

U-Pb zircon (CA-ID-TIMS) age supports globally synchronous Sturtian deglaciation

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Abstract

A central prediction of the Snowball Earth hypothesis is that glacial onset should be synchronous at low latitudes, and its termination should be rapid and synchronous globally. High precision U/Pb zircon ages provide supporting evidence for the synchronous onset (within error) of the Sturtian glaciation (ca. 716 Ma) on multiple continents. Successful application of Re-Os techniques on organic rich shales and carbonates allow for the possibility of a globally synchronous Sturtian deglaciation (ca. 660 Ma), but the sparse isotopic age constraints leave this open to debate. Here we report the first high precision U-Pb zircon age of 663.03 ± 0.11 Ma (2σ) for the end Sturtian recorded in the Bolla Bollana Formation of South Australia. This age supports previously published ages and is permissive with a globally synchronous deglaciation. In conjunction with the timing of glacial onset, this age reinforces the ca. 58 Myr duration of the Sturtian Snowball.

Keywords: Snowball Earth, Sturtian, Synchronous Deglaciation, CA-ID-TIMS

Introduction

Some of the most extreme climate episodes the Earth has experienced occurred during a protracted phase of near-global glaciations in the Cryogenian period (720-635 Ma). These “Snowball” glaciations (Hoffman et al., 1998; Kirschvink, 1992), or cryochrons (Hoffman et al., 2017) are evidenced by the wide distribution of glacial deposits, particularly those associated with carbonate platforms, with the implication that glaciers entered the marine realm

in the warmest areas (Harland, 1964; Kirschvink, 1992), along with sedimentary and palaeomagnetic evidence for grounded low latitude ice sheets (Evans, 2000; Hoffman and Schrag, 2002). While debate over the exact nature of these glaciations has ensued (e.g. Allen and Etienne, 2008; Eyles and Januszczak, 2004; Le Heron et al., 2011; Williams et al., 2016), the Snowball Earth hypothesis presents a holistic model when all relevant observations and modelling results are considered (Hoffman et al., 2017).

Many mechanisms have been proposed to explain the descent into the Snowball Earth climate state, including: enhanced silicate weathering associated with the emplacement of several large igneous provinces (Cox et al., 2016b; Godderis et al., 2003); nutrient induced enhancement in primary productivity (Cox et al., 2016b; Gernon et al., 2016; Horton, 2015); a reduction in volcanic degassing from continental arcs (McKenzie et al., 2016); and increased albedo due to volcanic sulphate aerosol emission (Macdonald and Wordsworth, 2017). Indeed, none of these mechanisms are mutually exclusive and long term global cooling could be due to feedbacks between these factors in conjunction with an a low-latitude continental configuration (Li et al., 2013; Merdith et al., 2017), which favours a cooler planet (Kirschvink, 1992). Regardless of the cause(s) of long term climate forcing, a variety of 1-D energy-balance models (Budyko, 1969; North et al., 1981; Sellers, 1969) and 3-D general circulation models (Abbot et al., 2013) exhibit an instability, or bifurcation: a point at which ice-albedo feedback becomes self-sustaining and ice lines progress rapidly and unstoppably to the equator.

Emergence from these global glaciations is linked to the sub-zero temperatures of the Snowball climate state and by consequence a restricted hydrological cycle that inhibits atmospheric CO₂ scrubbing through silicate weathering. Ultimately, this process allows atmospheric CO₂ to reach levels sufficient to overcome the ice-albedo feedback (Kirschvink, 1992).

As the Snowball Earth model is tied to ice-albedo feedback followed by extreme greenhouse warming, a central prediction of the hypothesis is that the onset should be synchronous at low latitudes, and its termination should be rapid and globally synchronous. Testing for synchronicity relies heavily on a scaffold of isotopic U-Pb ages obtained from zircon-bearing volcanic ash beds and Re-Os ages from black shales from multiple and globally distributed correlated sequences.

To date, high precision U-Pb zircon ages for the Sturtian glaciation have demonstrated the onset of glaciation on several continents occurs between ca. 719 to ca. 710 (Bowring et al., 2007; Cox et al., 2015; Lan et al., 2014; Macdonald et al., 2010; Macdonald et al., 2017; Song et al., 2017; Zhou et al., 2004) (Table 1a). Furthermore, lower precision Re-Os ages on black shales from Australia, North America and Mongolia (Kendall et al., 2006; Rooney et al., 2014; Rooney et al., 2015) and a low precision U-Pb zircon age for South China (Yu et al., 2017), hint at a globally synchronous termination (Table 1b).

Here we report the first high precision chemical abrasion isotope dilution thermal ionization mass spectrometry (CA-ID-TIMS) U-Pb zircon age for the termination of the Sturtian glaciation along with supporting detrital zircon ages. These ages establish the timing of deglaciation of the Sturtian glaciation in Australia and is consistent with a globally synchronous end of the Sturtian glaciation at ca. 660 Ma and its ca. 58 Myr duration (Macdonald et al., 2010).

Geologic Setting

The Sturtian and Marinoan Series, along with the older Torrensian and Willouran, were defined as chronostratigraphic subdivisions of the Adelaide System by Mawson and Sprigg (1950). The Sturtian was characterised by, but not restricted to, sediments of glacial origin first recognised by Howchin (1901) in Sturt Gorge near Adelaide; it also includes the underlying Belair Subgroup of the Burra Group and the overlying Tapley Hill Formation. Subsequent

mapping by the Geological Survey of South Australia established the wide distribution of glaciogenic sedimentary rocks of both Sturtian and Marinoan age. Together with an intervening succession of interglacial origin, these are grouped as the Umberatana Group of the Adelaide Rift Complex (Thompson et al., 1964), and are best exposed in the Flinders, Willouran and Gammon Ranges of South Australia (Fig. 1). However, Sturtian glacials are found discontinuously throughout the basin, including the first discovery (the Sturt Tillite) near Adelaide. Other lithostratigraphic names have been applied in other regions of the basin: Appila Formation in the Mid-North, Pualco Formation in the Olary Ranges and Bolla Bollana Formation in the northern Flinders Ranges. In the central Flinders and Olary Ranges, laminated siltstone (Benda Siltstone) and bedded ironstones (Holowilena and Braemar Ironstones) with dropstones are closely associated (Cox et al., 2016a; Le Heron et al., 2011; Lechte and Wallace, 2015; Lottermoser and Ashley, 2000). Basal Sturtian glacials rest on prominent erosional surfaces cut into a wide variety of older rocks (mostly Burra Group, but also Callanna Group and crystalline basement near the basin margins). The Wilyerpa Formation in the central Flinders and Olary Ranges overlies these rocks, locally unconformably, and is considered to represent the waning stages of the Sturtian glaciation, while the Lyndhurst Formation in the north-eastern Flinders Ranges is a likely equivalent (Preiss, 2000). The whole Sturtian glacial succession comprises the Yudnamutana Subgroup, the geology and tectonic setting of which was summarised by Preiss (2000) and Preiss et al., (2011).

The Yudnamutana Subgroup varies considerably in thickness from ~100m to over 5000m, reflecting local tectonic control. Extremely thick sections of the Yudnamutana Subgroup were deposited in actively extending grabens in the eastern (Baratta Trough) and north-eastern Flinders Ranges (Yudnamutana Trough) regions. The characteristic Cryogenian ironstones are found only in the Baratta Trough. The combination of ice rafted debris and diamictite with striated clasts overwhelmingly demonstrate the glaciogenic nature of the

Yudnamutana Subgroup.

In the northern extent of the rift complex, within the Gammon and Willouran Ranges, the Bolla Bollana Formation is ~1500m thick near Arkaroola and contain abundant ice-rafted debris along with rare striated clasts (Le Heron et al., 2014). To the west and north-west, near Copley and Termination Hill (Fig. 1a), the Bolla Bollana Formation is comparatively thin reaching a maximum thickness ~150 m (Fig. 1b/c), with direct evidence for glaciation limited to abundant ice-rafted debris (Fig. 2). The glacial deposits in these localities overlie the pre-Sturtian Myrtle Springs Formation (Fig. 1b/c).

Previous age constraints on the Sturtian glaciation in South Australia include a maximum depositional age on the much older Skillogalee Dolomite of 792 ± 6 Ma (2σ) (Preiss et al., 2009), and a Re-Os black shale age from the lower part of the Tapley Hill Formation of 645.1 ± 4.8 Ma (2σ) and 647 ± 10 Ma (2σ) (Kendall et al., 2006). A complementary Re-Os age 657.2 ± 5.4 Ma (2σ) (Kendall et al., 2006) for the basal Aralka Formation, which overlies Sturtian glacials in the Amadeus Basin of the Northern Territory of Australia, constrains the Sturtian glaciation in Australia to be between ca. 798 Ma and ca. 657 Ma. A U-Pb SHRIMP age of 659 ± 6 Ma (2σ) on the thin tuff band in the Bolla Bollana Formation at the Copley section has been reported in abstract (Fanning and Link, 2006, 2008).

Geochronology

Age of tuff in the Bolla Bollana Formation at Copley

Zircon crystals were analysed from a ~3 cm thick tuff horizon within the Bolla Bollana Formation (Fig. 2). This horizon occurs ~80 m below the contact with the non-glacial Tindelpina Shale Member of the Tapley Hill Formation (Fig. 2). These zircons were screened via cathodoluminescence imaging and in situ LA-ICP-MS analysis prior to selection and analysis by CA-ID-TIMS using methods described by Macdonald et al. (2017). All crystal

fragments yield concordant analyses with dates between ca. 663 Ma and ca. 667 Ma (Fig. 3). Several older crystals are interpreted to record crystal inheritance into the volcanic population, a common phenomenon in silicic volcanic eruptions. The sixteen youngest grains form a statistically equivalent cluster with a weighted mean age of 663.03 ± 0.11 Ma (2σ - MSWD = 0.52, probability of fit = 0.93), which is interpreted as the igneous crystallization age of the zircons and the depositional age of the tuff. When tracer calibration and ^{238}U decay constant uncertainties are propagated in quadrature, the total error is 663.03 ± 0.76 Ma (2σ), which should be considered when comparing our dates with those derived from the ^{187}Re - ^{187}Os decay scheme (i.e. as presented in Table 1).

Bolla Bollana detrital zircon ages and provenance

Pre-Neoproterozoic detrital zircons from the Bolla Bollana Formation (sample W16-09 and W16-10; Fig. 1c) which were analysed via LA-ICP-MS (see supplementary material for further details), preserve U-Pb ages consistent with a relatively local provenance from the west and northwest. The Neoarchean-Paleoproterozoic Sleaford/Mulgathing Complexes of the central and southern Gawler Craton (ca. 2.4-2.5 Ga) are consistent with sourcing zircons of this age in both samples. The rare earlier Neoarchean to Mesoarchean grains may be recycled from metasedimentary gneisses in the Mulgathing Complex, where detrital grains of Mesoarchean age are preserved (McAveney, 2012). The majority of detrital zircon grains in both samples yield ages between ca. 1.80 and 1.55 Ga, these ages being common in the eastern Gawler Craton where they overlap with the ca. 1.86-1.8 Ga Cornian Orogeny and Donington Suite, the extensive ca. 1.73-1.69 Ga Kimban Orogeny and associated granitoids, the ca. 1.62-1.61 Ga St Peter Suite and the voluminous ca. 1.59-1.57 Ga Hiltaba siliceous large igneous province (Hand et al., 2007). The latter province comprises the Hiltaba Suite granitoids and the bimodal Gawler Range Volcanics. Felsic protoliths of ca. 1.7-1.5 Ga age are also common in the eastern Musgraves Province to the northwest of the field site (Smits et al., 2014; Wade et

al., 2006). Mesoproterozoic zircon dated between ca. 1.2 Ga and 1.1 Ga are probably related to magmatism within the Musgrave Province (Smithies et al., 2011) and ca. 1.1-1.0 Ga zircon is coeval with the extensive Warakurna Large Igneous Province that forms part of the Musgraves region (Wingate et al., 2004). Early Neoproterozoic (ca. 0.85 Ga) zircons conceivably come from magmatism associated with the Wooltana Volcanics that are found near the base of the Adelaidean System (Preiss, 1993). More problematic are Cryogenian zircons as magmatism is absent in the region, making the ca. 650 to 690 Ma zircons exotic to eastern Australia, however, these young ages provide a maximum depositional age of 654 ± 11 Ma (2σ) for the top of the Bolla Bollana. Interestingly, the youngest zircon from the basal Bolla Bollana gives a maximum depositional age of 673 ± 14 Ma (2σ). This age implies that much of the stratigraphic history of the early Cryogenian is missing (i.e. ca. 40 Myr) in the northern extent of the Adelaidean succession.

Is deglaciation globally synchronous?

Our new U-Pb age of 663.03 ± 0.11 Ma (2σ) for the Sturtian tuff that is interbedded with diamictite of the Bolla Bollana Formation, is consistent with the previously published Re-Os ages of 645.1 ± 4.8 Ma (2σ) and 647 ± 10 Ma (2σ) on the disconformably overlying Tapley Hill Formation, as well as the 657 ± 5.4 Ma (2σ) age from the Aralka Formation, a correlative unit to the Tapley Hill Formation in the Amadeus Basin of the Northern Territory of Australia. Furthermore, the ca. 663 Ma age is supported by a maximum depositional age of 654 ± 11 Ma (2σ) derived from detrital zircons from the Bolla Bollana Formation. It also confirms, with greater precision and accuracy, the SHRIMP age on the tuff previously reported in abstract (Fanning and Link, 2006, 2008). Given that this tuff horizon lies ~80 m below the transgressive post-glacial Tapley Hill Formation, this age constrains termination of the Sturtian glaciation in Australia to be younger than 663.03 ± 0.11 Ma (2σ).

Within a global context, this new age constraint is entirely consistent with low-precision U-Pb zircon ages in South China (Zhou et al., 2004) and Re-Os isochron ages from organic-rich deposits in North America, South China and Mongolia (Rooney et al., 2014; Rooney et al., 2015) (Fig. 5).

Conclusion

A U/Pb zircon age of 663.03 ± 0.11 Ma (2σ) provides the first high precision age constraint on the termination of the Sturtian glaciation globally, furthermore, this result robustly demonstrates the consilience between U-Pb and Re-Os geochronology. While further high precision ages from globally correlatable sequences are required, a central prediction of the Snowball climate state is long lived glaciation and a rapid and globally synchronous termination. The combination of Re-Os and U-Pb ages provides supporting evidence for a synchronous deglaciation of a Sturtian Snowball and its ca. 58 Myr duration.

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Figure Captions

Figure 1. (a) Aerial view of the Gammon and Willouran Ranges. (b) Geology of the Sturtian succession exposed west of Copley township. (c) Geology of the Sturtian succession exposed along the western flanks of Termination Hill on Witchelina Station. Samples locations: P1634: 30° 33' 34.9" S, 138° 23' 24.5" E; W16-09: 138° 01' 32" E, 30° 14' 10" S; W16-10: 138° 01' 14" E, 30° 13' 56" S.

Figure 2. Stratigraphic log of the upper portion of the Yudnamutana Subgroup and basal Tapley Hill Formation west of Copley.

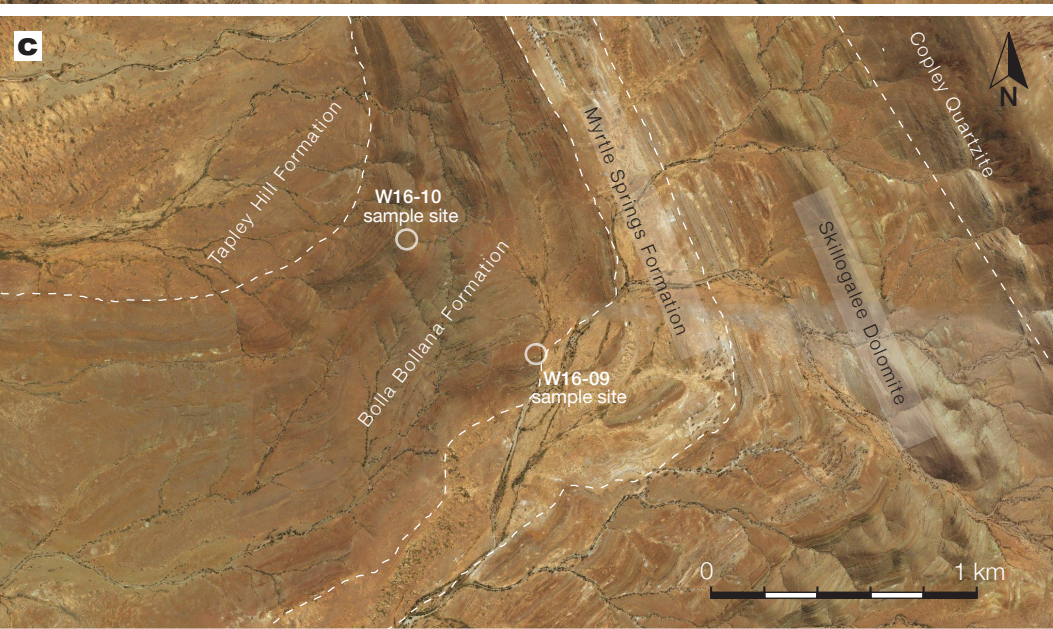
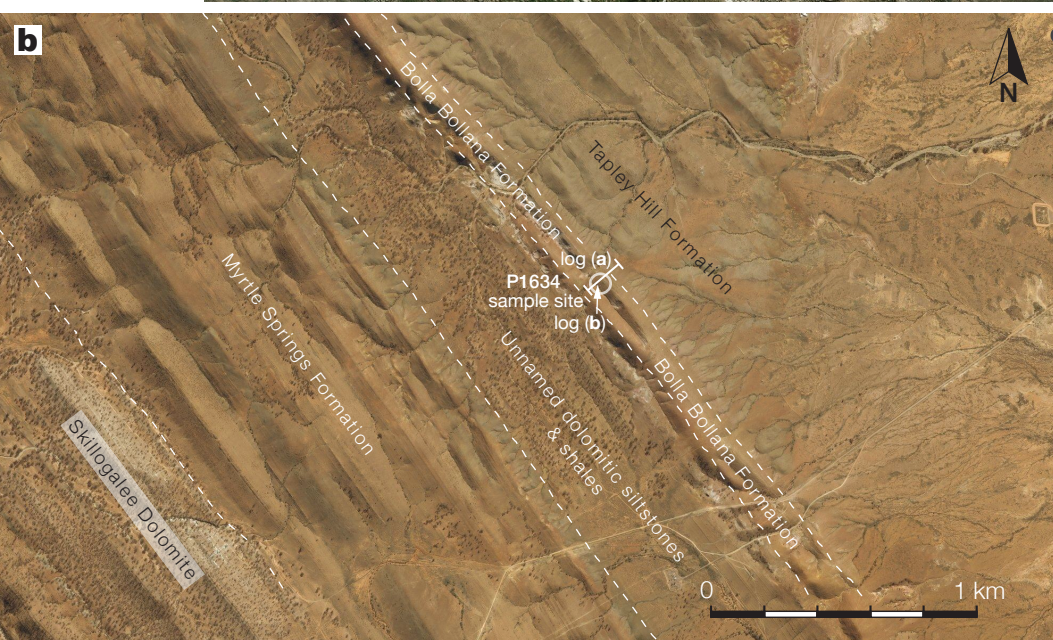
Figure 3. U/Pb Concordia (a) and individual zircon ages (b) along with their respective 2σ errors for tuff horizon within the Bolla Bollana Formation. Sample P1634: 30° 33' 34.9" S, 138° 23' 24.5" E (Figure 1b).

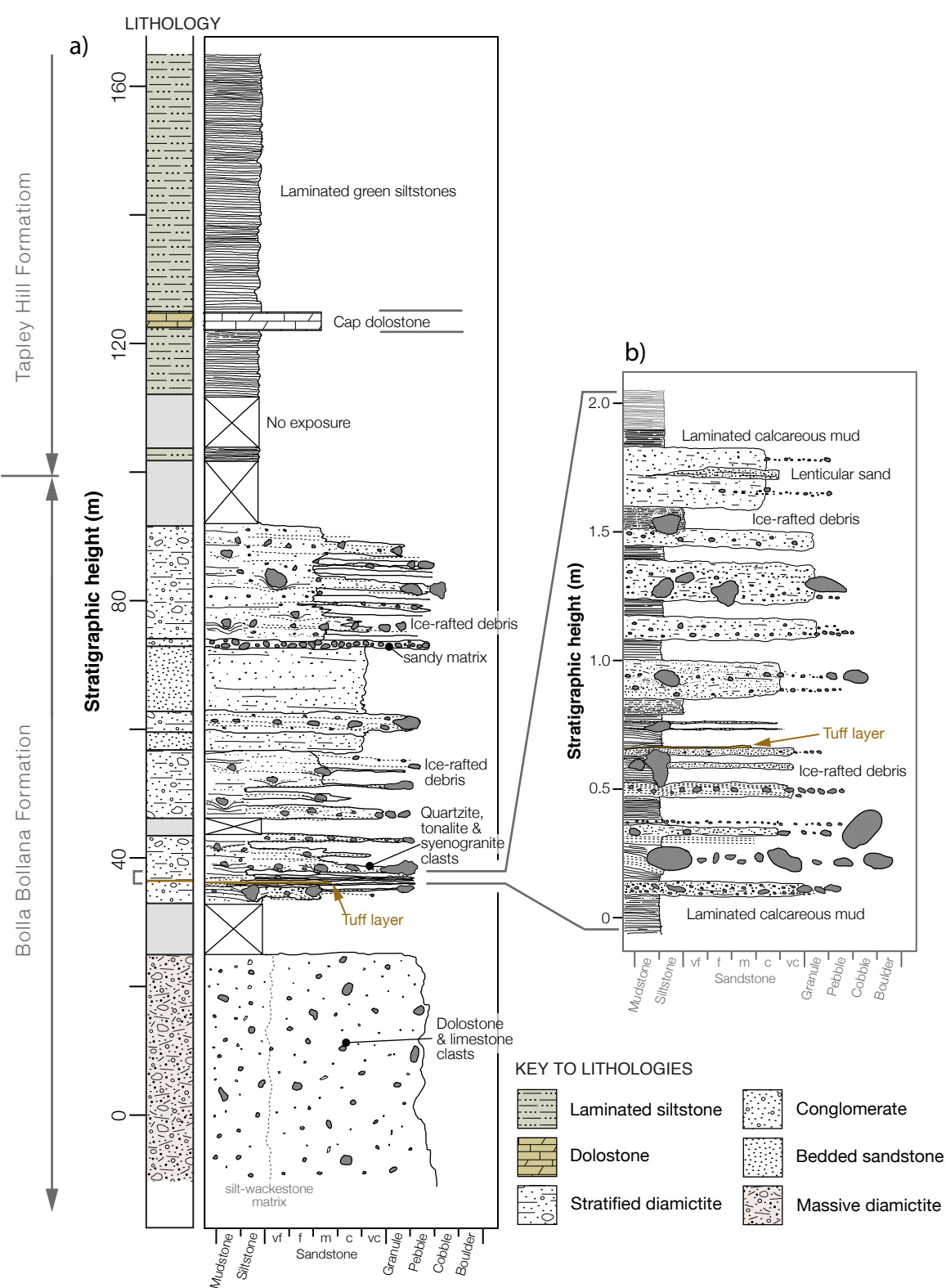
Figure 4. (a-b) Detrital zircon spectrum for samples from the Bolla Bollana Formation. W16-09: 138° 01' 32" E, 30° 14' 10" S; W16-10: 138° 01' 14" E, 30° 13' 56" S (Figure 1c). (c-d) Kernel density estimate of the same samples as (a) and (b). Plots constructed using DensityPlotter (Vermeesch, 2012). Grey curve is all zircons while black curve are zircons with concordancy between $> 95\%$ and $< 105\%$. Concordancy is defined as the ratio of the $^{207}\text{Pb}/^{206}\text{Pb}$ age to $^{206}\text{Pb}/^{238}\text{U}$ age.

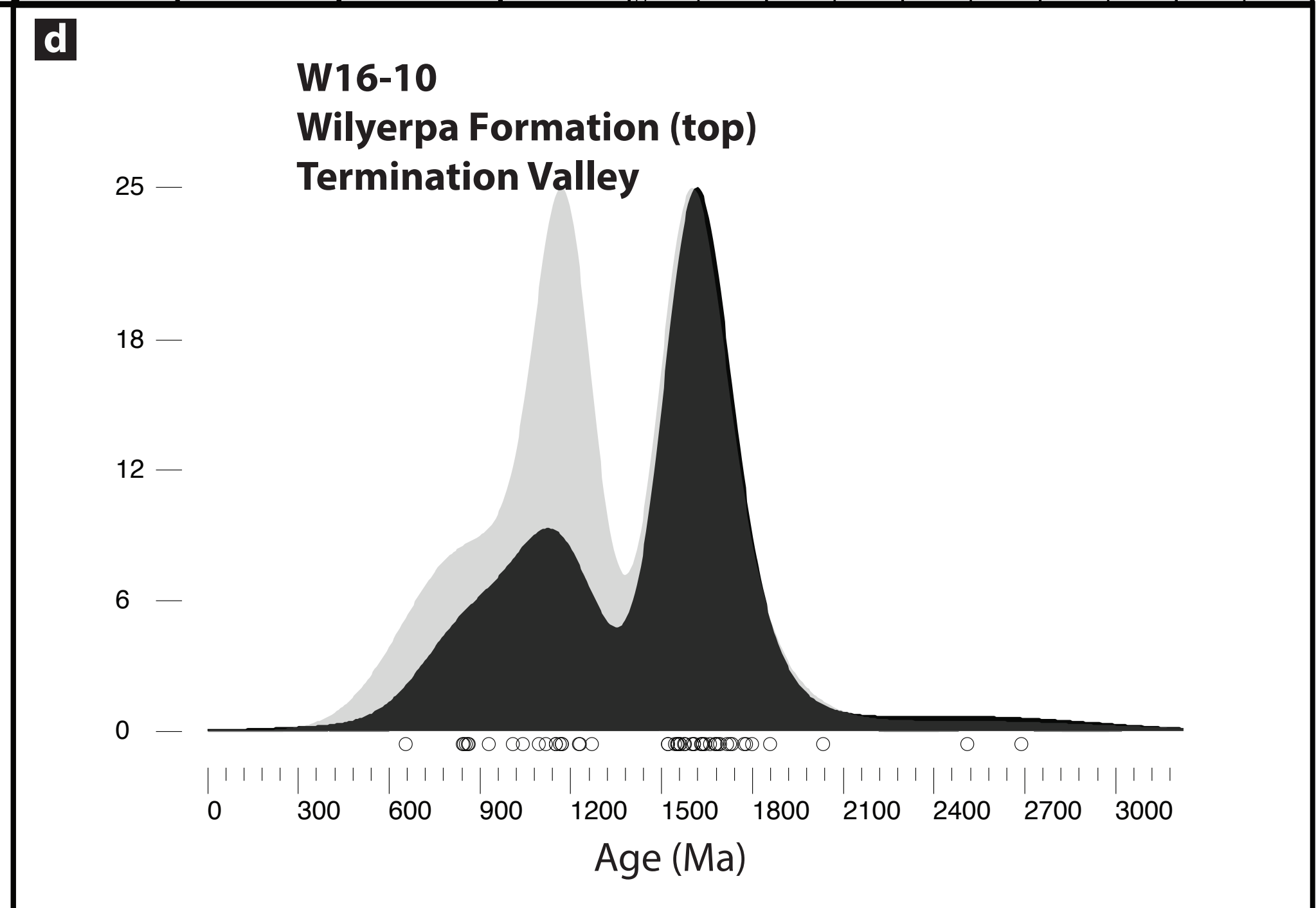
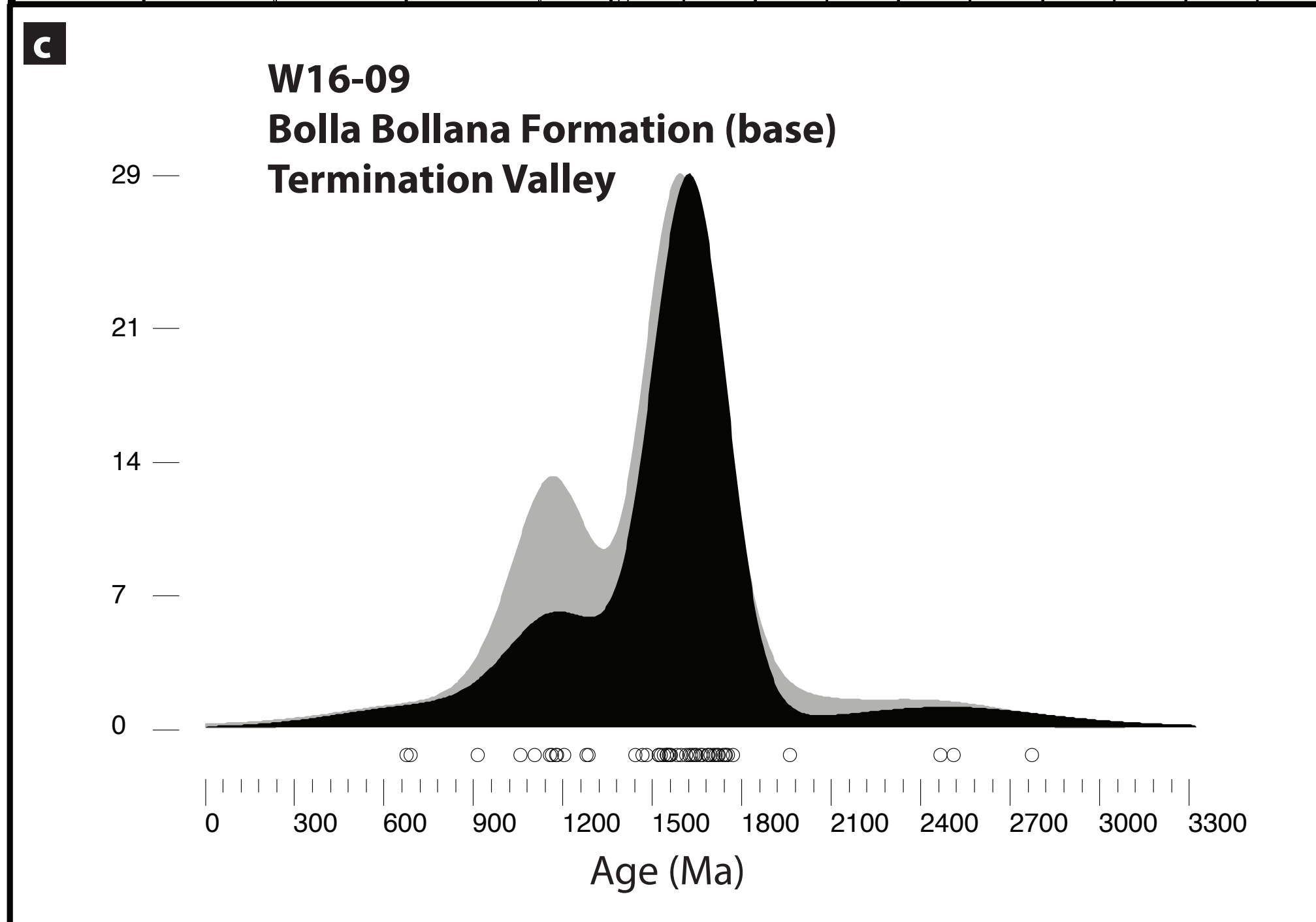
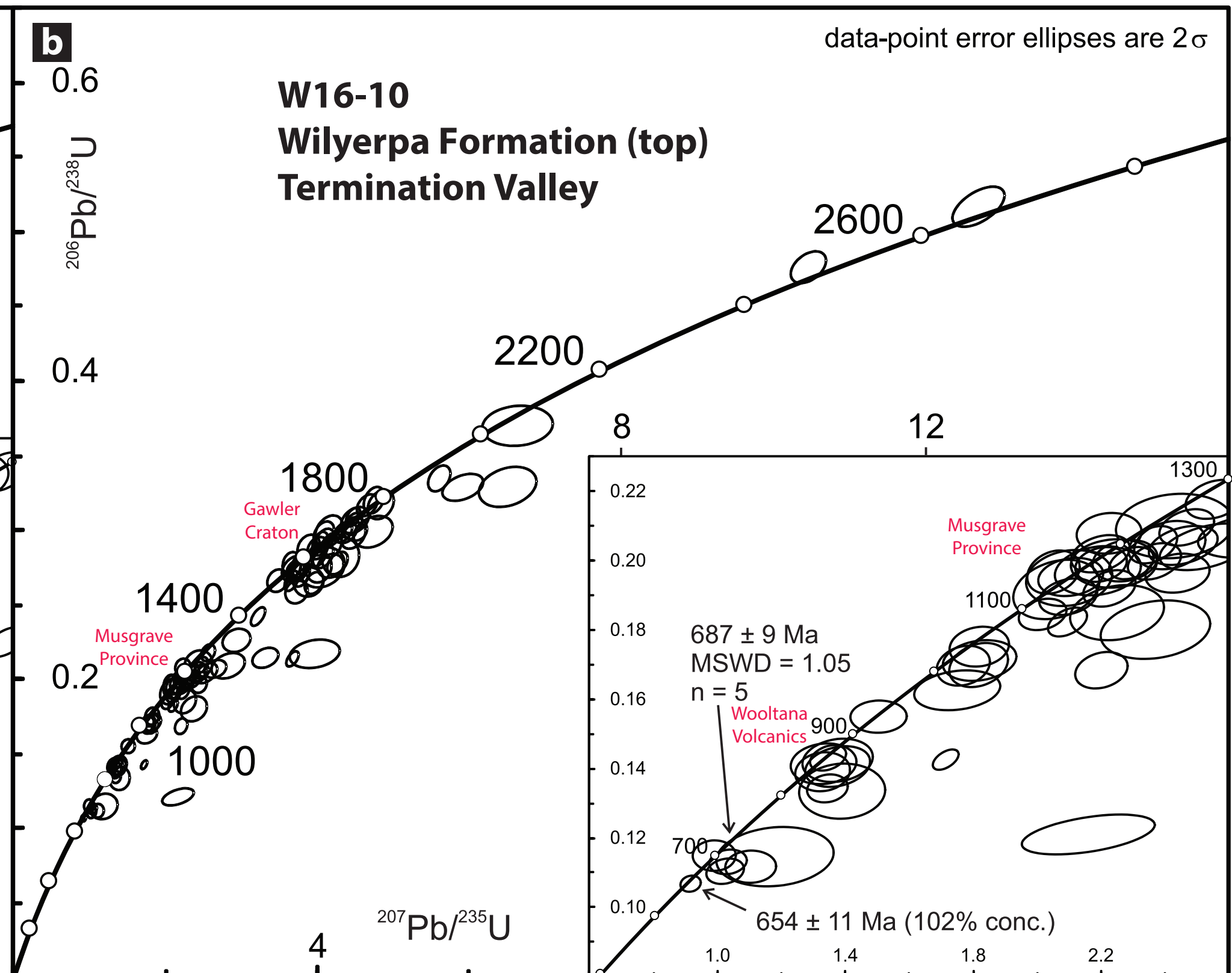
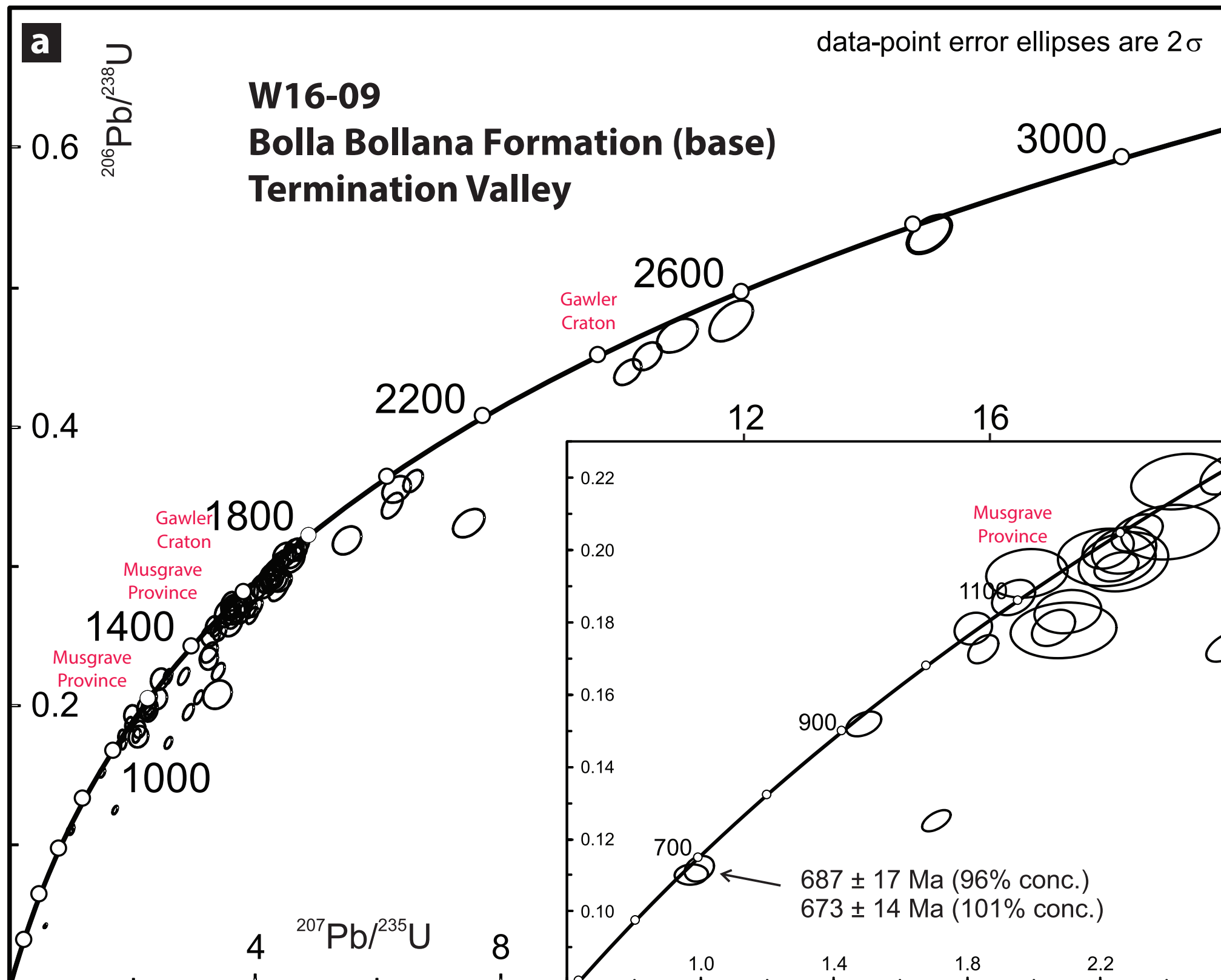
Figure 5. Compilation of U-Pb (Bowring et al., 2007; Cox et al., 2015; Fanning and Link, 2008; Lan et al., 2014; Macdonald et al., 2010; Macdonald et al., 2017; Song et al., 2017; Yu et al., 2017; Zhou et al., 2004) and Re-Os ages (Kendall et al., 2006; Rooney et al., 2014; Rooney et al., 2015) for the onset and termination of the Sturtian glaciation. For our new CA-ID-TIMS age, when tracer calibration and ^{238}U decay constant uncertainties are propagated in quadrature, the total error is 663.03 ± 0.76 Ma (2σ), which should be considered when comparing our dates with those derived from the ^{187}Re - ^{187}Os decay scheme. All errors are 2σ . For clarity some Laurentian ages listed in Table 1 are not reproduced here. Arrows indicate maximum and minimum age constraints.

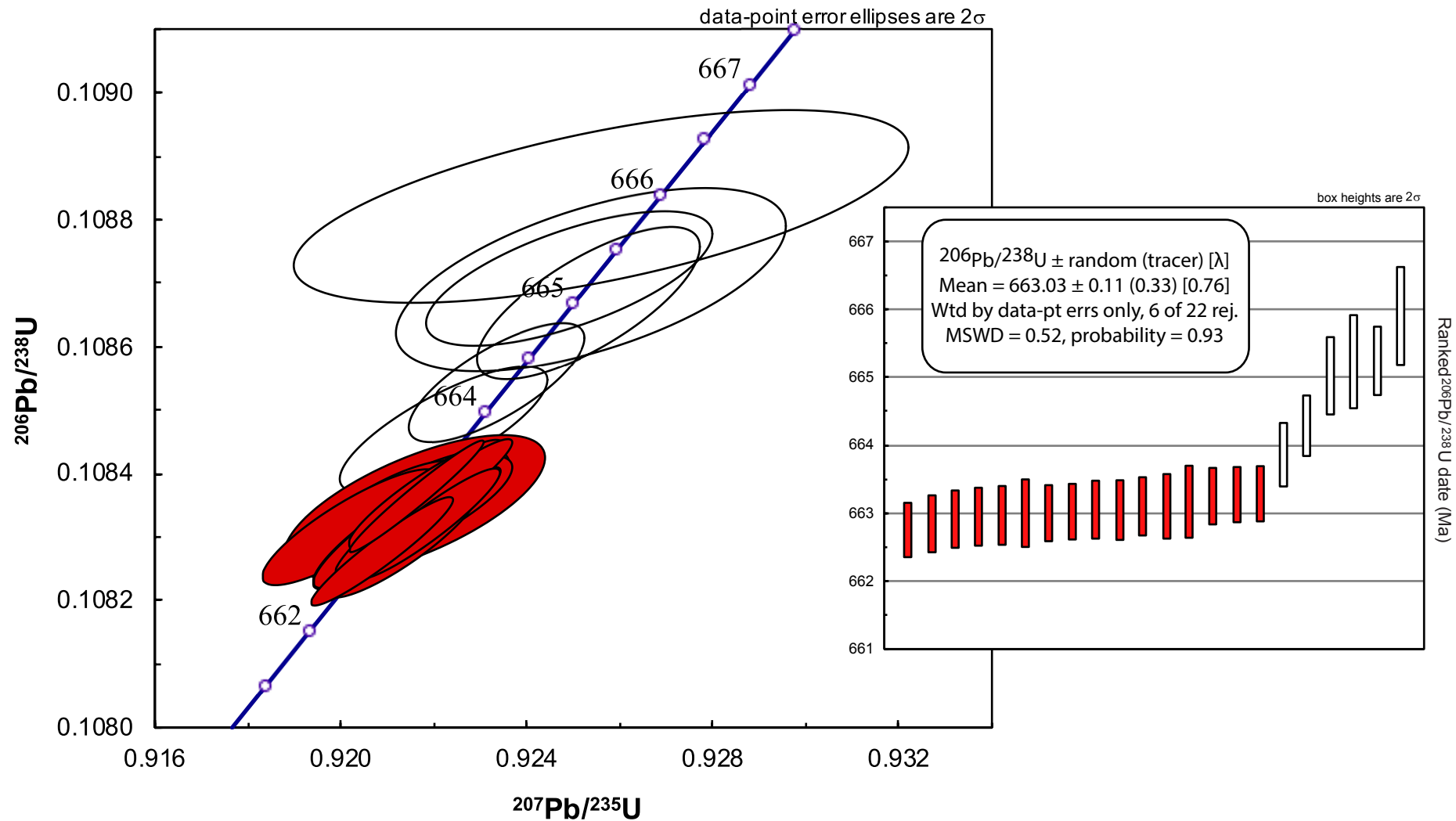
Table Captions

Table 1. Compilation of U-Pb (Bowring et al., 2007; Cox et al., 2015; Fanning and Link, 2008; Lan et al., 2014; Macdonald et al., 2010; Macdonald et al., 2017; Song et al., 2017; Yu et al., 2017; Zhou et al., 2004) and Re-Os ages (Kendall et al., 2006; Rooney et al., 2014; Rooney et al., 2015) for the onset and termination of the Sturtian glaciation. All U/Pb ages are $^{206}\text{Pb}/^{238}\text{U}$ ages except Zhou et al., 2004 which is a $^{207}\text{Pb}/^{206}\text{Pb}$ age. All errors are 2σ .





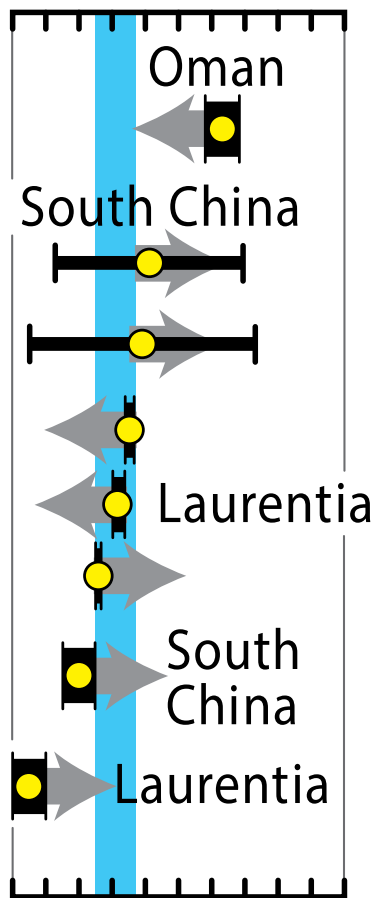




STURTIAN GLACIATION (58 Myr)

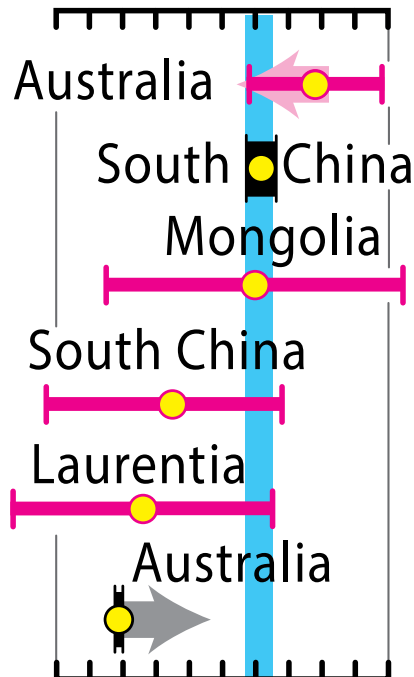
ONSET

720 715 710



TERMINATION

665 660 655



Sturtian epoch

720 700 680 660

U-Pb SIMS

U-Pb CA-ID-TIMS

Re-Os

max.

min.

| a) Age constraints for the onset of Sturtian glaciation | | | |
|--|-------------------|-----------------|-------------------------|
| Oman | 713.7 ± 0.5 | U-Pb CA-ID-TIMS | Bowring et. al., 2007 |
| South China | 714.6 ± 5.2 | U-Pb LA-ICP-MS | Song et. al., 2017 |
| South China | 715.9 ± 2.8 | U-Pb SIMS | Lan et. al., 2014 |
| Laurentia | 716.5 ± 0.2 | U-Pb CA-ID-TIMS | Macdonald et. al., 2010 |
| Laurentia | 717.4 ± 0.1 | U-Pb CA-ID-TIMS | Macdonald et. al., 2010 |
| Laurentia | 719.47 ± 0.29 | U-Pb CA-ID-TIMS | Cox et. al. 2015 |
| Laurentia | 716.5 ± 0.2 | U-Pb CA-ID-TIMS | Macdonald et. al., 2017 |
| Laurentia | 716.9 ± 0.4 | U-Pb CA-ID-TIMS | Macdonald et. al., 2017 |
| Laurentia | 717.4 ± 0.2 | U-Pb CA-ID-TIMS | Macdonald et. al., 2017 |
| Laurentia | 717.7 ± 0.3 | U-Pb CA-ID-TIMS | Macdonald et. al., 2017 |
| Laurentia | 717.8 ± 0.2 | U-Pb CA-ID-TIMS | Macdonald et. al., 2017 |
| b) Age constraints for the termination of Sturtian glaciation | | | |
| Australia | 645.1 ± 4.8 | Re-Os | Kendall et. al., 2006 |
| Australia | 647 ± 10 | Re-Os | Kendall et. al., 2006 |
| Australia | 657.2 ± 5.4 | Re-Os | Kendall et. al., 2006 |
| Tuva-Mongolia | 659.0 ± 4.5 | Re-Os | Rooney et. al., 2015 |
| Australia | 659.7 ± 5.3 | U-Pb SIMS | Fanning et. al., 2008 |
| Laurentia | 662.4 ± 3.9 | Re-Os | Rooney et. al. 2014 |
| South China | 662.7 ± 6.2 | U-Pb LA-ICP-MS | Yu et. al., 2017 |
| South China | 662.9 ± 4.3 | U-Pb ID-TIMS | Zhou et. al., 2004 |