
361 altered rock can therefore be determined from the intersection of its alteration trend and  
362 the fractionation trend, established using fresh samples from the same fractionated suite  
363 (MacLean, 1990) (Fig. 7b). Similarly the precursor's mobile compatible element  
364 concentrations can be determined from their fractionation trends with the immobile  
365 incompatible element, also defined using the fresh samples, and the calculated precursor  
366 immobile incompatible monitor element concentration.

367

368 Determining fresh precursor compositions for Macquarie Island rocks is challenging  
369 given the complex melting history of the ophiolite. For each immobile element pair the  
370 variable degree of partial melting produced a series of primitive parental melt  
371 compositions from which fractional crystallization yielded a large array of potential  
372 fresh precursor compositions (Fig 7c-d). Consequently, precursor immobile element  
373 concentrations are difficult to determine, unless the net mass change due to alteration is  
374 known. There is a weak correlation between the samples' specific gravity and LOI  
375 (Table S2;  $R^2 = 0.3$ ,  $n = 128$ ) indicating a slight decrease in density associated with  
376 hydration during alteration, which equates to less than a 5% decrease in mass assuming  
377 constant volume. In the following calculations we therefore assume that there is no net  
378 mass change during alteration, i.e.  $M^A/M^o = 1$ , where  $M^A$  and  $M^o$  are the total masses of  
379 an altered rock and its fresh precursor, respectively, but allow for 5% uncertainty in this  
380 ratio.

381

382 We also assume that: (i) Zr, Y, and Nb were immobile during alteration, and can be  
383 used as 'monitors' of magmatic fractionation to estimate precursor mobile element

384 concentrations; and (ii) fresh glasses and the least altered whole rock samples are  
385 representative precursors for the altered rocks. The Macquarie samples include eight  
386 pairs, each cut from variably altered portions of the same protolith. The more altered  
387 portions have the same Zr/Y and Nb/Y ratios as their less altered counterparts (Fig. 8)  
388 indicating that Y, Zr, and Nb were immobile during alteration. The least altered whole  
389 rock samples were identified petrographically and from their LOI, with low LOI  
390 indicating minimal addition of hydrous secondary minerals. The assumption that they  
391 represent fresh precursor compositions despite their slight alteration may result in the  
392 underestimation of some hydrothermal chemical changes.

393

394 The methods used to determine precursor compositions for different sample types are  
395 summarized in Figure 9. The precursors to samples crystallized from fractionated  
396 magmas (aphyric lavas and dikes, and non-cumulate plutonic rocks) are determined  
397 from the magmatic fractionation trends of compatible elements (e.g., Ti, Fe, Mn, Mg,  
398 Ca, Na and K) against immobile incompatible ‘monitor’ elements (Zr and Nb) defined  
399 by Macquarie glasses supplemented with the least altered dike, anorthosite, and gabbro  
400 samples to extend the range of precursor compositions. The interplay of varying degree  
401 of partial melting and extent of subsequent fractional crystallization produced a diverse  
402 range in Macquarie magma compositions (Fig 7). However, for each compatible  
403 element the extent to which the partial melting and fractional crystallization trends  
404 diverge against the immobile monitor elements varies. We have exploited this effect,  
405 selecting the immobile monitor element (Nb or Zr) that best constrains the primary  
406 magmatic trend for each compatible element (Table S5). Several elements (Si, Al, S, Zn  
407 and Lu) show no significant variation with Zr or Nb, and the precursor concentrations

408 adopted for each rock type are the average concentrations of the glasses and least altered  
 409 samples (Table S5). Phenocryst accumulation in phryic samples causes deviations from  
 410 primary magmatic fractionation trends. To determine precursor compositions for such  
 411 rocks we must account for: (i) the types, abundances and primary compositions of  
 412 phenocrysts; and (ii) primary magmatic variation of the groundmass. The former could  
 413 be achieved by comparing samples to the least altered samples with similar phenocrysts  
 414 abundances, but this does not compensate for primary variation in groundmass  
 415 composition. Here, we attempt to account for the variation in phryic sample precursor  
 416 composition using the average compositions of the least altered samples of similar Nb  
 417 content (Table S6). We make a first order estimate of the hydrothermal chemical  
 418 changes to cumulate olivine gabbros through comparison with the average composition  
 419 of the least altered cumulate olivine gabbro samples (Table S6). All troctolite, dunite  
 420 and harzburgite samples are significantly altered, and do not represent precursor  
 421 compositions. Consequently we have not calculated the chemical changes associated  
 422 with alteration of these rocks.

423

424 The change in mass of component  $i$  during alteration ( $\Delta m_i$ ) is the difference between  
 425 the mass of component  $i$  before ( $m_i^0$ ) and after ( $m_i^A$ ) alteration:

$$426 \quad \Delta m_i = m_i^A - m_i^0 \quad (1)$$

427 Where  $m_i^A = c_i^A M^A$  and  $m_i^0 = c_i^0 M^0$  ( $M^A$  and  $M^0$  are the total masses of the altered  
 428 rock and its precursor, respectively, and  $c_i^A$  and  $c_i^0$  are their concentrations of  
 429 component  $i$ ). The mass of the rock after alteration can be determined from the change  
 430 in concentration of an immobile element,  $c_x$

$$431 \quad M^A = \frac{c_x^0}{c_x^A} M^0 \quad (2)$$

432 So Equation 1 can be rewritten:

433 
$$\Delta m_i = M^o \left[ \left( \frac{c_x^o}{c_x^A} \right) c_i^A - c_i^o \right] \quad (3)$$

434 Here we present changes in mass of each element during hydrothermal alteration,  
 435 calculated from the analyzed whole rock compositions and their estimated precursor  
 436 compositions using Equation 3, for an arbitrary initial sample mass ( $M^o$ ) of 100 g (Fig.  
 437 10) and assuming the net mass change is less than 5% (i.e.  $c_x^o/c_x^A = 1 \pm 0.05$ ). The  
 438 errors in the calculated values of  $\Delta m_i$  are propagated on a sample-by-sample basis (Fig.  
 439 10) to include (i) analytical uncertainty, (ii) the uncertainty in the net change in mass;  
 440 and (iii) the uncertainty in estimated precursor compositions (see Supplementary  
 441 Material).

442

443 6. Hydrothermal changes in Macquarie Island whole rock chemistry  
 444 In the following discussion depths (m) refer to the crustal depth within the proto  
 445 Macquarie Island crust (Fig. 4). The calculated hydrothermal changes in whole rock  
 446 chemistry ( $\Delta m_i$ ) are presented as absolute values (per 100g of initial sample; Fig. 10), to  
 447 enable comparison between lithologies and evaluation of hydrothermal exchanges  
 448 within the crust. The crust is divided into 25-400 m thick intervals and the mean  $\Delta m_i$   
 449 within each interval are calculated assuming the samples are representative of the  
 450 lithologies and their proportions in that interval. The results (termed ‘depth-averaged’  
 451 data) are presented in Figure 11. The thicknesses of the selected intervals reflect  
 452 stratigraphic horizons in the section (e.g. the 25 m LDTZ) and the sampling frequency  
 453 versus depth, with depth intervals containing 2-27 samples. The average relative  
 454 changes in composition of each rock type are summarized in Table 1.

455

456 The most significant change in Macquarie Island whole rock chemistry is the increase in  
457 volatile content, which we have determined from altered samples' loss on ignition (LOI)  
458 and their estimated precursor H<sub>2</sub>O contents. A sample's LOI gives an indication of its  
459 volatile content, offset by the increase in mass due to oxidation of ferrous iron during  
460 combustion. Provided the Fe<sup>2+</sup>/Fe<sup>3+</sup> ratio is known and there was complete oxidation of  
461 the ferrous iron, this effect can be corrected for using LOI<sub>actual</sub> = LOI<sub>measured</sub> + 0.11 FeO  
462 (wt%) (Lechner and Desilets, 1987). Assuming that 90% of the total iron is Fe<sup>2+</sup> the  
463 average Fe contents of the Macquarie lavas, dikes and gabbros correspond to an  
464 additional LOI of 0.79, 0.82 and 0.41 wt% respectively. Consequently the hydrothermal  
465 increases in volatile content of Macquarie samples are likely underestimated here by  
466 ~0.4 – 0.8 g/100g. The increase in volatile content primarily reflects hydration and CO<sub>2</sub>-  
467 uptake during fluid-rock interaction. Given limited sample material the C content was  
468 determined for only 55% of samples and LOI data are not corrected for CO<sub>2</sub> content.  
469 However, assuming that all C present in the analyzed samples occurs as CO<sub>2</sub>, the CO<sub>2</sub>  
470 content of the majority of samples is less than 0.5 wt%, with CO<sub>2</sub> contents greater than  
471 1 wt% only observed in five CaCO<sub>3</sub> bearing dike samples. The calculated increases in  
472 volatile content (up to 8.3g/100g) therefore primarily reflect the formation of hydrous  
473 secondary phases, including clay minerals, zeolites, chlorite, amphiboles, epidote, talc  
474 and serpentine. On average all lithologies are moderately to highly enriched in volatiles,  
475 but the calculated enrichment is typically lower for porphyritic lava and dike samples  
476 compared to the aphyric samples (Table 1) because the least altered whole rock samples  
477 chosen to represent precursor compositions are themselves partially hydrated (LOI <2  
478 wt %) relative to the fresh Macquarie glasses (0.25 – 1.49 wt% H<sub>2</sub>O (Kamenetsky et al.,  
479 2000; Wertz, 2003)). Samples from throughout the Macquarie crust are hydrated, with

480 harzburgite-hosted dike and gabbro samples from the base of the section (3300 m)  
481 enriched in volatiles by 3.1 and 2.3 g/100g but CO<sub>2</sub> contents of <0.1 wt%, indicating  
482 that hydrous fluids penetrated through the entire crustal section and into the uppermost  
483 mantle (Fig. 10). The greatest average enrichment (~2.5 g/100g) occurs in the  
484 uppermost sheeted dikes, between 850 and 950 m, decreasing with depth through the  
485 sheeted dikes to ~0.4 g/100g at 1550 m (Fig. 11).

486

#### 487 6.1 Major elements

488 The changes in mass of the major element oxides are variable both within and between  
489 lithologies reflecting differences in secondary mineral assemblages, variations in the  
490 extent of fluid-rock interactions, and the thermal structure of the crust. SiO<sub>2</sub> was  
491 variably enriched or depleted throughout the crust, with ΔSiO<sub>2</sub> within error of zero for  
492 many samples due to the uncertainty in precursor compositions (Fig. 10). SiO<sub>2</sub> is on  
493 average depleted through most of the crust, but enrichments occur in the lavas (500-600  
494 m), LDTZ, uppermost dikes (825-1050 m), and the DGTZ (1550-1650 m), where SiO<sub>2</sub>  
495 enrichment is associated with secondary quartz formation.

496

497 Titanium is usually assumed to be immobile during hydrothermal processes (Bednarz  
498 and Schmincke, 1991; MacLean and Barrett, 1993; Teagle et al., 1996), consistent with  
499 many Macquarie whole rock samples recording enrichments or depletions in TiO<sub>2</sub>  
500 within error of zero change (Fig. 10). Depth-averaged changes in TiO<sub>2</sub> reveal that the  
501 upper 200m of gabbros gained on average ~0.2 g/100g TiO<sub>2</sub>, but Ti was immobile  
502 through the rest of the lower crust (Fig. 11). Ti was also enriched within the upper-

503 sheeted dikes (between 950-1050 m), but depleted or immobile through the lavas and  
504 lower sheeted dikes (Fig. 11). Enrichments likely reflect secondary titanite formation.

505

506 Each Macquarie lithology includes samples that were enriched and samples that were  
507 depleted in  $\text{Al}_2\text{O}_3$ ,  $\text{CaO}$  and  $\text{Na}_2\text{O}$  due to hydrothermal alteration (Fig. 10). The depth-  
508 averaged data indicate that whole rock samples were on average depleted in  $\text{CaO}$  and  
509  $\text{Al}_2\text{O}_3$  and enriched in  $\text{Na}_2\text{O}$  or experienced changes in these elements within error of  
510 zero throughout the majority of the Macquarie crust (Fig. 11). Depletions in  $\text{Al}_2\text{O}_3$  and  
511  $\text{CaO}$  associated with  $\text{Na}_2\text{O}$  enrichments reflect the replacement of calcic plagioclase by  
512 secondary albite, although the loss of  $\text{CaO}$  due to albitisation is partially compensated  
513 by the uptake of Ca in calcium carbonate minerals and prehnite. The plagioclase-phyric  
514 lavas are depleted in Al and Ca and enriched in Na, but to a lesser extent than the  
515 aphyric lavas and dikes (Fig. 10). This may be because the large plagioclase  
516 phenocrysts are relatively fresh in the porphyritic samples, except where they are  
517 intersected by veins. Alternatively, this difference could be an artifact of using 'least  
518 altered' whole rock samples to define the precursor compositions of porphyritic  
519 samples.

520

521 Changes in  $\text{K}_2\text{O}$  are variable throughout the Macquarie section (Fig. 10), but  $\text{K}_2\text{O}$  is on  
522 average enriched in all rock types except the aphyric dikes (Table 1).  $\text{K}_2\text{O}$  enrichment  
523 of the lavas (~0.2 g/100g on average) reflects potassium uptake from seawater-derived  
524 fluids into secondary celadonite and zeolite (phillipsite). The greatest average  $\text{K}_2\text{O}$   
525 enrichment (~0.5 g/100g) occurs at the LDTZ, with the extent of  $\text{K}_2\text{O}$  enrichment  
526 decreasing down through the sheeted dikes. The greatest average  $\text{K}_2\text{O}$  enrichment in the

527 lower crust (0.25 g/100g) occurs between 2100 and 2200 m due to K<sub>2</sub>O enrichment of  
528 anorthosites, in which plagioclase has been seritized.

529

530 Magnesium, iron and manganese were variably enriched or depleted throughout the  
531 Macquarie crust due to hydrothermal exchange, with changes reflecting the  
532 development of saponite, celadonite, Fe-oxyhydroxide and chlorite in the upper crust,  
533 and chlorite, amphibole (actinolite, tremolite and hornblende), talc and serpentine in the  
534 lower crust. There are significant enrichments in MgO, FeO, and MnO in the gabbros  
535 and olivine gabbros at ~2300 m of up to ~15, 3.5 and 0.07 g/100g, respectively (Fig.  
536 10). These calculated enrichments should be treated with caution as they may be  
537 artifacts of estimated precursor compositions that do not adequately account for the  
538 primary mineral modal variation of these samples. However, they are associated with  
539 hydration of up to 8.3 g/100g, consistent with significant fluid-rock interaction that  
540 resulted in the replacement of olivine by talc + amphibole + chlorite + clay ± serpentine.  
541 The depth-averaged  $\Delta$ MgO data indicate that MgO was enriched or within error of zero  
542 change throughout the crustal section, with the exception of the uppermost lower crust,  
543 which lost on average ~0.9 g/100g MgO (Fig. 11). In contrast FeO was on average  
544 depleted or within error of zero change throughout, with the only significant  
545 enrichments (<1.2 g/100g) occurring between 950 and 1050 m and below 2200 m.  
546 Depth averaged  $\Delta$ MnO reveal variable enrichments or depletions within error of zero  
547 change through much of the Macquarie crust. However, MnO is on average enriched in  
548 the lowermost lavas and upper sheeted dikes, with the extent of enrichment increasing  
549 with depth to ~0.05 g/100g at 1050 m, but depleted in the uppermost lower crust (Fig.  
550 11).

551

552 Changes in the sulfur content are typically small (< 0.1 g/100g) through the Macquarie  
553 crustal section and dominated by S-enrichment at the LDTZ (<3.2 g/100g) where  
554 secondary sulfides are most abundant (Davidson et al., 2004). Above the LDTZ, the  
555 majority of aphyric lavas have lost ~0.08 g/100g S, whereas porphyritic lavas show no  
556 change (Fig. 10). This difference likely reflects a very early stage of alteration resulting  
557 in S-depletion that is not observed when comparing to 'least altered' porphyritic  
558 precursors that also experienced S-depletion. The average changes in S content due to  
559 hydrothermal alteration may therefore be underestimated here by 0.04 to 0.08 g/100g,  
560 depending on the relative proportion of the lavas that are porphyritic in each depth  
561 interval (50 – 100%). Similarly, aphyric dikes show greater S-depletions compared to  
562 porphyritic dikes. The LDTZ and upper sheeted dikes are on average enriched in S by  
563 ~0.5 and 0.05 g/100g, respectively, whereas the lower sheeted dikes are on average  
564 depleted in S. This is consistent with leaching of sulfur from the lower sheeted dikes by  
565 high temperature hydrothermal fluids (Alt, 1994, 1995b). In contrast the lower crust is  
566 on average enriched in S (Fig. 11).

567

568 Phosphorous is variably enriched or depleted throughout the Macquarie crustal section  
569 (Fig. 10). On average the gabbros, olivine gabbros and anorthosites are enriched in P<sub>2</sub>O<sub>5</sub>  
570 (Table 1), but their absolute calculated ΔP<sub>2</sub>O<sub>5</sub> values are small (<0.01 g/100g) and  
571 generally within error of no change (Fig. 10) given the uncertainty in precursor P<sub>2</sub>O<sub>5</sub>  
572 concentrations.

573

## 574 6.2 Trace elements

575 The base metals Cu and Zn were highly mobile during alteration of the Macquarie  
576 Island crust (Fig. 10). Cu is accommodated in fresh rocks by primary sulfides that  
577 typically occur as 'blebs' in the groundmass (Doe, 1994). The estimated precursor Cu  
578 contents of the porphyritic lavas and dikes ( $\sim 25 \pm 7$ ) is lower than that of the aphyric  
579 rocks ( $\sim 100 \pm 20$  ppm), which may reflect the lower proportion of groundmass in  
580 porphyritic rocks. Alternatively it could be an artifact of using 'least altered samples' as  
581 precursors to these rocks if they were depleted in sulfur and consequently Cu during  
582 early hydrothermal alteration. Cu is on average moderately depleted in the aphyric lavas  
583 and dikes (Table 1). Cu-depletion of the aphyric rocks is less pronounced at the LDTZ  
584 (Fig. 10) coincident with the zone where secondary sulfides including chalcopyrite are  
585 most abundant, indicating that the depletion is partially compensated by secondary  
586 mineralization. Aphyric dikes from the sheeted dike complex experienced a greater  
587 extent of Cu-depletion (<15 mg/100g) than those from below the DGTZ (<10 mg/100g;  
588 Fig. 10). Consequently the greatest depth-averaged Cu depletions occur within the  
589 sheeted dikes (Fig. 11).

590

591 Zn is mildly incompatible during magmatic fractionation and is concentrated in the  
592 glass phase, but can be incorporated into olivine, spinel, magnetite, titanomagnetite and,  
593 to a lesser extent, pyroxene (Doe, 1994). Zn behaves more variably than Cu, and  
594 absolute changes in Zn content are smaller (<6 mg/100g). The lavas are variably  
595 enriched or depleted in Zn, with no systematic difference in  $\Delta\text{Zn}$  of aphyric and  
596 porphyritic lava samples (Fig. 10). Within the LDTZ there is variable Zn-enrichment (<  
597 5 mg/100g) consistent with quartz-sphalerite-chalcopyrite mineralization in this zone

598 (Davidson et al., 2004). Aphyric dikes record increasing Zn depletion with depth  
599 through the sheeted dike complex, to ~5 mg/100g at the DGTZ, whereas Zn is variably  
600 enriched or depleted in porphyritic dikes and gabbros in this interval (Fig. 10). Although  
601 there are no correlations between  $\Delta S$  and either  $\Delta Cu$  or  $\Delta Zn$  within the Macquarie crust,  
602  $\Delta Cu$  and  $\Delta Zn$  are weakly correlated in dike samples ( $R^2 = 0.44$ ,  $n=59$ ). The depletion of  
603 Cu and Zn from the aphyric dikes is consistent with previously observed base-metal  
604 losses from the lower sheeted dikes of ODP Holes 504B and 1256D due to the  
605 breakdown of sulfide minerals and titanomagnetite under greenschist facies conditions  
606 (Alt et al., 2010; Alt et al., 1996a).

607

608 The alkali elements Cs and Rb are incorporated into minerals by cation-substitution,  
609 with preferential substitution for ions of similar ionic potential. They therefore have an  
610 affinity for K-rich phases including clays and feldspars (Hart, 1969). Consequently  $\Delta Rb$   
611 and  $\Delta K_2O$  of Macquarie samples are correlated ( $R^2 = 0.82$ ,  $n= 203$ ). Rb and Cs display  
612 similar average behavior in most Macquarie lithologies except olivine gabbros, which  
613 are on average highly enriched in Rb but slightly depleted in Cs (Table 1). The  
614 anorthosites are highly enriched in Cs and Rb, which may reflect the seritization of  
615 plagioclase. The Cs and Rb enrichment of porphyritic samples relative to aphyric  
616 samples at a given depth may reflect the greater proportion of plagioclase feldspar in the  
617 phryic samples. The observed average enrichments of K, Rb and Cs in the extrusive  
618 section are consistent with uptake from seawater during low-temperature hydrothermal  
619 alteration of in-situ upper crust (Staudigel and Hart, 1983; Teagle et al., 1996). The  
620 depletion of Rb and Cs in the lower sheeted dikes and upper gabbros (Fig. 10) is

621 consistent with their enrichment in hydrothermal fluids relative to both MORB and  
622 seawater (Palmer and Edmond, 1989; Von Damm, 1995).

623

624 The alkali earth elements Sr and Ba have an affinity for Ca, and Sr is therefore  
625 predominantly incorporated into Ca-rich phases such as feldspar, calcium carbonate,  
626 prehnite, and epidote. Sr is variably enriched or depleted throughout the Macquarie  
627 crust (Fig. 10). This variability reflects differences in the secondary mineral  
628 assemblages, with samples that contain significant volumes of calcium carbonate or  
629 epidote having gained Sr. On average Sr is slightly depleted in the aphyric lavas and  
630 dikes, but slightly to highly enriched in gabbro, olivine gabbro and anorthosite.  
631 Consequently, Sr is on average depleted or within error of zero change through the  
632 extrusive section and sheeted dikes, but enriched or within error of zero change through  
633 the lower crust (Fig. 11). This indicates that hydrothermal circulation caused a net  
634 transfer of Sr downward within the crust. Ba displays similar behavior to Sr on average  
635 in most Macquarie lithologies (Table 1) but unlike Sr, Ba is moderately enriched in  
636 porphyritic lavas and is therefore on average enriched through the upper 500 m of the  
637 crust (Fig. 11).

638

639 Uranium is variably enriched or depleted throughout the Macquarie crust (Fig. 10), and  
640 the depth-averaged changes in U content are within error of zero throughout most of the  
641 crustal section, with significant U-enrichment (<0.025 mg/100g) occurring only in the  
642 lavas (0-100 m and 400-500 m) as a result of U-uptake from cold seawater. U is  
643 depleted on average in the Macquarie lower lavas and upper sheet dikes, by up to 0.02  
644 mg/100g, indicating U-loss during higher temperature fluid-rock reaction. In contrast,

645 the lower lavas and upper sheeted dikes of ODP Hole 504B were enriched in U under  
 646 greenschist facies conditions (Bach et al., 2003). Calculated values of  $\Delta\text{La}$  and  $\Delta\text{Lu}$  are  
 647 typically small relative to the computational uncertainty for the majority of samples  
 648 (Fig. 10), consistent with previous observations that the rare earth elements are  
 649 generally immobile during hydrothermal alteration of in-situ ocean crust (Bach et al.,  
 650 2003; Teagle et al., 1996).

651

652 7. Elemental exchange fluxes for Macquarie Island

653 To assess the contribution hydrothermal alteration of the Macquarie crust made to  
 654 global geochemical cycles the calculated changes in composition of the Macquarie crust  
 655 due to hydrothermal alteration are converted into net fluxes to or from the crust,  
 656 following:

$$F_i = \sum_t^T \Delta\bar{m}_{i-t} z_t \rho_c \quad (4)$$

657 Where  $F_i$  is the mass flux of component  $i$  through  $1 \text{ m}^2$  of seafloor due to alteration of a  
 658 section of crust of thickness  $T$  and  $\Delta\bar{m}_{i-t}$  is the average change in mass of component  $i$   
 659 per unit mass of rock in each sub-interval  $t$ ,  $z_t$  is the thickness of each sub-interval  $t$  and  
 660  $\rho_c$  is the density of the crust ( $2900 \text{ kg/m}^3$ ). The net fluxes due to hydrothermal alteration  
 661 of the Macquarie lavas, LDTZ, sheeted dikes and gabbros, are determined from the  
 662 depth-averaged  $\Delta m_i$  (Table S7), with uncertainty propagated from the standard error of  
 663 the mean chemical change of each sub-interval (Supplementary Material). However, to  
 664 fully account for the fluxes associated with hydrothermal alteration we also need to  
 665 include the contribution from the veins that formed from fluids circulating through  
 666 fractures in the crust.

667

668 Veins were deliberately excluded from the whole rock samples, as they need to be  
 669 included in a manner representative of their true proportion within the crust, rather than  
 670 individual samples. Unfortunately it is not possible to accurately estimate the proportion  
 671 of the Macquarie crust that veins make up without drill core. However the volume and  
 672 mineralogy of hydrothermal veins have been quantitatively logged in sections of ocean  
 673 crust recovered by scientific ocean drilling. We therefore estimate the volume  
 674 proportions of the various secondary vein-filling minerals through the Macquarie crust  
 675 from their occurrence in representative drilled sections of ocean crust (Table 2). These  
 676 data are combined with published major element analyses of secondary mineral  
 677 compositions (Table 2) to estimate the net effect of hydrothermal vein formation in the  
 678 Macquarie crust on global geochemical cycles. Assuming that the veins formed in open  
 679 fractures, the net mass flux of component  $i$  through  $1\text{m}^2$  of seafloor due to hydrothermal  
 680 vein formation in crustal interval  $T$  is given by:

$$F_{i-v} = z_T \sum_x V_x \rho_x M_{i-x} \quad (5)$$

681 where  $z_T$  is the thickness of interval  $T$ ,  $V_x$  is the volume proportion of the rock filled  
 682 with vein mineral  $x$ ,  $\rho_x$  is the specific gravity of mineral  $x$ , and  $M_{i-x}$  is the mass  
 683 proportion of component  $i$  in vein mineral  $x$ . The net uptake of major elements as a  
 684 result of vein formation in the Macquarie lavas, LDTZ, sheeted dikes and gabbro are  
 685 calculated using equation 5. The net mass flux of component  $i$  through  $1\text{m}^2$  of seafloor  
 686 due to hydrothermal alteration of crustal interval  $T$  is given by combining Equations 4  
 687 and 5:

$$F_{i-TOTAL} = F_{i-v} + (1-V).F_i \quad (6)$$

689 where  $V$  is the total volume proportion of veins in interval  $T$ . The hydrothermal fluxes  
690 due to alteration of the Macquarie Island lavas, LDTZ, sheeted dikes and gabbros  
691 calculated using Equation 6 (Table S7) are extrapolated to global annual fluxes (Table  
692 3), assuming a global ocean crustal production rate of  $\sim 3 \text{ km}^2/\text{yr}$  (Müller et al., 2008).

693

#### 694 8. Global Hydrothermal Fluxes

695 The ultimate goal of this and similar studies is to quantify the hydrothermal  
696 contributions from seawater-ocean crust exchange to global biogeochemical cycles, and  
697 assess how they have varied in the past. Hydrothermal fluxes between a given section of  
698 crust and the overlying ocean depend on the crust's age, architectural and thermal  
699 history, and the spreading rate. Consequently hydrothermal contributions to global  
700 geochemical cycles depend on the global length of slow, intermediate and fast spreading  
701 ridges and the age-area distribution of the ridge flanks, all of which have varied  
702 significantly throughout the Phanerozoic (Müller et al., 2008). To achieve this goal we  
703 therefore require complete sections of altered ocean crust produced at different  
704 spreading rates and at different times.

705

706 Several studies have quantified the chemical changes associated with hydrothermal  
707 alteration from sections of ocean crust, recovered through scientific ocean drilling of in-  
708 situ crust or tectonically uplifted lower crust, and sampling of crust tectonically exposed  
709 on the seafloor or sub-aerially exposed in ophiolites (for example, Bach et al., 2003;  
710 Bednarz and Schmincke, 1989; Coogan and Dosso, 2012; Gillis and Robinson, 1990;  
711 Staudigel, 2014). These studies vary in many ways, including: (i) the properties of the  
712 section studied, including: spreading rate, age, depth interval, and its thermal,

713 architectural, and hydrogeologic evolution; (ii) the elements investigated; (iii) the  
714 assumptions and numerical approaches used to compute chemical changes; (iv) the  
715 parameters used to extrapolate calculated chemical changes to global hydrothermal  
716 fluxes; and (v) their assessment of the uncertainties involved. Unfortunately our current  
717 sampling of in-situ ocean crust is too sparse to make a detailed assessment of the  
718 variations in hydrothermal fluxes with respect to spreading rate and crustal age, and a  
719 full comparison of the results of the sections sampled to date is beyond the scope of this  
720 investigation.

721

722 Here we compare the global hydrothermal fluxes extrapolated from chemical changes  
723 through the Macquarie crust with: (i) fluxes determined from hydrothermal chemical  
724 changes through a composite section of ocean crust recovered by scientific ocean  
725 drilling (Staudigel, 2014); and (ii) fluxes extrapolated from element/heat ratios of  
726 sampled hydrothermal fluids and estimates of the total convective heat loss (Elderfield  
727 and Schulz (1996) (Table 4). The altered sections of ocean crust provide time-integrated  
728 records of hydrothermal alteration, and hence should be comparable to the net axial and  
729 ridge flank fluxes (Fig. 12).

730

731 The chemical changes for the composite section of ocean crust are primarily based on  
732 analyses of in-situ lavas from DSDP Holes 417D and 418A (120 Ma, slow-spread crust;  
733 (Donnelly et al., 1979)) and tectonically uplifted gabbros from ODP Hole 735B; (9 Ma,  
734 ultraslow-spread crust (Dick et al., 1999)). Although in-situ sheeted dikes were  
735 recovered from in-situ crust at ODP Site 504 (6.9 Ma, intermediate-spread crust (Alt et  
736 al., 1996b)), the chemical changes at intermediate depths in the composite section were

737 in most cases extrapolated from analyses of Site 417/418 lavas and 735B gabbros  
738 because Site 504 was considered to be too young to represent hydrothermally mature  
739 ocean crust (Staudigel, 2014). Such an approach ignores the significant differences  
740 between the hydrothermal processes and hence chemical reactions occurring in the  
741 dikes compared to both the lavas and the lower crust. The composite section was  
742 therefore produced at a similar (slow) spreading rate to the Macquarie crust.  
743 Uncertainties in the estimated chemical changes are only quoted for the upper 600 m of  
744 lavas in the composite section (Staudigel, 2014), hence the uncertainties in the  
745 extrapolated annual hydrothermal fluxes (Table 4; Fig. 12) are not known.

746  
747 Differences between the average chemical changes recorded by the two crustal sections  
748 likely reflect (i) differences in their stratigraphy, predominantly due to the greater  
749 thickness of gabbro in the composite section (5000 m; (Staudigel, 2014)) compared to  
750 the Macquarie crust (1150 m; Fig. 4); (ii) the longer duration of hydrothermal exchange  
751 within the lavas of the composite section, given their greater age (120 vs 10 Ma); (iii)  
752 their differing geologic histories; and (iv) and the approaches used to calculate the  
753 chemical changes. In contrast to the Macquarie crust and sampled vent fluids, the  
754 composite crust records net fluxes of Si, Al and Ca to the crust and a net flux of Mg to  
755 the oceans (Fig. 12). The discrepancies between the two studies of slow-spread crust  
756 emphasize the need for more thorough sampling of in-situ ocean crust, a consistent  
757 approach to calculating chemical changes, and full consideration of the associated  
758 uncertainties.

759

760 Assuming that hydration is the primary cause of volatile enrichment, Macquarie-style  
761 alteration results in uptake of  $1.9 \times 10^{13}$  mol/year of water. In contrast hydrothermal  
762 alteration of the Macquarie crust resulted in a net flux of Si to the oceans, consistent  
763 with the observed Si-enrichment of black smoker and ridge flank vent fluids relative to  
764 seawater (Butterfield et al., 2003; Wheat and Mottl, 2000). Macquarie whole rock  
765 alteration extrapolates to a Si flux to the oceans of  $9.7$  to  $25 \times 10^{11}$  mol/year but a large  
766 proportion is compensated by vein formation, resulting in a net flux to the oceans of  $3.7$   
767  $\pm 7.3 \times 10^{11}$  mol/year. This is less than the estimated combined Si flux from axial vents  
768 and ridge flank exchange (Elderfield and Schultz (1996); Fig. 12).

769

770 Ti was assumed to be immobile during hydrothermal alteration of the composite crustal  
771 section ((Staudigel, 2014) Table 4). However, our results indicate that hydrothermal  
772 alteration causes a net Ti flux to the oceans of  $0.8$ – $2.5 \times 10^{11}$  mol/year, equivalent to 3–  
773 8% loss of Ti from the full crustal section (Table 3). Similar Ti losses (15%) from upper  
774 crustal samples from ODP Hole 504B cannot solely be attributed to dilution through  
775 secondary mineral infilling and require Ti mobility during hydrothermal alteration  
776 (Bach et al., 2003). However, titanite is common in small amounts in chlorite veins in  
777 ODP Holes 504B and 1256D (Alt et al., 1996a; Teagle et al., 2006) and the Macquarie  
778 crust, indicating that some of the mobilized Ti is re-incorporated into veins. The titanite  
779 abundance in the drilled crustal sections is not quantified and titanite-uptake is not  
780 included in our hydrothermal flux estimates. Consequently the estimated Ti-loss from  
781 the Macquarie crust is a maximum estimate.

782

783 The hydrothermal changes in Macquarie crustal composition extrapolate to a global Mg  
784 flux of  $3.3 \pm 1.1 \times 10^{12}$  mol/year into the crust, but indicate that there is no net flux of  
785 Fe or Mn into or out of the crust (Fig. 12). Our extrapolated Mg flux is comparable to  
786 the combined estimates of the axial and ridge flank Mg flux ( $2.5 \pm 0.2 \times 10^{12}$  mol/yr;  
787 Elderfield and Shultz (1996)). Approximately half the Macquarie Mg-uptake occurs in  
788 the lower crust (Table 3). Most likely due to tectonic exposure of the lower crust to  
789 seawater with the greatest Mg-enrichment observed 1m from a chloritic fault zone. The  
790 timing of this alteration is uncertain, and could have occurred on axis, off axis, or  
791 during uplift. Significant Mg-enrichment of the lower crust maybe therefore be  
792 restricted to crust produced at slow spreading ridges, where amagmatic extension results  
793 in uplift and exposure of the lower crust (MacLeod et al., 2009). However, our results  
794 contrast with those from ODP Hole 735B, where low-temperature alteration of uplifted  
795 lower crust acts as a source of Mg to the oceans, rather than a sink (Bach et al., 2001).  
796 The behavior of Mg during uplift of the lower crust may depend on the timing and rate  
797 of exhumation, which affect the thermal and chemical conditions of fluid-rock  
798 interaction as indicated by differing clay mineral distributions in Holes 735B and  
799 U1309D (Nozaka et al., 2008). Penetration of cold seawater causes oxidation and Mg  
800 removal but reaction with warmer fluids leads to chlorite and smectite precipitation and  
801 Mg uptake.

802

803 The modern oceans contain  $\sim 75 \times 10^{18}$  moles Mg, but the Mg concentration of seawater  
804 has increased from  $\sim 35$  to  $55$  mmol/kg since the Neogene (Horita et al., 2002). Such an  
805 increase requires an average net Mg flux of  $\sim 1 \times 10^{12}$  mol/yr to the oceans since 35 Ma.  
806 Global hydrothermal Mg-uptake at the upper end of our estimated range ( $4.3 \times 10^{12}$

807 mol/year) is consistent with the observed increase in Mg in seawater since the Neogene,  
808 given an estimated riverine Mg input from the continents of  $5.2 \times 10^{12}$  mol/year  
809 (Edmond et al., 1979). If the Mg-uptake in the Macquarie lower crust is globally  
810 representative, our results therefore indicate that Mg uptake during axial and ridge flank  
811 hydrothermal alteration is sufficient to balance the Mg-budget of the oceans, given the  
812 uncertainties involved.

813

814 Hydrothermal alteration of the Macquarie crust supplied Ca to the oceans, but provided  
815 a net sink for Na. Given our estimated Ca flux to the oceans ( $4.2\text{--}5.5 \times 10^{12}$  mol/year)  
816 and a maximum axial Ca flux of  $1.3 \times 10^{12}$  mol/year (Elderfield and Shultz, 1996) at  
817 least two thirds of the Ca removal must occur on the ridge flanks. We estimate that  
818 hydrothermal exchange removes  $3.3\text{--}5.2 \times 10^{11}$  mol/yr of K from the oceans, but it is  
819 neither a net sink nor source for Cs and Rb because their removal during high  
820 temperature reactions is compensated by uptake during low temperature alteration.

821

822 Our extrapolated flux of S to the crust ( $1.9\text{--}2.7 \times 10^{11}$  mol/yr) is sensitive to the  
823 volume of vein pyrite, which accounts for three quarters of this uptake (Table 3). Bulk  
824 rock S contents combined with pyrite and anyhydrite vein abundances in crust from  
825 ODP Holes 504B and 735B indicate that the S flux from the volcanic crust to the oceans  
826 ( $2.5 \times 10^{12}$  g/yr) is approximately compensated by S uptake in the lower crust ( $2.1 \times$   
827  $10^{12}$  g/yr) (Alt, 1995b). Elderfield and Shultz (1996) estimate axial fluxes of  $0.85\text{--}9.6 \times$   
828  $10^{11}$  mol/year H<sub>2</sub>S to the oceans and  $8.4 \times 10^{11}$  mol/year SO<sub>4</sub> to the crust, resulting in a  
829 net axial S flux of  $3.2 \pm 4.4 \times 10^{11}$  mol/year of S to the crust. The discrepancy between  
830 the estimated axial and time integrated hydrothermal S fluxes indicate that much of the

831 anhydrite is either dissolved at lower temperatures on the ridge flanks, or that the  
832 paucity of anhydrite in drilled crust reflects a sampling bias due to incomplete core  
833 recovery (Alt, 1995b). Although anhydrite was not recovered from the sampled  
834 Macquarie section, gypsum (formerly anhydrite) is locally present elsewhere in the  
835 Macquarie crust (Alt et al., 2003).

836

837 For the majority of the trace elements (Zn, Sr, Rb and Cs) we find no conclusive  
838 evidence that the hydrothermal alteration crust results in a net flux either to or from the  
839 oceans. The Macquarie-based estimates of global Si and S fluxes discussed above  
840 confirm the importance of determining the contribution of hydrothermal veins (Alt,  
841 1995a; Alt et al., 1986; Alt and Teagle, 2000). Unfortunately we do not have trace  
842 element analyses of all the hydrothermal vein minerals, so the extent to which any trace  
843 element loss from the whole rock samples is compensated by uptake in veins is not  
844 determined. In the absence of uptake by veins our results indicate a net flux of  $\sim 6.5 \times$   
845  $10^9$  mol/year of Cu to the oceans and  $\sim 1.3 \times 10^9$  mol/year of Ba to the crust. However,  
846 given the occurrence of minor chalcopyrite ( $\text{CuFeS}_2$ ) veins at least some of the Cu lost  
847 from the bulk rock is re-incorporated into veins.

848

## 849 9. Conclusions

850 Most elements were variably enriched or depleted through the Macquarie crust during  
851 hydrothermal alteration. The changes in bulk rock composition (enrichment or  
852 depletion) depend upon the secondary mineral assemblages developed, and are  
853 controlled by: (i) the modal abundances of the primary minerals in the rocks; (ii) the  
854 alteration conditions such as temperature, fluid composition, or water: rock ratios; and

855 (iii) the chemical behavior of the elements, such as their mobility in fluid. Consequently  
856 there are variations with depth, most notably an interval of greater fluid-rock reaction at  
857 the lava-dike transition zone where lavas and dikes are enriched in K, S, Rb, Ba, and  
858 Zn. Since the rocks provide a time-integrated record of alteration, the behavior of some  
859 elements appears complex, and initial changes during high temperature alteration may  
860 be partially or completely compensated for during later, low-temperature alteration.

861

862 The hydrothermal changes in Macquarie crustal composition are used to estimate net  
863 elemental fluxes to and from the crust. Our results indicate that hydrothermal alteration  
864 results in a net flux of Si, Ti, Al, and Ca, to the oceans, whereas the crust is a net sink  
865 for H<sub>2</sub>O, Mg, Na, K, and S. Our results also demonstrate the importance of accounting  
866 for hydrothermal uptake in veins, which affects the seawater-ocean crust exchange  
867 budgets of Si, Fe, Mg and S. The extrapolation of the hydrothermal changes through a  
868 section of ocean crust to global hydrothermal fluxes is limited by how representative  
869 that section of crust is. The relationship between spreading rate and hydrothermal flux  
870 remains poorly known. Consequently, the approach described here needs to be applied  
871 to ocean crustal sections produced at a range of spreading rates to refine global  
872 hydrothermal flux estimates.

873

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890 **Figure Captions**

891 **Figure 1.** Summary geological map of Macquarie Island, after Goscombe and Everard  
892 (1998a) and Varne et al. (2000). Major faults interpreted to be seafloor-spreading  
893 structures are distinguished, following Wertz et al. (2003) and Daczko et al. (2005).  
894 Inset map shows the main regional tectonic features, after Kamenetsky et al. (2000).

895

896 **Figure 2.** Cross-sections parallel to: **(a)** the Mount Waite transect; and **(b)** the Double  
897 Point transect, after Davidson et al. (2004). Representative lava bedding and dike  
898 chilled margin orientations are indicated.

899

900 **Figure 3. (a).** Geological map of the Isthmus-Eagle Point transect area, after Goscombe  
901 and Everard (1998a), supplemented with our own structural measurements. Lithological  
902 and structural data are projected onto a line perpendicular to the paleo-vertical to  
903 construct a cross-section **(b)**, with reconstructed proto-Macquarie Island crustal depths  
904 indicated along this line.

905

906 **Figure 4.** Stratigraphic reconstruction of the proto-Macquarie Island ocean crust and  
907 depth distribution of secondary minerals. Details within the extrusive section after  
908 Goscombe and Everard (1998a). Secondary mineral distribution in upper crust after  
909 Griffin (1982), supplemented by our own observations.

910

911 **Figure 5 (a)** Y and Zr concentrations of Macquarie Island glasses, after Kamenetsky et  
912 al. (2000) and Wertz (2003). Group 1 glasses (black circles) comprise a suite of  
913 primitive parental magmas from which Group 2 melts (open circles) evolved as a result

914 of fractional crystallization. MORB glasses from a global database of spreading ridges  
915 (grey diamonds) and fracture zones (open diamonds) are shown for comparison (Jenner  
916 and O'Neill, 2012). **(b)** Comparison of Y and Zr concentrations of Macquarie Island  
917 glasses (Kamenetsky et al., 2000; Wertz, 2003) and whole rock samples.

918

919 **Figure 6.** Whole rock chemistry of the proto-Macquarie Island crust. Shaded area  
920 indicates the range of fresh Macquarie glass compositions, after Kamenetsky et al.  
921 (2000) and Wertz (2003).

922

923 **Figure 7.** Immobile element behavior during alteration. During alteration, immobile  
924 element concentrations change only in passive response to net mass loss or gain (black  
925 arrows). **(a)** If two immobile elements are similarly incompatible the magmatic  
926 fractionation crystallization trend (blue arrow) parallels the alteration trend. **(b)** If one  
927 element is compatible the magmatic and alteration trends are distinct, and the precursor  
928 composition to an altered rock is given by their intersection (e.g. altered rocks 1 and 2  
929 have precursor compositions  $1^i$  and  $2^i$ , respectively). The variation of immobile  
930 elements in Macquarie magmas is more complex: **(c)** if both elements are incompatible  
931 the fractional crystallization of a series of primitive parental melts (produced by  
932 differing extents of partial melting) yields an array of precursor compositions (shaded  
933 region). **(d)** A similar ‘precursor array’ is produced if one element is compatible. Here  
934 altered rock 1 must have lost mass to concentrate the immobile elements, but rock 2  
935 could be fresh or may have altered from a wide range of precursors spanning from  $2'$  to  
936  $2''$ .

937

938 **Figure 8.** Comparison of (a) Zr/Y and (b) Nb/Y ratios for pairs of more- and less-  
939 altered rocks from the same precursor; glass compositions after Wertz (2003).

940

941 **Figure 9.** Flow chart showing the methods used to calculate precursor compositions for  
942 the different types of altered whole rock samples: (i) aphyric lavas and dikes, and  
943 plutonic rocks with fractionated magma compositions; (ii) phryic lavas and dikes; and  
944 (iii) cumulate plutonic rocks.

945

946 **Figure 10.** Calculated hydrothermal changes in mass of major element oxides and trace  
947 elements through the proto-Macquarie Island crust, for 100 g of precursor. Errors  
948 propagated on a sample-by-sample basis (see Supplementary Material).

949

950 **Figure 11.** Average hydrothermal changes in mass of major element oxides (g per 100g  
951 precursor) and trace elements (mg per 100g precursor), through the proto-Macquarie  
952 Island crust. Error bars show the standard error of the mean change in mass in each  
953 depth interval.

954

955 **Figure 12.** Global hydrothermal fluxes (mol/year) extrapolated from calculated  
956 hydrothermal changes in Macquarie crustal whole rock composition, excluding (white  
957 bars) and including (grey bars) estimated contributions from veins. Results are  
958 compared to: global hydrothermal fluxes extrapolated from chemical changes within a  
959 composite crustal section recovered by scientific ocean drilling (green ovals; (Staudigel,  
960 2014)); and global fluxes through axial vents (red bars), ridge flanks (blue bars), and  
961 both combined (purple bars), after (Elderfield and Schultz, 1996).

962 **Table Footnotes:**963 **Table 1:**

964 Blank = variable behavior; (+) < 10% enriched; + slightly (10-50%) enriched; ++ moderately (50-  
 965 100%) enriched; +++ highly (>100%) enriched; (-) = < 10% depleted; - slightly (10-50%)  
 966 depleted; -- moderately (50-100%) depleted.

967

968 **Table 2:**

969 a: average vol% of Hole 1256D inflated flows and sheet and massive flows (Wilson et al., 2003);  
 970 b vol% in Hole 1256D (Teagle et al., 2006); c: vol% in Hole 735B (Dick et al., 1999); d: vol%  
 971 estimated from Macquarie Island sample petrographic observations; e: median Hole 1256D vein  
 972 mineral composition (Alt et al., 2010); f: mean Hole 504B secondary mineral composition  
 973 (Laverne et al., 1995); g: mean Hole 504B vein and halo amphibole composition (Vanko et al.,  
 974 1996); h: mean Hole 735B vein feldspar composition (Robinson et al., 2002); i: mean Hole 896A  
 975 secondary Na-zeolite composition (Laverne et al., 1996); j: mean of Holes 1274A, 1268A and  
 976 1272A vein serpentine compositions (Moll et al., 2007).

977

978 **Table 4:**

979 a: average bulk chemical changes for a composite drilled ocean crustal section (Staudigel, 2014);  
 980 b: average bulk chemical changes through the Macquarie crust, including vein-contributions for  
 981 major elements; c: after Staudigel (2014), total crustal thickness = 7000 m; d: total crustal  
 982 thickness = 2700 m; e: after Elderfield and Schulz (1996); f: Ti assumed to be immobile; - not  
 983 determined.

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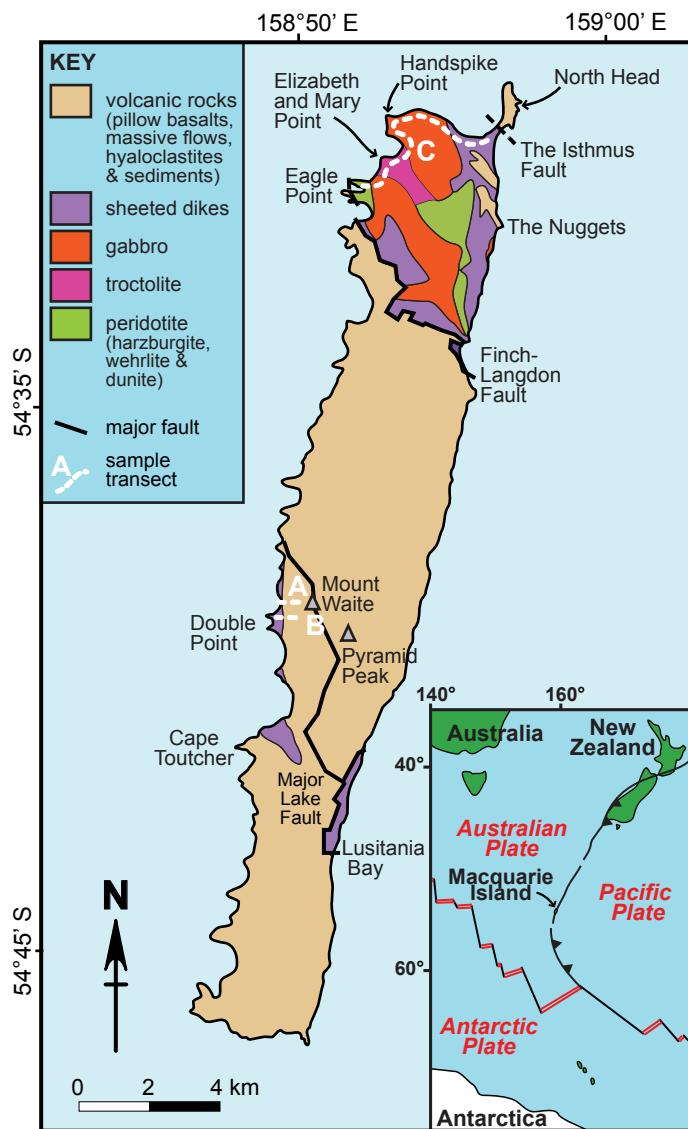
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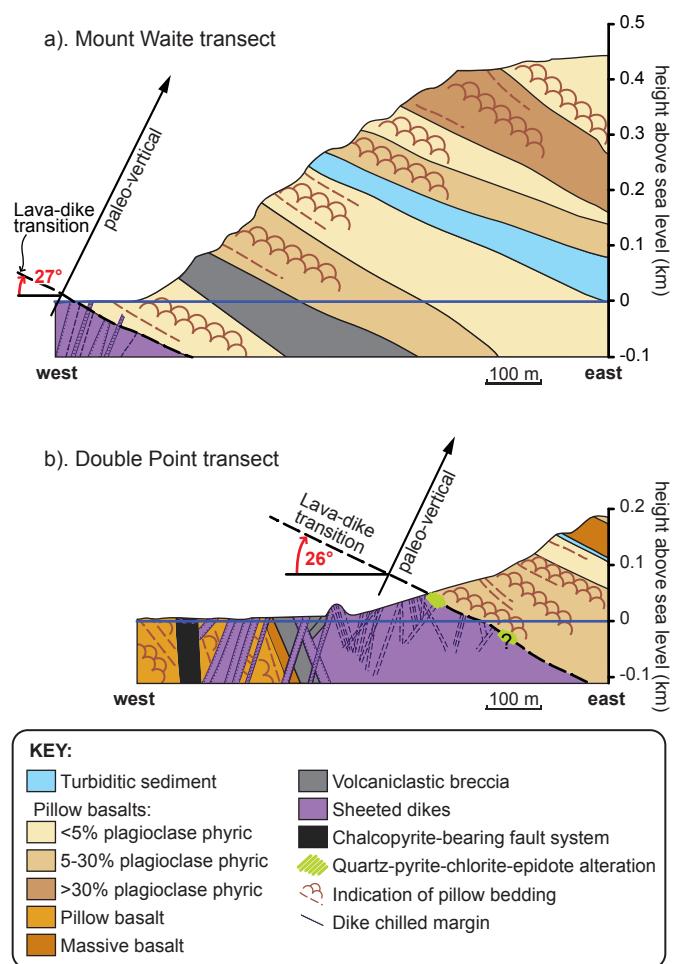
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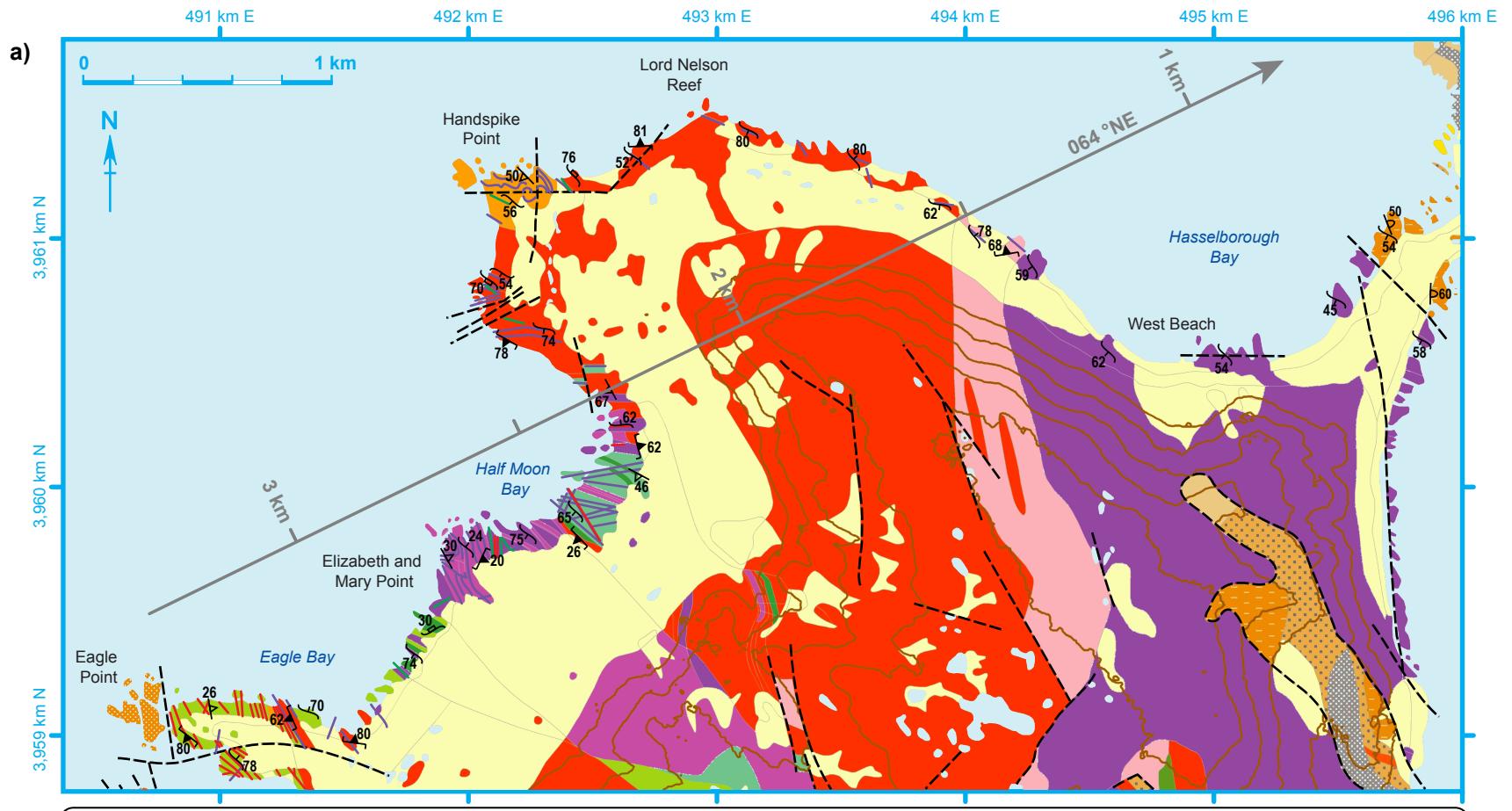


**Figure 1**



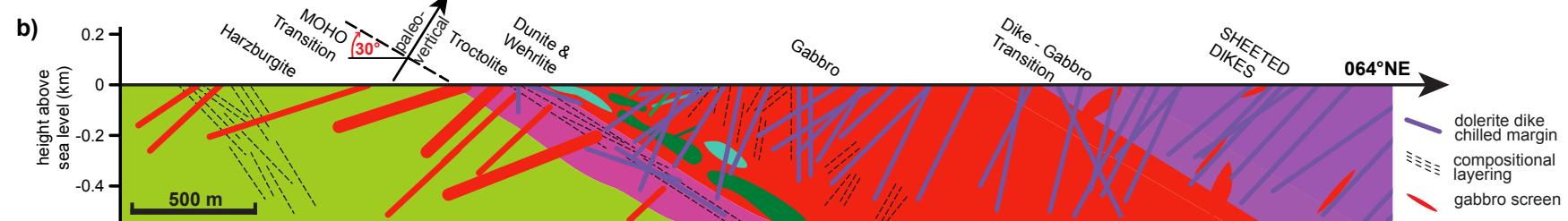
**Figure 2.**

Figure 3.

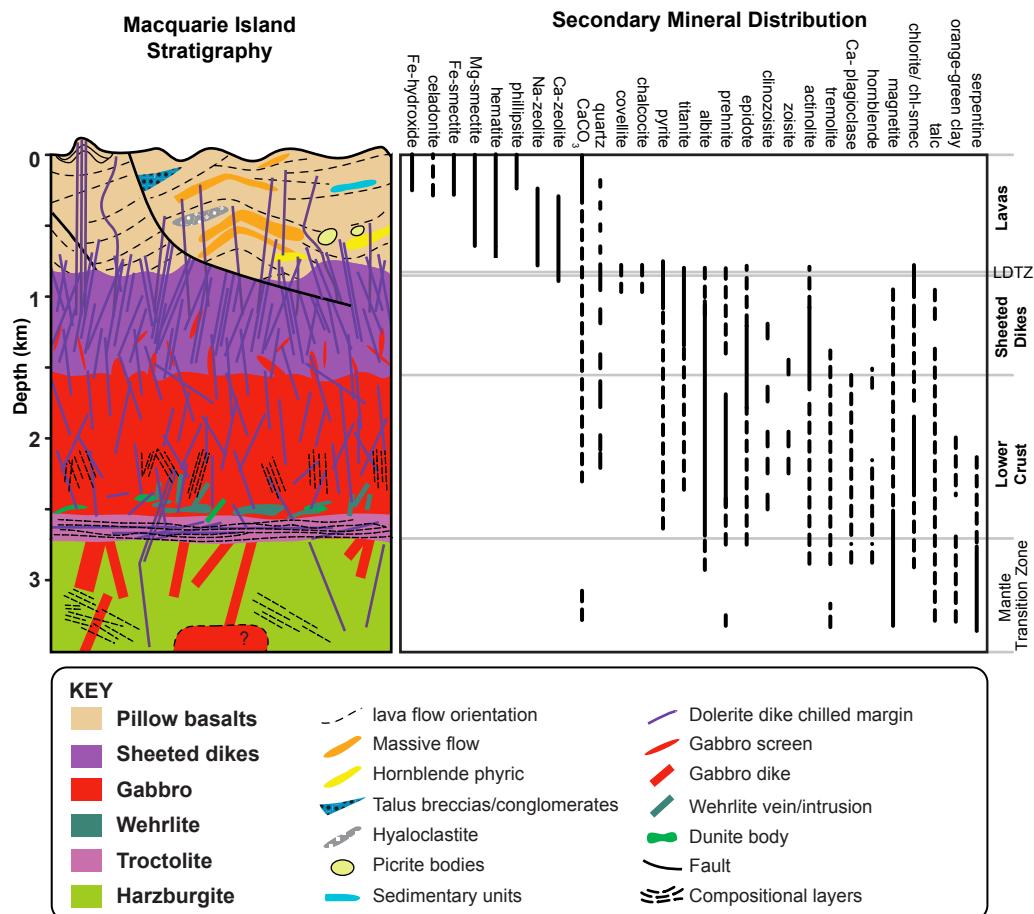


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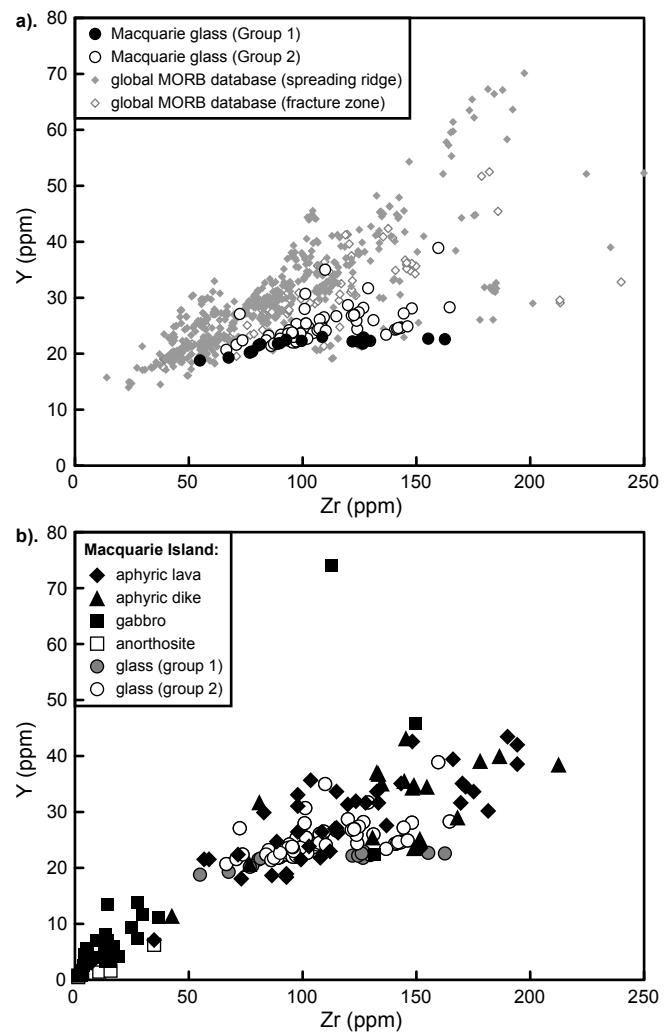
cover	pillow basalt (>30% plag phric)	troctolite	50 m contours	bedding (flattened pillows)
volcaniclastic breccia	sheeted dikes (<30% gabbro)	plagioclase wehrlite + dunite	-	strike and dip (dolerite dike)
pillow basalt	dike - gabbro transition zone	dunite + plagioclase dunite	-	strike and dip (screen)
pillow basalt (<5% plag phric)	gabbro & olivine gabbro	harzburgite + dunite + wehrlite	-	gabbro screen trace
pillow basalt (5-30% plag phric)	layered gabbro	harzburgite	-	ultramafic screen trace



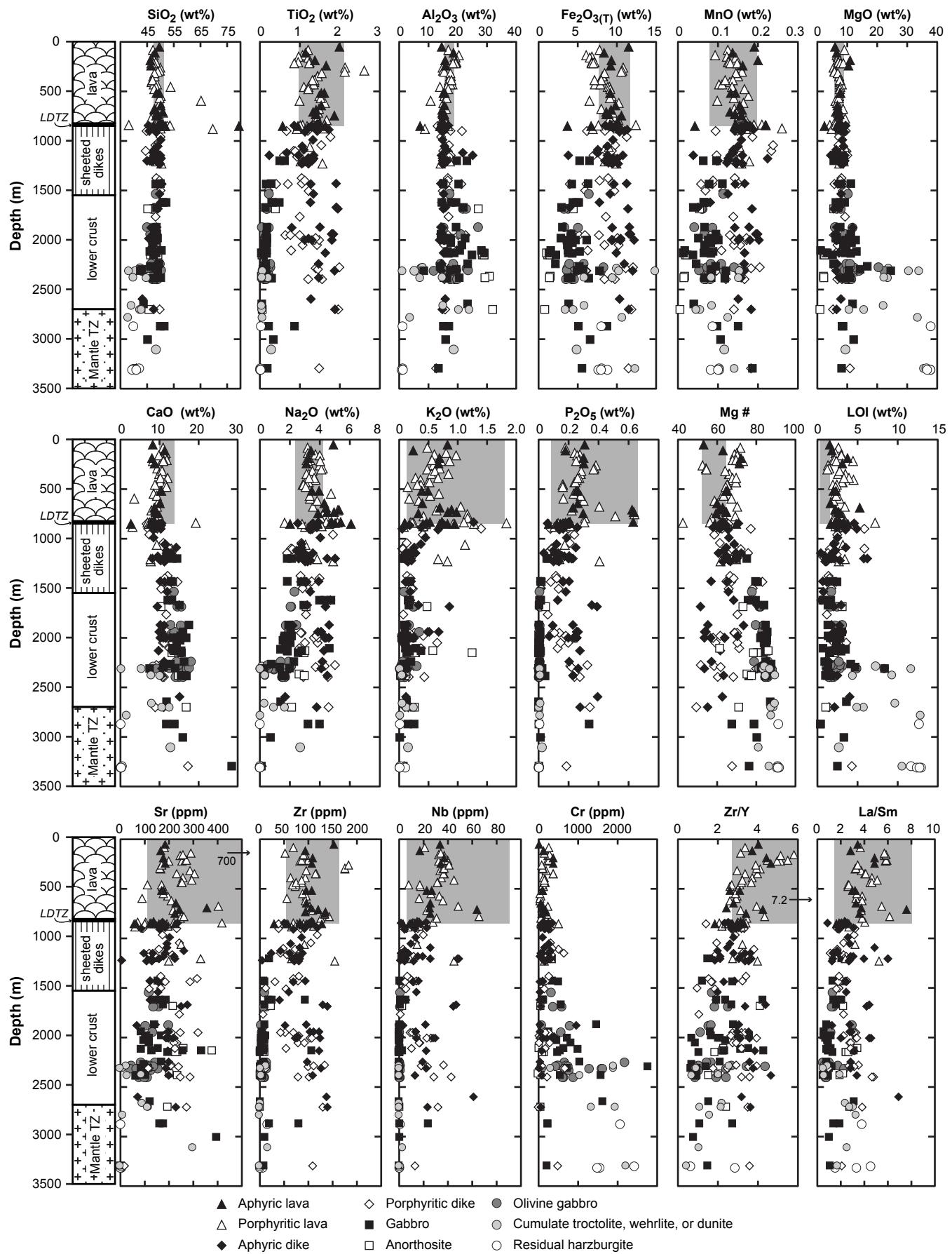
**Figure 4.**



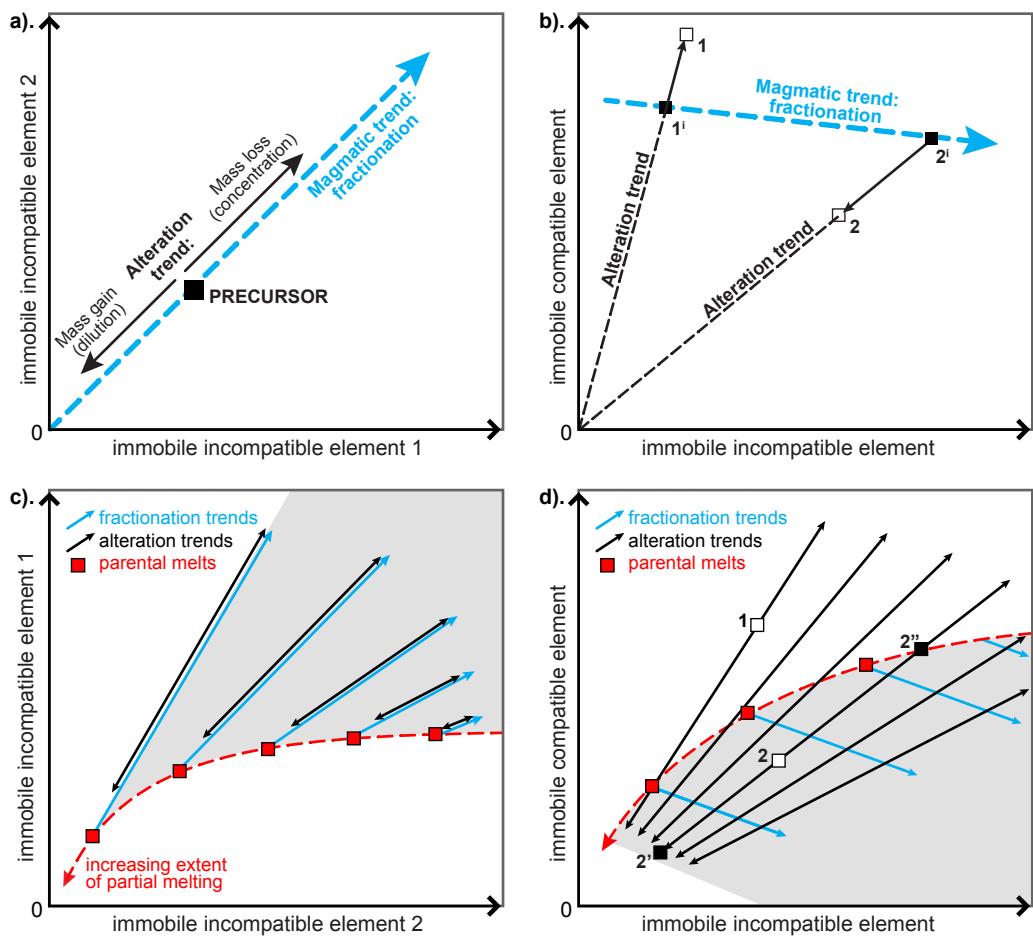
**Figure 5.**



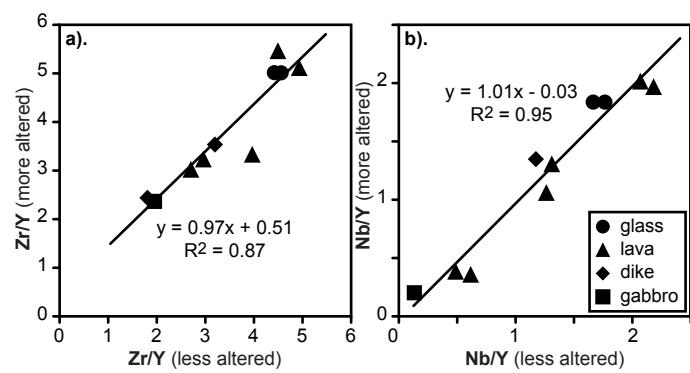
**Figure 6.**



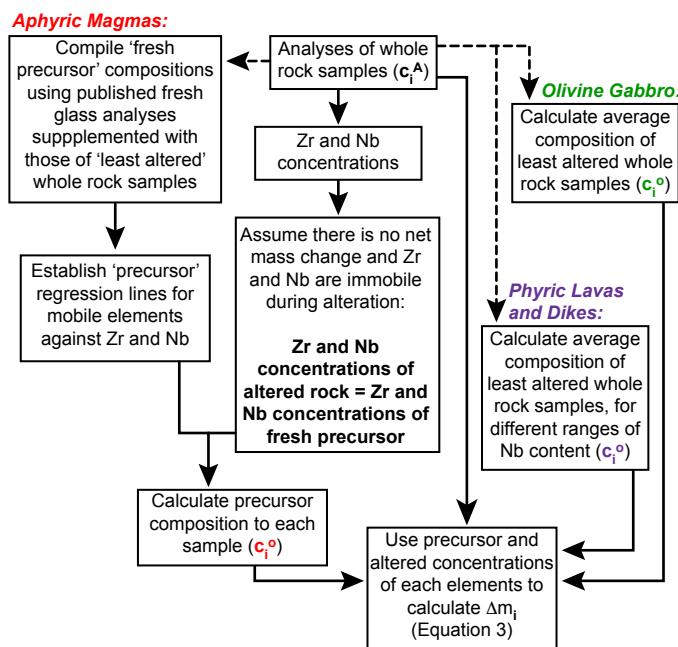
**Figure 7.**



**Figure 8.**



**Figure 9.**



**Figure 10.**

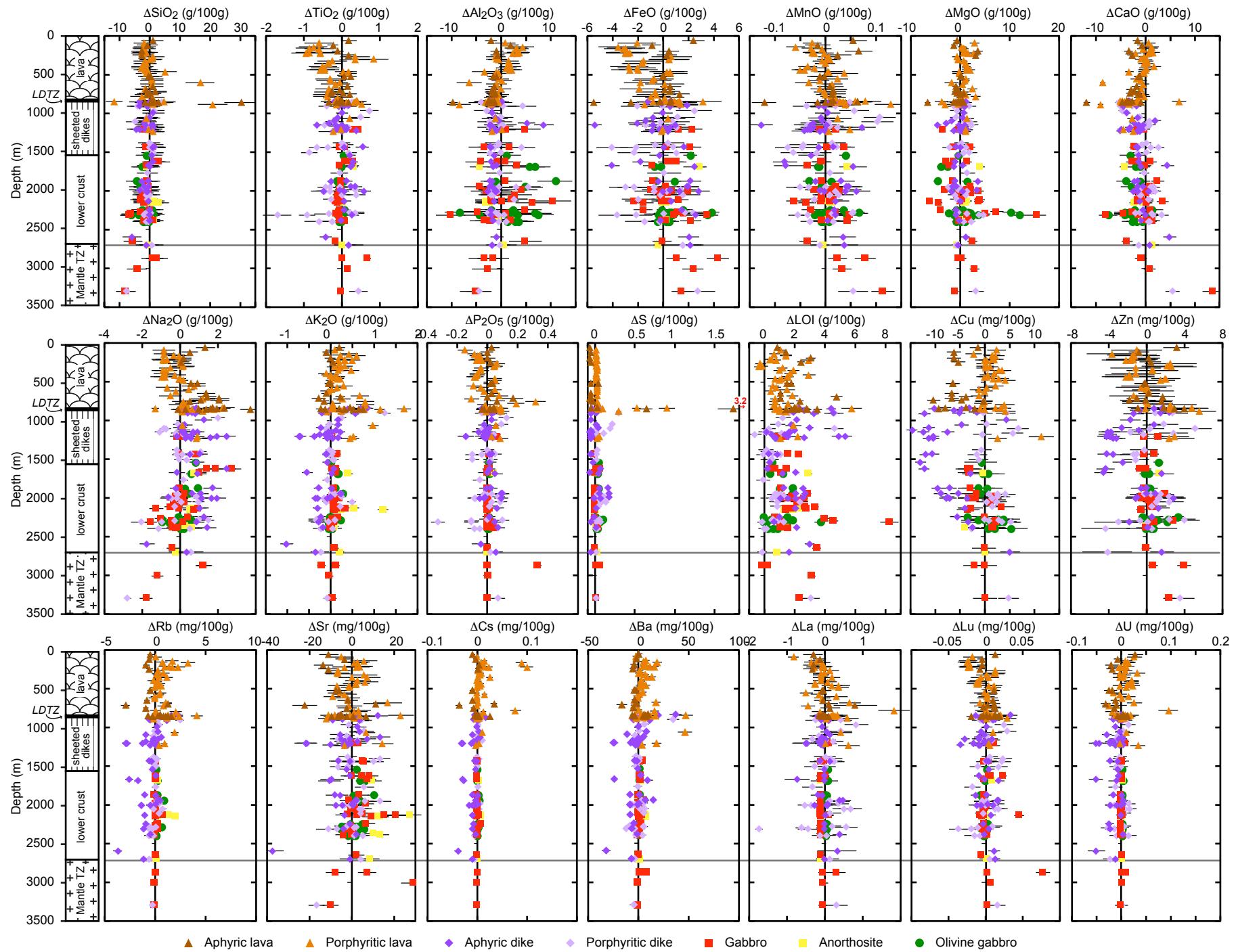
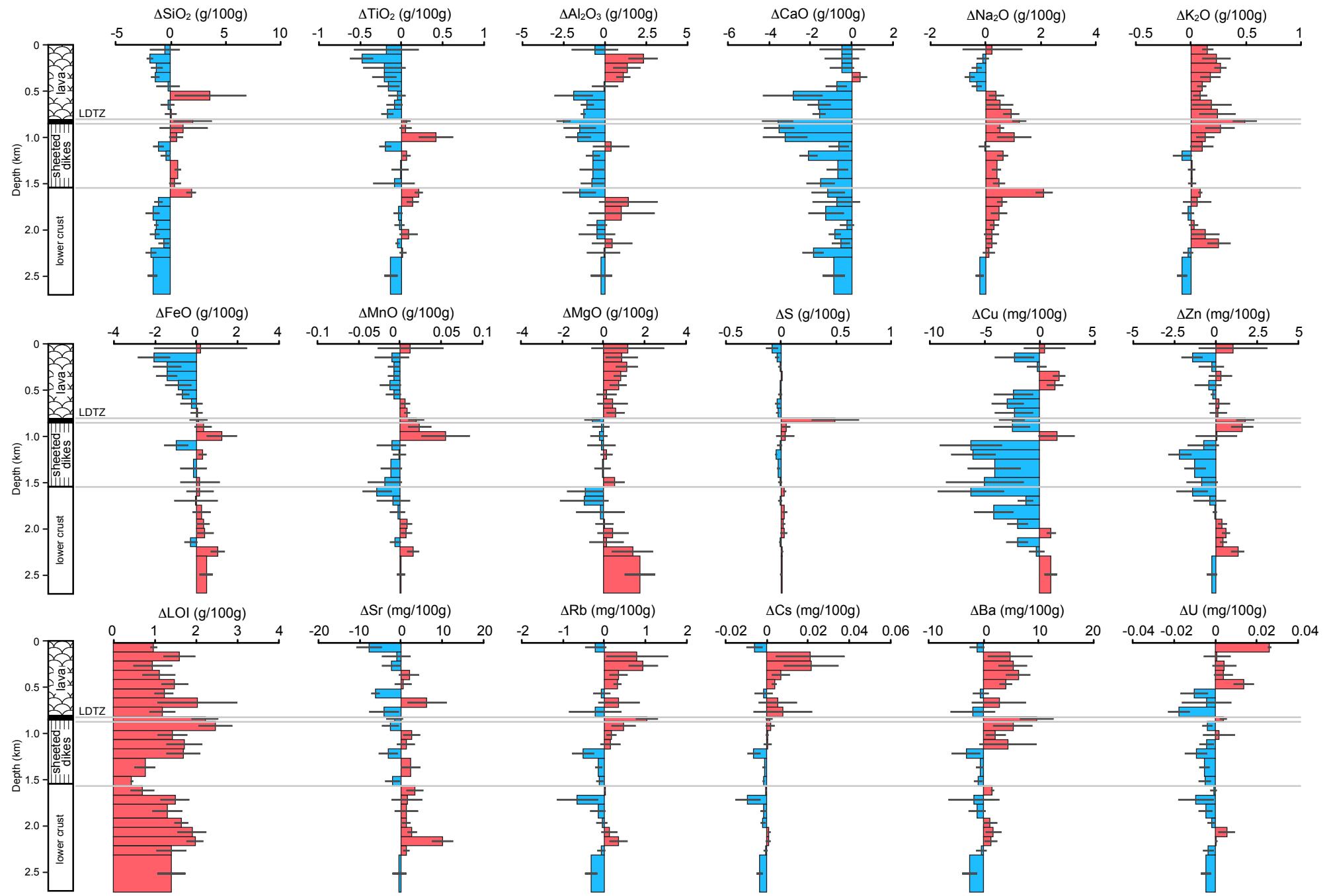
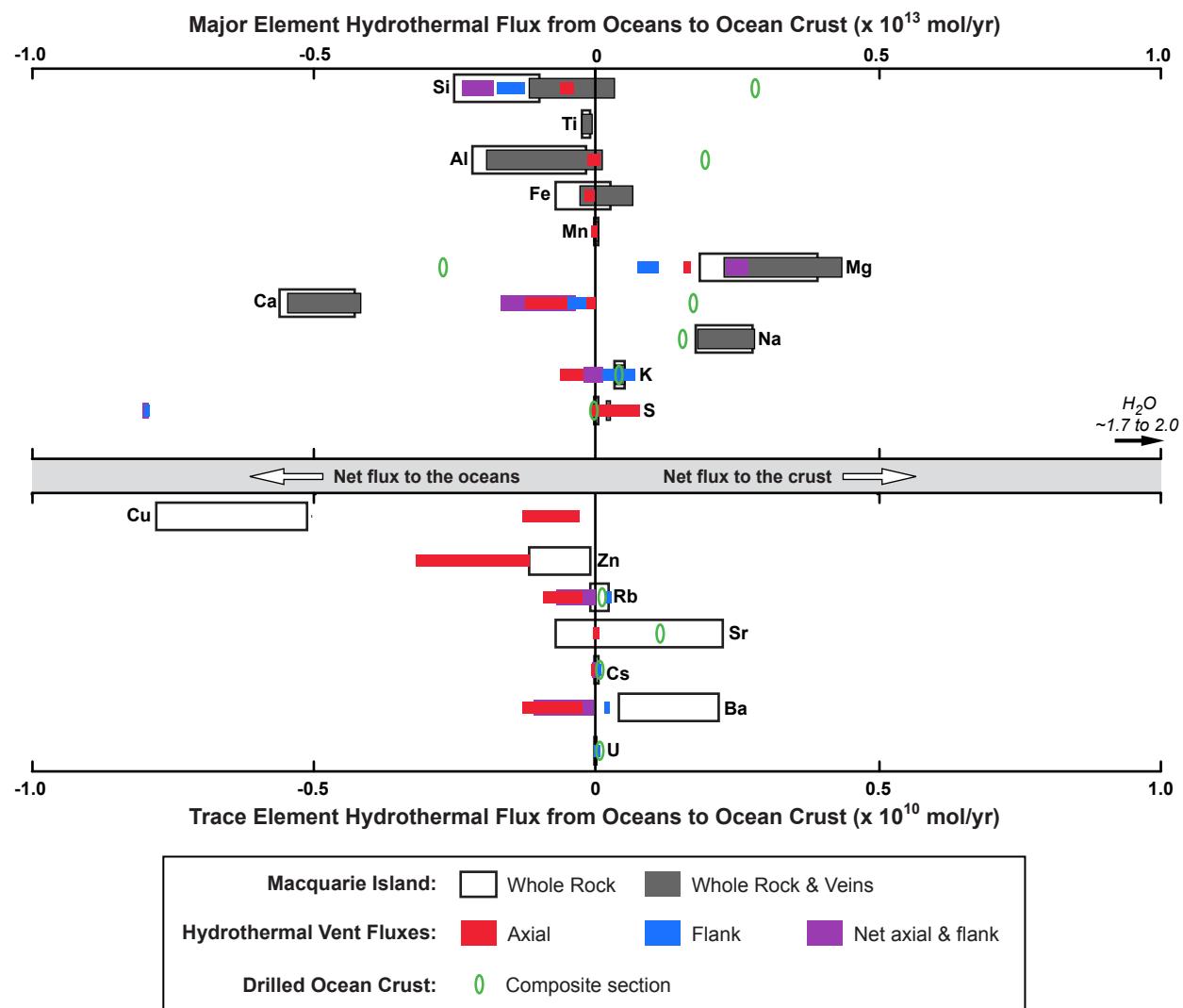


Figure 11.



**Figure 12.**



**Table 1. Summary of chemical changes for each Macquarie lithology**

	Lavas		Dikes		Gabbro	Anorthosite	Olivine Gabbro
	Aphyric	Porphyritic	Aphyric	Porphyritic			
SiO <sub>2</sub>	(+)		(-)	(-)	(-)	(+)	(-)
TiO <sub>2</sub>		-	(+)	(-)	+	+++	(-)
Al <sub>2</sub> O <sub>3</sub>	-		(-)	(-)		(-)	+
FeO <sub>T</sub>		(-)	(+)		+	+	+
MnO	(+)	(+)			(+)	++	
MgO	(+)			(+)	(+)	++	
CaO	-	-	-			(-)	-
Na <sub>2</sub> O	+			+		+	
K <sub>2</sub> O	+	+++	-	+	+++	+++	+
P <sub>2</sub> O <sub>5</sub>	+				+++	+++	+++
S	+++	+++	-	+++	-		++
LOI	+++	++	+++	++	+++	+++	++
Cu	--	++	--	++	-	-	+
Zn	+	+	-	+	+	++	+
Rb	-	+++	--		++	+++	+++
Sr	-		-		+	+++	+
Cs	-	+++	--	-	+	+++	-
Ba	(-)	++	-		+	+++	+++
La			(-)		-	--	+
Lu	+				+	+++	-
U	-	+	-	-	+	++	++

**Table 2. Summary of average vein mineral abundances and compositions in drill core**

Vein Mineral	Specific Gravity	Vein Mineral Abundance (vol%)				Vein Mineral Composition (wt%)											Total	Reference
		Lavas <sup>a</sup>	LDTZ <sup>b</sup>	Sheeted Dikes <sup>b</sup>	Gabbro <sup>c</sup>	SiO <sub>2</sub>	TiO <sub>2</sub>	Al <sub>2</sub> O <sub>3</sub>	FeO	Fe <sub>2</sub> O <sub>3</sub>	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O			
<b>Mg-saponite</b>	2.27	0.85	0.6	0.01	0.14	47.71	0.04	5.64	15.43		0.05	16.07	0.82	0.11	0.44	86.31	e	
<b>Celadonite</b>	3.00	0.12				51.24	0.05	3.57	23.56		0.02	4.93	0.50	0.01	7.19	91.06	e	
<b>Prehnite</b>	2.88	0.01 <sup>d</sup>	0.01	0.01	0.01 <sup>d</sup>	42.67	0.01	22.83		2.07	0.07	0.78	24.87	0.03	0.03	93.35	f	
<b>Epidote</b>	3.45			0.01	0.008	37.95	0.08	23.47		11.74	0.11	0.27	23.07	0.02	0.01	96.73	f	
<b>Chlorite</b>	2.95		0.3	0.44	0.007	28.52	0.02	17.59	21.89		0.26	14.45	0.20	0.01	0.01	82.94	e	
<b>Amphibole</b>	3.00			0.01	0.12	51.19	0.46	4.07	14.38		0.22	14.51	11.01	0.67	0.02	96.52	g	
<b>Feldspar</b>	2.63				0.11	63.44	0.01	22.85		0.15			4.15	8.95	0.10	99.64	h	
<b>Na-zeolite</b>	2.20	0.01	0.05			59.70		18.58	0.65			0.69	1.66	4.69	5.81	91.77	i	
<b>Serpentine</b>					0.005 <sup>d</sup>	40.32	0.01	0.56	5.18			37.33	0.26			96.01	j	
<b>Mineral Formulae</b>																		
<b>Calcite</b>	2.85	0.01	0.01	0.01	0.02	CaCO <sub>3</sub>												
<b>Fe(O,OH)x</b>	2.71	0.01			0.01 <sup>d</sup>	Fe(O,OH)												
<b>Quartz/SiO<sub>2</sub></b>	2.65	0.15	0.03	0.13	0.02	SiO <sub>2</sub>												
<b>Ca-zeolite</b>	2.29		0.05	0.01	0.009	CaAl <sub>2</sub> Si <sub>4</sub> O <sub>10</sub> (OH) <sub>2</sub>												
<b>Pyrite</b>	5.05	0.02	0.01	0.1		Fe <sub>2</sub> S												
<b>Talc</b>	2.75			0.1	0.01 <sup>d</sup>	Mg <sub>3</sub> Si <sub>4</sub> O <sub>10</sub> (OH) <sub>2</sub>												
<b>TOTAL:</b>	<b>1.17</b>	<b>1.06</b>	<b>0.83</b>	<b>0.47</b>														

**Table 3. Hydrothermal fluxes**

Unit factor (mol/yr)	LAVAS		LAVA-DIKE TZ		SHEETED DIKES		LOWER CRUST		FULL CRUST		% of Macquarie crustal budget	
	WR	WR + Veins	WR	WR + Veins	WR	WR + Veins	WR	WR + Veins	WR	WR + Veins		
Si	$\times 10^{11}$ 3.2	-8.1 to 9.5	-1.8 to 1.4	0.1 to 1.5	0.2 to 6.6	-0.8 to 9.9	2.5 to -16	-22 to -12	-18 to -10	-25 to -10	-12 to 3.3	-0.6 to 0.2
Ti	$\times 10^{10}$ -11	-23 to -11	-23 to 0.3	0.04 to 0.3	0.04 to 6.9	-2.0 to 6.9	-2.1 to 1.7	-5.9 to 1.8	-5.8 to -8.4	-25 to -25 to -8.4	-25 to -8.1	-3 to -8
Al	$\times 10^{11}$ 3.4	-4.9 to 4.1	-4.3 to -0.9	-1.3 to -0.8	-1.2 to -5.7	-14 to -4.6	-12.9 to 8.2	-8.4 to 9.0	-7.6 to -1.6	-22 to -19 to 1.0	-19 to -2 to 0.1	-2 to 0.1
Fe	$\times 10^{11}$ -4.5	-11 to -2.9	-9.5 to 0.2	-0.1 to 0.2	-0.1 to 3.2	-1.4 to 5.0	0.3 to 7.2	2.3 to 7.8	2.9 to 7.8	-7.0 to 2.5	-2.8 to 6.6	-1 to 3
Mn	$\times 10^9$ 4.3	-7.9 to 4.7	-7.5 to 0.9	0.3 to 0.9	0.3 to 9.7	-2.7 to 10	-1.7 to 3.8	-5.8 to 4.3	-5.2 to 11	-8.5 to 11	-6.6 to 13	-2 to 3
Mg	$\times 10^{11}$ 18	8.2 to 20	10 to -0.1	-0.5 to 0	-0.4 to 3.3	-2.5 to 4.8	-1.1 to 24	6.7 to 25	7.7 to 25	18 to 39	23 to 44	5 to 10
Ca	$\times 10^{12}$ -0.9	-1.6 to -0.9	-1.6 to -0.1	-0.2 to -0.1	-0.2 to -1.6	-2.2 to -1.6	-2.2 to -1.2	-2.1 to -1.2	-2.0 to -1.1	-5.6 to -4.3	-5.5 to -4.2	-8 to -10
Na	$\times 10^{11}$ 6.4	-1.2 to 6.3	-1.2 to 1.0	0.7 to 1.0	0.7 to 12	7.4 to 12	7.3 to 12	7.1 to 12	7.3 to 13	18 to 28	18 to 28	8 to 13
K	$\times 10^{10}$ 34	21 to 35	23 to 2.8	1.8 to 2.8	1.8 to 12	4.1 to 12	4.1 to 12	-2.4 to 8.6	-2.3 to 8.7	32 to 50	36 to 52	27 to 42
S	$\times 10^{10}$ -7.4	-3.5 to 0.7	-3.2 to 4.8	1.9 to 4.8	1.9 to 2.0	-3.3 to 20	14 to 20	1.8 to 5.1	1.7 to 5.1	-3.3 to 4.7	19 to 27	76 to 108
H <sub>2</sub> O	$\times 10^{12}$ 5.9	4.6 to 6.2	4.9 to 0.31	0.23 to 0.32	0.24 to 4.9	4.0 to 5.2	4.3 to 8.8	7.3 to 8.8	7.4 to 8.9	17 to 19	18 to 20	195 to 220
Cu	$\times 10^8$ -4.0	-16 to -0.4	-1.3 to -0.4	-46 to -26		-25 to -11			-78 to -51			-25 to -38
Zn	$\times 10^8$ 2.8	-4.3 to 0.8	0.4 to 0.8	-9.4 to -2.9		-2.8 to 2.5			-12 to -1.0			-1 to -9
Rb	$\times 10^8$ 3.6	0.9 to 0.3	0.2 to 0.3	-0.6 to 0.4		-2.8 to -1.0			-1.2 to 2.2			-8 to 15
Sr	$\times 10^8$ -4.5	-22 to 0.1	-0.9 to 0.1	-5.5 to 7.5		11 to 30			-7.1 to 22			-2 to 5
Cs	$\times 10^6$ 5.9	2.0 to 0.05	-0.01 to 0.05	-0.8 to -0.2		-2.4 to -1.2			-0.5 to 3.7			-5 to 42
Ba	$\times 10^8$ 18	6.4 to 2.0	1.0 to 2.0	-0.5 to 8.7		-10 to -0.2			3.8 to 22			3 to 17
U	$\times 10^6$ 1.1	-0.3 to 0.05	0.02 to 0.05	-1.6 to -0.7		-1.8 to -0.7			-3.0 to -1.0			-12 to -4

**Table 4. Comparison of Estimated Hydrothermal Chemical Changes and Fluxes**

Average Bulk Crustal Chemical Changes ( $\Delta m_i$ ): gains (+) and losses (-)		Extrapolated Global Hydrothermal Element Fluxes: (+ve, net flux to crust; -ve net flux to the oceans)				
ODP <sup>a</sup>	MQ <sup>b</sup>	ODP <sup>c</sup>	MQ <sup>d</sup>	Axial Fluids <sup>e</sup>	Flank Fluids <sup>e</sup>	
<b>MAJOR ELEMENTS</b>						
	g/100g	g/100g		$10^{12}$ mol/yr	$10^{12}$ mol/yr	$10^{12}$ mol/yr
<b>SiO<sub>2</sub></b>	0.237	-0.30 to 0.08	<b>Si</b>	2.78	-1.17 to 0.33	-0.66 to -0.43
<b>TiO<sub>2</sub></b>	[0] <sup>f</sup>	-0.08 to -0.03	<b>Ti</b>	[0] <sup>f</sup>	-0.25 to -0.08	-
<b>Al<sub>2</sub>O<sub>3</sub></b>	-	-0.42 to 0.02	<b>Al</b>	1.89	-1.94 to 0.10	-0.0006 to -0.0001
<b>FeO<sub>T</sub></b>	-	-0.09 to 0.20	<b>Fe</b>	-	-0.28 to 0.66	-0.19 to -0.02
<b>MnO</b>	-	-0.002 to 0.004	<b>Mn</b>	-	-0.01 to 0.01	-0.034 to -0.011
<b>MgO</b>	-0.157	0.39 to 0.75	<b>Mg</b>	-2.75	2.26 to 4.35	1.6
<b>CaO</b>	0.134	-1.3 to -1.0	<b>Ca</b>	1.68	-5.48 to -4.16	-1.3 to -0.009
<b>Na<sub>2</sub>O</b>	0.0655	0.24 to 0.37	<b>Na</b>	1.49	1.79 to 2.81	-
<b>K<sub>2</sub>O</b>	0.0485	0.07 to 0.1	<b>K</b>	0.36	0.33 to 0.52	-0.69 to -0.23
<b>H<sub>2</sub>O</b>	0.449	1.35 to 1.52	<b>H<sub>2</sub>O</b>	17.6	17.6 to 19.8	-
<b>S</b>	-0.004	0.03 to 0.04	<b>S</b>	-0.08	0.19 to 0.27	-0.12 to 0.76
<b>TRACE ELEMENTS</b>						
	mg/kg	mg/kg		$10^8$ mol/yr	$10^8$ mol/yr	$10^8$ mol/yr
<b>Cu</b>	-	-21 to -14		-	-78 to -51	-13 to -3
<b>Zn</b>	-	-3.3 to -0.3		-	-12 to -1.0	-32 to -12
<b>Rb</b>	0.927	-0.4 to 0.8		0.77	-1.2 to 2.2	-9.5 to -2.6
<b>Cs</b>	0.0172	-0.003 to 0.02		0.01	-0.005 to 0.04	-0.06 to -0.03
<b>Sr</b>	1.4	-2.6 to 8.3		11	-7.1 to 22	0
<b>Ba</b>	-	2.2 to 13		-	3.8 to 22	-13 to -2.4
<b>U</b>	0.0375	-0.03 to -0.01		0.11	-0.03 to -0.01	-

## Supplementary Material for ‘Hydrothermal contributions to global biogeochemical cycles; insights from the Macquarie Island ophiolite’.

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### 1. Uncertainties

#### 1.1 Uncertainties in Calculated Chemical Changes ( $\Delta m_i$ )

The methods for calculating precursor compositions are subject to the following errors:

(i) analytical errors, (ii) the errors that result from any unaccounted for net changes in mass; and (iii) the errors in calculated precursor-mobile element concentrations due to the broad range of fresh sample compositions at a given immobile monitor element concentration, which we term the primary magmatic error. The errors in estimating precursor compositions and elemental mass changes  $\Delta m_i$  are propagated on a sample-by-sample basis, as follows:

$$\delta(\Delta m_i) = \sqrt{\left(M^o \frac{c_x^o}{c_x^A} c_i^A\right)^2 \left[\left(\frac{\delta(c_x^o/c_x^A)}{(c_x^o/c_x^A)}\right)^2 + \left(\frac{\delta c_i^A}{c_i^A}\right)^2\right] + (M^o \delta c_i^o)^2}$$

where:  $\delta(\Delta m_i X)$  is the error in the calculated elemental mass change;  $M^o$  is the total mass of the precursor,  $c_i^A$  and  $c_i^o$  are the concentrations of component  $i$  in the altered rock and the precursor, respectively;  $c_x^A$  and  $c_x^o$  are the concentrations of an immobile element in the altered rock and the precursor, respectively;  $\delta(c_x^o/c_x^A)$  is the error in the ratio of the concentration of the immobile monitor element in the precursor and altered rock due to any unaccounted for net changes in mass;  $\delta c_i^A$  is the analytical

uncertainty in the measured concentration of component i, and  $\delta c_i^0$  is the primary magmatic error.

The accuracy and precision of major and trace element analyses are provided in Tables S3, S4 and S5. For the purposes of error propagation we assume that the analytical uncertainty in  $c_i^A$  is 5% for major elements and 10% for sulfur and trace elements.

Given the evidence that net mass changes were typically <5% we assume that  $(c_x^0 / c_x^A) = 1 \pm 0.05$ .

The primary magmatic error in the precursor concentration of an element determined from its magmatic trend with an immobile monitor element (Zr on Nb) is given by:

$$\delta c_i^0 = \sqrt{(STE c_i^0)^2 + (A \delta c_x^0)^2}$$

where  $STE c_i^0$  is the standard error of the estimate from the regression (i.e one standard deviation of the residuals of fresh rock concentrations from the regression line, presented in Table S5), A is the slope of the regression line and  $\delta c_x^0$  is the analytical uncertainty in the immobile monitor element concentration (10%). Precursor-concentrations of elements that show no significant correlation with immobile monitor element concentrations are calculated as the mean concentration of the fresh and/or least altered samples, and the primary magmatic error is taken as one standard deviation of the mean (Table S6). Where altered rocks are compared to a single least-altered sample a primary magmatic error of  $\pm 10\%$  is assumed.

## 1.2 Uncertainties in Average Hydrothermal Fluxes.

The calculated changes in composition of the Macquarie crust due to hydrothermal alteration are converted into net fluxes to or from the crust, using the equation:

$$F_i = \sum_t^T \Delta \bar{m}_{i-t} z_t \rho_c$$

Where:  $F_i$  is the mass flux of component  $i$  through  $1 \text{ m}^2$  of seafloor due to alteration of a section of crust of thickness  $T$ ;  $\Delta \bar{m}_{i-t}$  is the mean change in mass of component  $i$  per unit mass of rock in each sub-interval  $t$ ;  $z_t$  is the thickness of each sub-interval  $t$ ; and  $\rho_c$  is the density of the crust ( $2900 \text{ kg/m}^3$ ). The uncertainty in the net flux ( $\delta F_i$ ) is propagated from the standard error of the mean chemical change of each sub-interval ( $\delta \Delta \bar{m}_{i-t}$ ) as follows:

$$\delta(F_i) = \sqrt{\sum_t^T [(\delta \Delta \bar{m}_{i-t}) z_t]^2}$$

## References:

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**Table S1. Summary of the primary mineralogy and textures and the metamorphism of the igneous lithologies on Macquarie Island.**

LITHOLOGY	TEXTURES	PRIMARY MINERALOGY	SECONDARY MINERALOGY		ALTERATION INTENSITY
			PRIMARY MINERAL REPLACEMENT	VEINS	
<b>BASALT LAVAS</b>  Pillow lavas, massive basalt, hyaloclastite breccia and feeder dikes	Glassy to fine-grained; sub-ophitic to interstitial; aphyric to highly phryic	Phenocrysts: <ul style="list-style-type: none"><li>■ Plagioclase: &lt;25-30 vol%; on average 1-3 mm (up to 30 mm)</li><li>■ Olivine: &lt; 3 vol%, 1-3 mm</li><li>■ Clinopyroxene: &lt;2 vol%, 1-3 mm</li><li>■ Spinel (reddish-brown): &lt; 1 vol%</li></ul>	OFW: <ul style="list-style-type: none"><li>■ Olivine: calcite + smectite</li><li>■ Interstitial glass: smectite + phillipsite + calcite</li></ul> ZF: <ul style="list-style-type: none"><li>■ Olivine: smectite + calcite ± chlorite</li><li>■ Plagioclase: zeolite ± sericite ± k-feldspar</li><li>■ Interstitial glass: smectite + calcite ± chlorite</li></ul> LGF: <ul style="list-style-type: none"><li>■ Olivine: chlorite</li><li>■ Plagioclase: albite ± sericite ± chlorite ± prehnite ± epidote ± titanite</li><li>■ Spinel: titanite</li></ul>	OFW: calcite, smectite, celadonite, phillipsite, saponite, hematite, iron-oxyhydroxides Amygdales filled with: smectite + calcite  ZF: Na, K, Ca zeolites ± smectite ± calcite ± chlorite ± albite Amygdales filled with: smectite + calcite  LGF: Chlorite ± albite ± k-feldspar ± prehnite ± epidote Amygdales filled with: albite ± k-feldspar ± prehnite	Alteration pervasive, extent highly variable (primarily controlled by abundance of olivine and glass)
<b>DOLERITE DIKES</b>	Crypto-crystalline to fine-grained; sub-ophitic to intergranular, aphyric to highly phryic	Groundmass: plagioclase + clinopyroxene ± olivine + oxides  Phenocrysts: <ul style="list-style-type: none"><li>■ Plagioclase: typically &lt;35 vol% (up to 85 vol%); 0.5-15 mm</li><li>■ Olivine: ~2.5 vol% (up to 10 vol%); 0.2-4 mm (up to 10 mm)</li><li>■ Clinopyroxene: ~ 2 vol% (up to 9 vol%); 0.5-5 mm</li><li>■ Spinel (reddish-brown): &lt; 1 vol%; 0.1-0.6 mm</li></ul>	LGF: <ul style="list-style-type: none"><li>■ Plagioclase: plagioclase ± sericite ± prehnite ± epidote ± titanite ± chlorite/smectite</li><li>■ Olivine: chlorite ± smectite ± calcite ± sulphides</li><li>■ Clinopyroxene and spinel: relatively fresh</li></ul> UGF/LAF: <ul style="list-style-type: none"><li>■ Plagioclase: as in LGF dikes</li><li>■ Olivine: smectite/chlorite ± sulphides; tremolite + chlorite ± talc ± magnetite</li><li>■ Clinopyroxene: (uralitised); actinolite ± tremolite ± hornblende</li></ul>	LGF: Epidote, CaCO <sub>3</sub> , chlorite, smectite, secondary plagioclase, pyrite, quartz, zeolite, prehnite  UGF/LAF: amphibole, epidote, CaCO <sub>3</sub> , chlorite, smectite, secondary plagioclase, pyrite, quartz, zeolite, prehnite	Moderately to highly altered: greatest in the upper greenschist/lower amphibolite facies dikes

LITHOLOGY	TEXTURES	PRIMARY MINERALOGY	SECONDARY MINERALOGY		ALTERATION INTENSITY
			PRIMARY MINERAL REPLACEMENT	VEINS	
<b>GABBROS</b>  Gabbro, anorthosite and olivine gabbro	Fine-grained porphyritic; medium-grained to pegmatitic hypidiomorphic granular to poikilitic	<ul style="list-style-type: none"> <li>■ Plagioclase: anhedral to subhedral, tabular to equant; euhedral to sub-rounded chadacrysts</li> <li>■ Clinopyroxene: medium-grained anhedral; coarse-grained to pegmatitic anhedral to subhedral poikilitic</li> <li>■ Olivine: anhedral; sub-rounded chadacrysts</li> </ul>	<ul style="list-style-type: none"> <li>■ Plagioclase: secondary plagioclase <math>\pm</math> prehnite <math>\pm</math> epidote <math>\pm</math> <math>\text{CaCO}_3</math> <math>\pm</math> clay</li> <li>■ Clinopyroxene: amphibole <math>\pm</math> chlorite <math>\pm</math> clay <math>\pm</math> <math>\text{CaCO}_3</math> <math>\pm</math> magnetite</li> <li>■ Olivine: amphibole <math>\pm</math> chlorite <math>\pm</math> talc <math>\pm</math> sulphides <math>\pm</math> serpentine <math>\pm</math> magnetite <math>\pm</math> clay</li> </ul>	Gabbro and olivine gabbro: amphibole, chlorite, epidote, clinozoisite, zoisite, secondary plagioclase, prehnite, quartz  Anorthosite: secondary plagioclase, chlorite, amphibole, epidote, clinozoisite, prehnite, $\text{CaCO}_3$ , quartz	Highly variable; slightly to highly altered under greenschist to amphibolite facies conditions
<b>Troctolite, olivine gabbro, wehrlite, plagioclase dunite and dunite</b>	Medium-grained to pegmatitic; ad- to ortho-cumulates	<ul style="list-style-type: none"> <li>■ Olivine: anhedral to euhedral; cumulus phase</li> <li>■ Clinopyroxene: subhedral to anhedral; post cumulus phase</li> <li>■ Plagioclase: subhedral to anhedral; cumulus and/or post cumulus phase</li> <li>■ Spinel: &lt; 1 vol%, dark brown</li> </ul>	<ul style="list-style-type: none"> <li>■ Olivine: serpentine + magnetite; or amphibole <math>\pm</math> chlorite <math>\pm</math> talc <math>\pm</math> sulphides <math>\pm</math> serpentine <math>\pm</math> magnetite <math>\pm</math> clay</li> <li>■ Clinopyroxene: amphibole <math>\pm</math> chlorite <math>\pm</math> clay <math>\pm</math> magnetite</li> <li>■ Plagioclase: prehnite + chlorite + serpentine + secondary plagioclase</li> </ul>	Serpentine $\pm$ magnetite, chlorite, talc, prehnite, $\text{CaCO}_3$	Highly to completely altered; serpentinised
<b>HARZBURGITE</b>	Medium- to coarse-grained Cumulates?	<ul style="list-style-type: none"> <li>■ Olivine: anhedral; cumulus phase?</li> <li>■ Orthopyroxene: irregular, &lt;40 vol%</li> <li>■ Spinel: ~2 vol%, dark brown, irregular</li> </ul>	<ul style="list-style-type: none"> <li>■ Olivine: serpentine + magnetite; talc, clay, amphibole</li> <li>■ Orthopyroxene: serpentine + talc (bastite); amphibole</li> </ul>	Serpentine $\pm$ magnetite, chlorite, talc, prehnite, $\text{CaCO}_3$	Highly to completely altered; serpentinised

OFW = ocean floor weathering; ZF = zeolite facies; LGF = lower greenschist facies; UGF = upper greenschist facies; LAF = lower amphibolite facies.

**Table S2. Macquarie Island whole rock geochemistry**

- see excel sheet.

**Table S3. Accuracy and precision for XRF major and trace elements analyses at the University of Leicester.**

Standard	NIM-G [n=4]				W-1 [n=3]				BCS372/1 [n=4]			
	Mean Value	Preferred Value	Precision (% RSD)	Accuracy (% RMSD)	Mean Value	Preferred Value	Precision (% RSD)	Accuracy (% RMSD)	Mean Value	Preferred Value	Precision (% RSD)	Accuracy (% RMSD)
<b>SiO<sub>2</sub></b>	75.87	75.70	0.41	0.32	52.36	52.46	0.06	0.10	20.23	20.50	0.34	0.28
<b>TiO<sub>2</sub></b>	0.10	0.09	12	0.01	1.01	1.07	6.5	0.08	0.16	0.16	8.4	0.01
<b>Al<sub>2</sub>O<sub>3</sub></b>	11.95	12.08	0.44	0.14	15.04	15.00	0.60	0.08	5.03	3.77	2.3	1.3
<b>Fe<sub>2</sub>O<sub>3</sub></b>	2.06	2.02	5.0	0.10	11.13	11.11	0.77	0.07	3.35	4.82	1.5	1.5
<b>MnO</b>	0.02	0.02	11	0.00	0.17	0.17	1.6	0.00	0.06	0.06	9.0	0.01
<b>MgO</b>	0.02	0.06	299	0.07	6.60	6.62	3.0	0.16	1.48	2.42	9.8	0.95
<b>CaO</b>	0.75	0.78	5.7	0.05	10.93	11.00	1.3	0.14	64.75	64.80	0.45	0.26
<b>Na<sub>2</sub>O</b>	3.49	3.36	2.9	0.16	2.23	2.16	3.8	0.10	0.27	0.10	10	0.17
<b>K<sub>2</sub>O</b>	5.01	4.99	0.91	0.04	0.63	0.64	4.9	0.03	0.73	0.49	0.94	0.24
<b>P<sub>2</sub>O<sub>5</sub></b>	0.01	0.01	13	0.00	0.13	0.13	1.3	0.00	0.08	0.08	3.1	0.004
<b>Total</b>	99.63	99.11	0.28	0.57	100.18	100.36	0.33	0.33	96.72	97.14	0.45	0.57

Standard	BIR-1 [n=7]				BCR-1 [n=7]				JR-1 [n=7]			
	Mean Value	Preferred Value	Precision (% RSD)	Accuracy (% RMSD)	Mean Value	Preferred Value	Precision (% RSD)	Accuracy (% RMSD)	Mean Value	Preferred Value	Precision (% RSD)	Accuracy (% RMSD)
<b>Cr</b>	394	382	2.7	16	19	16	23	5.4	7.8	2	57	7.1
<b>Cu</b>	126	126	1.9	2.3	28	19	12	9.9	0.40	1	75	0.7
<b>Ni</b>	166	166	2.2	3.4	13	13	13	1.5	2.6	1	56	2.1
<b>Sc</b>	41	44	4.8	3.7	33	33	3.9	1.2	4.0	5	19	1.2
<b>V</b>	301	313	8.2	26	395	407	1.4	13	8.5	-	40	-
<b>Zn</b>	70	71	1.1	0.90	128	130	0.90	2.0	30.1	30	3.0	0.8
<b>Zr</b>	17	22	29	6.8	193	190	1.1	3.2	99.0	102	5.7	6.0

Accuracy is reported as the relative root-mean square deviation (%RMSD) compared to the preferred reference value.

**Table S4.** Accuracy, precision and limits of detection for trace elements analyzed by ICP-MS at the University of Southampton.

	JB-1a [n=8]				JA-1 [n=6]				BHVO-1 [n=3]				Average precision (%RSD)	Average accuracy (%RMSD)
	Mean Value	Preferred Value <sup>1</sup>	Precision (%RSD)	Accuracy (%RMSD)	Mean Value	Preferred Value <sup>1</sup>	Precision (%RSD)	Accuracy (%RMSD)	Mean Value	Preferred Value <sup>1</sup>	Precision (%RSD)	Accuracy (%RMSD)		
Rb	40.9	39	0.8	5.1	11.5	10.65	8.9	11.8	9.76	9.19	9.3	10.2	6.3	9.0
Sr	444	444	1.9	1.8	254	264	3.3	4.7	391	396	1.3	1.7	2.2	2.7
Y	23.9	24	3.8	3.6	30.0	29	3.1	4.5	27.3	26	1.3	5.1	2.7	4.4
Zr	145	142	1.3	2.5	89.0	84	3.1	6.6	179	174	1.3	2.8	1.9	4.0
Nb	27.2	28	1.7	3.3	1.47	1.4	18.4	18.4	18.1	18.6	4.8	4.7	8.3	8.8
Cs	1.31	1.2	4.4	10.0	0.64	0.64	5.1	4.7	0.11	0.101	4.1	12.4	4.5	9.0
Ba	509	489	2.0	4.6	297	303	3.4	3.6	128	133	1.8	3.7	2.4	4.0
La	36.8	38	1.3	3.4	4.6	5	2.7	8.6	14.7	15.5	2.5	5.2	2.2	5.7
Ce	68.4	66	6.7	7.5	13.0	13.5	2.5	4.4	38.0	38.1	1.0	0.9	3.4	4.3
Pr	6.76	7.2	3.3	6.8	1.98	2.08	1.8	5.0	5.08	5.42	1.5	6.3	2.2	6.0
Nd	25.8	26.3	3.5	3.8	10.5	10.9	1.3	4.2	24.6	24.7	2.9	2.3	2.6	3.4
Sm	4.96	5.10	4.4	4.8	3.27	3.4	2.4	4.5	6.02	6.12	1.6	2.0	2.8	3.8
Eu	1.45	1.46	3.6	3.5	1.08	1.12	1.1	3.5	2.06	2.09	1.7	2.1	2.1	3.0
Gd	4.71	4.70	12.8	12.0	4.01	4.20	2.1	5.0	6.00	6.33	2.5	5.6	5.8	7.5
Tb	0.69	0.72	7.8	8.3	0.71	0.73	1.6	3.4	0.92	0.96	2.9	4.3	4.1	5.3
Dy	4.02	4.10	4.5	4.6	4.66	4.80	1.5	3.3	5.22	5.31	2.3	2.4	2.8	3.4
Ho	0.82	0.83	5.1	4.9	1.01	1.05	1.6	3.8	0.97	0.98	2.6	2.2	3.1	3.7
Er	2.22	2.20	5.2	5.0	2.92	3.00	1.1	3.0	2.47	2.55	1.0	3.4	2.4	3.8
Tm	0.30	0.30	4.3	4.1	0.44	0.44	8.7	7.9	0.31	0.33	2.4	6.1	5.2	6.1
Yb	2.02	2.16	3.9	7.4	2.86	3.00	1.9	4.9	1.92	2.00	0.7	4.3	2.2	5.5
Lu	0.31	0.31	5.5	5.3	0.47	0.45	4.0	5.2	0.28	0.27	1.3	5.4	3.6	5.3
Hf	3.40	3.50	6.1	6.2	2.45	2.50	1.6	2.4	4.09	4.46	2.2	8.5	3.3	5.7
Pb	7.16	6.30	4.3	14.4	6.43	5.80	1.6	10.9	2.51	2.40	11.9	11.1	5.9	12.1
Th	8.83	9.20	0.7	4.1	0.77	0.76	5.8	5.3	1.25	1.23	2.2	2.4	2.9	3.9
U	1.52	1.60	9.4	9.7	0.32	0.35	6.9	10.5	0.38	0.41	5.2	8.7	7.1	9.6

1: preferred reference values from GeoReM (Jochum et al., 2005). Accuracy is reported as the relative root-mean square deviation (%RMSD) compared to the preferred reference value. A synthetic calibration was used for Ta.

**Table S5. Precursor compositions of whole rock samples from fractionated melts**

LAVAS AND DIABASE DIKES		GABBRO	ANORTHOSITE
<b>Major elements (Wt %)</b>			
SiO <sub>2</sub>	49.5 ± 0.7 [n = 96]		46.2 ± 0.4 [n = 2]
TiO <sub>2</sub>	0.007 (Zr) + 0.7 ± 0.1 [R <sup>2</sup> = 0.78; n = 92]	0.19 ± 0.04 [n = 5]	0.05 ± 0.01 [n = 2]
Al <sub>2</sub> O <sub>3</sub>	16.8 ± 0.7 [n = 98]	18.7 ± 2.9 [n = 5]	31.5 ± 0.8 [n = 2]
FeO <sub>T</sub>	8.98 – 0.028 (Nb) ± 0.52 [R <sup>2</sup> = 0.41; n = 93]	3.52 ± 0.74 [n = 5]	1.02 ± 0.45 [n = 2]
MnO	0.17 – 0.0008 (Nb) ± 0.02 [R <sup>2</sup> = 0.22; n = 93]	0.08 ± 0.02 [n = 5]	0.01 ± 0.01 [n = 2]
MgO	9.42 – 0.018 (Zr) ± 0.48 [R <sup>2</sup> = 0.59; n = 97]		1.5 ± 0.8 [n = 2]
CaO	15.7 – 0.037 (Zr) ± 0.55 [R <sup>2</sup> = 0.84; n = 99]		
Na <sub>2</sub> O	0.012 (Zr) + 1.79 ± 0.33 [R <sup>2</sup> = 0.59; n = 97]		2.4 ± 0.4 [n = 2]
K <sub>2</sub> O	0.018 (Nb) + 0.023 ± 0.09 [R <sup>2</sup> = 0.93; n = 99]		
P <sub>2</sub> O <sub>5</sub>	0.006 (Nb) + 0.085 ± 0.04 [R <sup>2</sup> = 0.83; n = 93]	0.003 ± 0.001 [n = 5]	0.003 ± 0.005 [n = 2]
H <sub>2</sub> O	0.015 (Nb) + 0.18 ± 0.10 [R <sup>2</sup> = 0.85; n = 89] <sup>a</sup>		
S	0.08 ± 0.01 [n = 98]	0.03 ± 0.01 [n = 5]	0.01 ± 0.007 [n = 2]
<b>Trace elements (ppm)</b>			
Cu	183 – 0.8 (Zr) ± 10 [R <sup>2</sup> = 0.58; n = 33] <sup>a</sup>	40 ± 24 [n = 5]	56
Zn	56 ± 10 [n = 37]	13 ± 5 [n = 5]	11
Rb	0.68 (Nb) – 3.1 ± 1.6 [R <sup>2</sup> = 0.98; n = 89] <sup>a</sup>	0.87 ± 0.55 [n = 5]	2.3 ± 1.8 [n = 2]
Sr	5.7 (Nb) + 100 ± 38 [R <sup>2</sup> = 0.87; n = 99]		
Cs	0.008 (Nb) – 0.067 ± 0.03 [R <sup>2</sup> = 0.95; n = 59]	0.01 ± 0.005 [n = 4]	0.008 ± 0.001 [n = 2]
Ba	6.5 (Nb) – 20 ± 19 [R <sup>2</sup> = 0.97; n = 93]	16 ± 10 [n = 5]	18 ± 12 [n = 2]
La	0.505 (Nb) + 1.38 ± 1.1 [R <sup>2</sup> = 0.98; n = 99]		
Lu	0.38 ± 0.06 [n = 93]	0.08 ± 0.03 [n = 5]	0.009 ± 0.005 [n = 2]
U	0.022 (Nb) – 0.03 ± 0.08 [R <sup>2</sup> = 0.95; n = 93]	0.011 ± 0.010 [n = 4]	0.004 ± 0.003 [n = 2]

Precursor compositions determined from analyses of Macquarie glasses (Kamenetsky et al., 2000; Wertz, 2003), supplemented with analyses of the least altered whole rock samples, including: aphyric dike chilled margins (MQ 47A and 105B); aphyric dikes (MQ 61 and 69A); gabbros (MQ 40, 45A, 85, 87 and 115A); and anorthosites (MQ 59 and 92). <sup>a</sup> Glass samples only used to determine precursor trend. Major elements: wt%; trace elements (including Zr and Nb): ppm. R<sup>2</sup> = regression coefficient; n = number of

fresh precursor samples that the regression line or average composition was determined from. Primary magmatic errors (red type) were calculated as one standard deviation of the residuals of precursor compositions from regression lines, or one standard deviation of precursor compositions for elements that show no trend with Nb or Zr. For some elements magmatic trends could not be extended to anorthosite ± gabbro compositions; for these elements average values of least-altered samples are used. Fe is recalculated to total FeO ( $\text{FeO}_T = 0.8998 \text{ Fe}_2\text{O}_3_T$ ) to allow comparison with analyzed glass Fe concentrations.

**Table S6. Precursor compositions of porphyritic lavas and dikes, and olivine gabbro samples.**

PORPHYRITIC LAVAS AND DIKES:				
Nb <20 ppm:		Nb > 20 ppm:		OLIVINE GABBRO:
MAQ 222B		MQ58		MQ 68A
MQ12		MAQ119		MQ 74B
MQ98		MAQ 134		MQ 75
MQ102		MAQ 139B		
MAQ 126B		MAQ 136, MQ 89		
<b>Major elements (Wt %)</b>				
SiO <sub>2</sub>	<b>49.2 ± 0.7</b>	<b>49.7 ± 1.1</b>	<b>49.5 ± 1.3</b>	
TiO <sub>2</sub>	<b>1.05 ± 0.22</b>	<b>1.80 ± 0.36</b>	<b>0.17 ± 0.08</b>	
Al <sub>2</sub> O <sub>3</sub>	<b>17.4 ± 2.4</b>	<b>16.2 ± 1.8</b>	<b>16.4 ± 2.9</b>	
FeO <sub>T</sub>	<b>7.70 ± 1.26</b>	<b>9.16 ± 1.5</b>	<b>3.61 ± 0.38</b>	
MnO	<b>0.13 ± 0.03</b>	<b>0.15 ± 0.02</b>	<b>0.08 ± 0.01</b>	
MgO	<b>7.83 ± 1.31</b>	<b>6.18 ± 0.71</b>	<b>10.5 ± 0.4</b>	
CaO	<b>12.2 ± 0.6</b>	<b>10.42 ± 0.99</b>	<b>16.87 ± 1.81</b>	
Na <sub>2</sub> O	<b>2.93 ± 0.17</b>	<b>4.08 ± 0.55</b>	<b>1.52 ± 0.12</b>	
K <sub>2</sub> O	<b>0.16 ± 0.09</b>	<b>0.38 ± 0.19</b>	<b>0.08 ± 0.02</b>	
P <sub>2</sub> O <sub>5</sub>	<b>0.12 ± 0.05</b>	<b>0.33 ± 0.05</b>	<b>0.002 ± 0.002</b>	
S	<b>0.008 ± 0.006</b>	<b>0.005 ± 0.003</b>	<b>0.02 ± 0.02</b>	
H <sub>2</sub> O	<b>1.27 ± 0.48</b>	<b>1.60 ± 0.31</b>	<b>1.22 ± 0.13</b>	
<b>Trace elements (ppm)</b>				
Cu	<b>19.0 ± 19.5</b>	<b>32.1 ± 25.9</b>	<b>39.9 ± 32.5</b>	
Zn	<b>27.9 ± 14.4</b>	<b>55.4 ± 2.1</b>	<b>13.4 ± 4.1</b>	
Rb	<b>2.66 ± 3.14</b>	<b>6.56 ± 5.49</b>	<b>0.45 ± 0.18</b>	
Sr	<b>182 ± 38</b>	<b>230 ± 54</b>	<b>92.6 ± 10.0</b>	
Cs	<b>0.021 ± 0.031</b>	<b>0.055 ± 0.051</b>	<b>0.015 ± 0.013</b>	
Ba	<b>49.4 ± 41.5</b>	<b>131 ± 31</b>	<b>6.57 ± 4.29</b>	
La	<b>5.03 ± 2.72</b>	<b>17.8 ± 1.0</b>	<b>0.26 ± 0.05</b>	
Lu	<b>0.28 ± 0.11</b>	<b>0.45 ± 0.13</b>	<b>0.09 ± 0.04</b>	
U	<b>0.143 ± 0.138</b>	<b>0.46 ± 0.13</b>	<b>0.004 ± 0.001</b>	

**Table S7. Hydrothermal fluxes (kg/m<sup>2</sup>)**

Unit factor	LAVAS		LAVA-DIKE TZ		SHEETED DIKES		LOWER CRUST		UPPER CRUST		FULL CRUST		
	WR	WR + Veins	WR	WR + Veins	WR	WR + Veins	WR	WR + Veins	WR	WR + Veins	WR	WR + Veins	
<b>SiO<sub>2</sub></b>	x10 <sup>4</sup>	-1.6 to 0.6	-0.4 to 1.9	0.02 to 0.27	0.04 to 0.30	-0.2 to 1.3	0.5 to 2.0	-3.1 to -4.4	-2.4 to -3.7	-1.1 to 1.6	0.8 to 3.5	-2.0 to -5.0	-2.3 to 0.7
<b>TiO<sub>2</sub></b>	x10 <sup>3</sup>	-3.0 to -6.2	-2.9 to -6.2	0.01 to 0.08	0.01 to 0.08	-0.5 to 1.8	-0.5 to 1.8	-1.6 to 0.4	-1.6 to 0.5	-1.9 to -5.9	-1.9 to -5.9	-2.2 to -6.7	-2.2 to -6.6
<b>Al<sub>2</sub>O<sub>3</sub></b>	x10 <sup>3</sup>	-8.4 to 5.8	-7.3 to 6.9	-1.5 to -2.2	-1.4 to -2.1	-9.7 to -24	-7.8 to -22	-14 to 14	-13 to 15	-10 to -30	-6.8 to -27	-2.8 to -37	-33 to 1.7
<b>FeO</b>	x10 <sup>3</sup>	-11 to -27	-6.8 to -23	-0.23 to 0.39	-0.12 to 0.50	-3.5 to 7.7	0.7 to 12	5.4 to 17	7.0 to 19	-6.9 to -26	-18 to 1.4	-17 to 6	-6.8 to 16
<b>MnO</b>	x10 <sup>2</sup>	-1.9 to 1.0	-1.8 to 1.1	0.07 to 0.22	-0.20 to -0.31	-0.6 to 2.3	-0.4 to 2.5	-1.4 to 0.9	-1.2 to 1.0	-1.5 to 2.6	-1.2 to 3.0	-2.0 to 2.7	-1.6 to 3.1
<b>MgO</b>	x10 <sup>4</sup>	1.1 to 2.4	1.3 to 2.7	-0.01 to -0.07	-0.06 to 0	-0.3 to 0.4	-0.1 to 0.6	0.9 to 3.2	1.0 to 3.4	1.0 to 2.6	1.5 to 3.0	2.5 to 5.3	3.0 to 5.9
<b>CaO</b>	x10 <sup>4</sup>	-1.7 to -3.0	-1.6 to -3.0	-0.20 to -0.31	-0.20 to -0.31	-3.0 to -4.1	-3.0 to -4.1	-2.2 to -3.9	-2.1 to -3.8	-5.3 to -7.1	-5.2 to -6.9	-8.0 to -10	-7.8 to -10
<b>Na<sub>2</sub>O</b>	x10 <sup>3</sup>	-1.2 to 6.6	-1.2 to 6.6	0.71 to 1.1	0.70 to 1.1	7.7 to 12	7.6 to 12	7.3 to 13	7.6 to 13	9.0 to 18	8.9 to 18	18 to 29	18 to 29
<b>K<sub>2</sub>O</b>	x10 <sup>3</sup>	3.4 to 5.3	3.6 to 5.5	0.28 to 0.44	0.27 to 0.44	0.65 to 1.9	0.64 to 1.9	-0.38 to 1.4	-0.36 to 1.4	4.8 to 7.1	5.0 to 7.4	5.0 to 7.9	5.3 to 8.2
<b>S</b>	x10 <sup>2</sup>	-3.7 to -7.9	-3.4 to 0.8	2.0 to 5.1	2.0 to 5.2	-3.5 to 2.2	15 to 21	1.9 to 5.5	1.9 to 5.5	-6.8 to 0.9	17 to 24	-3.5 to 5.0	20 to 28
<b>LOI</b>	x10 <sup>4</sup>	2.7 to 3.5	3.0 to 3.7	0.14 to 0.18	0.14 to 0.19	2.4 to 3.0	2.6 to 3.1	4.4 to 5.3	4.4 to 5.4	5.5 to 6.5	5.9 to 6.8	10 to 11	11 to 12
<b>Cu</b>		-8.4 to -33		-1.0 to -2.7		-56 to -98		-23 to -52		-75 to -124		-109 to -165	
<b>Zn</b>		-9.5 to 6.2		0.89 to 1.7		-6.4 to -21		-6.0 to 5.5		-3.3 to -24		-2.1 to -26	
<b>Rb</b>		2.6 to 10		0.55 to 0.94		-1.9 to 1.2		-2.7 to 7.9		2.7 to 11		-3.3 to 6.4	
<b>Sr</b>		-13 to -65		-2.5 to 0.35		-16 to 22		31 to 88		-5.2 to -70		-21 to 65	
<b>Cs</b>		0.09 to 0.26		0 to 0.002		-0.01 to -0.04		-0.05 to -0.11		0.06 to 0.24		-0.02 to 0.17	
<b>Ba</b>		29 to 82		4.7 to 9.3		-2.4 to 40		-1.0 to -46		48 to 116		18 to 99	
<b>U</b>		-0.02 to 0.09		0.001 to 0.004		-0.06 to -0.13		-0.06 to -0.15		-0.12 to 0.01		-0.08 to -0.24	