

Organic geochemical, palynofacies, and petrographic analyses examining the hydrocarbon potential of the Cretaceous (Albian) Kharita Formation in the Matruh Basin, northwestern Egypt

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Abstract

A recent study of selected samples from the Cretaceous (Albian) Kharita Formation of Egypt revealed very good to excellent source rock (SR) potential for six intraformational, organic-rich intervals. This work investigates the SR potential of the entire Kharita Formation across the Matruh Basin, using samples from two wells: the Abu Tunis 1X well from the central part of the basin, and the Siqeifa 1X well drilled on the eastern margin of the basin. More strongly reducing conditions were developed in the centre of the basin, and resulted in the deposition of more organic-rich shales by comparison to the less reducing conditions that prevailed on the eastern basin margin, where the shales contain less organic matter. Deltaic intraformational shales and carbonates in the Kharita Formation of Abu Tunis 1X constitute a significant 120 m net of the potential SR. The lower Kharita Formation contains 34 m net shale SR of good to very good/excellent organic richness, yielding values of 1.14–11.59 wt % total organic carbon (TOC). The organic matter has low Hydrogen Index (HI) values (184–389 mg HC/g TOC) and amorphous organic matter (AOM) and relatively high non-opaque phytoclast frequencies indicating mainly gas/oil-prone organofacies (kerogen types II/III). The upper Kharita is more important, containing 86 m net shale/carbonate SR that has fair to good organic richness (0.8–1.8 wt % TOC), and lower HI (126–250 mg HC/g TOC), a dominance of non-opaque phytoclasts, and subordinate AOM frequencies, which together

36 indicate gas/oil-prone organofacies (kerogen Types III/II). In the Sikeifa 1X well, Kharita deltaic
37 intraformational shales and shaley dolostones comprise 80 m net SR, which has mainly fair to
38 good to less very good organic richness (0.8–2.1 wt % TOC), whilst low HI (93–220 mg HC/g
39 TOC), dominance of non-opaque phytoclasts and subordinate AOM indicate gas-prone
40 organofacies (kerogen Type III). A relative upward increase in deposition of lignite and coaly
41 carbonaceous material supports a gas-prone organofacies. Whilst thermal maturity indices only
42 point to immature to early mature (pre- to early oil-window) SRs in both the Abu Tunis 1X and
43 Sikeifa 1X wells, hydrocarbon exploration focussing on this potential source rock may be justified
44 in areas to the southeast of the Matruh Basin, where modelling indicates this unit may have
45 reached the late mature oil to main gas-generation window.

46
47 **Keywords:** Palynofacies, Organic petrology; Rock-Eval pyrolysis; Organofacies; Hydrocarbon
48 potential; Kharita Formation; Matruh Basin; Egypt

50 1. Introduction

51 Lower Cretaceous Albian strata in the northern and southern Mediterranean regions
52 contain intraformational organic-rich deposits which show significant hydrocarbon source rock
53 potential, such as in SE France (Herrle et al., 2003; Bornemann et al., 2005), Italy (Katz et al.,
54 2000), and Tunisia (e.g., Ben Fadhel et al., 2011; Khalifa et al., 2018). However, information on the
55 organic richness and petrography, and hence the hydrocarbon source rock potential of the Albian
56 deposits of the southeastern Mediterranean area, specifically of Egypt, is still fragmentary. The
57 only recent hydrocarbon study was carried out by Gentzis et al. (2019) on the Albian Kharita
58 Formation in the Abu Gharadig Basin in the north Western Desert of Egypt. Within the north
59 Western Desert, the Matruh area is typically regarded as one of the important hydrocarbon
60 producing basins amongst the other costal basins, namely the Shushan, Dahab–Mireir (= Alamein), and Natrun (Fig. 1). The Matruh Basin holds about 23 BBOE of known oil (Shahin, 1992)
62 and about 3 TCF of gas reserves (Metwalli et al., 2018), and the main Mesozoic hydrocarbon
63 source rocks are the Middle Jurassic upper Khatatba, the Lower Cretaceous (Berriasian–

64 Barremian) Alam El Bueib, and the Upper Cretaceous (middle–upper Cenomanian) “G” Member of
65 the Abu Roash formations (Fig. 2). The hydrocarbon reservoirs are represented by the lower
66 Khatatba, Lower Cretaceous (Aptian) Alamein, (Albian) Kharita, Upper Cretaceous (Cenomanian)
67 Bahariya, and (Turonian) “D” Member of the Abu Roash formations (e.g., Aram et al., 1988; EGPC,
68 1992; Shalaby et al., 2011; Abrams et al., 2016; Tahoun and Deaf, 2016; Deaf and Tahoun, 2018).

69 The Kharita Formation of the north Western Desert of Egypt is a classical hydrocarbon
70 reservoir (e.g., Meshref, 1996), deposited in a deltaic setting during a regressive phase (Said,
71 1990; Deaf, 2009; Mahmoud et al., 2019), where fine to coarse sandstone is the dominant lithology
72 but subordinate shale intercalations also occur (Hantar, 1990). This resulted in the Kharita
73 Formation being largely organic-poor, and having better reservoir quality rather than source rock
74 potential in most of the north Western Desert (EGPC, 1992). Indeed, in the neighbouring coastal
75 Shushan and Sidi Barani basins (Fig. 1), the Kharita Formation has been demonstrated to possess
76 only poor to fair gas source rock potential (avg. 0.83 wt % TOC, avg. HI 44.5 mg HC/g TOC, and
77 Tmax 429–435 °C) in the El Noor well of the Sidi Barani Basin (Aboul Ela et al., 2018). Similarly, to
78 the south of the costal basins, the Kharita Formation in the Abu Gharadig Basin has a very low gas
79 source potential (avg. 0.71 wt % TOC, avg. HI 149 mg HC/g TOC, and Tmax 426–438 °C), and
80 better reservoir properties (Gentzis et al., 2019). These properties are related to the high
81 sandstone volume and very low volumes of (organic-poor) shale (Metwalli and Pigott, 2005; Aboul
82 Ela et al., 2018). However, a quantitative organic petrographic analysis carried out by Deaf (2009)
83 on thirty-nine palynological samples of the Kharita Formation in the Abu Tunis 1X well in the
84 Matruh Basin identified the presence of six organic-rich horizons that contain exceptional
85 abundances of AOM (500,000–700,000, avg. 572,000 particles/gram of dry sediments; Fig. 3).
86 Geochemical evaluation by Tahoun et al. (2017) of these samples has indicated very good to
87 excellent hydrocarbon source potential for these six fine clastic (mainly shales), organic-rich
88 intervals. The total organic carbon content of these Albian samples ranged from 2 to 11.6 wt %
89 with an average of 3.4 wt %, and S₂ values which ranged from 1.1 to 23.5 mg HC/g dry rock (avg.
90 8 mg HC/g dry rock). The calculated HI (165–318, avg. 228 mg HC/g TOC) and organic
91 petrographic analyses (light microscopic palynofacies and UV fluorescence) indicated a mixture of

kerogen Types III and II. Thermal maturation indices indicated immature to early mature (early oil-window) stage of maturity for these sample, with values for Tmax of 416–428 °C (avg. 421 °C), vitrinite reflectance of 0.49–0.58 % Rv, and a thermal alteration index (TAI) of 2 to 2+.

To our knowledge, the current investigation presents the first organic petrographic analysis of the Albian Kharita Formation deposits in Egypt from the Matruh Basin, integrated here with organic geochemical and palynological analyses to evaluate the hydrocarbon potential of this historically neglected formation.

The hydrocarbon potential of source rocks has conventionally been assessed using organic geochemical parameters alone (e.g., Shalaby et al., 2011, 2012; Makky et al., 2014). Other investigations combined reflected white light (RWL) petrography and organic geochemical analyses of the organic matter to study the hydrocarbon potential of source rocks (e.g., Espitalié et al., 1985; Mukhopadhyay et al., 1989; Powell and Boreham, 1994; Shalaby et al., 2012; Hazra et al., 2015). Other studies have emphasised organic facies analysis, based on the RWL examination of the organic matter (Suárez-Ruiz, et al. 2012; Mendonça Filho et al., 2017), while the role of transmitted white light (TWL) petrography has been minor. Here, we propose a practical and productive protocol, based on the integration of organic petrographic (TWL, RWL, and ultraviolet fluorescence = UVF), palynofacies, and geochemical analyses of the organic matter that can be applied to any formation being studied in any basin, worldwide (Fig. 4). This protocol provides visual characterization and quantification of the kerogen constituents via TWL petrography, which is then correlated to kerogen characteristics defined using RWL. Moreover, this method overcomes the limitations of Rock-Eval analysis, such as the problems with identifying kerogen types due to mixing/averaging of reactive and inert kerogen types. RWL analysis does have certain limitations, whereby it cannot provide a detailed visual characterization of the different maceral constituents and cannot infer the botanical precursors of the organic matter in detail (Suárez-Ruiz et al., 2012; Mendonça Filho et al., 2017). The above problem was largely solved by Mendonça Filho (2012) who proposed palynological working groups (WG), a scheme which permits kerogen/maceral types to be correlated between TWL and RWL techniques. This cross correlation was successfully applied in the current study, and provided a good means of calibration between the TWL-based

(i.e., spore colouration) and the RWL-based (i.e., vitrinite reflectance/UV fluorescence) thermal maturity analyses. Furthermore, most hydrocarbon exploration studies have identified the environmental settings of the organic facies mainly based on RWL organic petrography and geochemical analyses (e.g., Singh et al., 2017b). In this study, we will present an interpretation of the depositional environments based mainly on integrating palynological and sedimentological data. This approach has been widely used in several palynological studies and provides important information on the source and the environmental parameters that controlled the accumulation and preservation of organic matter (e.g., Tyson, 1996; El-Soughier et al., 2014; Tahoun and Deaf, 2016; Mahmoud et al., 2017; Tahoun et al., 2017; Deaf and Tahoun, 2018; Deaf et al., 2019). Palynological analysis enjoys the merit of being simple and cost effective as it needs no further laboratory preparations, uses the same slide-mounted organic matter used for the thermal maturity analysis (i.e. TAI), and provides rapid but informative interpretations. On the other hand, several hydrocarbon investigations were based mainly on TWL (i.e., palynofacies) and geochemical analyses (e.g., Alaug et al., 2014; Tahoun and Deaf, 2016; Deaf and Tahoun, 2018). Recently, Gentzis et al. (2018) used an approach, similar to the one we propose here, to study the hydrocarbon potential of some Egyptian Upper Jurassic sediments, but without employing the organic facies concept. Thus, the current investigation aims to present a balanced and integrated TWL/RWL organic petrographic, palynofacies, and geochemical identification of the organic facies of the Egyptian Albian Kharita Formation to explore the hydrocarbon potential of this neglected unit at both the basin centre and the eastern margin of the Matruh Basin.

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141 **2. Geological setting and stratigraphy**

142 The Matruh Basin is a Late Jurassic–Early Cretaceous graben with a NNE–SSW orientation, and
143 lies approximately between longitudes 26° and 27° 30' E and latitudes 31° and 31° 17' N (EGPC,
144 1992; Meshref, 1996; Guiraud and Bosworth, 1999; Shalaby et al., 2012). It is bounded to the
145 northwest by the Sidi Barani High and to the South by the Ras Qattara High (Fig. 1). The Matruh
146 basin was originally formed as one large continental basin made of the Matruh-Shushan Basin
147 during the Permo-Triassic (Meshref and Hammouda 1990; Abdel Halim and Moussad, 1992). It

148 developed later as a rift basin during the late Cimmerian Orogeny as a result of the separation of
149 the northern African plate from the European plate (Meshref, 1996). In most of the basin area, the
150 sedimentary sequence ranges in age from the Cambrian to the Miocene (Fig. 2), while in other
151 parts of the basin the Cambrian to Middle Jurassic is represented by basement (Aram et al., 1988).

152 Sedimentation in the Matruh Basin was controlled largely by tectonics, where four major
153 sedimentary cycles are separated by regional unconformities (Sultan and Halim, 1988). The first
154 cycle was developed during the Early-Mid Jurassic, where the fluvio–lacustrine sediments of the
155 Ras Qattara and Yakout formations and the deltaic and shallow marine sediments of the Khatatba
156 and Masajid formations were deposited (Fig. 2). The Lower–Upper (Albian–middle Cenomanian)
157 Cretaceous fluvio–deltaic to shallow marine sediments of the Alam El Bueib, Alamein, Dahab,
158 Kharita, and Bahariya formations represent the second cycle. The uppermost (upper Cenomanian–
159 Campanian) Cretaceous open marine shales of the Abu Roash and carbonates of the Khoman
160 formations represent the third cycle, whilst the fourth cycle consists of the upper Paleogene–lower
161 Neogene (Eocene–Miocene) open marine shale and carbonates of the Apollonia, Dabaa, Moghra,
162 and Marmarica formations.

163 Lithostratigraphically, the Kharita Formation was introduced by Norton (1967) as a member
164 of the Burg El Arab Formation, Ghorab et al. (1971) later reclassifying it as a formation. According
165 to Said (1990) the Kharita is a clastic unit of Albian age, comprised of fine- to coarse-grained
166 sandstones with subordinate shale and carbonate horizons (Hantar, 1990; Kerdany and Cherif,
167 1990), and its type locality located in the interval between 2501 to 2890 m in the Kharita-1 well in
168 the Abu Dahab-Mireir Basin (Fig. 1). In the north Western Desert area, including the Matruh Basin,
169 the Kharita Formation conformably overlies upper Aptian shales and fine sandstones of the Dahab
170 Formation and conformably underlies the middle Cenomanian clastic and carbonate units of the
171 Bahariya Formation. The Kharita has been interpreted as being deposited in fluvial to shallow
172 marine settings during a major marine regressive episode (Hantar, 1990; Said, 1990).
173 Palynological investigations (e.g., El-Soughier et al., 2014; Aboul Ela et al., 2018) have also
174 suggested deltaic to marginal marine environments for the Kharita Formation in several areas of
175 the north Western Desert of Egypt (e.g., Abu Dahab–Mireir and Abu Gharadig). In the Matruh

176 Basin, the Kharita Formation was deposited during a pronounced regressive phase in marginal
177 marine settings in the area of the Siqefa 1X well (Mahmoud and Deaf, 2007) and in deltaic to
178 shallow marine settings in the area of the Abu Tunis 1X well (Deaf, 2009; Deaf et al., 2019).

179 In the Abu Tunis 1X well, the Kharita Formation is composed of medium- to fine-grained
180 argillaceous sandstones with a carbonate and/or silicic matrix and contain traces of pyrite, bulk
181 coaly carbonaceous material and infrequent traces of anhydrite. These argillaceous sands are
182 intercalated with thin light grey to green fissile shales that contain traces of macroscopic black
183 coaly carbonaceous material. The Kharita Formation in the Abu Tunis 1X attains a thickness of
184 1700 ft (518.2 m). In the Siqefa 1X well, the Kharita Formation is 1550 ft (472.4 m) thick and is
185 mainly composed of fine- to coarse-grained argillaceous sandstones that are intercalated with thin,
186 light to medium grey fissile to massive shale horizons. The sandstones possess dolomitic cement
187 and contain traces of low-rank coal (lignite) and/or bulk coaly carbonaceous material, in addition to
188 frequent traces of pyrite and glauconite. Shale intercalations contain traces of pyrite and glauconite
189 and infrequent traces of lignite.

190 Mahmoud and Deaf (2007) and Deaf et al. (2014) studied the palynostratigraphy of the
191 Albian Kharita Formation in the Siqefa 1X and Abu Tunis 1X wells respectively. In the Abu Tunis
192 1X well, the drilling company WEPCO (1968) did not identify the lower part of the unit being
193 studied here with any particular formation, and described its upper part as simply “Cenomanian
194 Clastics”. However, based on palynological and lithological data, Deaf et al. (2014) identified this
195 section of the Abu Tunis 1X stratigraphy as the Kharita Formation, and proposed an Albian age. In
196 the Siqefa 1X well, WEPCO (1970) did not identify the lower part of the clastic sequence we study
197 here, and identified the upper part of this sedimentary sequence as “Cenomanian Carbonates”.
198 Mahmoud and Deaf (2007) suggested an Albian age for this clastic unit, and combining this age
199 with the dominant sandstone lithology, we believe that the clastic unit in the Siqefa 1X well also
200 represents the Kharita Formation, which is known to have an Albian age across the north Western
201 Desert of Egypt (Said, 1990).

202

203 **3. Materials and methods**

204 **3.1. Open-system programmed pyrolysis (Rock-Eval pyrolysis)**

205 Twenty-eight samples from the Abu Tunis 1X well and twenty-five samples from the
206 Siqueifa well were analyzed by open-system programmed pyrolysis analysis (Tables 1 and 2). The
207 instrument utilized was a Rock-Eval® 6 Turbo unit (RE6) made by Vinci Technologies, France. The
208 pyrolysis method used was the Basic/Bulk-Rock method (IFP Rock-Eval methods®) typically
209 utilized for source rock evaluation. Briefly, 60 mg of pulverized sample were weighed and placed
210 inside a stainless-steel crucible. The sample was first decomposed in the pyrolysis oven under a
211 nitrogen atmosphere to obtain the weight % of pyrolyzable carbon (PC) and pyrolyzable mineral
212 carbon. Hydrocarbons and both CO and CO₂ were detected simultaneously by a flame ionization
213 detector (FID for hydrocarbons) and infrared cell (IR for CO₂ and CO). Subsequently, each sample
214 was combusted in the oxidation oven to obtain the weight % of residual carbon (RC) and oxidized
215 mineral carbon (oxiMinC). The temperature program for the pyrolysis cycle was 300 °C isothermal
216 for 3 min followed by a 25 °C/min ramping from 300 °C to 650 °C. The oxidation program was 300
217 °C isothermal for 60 s followed by a 25 °C/min ramping from 300 °C to 850 °C, held isothermal for
218 5 min at 850 °C. For more details, the reader is referred to Behar et al. (2001).

219

220 **3.2. Transmitted light microscopy: palynology and palynofacies analysis**

221 The standard HCl/HF maceration techniques of Phipps and Playford (1984) and Green
222 (2001) were followed to extract the palynological matter (PM) as it is referred to by the
223 palynological community for the transmitted white light (TWL) petrographic and palynofacies
224 analyses (e.g., Batten and Stead, 2005; Traverse, 2007). A sum total of fifty-nine ditch-cutting rock
225 samples (34 from Abu Tunis 1X and 25 from Siqueifa 1X) were collected. Three grams of each
226 sample from the Abu Tunis 1X well were processed, and spiked with one tablet of modern
227 *Lycopodium* spores (12,542 grains/tablet with $V \pm 3.3\%$) by Deaf (2009) at the School of Ocean
228 and Earth Science, National Oceanography Centre, University of Southampton, UK. Samples were
229 sieved through a 15 µm mesh and organic residue from each sample was strewn onto two cover
230 slips, dried and later mounted on microscope slides using Elvacite 2044. Ten grams of each
231 sample from the Siqueifa 1X well were processed by Deaf (2002) using the standard HCl/HF

232 maceration at the Geology Department, Faculty of Science, Assiut University, Egypt. The organic
233 residues stored at the Geological Museum of the Geology Department (Faculty of Science, Assiut
234 University) were resuspended with distilled water and few drops of each residue were mounted
235 onto each of two microscope slides using Canada Balsam.

236 Palynofacies analysis was carried out on the PM residues recovered from 34 samples of
237 Abu Tunis 1X and from 16 out of 25 samples available after the Rock-Eval and vitrinite reflectance
238 analyses of Siqueifa 1X using the TWL Olympus (BX41) microscope at the Geology Department,
239 Faculty of Science, Assiut University. For qualitative palynofacies analyses, the identification and
240 classification of the PM (kerogen) constituents were made following Tyson (1995) and the recently
241 updated classification presented by Mendonça Filho et al. (2012) and Mendonça Filho and
242 Gonçalves (2017). Quantitatively, separate counts of 500 particles of the kerogen constituents
243 ('palynodebris' or macerals), i.e., palynomorphs, translucent and opaque phytoclasts and AOM
244 (Tables 5 and 7), and of 300 palynomorphs (Tables 3 and 4) were carried out for each sample,
245 following Tyson (1995, p. 439) and Deaf and Tahoun (2018). The separate palynomorph count was
246 undertaken to counter the dilution effect of other "kerogen" constituents to the palynomorphs
247 (Tyson, 1995) and to obtain a statistically meaningful representation of different organic matter
248 entities (e.g., Tyson, 1995). For the interpretation of depositional environments of the Siqueifa 1X
249 well rocks, eight palynological categories based on two subsets of the organic matter were used,
250 namely the total palynomorphs and the total PM (see tables 4 and 8; e.g., Tyson, 1995; Mahmoud
251 et al., 2017; Deaf and Tahoun, 2018). Thus, the percentage frequencies of the terrestrial and
252 marine palynomorphs are derived from total palynomorphs, while those of the kerogen constituents
253 are derived from the total kerogen. Furthermore, the percentage frequency data derived from the
254 raw count of 300 palynomorphs was normalized to 100% (Suárez-Ruiz, et al., 2012) and plotted in
255 the Spores–Microplankton–Pollen (SMP) ternary diagram of Federova (1977) and Düringer and
256 Döbinger (1985) to interpret the depositional environment (tables 3 and 4). Similarly, the
257 normalized percentage frequency data derived from the total (kerogen) count of 500 particles
258 (Tables 5 and 7) was plotted on the Liptinite–Vitrinite–Inertinite (LVI) ternary diagram of Dow
259 (1982) and Tyson (1995) to identify hydrocarbon generation potential, and on the AOM–

Phytoclasts–Palynomorphs (APP) ternary plot of Tyson (1995) to identify the depositional environment and redox conditions (see Table 6; Tyson, 1995). It is important to note that the maceral analysis made for the LVI plot was performed on the PM (kerogen) concentrates through a series of non-overlapping traverses across the strew slides using the TWL microscope. The 500 particles of the three main groups of the PM (palynomorphs, phytoclasts, and AOM) were counted closest to the centre of the field of view according to Tyson (1995) and Mendonça Filho (2012). The counted particles were then related to their macerals and maceral groups as identified under RWL microscopy according to the International Committee for Coal and Organic Petrology (ICCP) palynofacies working group (see Mendonça Filho et al., 2012). Under the latter scheme, opaque phytoclasts are not assigned a specific maceral type but have been allocated to the maceral group inertinite; non-opaque, biostructured phytoclasts were identified as belonging to either the maceral telinite or to the maceral group vitrinite. Other non-opaque, membranous and cuticular phytoclasts have been allocated to the maceral cutinite and the maceral group liptinite. Sporomorphs and dinoflagellate cysts have been identified as the macerals sporinite and lamalginite respectively, and to the maceral group liptinite. Finally, AOM and highly degraded macrophyte tissues are equated with the maceral bituminite and the maceral group liptinite.

The palynological marine index (PMI) of Helenes and Somosa (1999) was calculated according to the following formula:

$$PMI = Rm (1+Rt) * 100$$

where,

Rm = the counted number of marine palynomorphs (dinoflagellate cysts and microforaminiferal test linings) per sample

Rt = the counted number of terrestrial palynomorphs (spores, pollen grains, and freshwater algae) per sample.

According to Helenes and Somosa (1999), a zero value of the PMI indicates a terrestrial environment without any form of marine signal. Low PMI values (~ 50–100) indicate marginal marine environments with brackish water conditions, whilst a PMI > 200 indicates deeper

288 marine/offshore environments. It should be borne in mind that Helenes and Somosa (1999) used
289 this formula to help interpret depositional environments of samples showing low to moderate
290 palynomorph recovery, where they used total counts of 100 grains, so that counts of 300 grains will
291 be more statistically meaningful (see Traverse, 1988, p. 490; Tyson, 1995, p. 433).

292 The quantitative absolute abundance analysis (grains/gram) of the Abu Tunis 1X PM
293 constituents was completed by Deaf (2009) to counter the data closure effect of the relative
294 abundance (percentage) analyses, and to identify any differences between the two datasets.
295 However, it was found that the vertical percentage distribution of single organic matter entities is
296 consistent with that of the absolute abundance data trends, especially for samples are dominated
297 by AOM and phytoclasts, although certain palynomorphs show minor differences in trends (see
298 Fig. 3 and Table 5).

299 The visual identification of spore colour in terms of the thermal alteration index (TAI) of
300 Pearson (1990) was undertaken on selected smooth spore grains from the total 300 palynomorphs
301 count to identify the degree of thermal maturation of the organic matter.

302

303 **3.3. Reflected light microscopy: vitrinite reflectance (*R_v*) and fluorescence**

304 The detailed sample preparation and analysis procedures employed here are described in
305 the ASTM D7708 standard test method (2014) and also by Hackley et al. (2015). Briefly, whole-
306 rock (WR) cuttings samples were crushed to -20 mesh size (0.85 mm) particles. Ground particles
307 were placed in specially designed plastic moulds (3.2 cm in diameter), mixed with Epo-Thin epoxy
308 resin and hardener (ratio of 2:1) and left to cure overnight. The resulting pellets were ground and
309 polished using a Buehler EcoMet/AutoMet 250 automated polisher using a combination of 320 µm
310 and 600 µm cloths and polished using a combination of two stages of a slurry of alumina powder
311 (0.3 and 0.05 µm) and water. Vitrinite random reflectance (%*R_v*) and fluorescence analyses were
312 performed on the dispersed organic matter (DOM) using a Carl Zeiss A2m Axio Imager RWL
313 microscope equipped with white (halogen) light source (12V/100W). A 50x oil immersion objective
314 (noil = 1.514 at 23°C) was used for a combined magnification of 500x. Qualitative UV analysis was
315 performed using the same microscope with high-pressure Hg lamp (HBO 100W) in the

316 fluorescence mode (FM). The excitation filter had a wavelength of 465 nm and the combined
317 dichroic mirror and barrier filter a wavelength of 515 nm. For additional information concerning
318 sample preparation protocol, see ASTM D 2797/D2979M-09 (2007). The description and
319 classification of the different kerogen constituents under RWL and FM were made following the
320 standard terminologies presented by the ICCP (1998, 2001) and Pickel et al. (2017).

321

322 **4. Results and discussion**

323 **4.1. Source rock potential**

324 The combined TOC (wt %) and S_2 (mg of HC per gram of rock) data were used together in
325 order to determine the source rock potential of the Kharita Formation, as suggested by Peters and
326 Cassa (1994) and Dembicki (2009), for the reason that elemental analysis of TOC takes into
327 account the “non-reactive organic carbon” (i.e., inertinite material), which is not capable of
328 generating hydrocarbons (Peters and Cassa, 1994; Hart and Steen, 2015). The S_2 parameter
329 measures the remaining (i.e., present-day) hydrocarbon generative potential, from the pyrolytic
330 degradation of the kerogen and heavy hydrocarbons. Thus, the S_2 parameter is regarded as a
331 more informative indicator of the source rock potential when compared to the TOC parameter
332 (Peters and Cassa, 1994; Dembicki, 2009, 2016). In the Abu Tunis 1X well, the deltaic deposits of
333 the lower Kharita Formation from 7300–6450 ft (2225–1966 m) are characterized mainly by
334 organic-rich intraformational shale intervals, and such lithologies have been identified as potential
335 source rocks in previous studies, such as Shalaby et al. (2011) and Abrams et al. (2016), which
336 comprise a net shale content of about 13.1 % (WEPCO, 1968). The lower Kharita Formation
337 contains several shale intervals ranging in thickness from 1.5 to 9.8 m. These shale intervals show
338 notable intraformational variations in organic richness, where they show good to very
339 good/excellent hydrocarbon source potential (1.14–11.59, avg. 4.19 TOC wt %; S_2 : 2.72–23.53,
340 avg. 10.41 mg HC/g; Figs. 5a, b, Table 1; see Baskin, 1997). One intraformational sandstone
341 sample was analysed from 6850 ft/2087.9 m and yielded 0.52 wt % TOC, and thus it does not
342 possess source rock characteristics, but would be better regarded as a reservoir unit. The above-

343 mentioned geochemical and lithological data indicate that the lower Kharita Formation contains
344 about 34 m net source rock.

345 The upper Kharita Formation (6449.9–5600 ft/1965.9–1707 m) is composed mainly of less
346 organic-rich argillaceous and/or carbonaceous sandstones and intercalations of limestone,
347 dolostone, and fewer shales. Shale intervals comprise 9.8 % (25.3 m) of the upper Kharita and
348 range in thickness from 3.7 to 12.2 m, while limestone and dolostone intervals comprise 23.4 %
349 (60.6 m) and range in thickness from 6.7 to 39.6 m (WEPCO, 1968). Shale and carbonate intervals
350 show little in the way of intraformational variation in organic richness and has fair to good
351 hydrocarbon source potential (0.82–1.78, avg. 1.10 TOC wt %; S_2 : 1.07–4.02, avg. 2.0 mg HC/g).
352 Sandy intervals (three samples analysed from 6100 ft/1859.3 m, 6050 ft/1844 m and 5800
353 ft/1767.8 m) are organic-lean (0.53–0.78, avg. 0.64 TOC wt %) and are considered reservoirs.
354 Based on the geochemical and lithological data advanced above, it is shown that shale and
355 carbonate intervals comprise a net source rock interval of about 33.2 % (86 m) of the upper Kharita
356 Formation. The cross-pot of S_2 against TOC indicates that the Kharita Formation has a fair to
357 excellent source rock potential (Fig. 5c).

358 In the eastern margin of the Matruh Basin in the Siqueifa 1X well, the Kharita Formation
359 shows a fair to good source rock potential (Figs. 5a–c, Table 2). The lower Kharita Formation from
360 7400–7000 ft (2255.5–2133.6 m) contains a higher intraformational shale content at 23.8 %, which
361 comprises 29 m net shale content of the lower Kharita interval (121.9 m) and ranges in thickness
362 from 3.4 to 9.1 m. The lower Kharita Formation also contains a shaley dolostone unit, which
363 represents about 7 % (8.5 m) net fine deposits of that part of the formation according to WEPCO
364 (1970). Shale and shaley dolostone intervals show little intraformational variation in organic
365 richness and have fair to good source rock potential (0.81–1.14, avg. 1.0 TOC wt %; S_2 : 1.06–1.69,
366 avg. 1.26 mg HC/g). The argillaceous and carbonaceous sandstones (three samples analysed
367 from 7250 ft/2209.8m, 7200 ft/2194.6 m and 6850 ft/2087.9 m) are less organic-rich (0.55–0.73,
368 avg. 0.66 TOC wt %), and, thus are considered as reservoir intervals. The combined geochemical
369 and lithological data indicates that the lower Kharita interval contains a total of about 30.7 % (37.5
370 m) net source rock.

371 The upper Kharita Formation (6999.9–5850 ft/2133.5–1783.1 m) shows a slight upward
372 increase in organic richness (Figs. 5a, b, Table 2), which is related to the deposition of more
373 frequent strings of lignitic and coaly carbonaceous material (Fig. 9) as was recorded by WEPCO
374 (1970). Shale intervals are more frequent but comprise a net shale content of about 12.1 % (42.4
375 m) of the upper Kharita interval (350.5 m) and range in thickness from 2.7 to 6.7 m (WEPCO,
376 1970). Shale and coaly shale intervals show fair to good gas-prone source rock potential (0.82–
377 2.14, avg. 1.33 TOC wt %; S₂: 0.88–4.70, avg. 2.10 mg HC/g). Argillaceous sandstones and
378 dolostones (three samples analysed from 6800 ft/2073 m, 6700 ft/2042.2 m and 6000 ft/1829 m)
379 show only poor to fair organic richness (0.35–0.67, avg. 0.49 TOC wt %). Thus, the upper Kharita
380 Formation is shown to contain 42.4 m (or 12.1%) net source rock.

381 As demonstrated here, the intraformational variations in organic richness require the
382 application of integrated lithofacies, palynofacies, and petrographic (TWL and RWL) analyses to
383 understand the environmental settings and the controlling factors acting on the organic richness,
384 and to overcome the averaging problem of the TOC and S₂ organic richness parameters.

385

386 **4.2. Environmental settings and their impact on kerogen preservation and quality**

387 Deaf (2009) and Deaf et al. (2019) interpreted the clastic deposits of the Kharita Formation
388 in the Abu Tunis 1X well as being deposited during a regressive phase of sedimentation in deltaic
389 to shallow marine environments. A plot of the main terrestrial and marine palynomorphs using the
390 SMP ternary diagram of Tyson (1995) also supports a deltaic to shallow marine depositional
391 setting for the Kharita deposits in the Abu Tunis 1X well (Fig. 6a). The proximal setting and the
392 development of regional humid climatic conditions during the Albian–Cenomanian (Deaf et al.,
393 2019) likely promoted flourishing of continental vegetation, associated with high run-off and
394 delivery of the large influxes of terrestrial organic matter into the deltaic environment in which the
395 Kharita Formation was deposited (Lamberson et al., 1991; Tyson, 1995). A plot of the main PM
396 constituents in the APP ternary diagram of Tyson (1995) indicates prevalence of reducing
397 (suboxic–anoxic) conditions during deposition of the Kharita Formation in the vicinity of Abu Tunis
398 1X (Figs. 7a and 8). Thus, favourable reducing conditions resulted in the deposition of organic-rich

399 clastics within the Kharita Formation at Abu Tunis, in marked contrast to the widely distributed
400 organic-poor clastics of the Kharita Formation across most of the rest of the north Western Desert
401 (Meshref, 1996; Metwalli and Pigott, 2005). The occurrence of lower quantities of terrestrial organic
402 matter in the marginal and shallow marine deposits of the underlying Alam El Bueib and Alamein
403 formations (Fig. 3), when regional climate was relatively drier in comparison to that of the Albian–
404 Cenomanian (Deaf, 2009; Deaf et al., 2019), supports our contention that humid climatic conditions
405 had an effect on the organic richness in the Kharita of Abu Tunis 1X. Adding to this idea, the
406 depositional cyclicity of the deltaic environment, as reflected in the sedimentary facies (sandstone
407 and shale alternations) and the self-potential (SP) log of the Kharita Formation (Deaf et al., 2019,
408 fig. 6) also affected kerogen quality, where kerogen Type III is alternates with and/or is mixed with
409 Type II (Fig. 8). The coarse-grained sandstones are dominated by vitrinite (kerogen Type III), whilst
410 the fine-grained siltstones and shales are dominated by the liptinite (kerogen Type II). This is
411 explained by the fact that high abundances of brown wood (Type II) tend to concentrate in coarse
412 siltstones and very fine sandstones of the proximal depositional settings that are close to land
413 vegetation (e.g., Habib, 1983; Firth, 1993; Tyson, 1995), whereas the finer particulate organic
414 matter (i.e., AOM) and miospores, typically smaller in size by comparison to the brown wood
415 particles found here, tend to concentrate in the fine sandstone, siltstone, and shale lithologies
416 (Batten, 1974; Bujak et al., 1977; Tyson, 1995). Furthermore, those shale intervals within the
417 Kharita Formation of the Abu Tunis 1X well (Fig. 5a, Table 1) that are exceptionally organic-rich
418 (from 4.41-11.60 wt % TOC) may be related to the Albian Oceanic Anoxic Event OAE 1b (Tahoun
419 et al., 2017). On the eastern margin of Matruh Basin, the Kharita Formation in the Siqueifa 1X well
420 shows lower organic richness and quality (Table 2). This is related to the more marginal marine
421 depositional setting of the formation (Mahmoud and Deaf, 2007). Detailed analysis of the
422 allochthonous (sporomorphs, non-opaque and opaque phytoclasts) and autochthonous
423 (dinoflagellate cysts and microforaminiferal test linings = MFTLs) organic matter again suggests
424 deposition in a deltaic environment. This is based on the very high abundance of terrestrially
425 derived organic matter (miospores and phytoclasts) and the very low abundance of dinoflagellate
426 cysts (Fig. 9). The rare occurrence and low species diversity (a total of 10 species) of the

dinoflagellate cysts, which are mainly dominated by the restricted, low salinity genera *Subtilisphaera*, *Odontochitina*, and *Cribroperidinium* (Mahmoud and Deaf, 2007, p. 216, fig. 4) also suggest deposition in proximal marginal marine (brackish–coastal) conditions (e.g., Mutterlose and Harding, 1987; Tahoun and Deaf, 2016; Deaf and Tahoun, 2018; Deaf et al., 2019). The high influxes of spores at the expense of spherical pollen grains (mainly *Arucariacites*) also suggests sedimentation in nearshore settings during a regressive episode, as high frequencies of spores are known to characterize delta-top and delta-front sand, silt and shale deposits (e.g., Degens and Mopper, 1976; Batten, 1982; Tyson, 1995). Similarly, the very high abundances of non-opaque phytoclasts indicate deposition in proximal marginal environments that were close to the fluvio-deltaic systems (e.g., Pocklington and Leonard, 1979; Tyson, 1995). The dominance of sandstones in which the few shale horizons are intercalated and high occurrences of non-opaque and opaque phytoclasts suggests deposition in proximal, shallow water marginal settings under relatively high-energy conditions, probably in the partly submerged bioturbated delta-front setting, an environment characterized by strong influxes of non-opaque and opaque phytoclasts (Tyson, 1995; Deaf et al., 2019). A deltaic setting is also indicated by the SMP ternary plot of the Siqueifa 1X well data (Fig. 6b). This type of setting is characterized by a highly fluctuating water table and strong water circulation (Boggs, 2006), where short-lived reducing (dysoxic–anoxic) conditions, as indicated by the APP ternary plot (Fig. 7b), and reflected by the frequent traces of pyrite, resulted in preservation of low concentrations of AOM mainly in the laminated Kharita shales (see Tyson, 1995). The lignite and bulk coaly carbonaceous material and pyrite are mainly associated with the sandstone units (Fig. 9), interpreted by Singh et al. (2017a) as indicating brackish water conditions at the depositional site of the Siqueifa 1X well, which fits with our interpreted deltaic setting. In the dysoxic–anoxic marginal marine facies, the formation of pyrite within lignite/carbonaceous material is related to the fluvial supply of detrital iron minerals commonly associated with high influxes of terrestrial organic matter into the brackish water, where bacterial sulphate reduction is an active process (Einsele, 1992; Tyson, 1995; Singh et al., 2017a).

The occurrence of glauconite along with carbonaceous material in samples from 7250 ft (2209.8 m) and 6050 ft (1844 m) of the Kharita Formation also indicates a marine influence

(Einsele, 1992; Dooley, 2006), which is further supported by the occurrence of open marine dinoflagellate cysts, for example *Oligosphaeridium* and *Florentinia* (Prauss, 2006; Carvalho et al., 2016; Deaf and Tahoun, 2018). Although the Abu Tunis 1X and Siqueifa 1X wells both being characterized by deltaic deposition, the Kharita Formation in Abu Tunis was deposited in relatively deeper deltaic to more distal shallow marine settings, as indicated by the open marine, middle shelf dinoflagellate cyst *Oligosphaeridium* in Abu Tunis, where the latter was recently found to tolerate low salinity conditions (Deaf and Tahoun, 2018). Conversely, the dinoflagellate cyst assemblages in Siqueifa 1X are composed mainly of the restricted (brackish) forms (*Subtilisphaera*, *Odontochitina*) with few open marine forms (*Oligosphaeridium* and *Florentinia*). The PMI values are higher in the Abu Tunis 1X well by comparison to those in the Siqueifa 1X well, which also supports an interpretation of relatively more offshore/deeper conditions in the Abu Tunis area (figs. 8 and 9). To summarize, it is suggested that more reducing conditions developed at the relatively more offshore/deeper settings of the Abu Tunis 1X well, resulting in preservation of more organic-rich deposits in comparison to the less reducing conditions that prevailed at the more marginal depositional setting of the Siqueifa 1X well on the eastern margin of Matruh Basin.

4.3. Organic geochemistry and petrography for organofacies detection

The combination of organic geochemical and TWL petrographic analysis (mainly palynofacies) is widely accepted among organic geochemists as a better tool used for recognizing organic facies and hydrocarbon source potential than using the kerogen type alone (Tyson, 1995; Mendonça Filho et al., 2012). This is because organic facies combined with the kerogen type reflect the source and depositional settings that control sedimentation and preservation of organic matter in a potential source rock (e.g., Tyson, 1995). Several definitions of 'organic facies' have been suggested (see Tyson, 1995; Mendonça Filho et al., 2012; Abarghani et al., 2018; and references therein). For example, Pepper and Corvi (1995, p. 297-298) defined an organofacies as "a collection of kerogens derived from common organic precursors, deposited under similar environmental conditions and exposed to similar to early diagenetic histories". They proposed six organofacies based on the kerogen type, principal biomass, sulphur content, and depositional

environment. However, Jones (1987) provided a classification of organic facies, which is primarily based on the organic geochemical parameters and takes into account the AOM as one of the main contributors to the organic matter, and hence provides a better identification of the organic facies. The above classification is recognized as a very practical tool because it discriminates the stratigraphic units successfully (e.g., Pasley, 1991; Tyson, 1995; Mendonça Filho et al., 2012). Later, Tyson (1995) provided a definition of the organic facies as “a body of sediment containing a distinctive assemblage of organic constituents, which can either be recognized by microscopy, or is associated with a characteristic bulk organic geochemical composition”. In this sense, Tyson (1995, p. 378) presented a modified model of Jones (1987), where he integrated the palynofacies kerogen parameters and the important environmental parameters with the pyrolysis-estimated organic geochemical parameters. The focus of the current study is to present detailed organic petrographic (TWL, RWL, and UV), palynofacies, and geochemical examinations of the organic matter of the Kharita Formation. Thus, the Jones (1987) model and the modified model of Tyson (1995) will be used here because they employ integrated organic petrographic and geochemical analyses, providing a better identification of the organic facies. In addition, these models provide supporting evidence on the source of the sedimentary organic matter and the environmental controls on sedimentation and preservation of the organic matter, which have been already interpreted in the previous section 4.2.

It is important to note that the deltaic deposits show strong vertical fluctuations in sedimentation (i.e., between regressive and transgressive), where repeated deposition of alternations of organic-rich, fine-grained and organic-poor, coarse-grained clastics is common (Einsele, 1992; Boggs, 2006). Here, source rock evaluation will be based primarily on the palynofacies units identified. As a result, the organic petrographic, geochemical and palynofacies units are used to define organofacies units according to the nature and type of organic matter and its organic geochemical parameters. This approach is consistent with the definitions of the organofacies of Jones (1987) and Tyson (1995).

Rock-Eval pyrolysis of the Abu Tunis 1X well samples indicates a mixture of kerogen types dominated by Type III along with subsidiary Type II for the entire Kharita Formation (HI: 126–389.3,

511 avg. 213.9 mg HC/g TOC; Fig. 10, Table 1). A plot of the different PM constituents in the Abu Tunis
 512 1X well in the LVI ternary diagram of Dow (1982) shows that the samples are rich in oil-prone
 513 liptinite macerals and gas-prone vitrinite macerals, and poor in inertinite macerals (Fig. 11a, Table
 514 5). TWL petrography and palynofacies analysis of the organic matter show a palynofacies (PF)
 515 unit, which is generally dominated by AOM and non-opaque phytoclasts (Fig. 8, Table 5). This
 516 palynofacies is subdivided into two sub-palynofacies units, namely PF-1A and PF-1B, based on the
 517 higher abundances of terrestrial or marine palynomorphs (Table 3). PF-1A is found in the frequent
 518 shale and argillaceous sandstone alternations of the lower Kharita Formation (depths 7300–6450
 519 ft/2225–1966 m), where the shale intervals are rich in organic matter, and is dominated by AOM
 520 (43.4–91, avg. 70.2 %) and subordinate translucent phytoclasts (3.8–47.6, avg. 21.4 %; Fig. 8, Fig.
 521 12a, and Table 5). PF-1A also contains high abundances of terrestrial palynomorphs (90.7–100,
 522 avg. 97.4 %), and very low frequencies of opaque phytoclasts (1.6–10.2, avg. 5 %) and marine
 523 palynomorphs (0.3–9.3, avg. 2.8 %). Palynofacies, TWL petrographic characterization, average HI
 524 (246.6 mg HC/g TOC) and OI (100 mg HC/g TOC) values of the shale intervals indicate a mixture
 525 of kerogen **T**ypes II and III (Baskin, 1997) in the organic-rich suboxic–anoxic deltaic facies of the
 526 lower Kharita. This generally corresponds to the organofacies BC (mixed kerogen **T**ypes II/III) of
 527 Jones (1987) and Tyson (1995), which produces mainly oil and less gas (Tyson, 1995; Baskin,
 528 1997). The low HI values (183.9–258.0, avg. 215.6 mg HC/g TOC) and AOM (60–87.6, avg. 66 %)
 529 and relatively high non-opaque phytoclast frequencies (3.8–33.2, avg. 25.2 %) of the lower PF-1A
 530 shale intervals at depths 7300–6850 ft/2225–2088 m (excepting at 6900 ft/2103.1 m) suggest a
 531 gas-prone organofacies C (kerogen Type III). However, the relatively higher HI values (228.4–
 532 389.3, avg. 282.8 mg HC/g TOC) and AOM (43.4–91, avg. 75.4 %) and the lower non-opaque
 533 phytoclast frequencies (5.4–47.6, avg. 16.6 %) of the shale intervals of the upper PF-1A (depths
 534 6849.9–6450 ft/2088–1966 m) suggest oil/gas-prone organofacies BC (mixed kerogen **T**ypes II/III).
 535 PF-1B is found in the shale and argillaceous sandstone alternations and the carbonate
 536 units of the upper Kharita Formation (depths 6449.9–5600 ft/1965.9–1707 m), which are notably
 537 less rich in organic matter in comparison to PF-1A (Fig. 8). PF-1B shows an increase in the
 538 frequencies of non-opaque phytoclasts (23.6–70, avg. 43.3 %) at the expense of AOM (19.6–69,

539 avg. 46.2 %; Fig. 8, Fig. 12b, and Table 5). PF-1B is relatively enriched in opaque phytoclasts (3–
540 12, avg. 7.4 %) and marine palynomorphs (0.2–16, avg. 6.8 %), and shows a relative decrease in
541 terrestrial palynomorphs (84–98, avg. 93.2 %). PF-1B also spans the stratigraphic interval with
542 frequent occurrences of bulk coaly carbonaceous material in the shale and argillaceous sandstone
543 intervals of the upper Kharita Formation (Fig. 8). Sampled shale and carbonate intervals of PF-1B
544 (excepting the organic-lean sand and carbonate samples at depths 6100 ft/1859.3 m, 6050 ft/1844
545 m, 5800 ft/1767.8 m and 5600 ft/1706.9 m) show comparatively lower HI (125.9–250, avg. 181 mg
546 HC/g TOC) and higher OI (104.8–201.2, avg. 158.5 mg HC/g TOC) values similar to those of PF-
547 1A. This indicates a mixture of kerogen Types II and III (Baskin, 1997) for the relatively less
548 organic, suboxic–anoxic deltaic–shallow marine facies of the upper Kharita Formation.
549 Palynofacies, organic petrographic and geochemical characterizations of the PM of the shale and
550 carbonate intervals of PF-1B show insignificant vertical changes and indicate organofacies C
551 (mixed kerogen Types II/III) of Jones (1987) and Tyson (1995), prone to produce mainly gas and
552 less oil (Tyson, 1995; Baskin, 1997).

553 RWL petrography of the kerogen constituents of PF-1A of the lower Kharita Formation
554 indicates that the vitrinitic maceral group (kerogen Type III) is represented by hydrogen-rich, dark-
555 grey vitrinite and telovitrinite (Figs. 13a, f). The liptinite maceral group (kerogen Type II) is
556 represented by marine unicellular telalginite (Fig. 13b), brown-coloured diffuse bituminite (Fig.
557 13c), bituminite lamellae (Fig. 13d), and minor amounts of cutinite, sporinite and lamalginite. The
558 inertinite maceral group (kerogen Type IV) is represented by rare fusinite (Fig. 13e). Incident UV
559 analysis of the DOM of the the lower Kharita Formation shows weak dull-yellow fluorescence of
560 telalginite (Fig. 14a, c), cutinite (Fig. 14d), and weak fluorescence of bituminite/AOM, visible at the
561 top of Fig. 14c. This maceral composition confirms the mixture of kerogen Types III and II (Tyson,
562 1995).

563 RWL petrography of the kerogen constituents of PF-1B of the upper Kharita Formation
564 shows a maceral composition similar to the lower Kharita Formation, including telovitrinite (Fig.
565 15a), H-rich vitrinite (Fig. 15b), telalginite (Fig. 15c), bituminite lamellae (Fig. 15d), and oval-

566 shaped brown-coloured bituminite (Fig. 15e). Incident UV light shows telalginite having a golden-
567 yellow fluorescence colour (Fig. 15f).

568 The LVI ternary plot of the different PM (kerogen) constituents in the Siqueifa 1X well shows
569 that the samples are rich in gas-prone vitrinite macerals, less rich in oil-prone liptinite macerals,
570 and very poor in inertinite macerals (Fig. 11b, Table 7). TWL petrography shows a palynofacies
571 different from that recorded from the Abu Tunis 1X well, which is dominated by non-opaque
572 phytoclasts and contains subordinate AOM frequencies (Fig. 9, Table 7), here referred to as PF-2
573 and subdivided into two sub-palynofacies - PF-2A and PF-2B - based on the frequency of
574 terrestrial and marine palynomorphs. PF-2A corresponds to the alternations of shale, bulk coaly
575 carbonaceous and argillaceous sandstone horizons, and the argillaceous dolostone interval of the
576 lower Kharita Formation (depths 7400–7000 ft/2255.5–2133.6 m), which are less rich in organic
577 matter in comparison to the Abu Tunis 1x well (Fig. 9). PF-2A is enriched in terrestrial organic
578 matter (non-opaque phytoclasts), and also has lower marine palynomorph frequencies (1.67–4,
579 avg. 2.67 %) when compared to those of PF-1A and PF-1B of Abu Tunis and to those of PF-2B of
580 Siqueifa 1X (figs. 8 and 9, tables 3 and 4). PF-2A is dominated by non-opaque phytoclasts (54.6–
581 63, avg. 59 %) and AOM (32–43, avg. 36 %), with only rare occurrences of opaque phytoclasts (1–
582 2.6, avg. 1.6 %; Fig. 9, Fig. 12c, and Table 7). Samples yielding PF-2A show minor vertical
583 variations in the palynofacies data and organic geochemical parameters (TOC, S₂, HI, OI; Fig. 9,
584 and Table 2) and are thus assigned to a single organofacies unit. The shales and argillaceous
585 dolostones that yield PF-2A show low HI (93–159.2, avg. 128 mg HC/g TOC) and high OI (118.5–
586 257, avg. 176.5 mg HC/g TOC) values, indicating kerogen Type III (Baskin, 1997) in the relatively
587 oxic (dysoxic–anoxic) deltaic shales and argillaceous dolostones of the lower Kharita Formation.
588 Integrated palynofacies, organic petrographic, and geochemical analyses suggest a gas-prone
589 organofacies C (kerogen Type III) of Jones (1987) and Tyson (1995).

590 PF-2B has been isolated from the upper Kharita Formation (depths 6999.9–5850 ft/2133.5–
591 1783 m), composed mainly of alternations of more frequent bulk coaly carbonaceous and/or
592 argillaceous sandstone and less frequent shale intervals (Fig. 9), and is relatively rich in organic
593 matter and PF-2B shows a relative increase in non-opaque phytoclasts (56–75.4, avg. 67.7 %) at

594 the expense of AOM (18.2–39.4, avg. 25.7 %). Minor enrichment in opaque phytoclasts (1.6–5.4,
595 avg. 3.1 %) and marine palynomorphs (3.3–6.3, avg. 4.8 %) is also recorded in PF-2B (Fig. 9, Fig.
596 12d, and tables 4 and 7). Samples containing PF-2B show a minor vertical variation in the
597 palynofacies and organic geochemical data (Fig. 9, and Table 2). The shales containing PF-2B
598 show relatively higher HI (107.3–292.7, avg. 170.5 mg HC/g TOC) and lower OI (97.6–199.1, avg.
599 144.4 mg HC/g TOC) values in comparison to those of PF2A. However, two samples (out of 17)
600 show an HI of 292.7 and 250.7 mg HC/g TOC (i.e., organofacies BC = mixed kerogen Types II/III).
601 These data indicate a dominance of kerogen Type III (Baskin, 1997) in the relatively oxic (dysoxic–
602 anoxic) deltaic shales of the upper Kharita Formation. Shale and argillaceous sandstone samples
603 yielding PF-2B are characterized by the occurrence of some stringers of lignite and frequent bulk
604 coaly carbonaceous material (WEPCO, 1970), which corroborates the presence of gas-prone
605 kerogen Type III material (Mendonça Filho et al., 2012; Singh et al., 2017b). In addition, at least
606 some of the AOM particles in PF-2A can be identified as originating from bacterially degraded non-
607 opaque phytoclasts, as some AOM shows relict structure of phytoclast outlines but with irregular
608 and diffuse edges, and also contain pyrite speckled and inclusions (Fig. 12d). Thus, these
609 terrestrially derived AOM particles can also be identified as Type III kerogen (e.g., Tyson, 1995;
610 Singh et al., 2017b). The combined palynofacies, organic petrography, and geochemistry data of
611 PF-2B indicate a gas-prone kerogen Type III-dominated organic facies for the shale intervals of the
612 upper Kharita Formation.

613 RWL analysis of kerogen constituents of PF-2A shows that the liptinite maceral group is
614 represented by telalginite (Fig. 16a–b) and sporinite (Fig. 16c–d). The vitrinite maceral group is
615 represented by telovitrinite (Fig. 16e) and H-rich vitrinite (Fig. 16f). The kerogen constituents of PF-
616 2B indicate that the liptinite maceral group is composed of *Tasmanites* telalginite (Fig. 17a),
617 sporinite (Fig. 17c, e) and cutinite (Fig. 17f). The vitrinite group consists of telovitrinite (Fig. 17e)
618 and the inertinite maceral group is represented by fusinite (Fig. 17d). Traces of spherical bituminite
619 are also present (Fig. 17b). Reflected UV light analysis shows the presence of sporinite (Fig. 18a,
620 b), which exhibits a yellow fluorescence colour.

621

622 **4.4. Organic maturation**

623 The organic matter of the Kharita Formation in the Abu Tunis 1X well is immature to
624 marginally mature, based on the Tmax values (all are in the 415–426 °C) and confirmed by
625 measured vitrinite reflectance (%Rv is in the 0.48–0.57 range) - although the sample at 7050 ft
626 (2149 m) may have reached the early oil-window stage (Fig. 19). It can therefore be anticipated
627 that intervals of the Kharita Formation present deeper in the Matruh Basin will have higher thermal
628 maturity and have evolved further into the oil-window (Metwalli et al., 2018). These data are in
629 good agreement with the TAI (2 to 2+) obtained from TWL petrography of smooth pteridophyte
630 spores. Examination of the UV fluorescence of the DOM shows orange-yellow and dull-yellow
631 colours of the liptinite macerals (sporinite and lamalginite) in the lower Kharita, to greenish-yellow
632 (lamalginite) in the upper Kharita, which indicates an immature to early upper mature phase
633 (Mukhopadhyay, 1994). The organic matter in the Kharita Formation of the Siqueifa 1X well is
634 slightly more mature than in the Abu Tunis 1X well based on the consistently higher Tmax (421–
635 432 °C) and measured %Rv values (0.48–0.59). The UV excitation of the DOM shows orange-
636 yellow sporinite and lamalginite colours, which indicates peripherally an immature to mainly early
637 upper mature phase (Mukhopadhyay, 1994).

638

639 **4.5. Hydrocarbon source rock evaluation**

640 The deltaic deposits of the lower Kharita Formation in the Abu Tunis 1X well contain considerable
641 organic-rich intraformational shale intervals, which comprise a total content of about 13.2 % and
642 range in thickness from 1.5 to 9.8 m, comprise about 34 m net source rock, and show remarkable
643 intraformational variations in organic richness, which range from good to very good/excellent.
644 Combining these factors indicates that the lower Kharita Formation should be considered as a
645 significant hydrocarbon source with very good to excellent hydrocarbon potential (Fig. 20a). This
646 interpretation is consistent with several hydrocarbon exploration studies, where organic-rich
647 intraformational intervals of similar net source thickness were deemed to have important
648 hydrocarbon source rock potential (e.g., Katz et al., 2000; Hakimi and Abdullah, 2014). Shale
649 intervals in the lower part of the lower Kharita Formation are particularly rich in terrestrially sourced

650 organic matter (i.e., gas-prone), while those in the upper part are rich in marine and terrestrially
651 sourced organic matter (i.e., oil and gas-prone). However, the immature to early upper thermal
652 maturity and the oil-cross over (Fig. 20b) indicate a low probability for the lower Kharita shales to
653 produce oil (avg. generated oil in rock from pyrolysis is 81.8 bbl oil/ac-ft; Table 1) with no gas
654 production (Fig. 8). The intraformational argillaceous sandstones are poor in organic matter and
655 regarded as reservoir intervals. In contrast, the deltaic-shallow marine deposits of the upper
656 Kharita Formation contain fewer organic-rich intraformational shale and carbonate intervals, with
657 the former (between 3.7 and 12.2 m thick) comprising a net source rock interval of about 9.8 %
658 (25.3 m), whereas limestone and dolostone intervals (6.7 to 39.6 m thick) comprise a greater net
659 source rock interval of about 23.4 % (60.6 m). The shale and carbonate intervals collectively
660 comprise in total a total net source rock interval of about 33.2 % (86 m) with little intraformational
661 variation in organic richness. The thicker organic-rich intervals show low potential as oil source
662 rocks but fair to good potential as gas sources due to enrichment of the gas-prone coaly
663 carbonaceous material, which could represent another important hydrocarbon source. However,
664 these deposits are still not sufficiently thermally mature to produce appreciable quantities of either
665 oil or gas. The argillaceous and/or carbonaceous sandstones of the upper Kharita Formation have
666 very low organic matter content and thus are defined as reservoir horizons.

667 Lateral intrabasinal variation in lithofacies, organic facies, and organic richness can be
668 demonstrated in the Kharita Formation, as on the eastern margin of the Matruh Basin in the Siqefia
669 1X area, there is a higher average shale content (34.2 %) than in the Abu Tunis 1X well (23 %)
670 according to WEPCO (1968, 1970). However, the deltaic shales of the lower Kharita Formation of
671 Siqefia 1X are less organic-rich (23.7 %, 29 m thick net source rock) as is the shaley dolostone (7
672 %, 8.5 m thick). However, these Siqefia 1X intervals are rich in terrestrially sourced organic matter
673 and show little intraformational variation in organic richness, and have been shown to have fair to
674 good gas-prone source rock potential (Fig. 20a). The argillaceous and carbonaceous sandstones
675 of the lower Kharita Formation are organic-lean and are considered reservoir intervals. Vertical
676 lithofacies and organic richness can be seen in the deltaic upper Kharita Formation as shale
677 intervals become less frequent (12.1 %) and their thicknesses reduce (2.7–6.7 m). Nevertheless,

678 these upper Kharita shales show relatively high organic richness due to deposition of coals and
679 coaly carbonaceous material, but only have a fair to good to less very good gas-prone source rock
680 potential. This lateral diminishing in the kerogen quality and richness of the Kharita Formation
681 between Abu Tunis 1X and Siqueifa 1X results from the latter being deposited in a relatively
682 oxygenated deltaic environment, where significant preservation of AOM and deposition of coal is
683 less likely (Tyson, 1995; Mendonça Filho, et al., 2012). The immature to early upper stage of
684 thermal maturity and the oil cross-over (Fig. 20b) indicates a very low probability for the shale
685 intervals of the entire Kharita Formation in the Siqueifa 1X to produce oil (avg. generated oil in rock
686 from pyrolysis is only 11.76 bbl oil/ac-ft) and would produce no gas (Table 2). The argillaceous
687 sandstone and dolostone deposits of the upper Kharita Formation have reservoir characteristics.

688 Against this, recent seismic structural analysis and basin modelling of the sedimentary
689 sequence in the Matruh Basin revealed more mature potential gas-prone source rocks of the
690 Kharita Formation to the southeast of the Abu Tunis 1X and Siqueifa 1X wells (Metwalli et al., 2018).
691 The latter authors constructed a maturity map and a thermal history model of the Kharita
692 Formation, based on measured well-bottom temperatures to the SE of our study wells. In this
693 maturity map, the Abu Tunis 1X and Siqueifa 1X wells showed relatively higher maturity levels (early
694 mature) of 0.60-0.70 %Rv in comparison to our current determination 0.48 and 0.57 %. Despite
695 this difference in measured and calculated values, the general increase in maturity in an eastward
696 direction consistent with maturation we record at Siqueifa 1X. The Kharita Formation in the Mideiwar
697 1X, Matruh 1-1, Matruh 2-1, and Matruh 3-1 wells in Matruh Basin show shale contents of 21.7 %,
698 22.1 %, 21.4 %, and 23.2 % similar to that recoded from the Abu Tunis 1X (23 %), but with a
699 relatively higher maturity level (Metwalli et al., 2018; see Fig. 21). The burial and thermal history
700 modelling of Metwalli et al. (2018) on the Kharita Formation indicates the presence of more mature
701 (late mature, oil-window and main gas generation, 1–1.3 %Rv) potential gas-producing zones to
702 the east and southeast of our studied area of the basin. This is specifically shown in the Darduma
703 1A, Matruh 1-1 and Ras Kanayes wells (Fig. 21).

704

705 **5. Conclusions**

706 Previously the Albian Kharita Formation has been understood to possess very little
707 hydrocarbon source rock potential in the Shushan and Sidi Barani basins, and was effectively
708 viewed as only a hydrocarbon reservoir unit. However, high organic matter influxes and reducing
709 conditions at the more offshore/deeper setting of the Abu Tunis 1X well in the Matruh Basin
710 resulted in better preservation of organic-rich, intraformational shale and carbonate deposits. In
711 this paper we have demonstrated that in the Abu Tunis 1X well the lower Kharita Formation is a
712 good to very good/excellent potential hydrocarbon source rock, although in the less reducing
713 conditions that prevailed at the eastern margin of Matruh Basin at the Siqueifa 1X well produced
714 lower organic content and lower source rock potential.

715 Our integrated organic geochemical, organic petrographic, and palynofacies analyses
716 indicate that the deltaic–shallow marine shales and minor carbonates of the Kharita Formation in
717 the Abu Tunis 1X well comprise significant net source rocks of about 120 m thickness. The lower
718 Kharita Formation in particular contains 34 m net shale source rock and shows good to very
719 good/excellent gas and oil/gas potential (1.14–11.59 TOC wt %; S_2 : 2.72–23.53 mg HC/g). The
720 shales and carbonates of the upper Kharita also comprise another important hydrocarbon source
721 with a thicker net source rock interval of 86 m, but they are less organic-rich and show only fair to
722 good) gas/oil source rock potential (0.82–1.78 TOC wt %; S_2 : 1.07–4.02 mg HC/g). Although the
723 entire Kharita Formation in the Abu Tunis area contains important source rock intervals, in terms of
724 thermal maturity these intervals are immature to marginally mature. In the Siqueifa 1X well, the
725 deltaic intraformational shale and shaley dolostone deposits are less organic-rich, but comprise a
726 considerable net source rock of about 80 m thickness, with an overall fair to good to less very good
727 gas source rock potential (0.81–2.14 TOC wt %; S_2 : 0.88–4.80 mg HC/g), and ranges from
728 immature to marginally mature. The frequent shale and minor carbonate intervals of the Kharita
729 Formation at the depocentre of the Matruh Basin (i.e., Abu Tunis 1X) are very rich in gas/oil-prone
730 organic matter, while the shale and minor shaley dolostone intervals of the Kharita Formation at
731 the basin margin (e.g., Siqueifa 1X) are rich in gas-prone organic matter. It has been shown that the
732 lithofacies, organic facies/richness and organic quality vary laterally across the basin, and therefore
733 so does the hydrocarbon potential of the Kharita Formation.

734 Previously often ignored, we have demonstrated that the shales within the Kharita
735 Formation possess suitable organic richness and source rock potential, and therefore in the future
736 this unit will be of more interest in terms of hydrocarbon prospectivity because of its regional
737 subsurface extent across northwestern Egypt (Tahoun et al., 2017). Although the Kharita
738 Formation at both of our study sites is too thermally immature to have produced oil or gas, this
739 formation may contain more mature organic matter that has reached the late mature oil- to main
740 gas-generation window in the more eastern and southeastern parts of the Matruh Basin where
741 higher thermal maturities have been modelled (Metwalli et al., 2018), and is thus a prime
742 exploration target.

743

744 **Figure captions**

745

746 Fig. 1: Map showing the location of the Matruh Basin, the Abu Tunis 1X and Sikeifa 1X wells, and
747 a simple structural map of the basin (modified after EGPC, 1992; Shalaby et al., 2012).

748

749 Fig. 2: Generalized lithostratigraphic column, showing chronostratigraphy, depositional
750 environments, and the petroleum system of the sedimentary sequences in the Matruh Basin, north
751 Western Desert, Egypt (after Aram et al., 1988; EGPC, 1992).

752

753 Fig. 3: Quantitative distribution of palynomorphs and particulate organic matter (in grains/gram),
754 showing six fine-grained clastic organic-rich intervals within the Kharita Formation (after Deaf,
755 2009).

756

757 Fig. 4: Scheme of the protocol employed in this paper for the analysis of hydrocarbon potential.

758

759 Fig. 5 (A) Plot of the total organic carbon content (TOC) against depth, showing the organic
760 richness of the Kharita Formation in the Abu Tunis 1X and Sikeifa 1X wells. (B) Plot of the oil
761 potential versus S_2 , showing the hydrocarbon potential of the Kharita Formation in the Abu Tunis
762 1X and Sikeifa 1X wells. (C) Plot of S_2 against TOC showing the hydrocarbon source rock potential
763 of the Kharita Formation in the Abu Tunis 1X and Sikeifa 1X wells.

764

765 Fig. 6: Spores–Microplankton–Pollen (SMP) ternary plots showing depositional environments of
766 the Kharita Formation (after Federova, 1977; Düringer and Doubinger, 1985). (A) SMP plot of the
767 Abu Tunis 1X well samples. (B) SMP plot of the Sikeifa 1X well samples.

768

769 Fig. 7: AOM–Phytoclasts–Palynomorphs (APP) ternary plots showing different palynofacies types
770 and redox conditions for the Kharita Formation (after Tyson, 1995). (A) APP plot of the Abu Tunis
771 1X well samples. (B) APP plot of the Siqueifa 1X well samples.

772
773 Fig. 8: Chart showing lithologic column, organic petrographic and geochemical data, and the
774 hydrocarbon potential of the Kharita Formation in the Abu Tunis 1X well.

775
776 Fig. 9: Chart showing lithologic column, vertical stratigraphic distributions of palynofacies
777 parameters, suggested depositional environments, organic petrographic and geochemical data,
778 and the hydrocarbon potential of the Kharita Formation in the Siqueifa 1X well.

779
780 Fig. 10: Pseudo-Van Krevelen diagram showing kerogen types in the Kharita Formation in the Abu
781 Tunis 1X and Siqueifa 1X wells.

782
783 Fig. 11: Liptinite–Vitrinite–Inertinite (LVI) ternary kerogen plots (after Dow, 1982; Tyson, 1995)
784 showing the hydrocarbon source rock potential of the Kharita Formation. (A) LVI plot of the Abu
785 Tunis 1X well samples, (B) LVI plot of the Siqueifa 1X well samples.

786
787 Fig. 12: Representative palynofacies assemblages from the Kharita Formation. Key to labels: Sp =
788 spore, Sp-Pyr = spore filled with pyrite, N-OP = non-opaque phytoclast, De N-OP = degraded non-
789 opaque phytoclast, OP = opaque phytoclast, AOM (De N-OP) = AOM of probably degraded N-OP
790 phytoclast origin, MbT = membranous tissue. (A) PF-1A (AOM/phytoclast-dominated) with
791 abundant spores from the lower Kharita Formation in the Abu Tunis 1X well, sample depth (7000a
792 ft/2133.6 m) at x250 magnification. (B) PF-1B (AOM/phytoclast-dominated) with abundant black
793 wood from the upper Kharita Formation in the Abu Tunis 1X well, sample depth (6300b ft/1920.2
794 m) at x250 magnification. (C) PF-2A (Phytoclast/AOM-dominated) with abundant spores from the
795 lower Kharita Formation in the Siqueifa 1X well, sample depth (7350a ft/2240.3 m) at x250
796 magnification. (D) PF-2B (Phytoclast/AOM-dominated) with abundant spores from the upper
797 Kharita Formation in the Siqueifa 1X well, sample depth (5850a ft/1783 m) at x250 magnification.

798
799 Fig. 13: Photomicrographs of dispersed organic matter (DOM) in PFA-1, Abu Tunis 1X well, taken
800 under reflected white light. (a) Hydrogen-rich vitrinite, depth 7000 ft (2133.6 m); (b) Telalginite,
801 depth 7000 ft; (c) Light-brown bituminite, depth 7000 ft; (d) Bituminite lamellae, depth 7000 ft; (e)
802 Fusinite with compressed cell lumens, depth 5750 ft (1752.6 m); and (f) Telovitrinite, depth 6600 ft
803 (2011.7 m).

804
805 Fig. 14: Photomicrographs of dispersed organic matter (DOM) in the Abu Tunis 1X well (6900
806 ft/2103 m) taken under reflected UV light. Excitation is at 465 nm; combined dichroic mirror and

807 barrier filter have a long pass at 515 nm. (a) Telalginite; (b) Sporinite; (c) Telalginite; and (d)
808 Cutinite.

809

810 Fig. 15: Photomicrographs of dispersed organic matter (DOM) in PFA-2, Abu Tunis 1X well, taken
811 under reflected white light. (a) Telovitrinite, depth 6400 ft; (b) Hydrogen-rich vitrinite, depth 5750 ft
812 (1752.6 m); (c) Telalginite being converted to bituminite, depth 5850 ft (1783 m); (d) Bituminite
813 lamellae, depth 6400 ft (1950.7 m); (e) Bituminite depth 6250 ft (1905 m); and (f) Lamalginite
814 depth 5850 ft (1783 m). Photomicrograph (e) was taken under reflected UV light using the same
815 filters as in Fig. 13.

816

817 Fig. 16: Photomicrographs of dispersed organic matter (DOM) in PF2-A, Siqeifa 1X well, taken
818 under reflected white light. (a) *Tasmanites* telalginite, depth 6450 ft (1966 m); (b) Telalginite, depth
819 6450 ft; (c) Sporinite showing the characteristic trilete mark, depth 7000 ft (2133.6 m); (d)
820 Sporinite, depth 7000 ft; (e) Telovitrinite, depth 7300 ft (2225 m); and (f) H-rich vitrinite, depth 7100
821 ft (2164 m).

822

823 Fig. 17: Photomicrographs of dispersed organic matter (DOM) in PF2-B, Siqeifa 1X well, taken
824 under reflected white light. (a) Telalginite, depth 6300 ft (1920 m); (b) Spherical bituminite, depth
825 6300 ft; (c) Sporinite (pollen grain showing two possible elater-like appendages), depth 5850 ft
826 (1783 m); (d) Inertinite, depth 5850 ft (1783 m); (e) Telovitrinite enclosing compressed sporinite,
827 depth 6150 ft (1874.5 m); and (f) Cutinite in transverse section showing the characteristic 'teeth'
828 (arrows), depth 6150 ft.

829

830 Fig. 18: Photomicrographs of dispersed organic matter (DOM) in the Siqeifa 1X well taken under
831 reflected UV light from a depth of 5700 ft (1737 m). (a) Sporinite (pollen grain showing at least two
832 elater-like appendages) arrows point to the two appendages (Ap); (b) Sporinite (spore).

833

834 Fig. 19: Plot of HI versus Tmax showing kerogen types and thermal maturation levels for the
835 Kharita Formation in the Abu Tunis 1X and Siqeifa 1X wells.

836

837 Fig. 20: (A) Kerogen quality diagram showing potential kerogen types of the Kharita Formation in
838 the Abu Tunis 1X and Siqeifa 1X wells. (B) Oil cross-over diagram showing a low probability of
839 both the Abu Tunis 1X and Siqeifa 1X wells to generate oil.

840

841 Fig. 21: Map showing the different maturity levels at the surface of the Kharita Formation across
842 Matruh Basin (modified after Metwalli et al., 2018), the recorded maturation of the Abu Tunis 1X
843 and Siqeifa 1X wells, and the potential promising hydrocarbon wells in the basin. Arrow points to
844 the potential promising mature source rocks in the basin.

845

846 **Disclosure of interest**

847 The authors declare that they have no competing interest.

848
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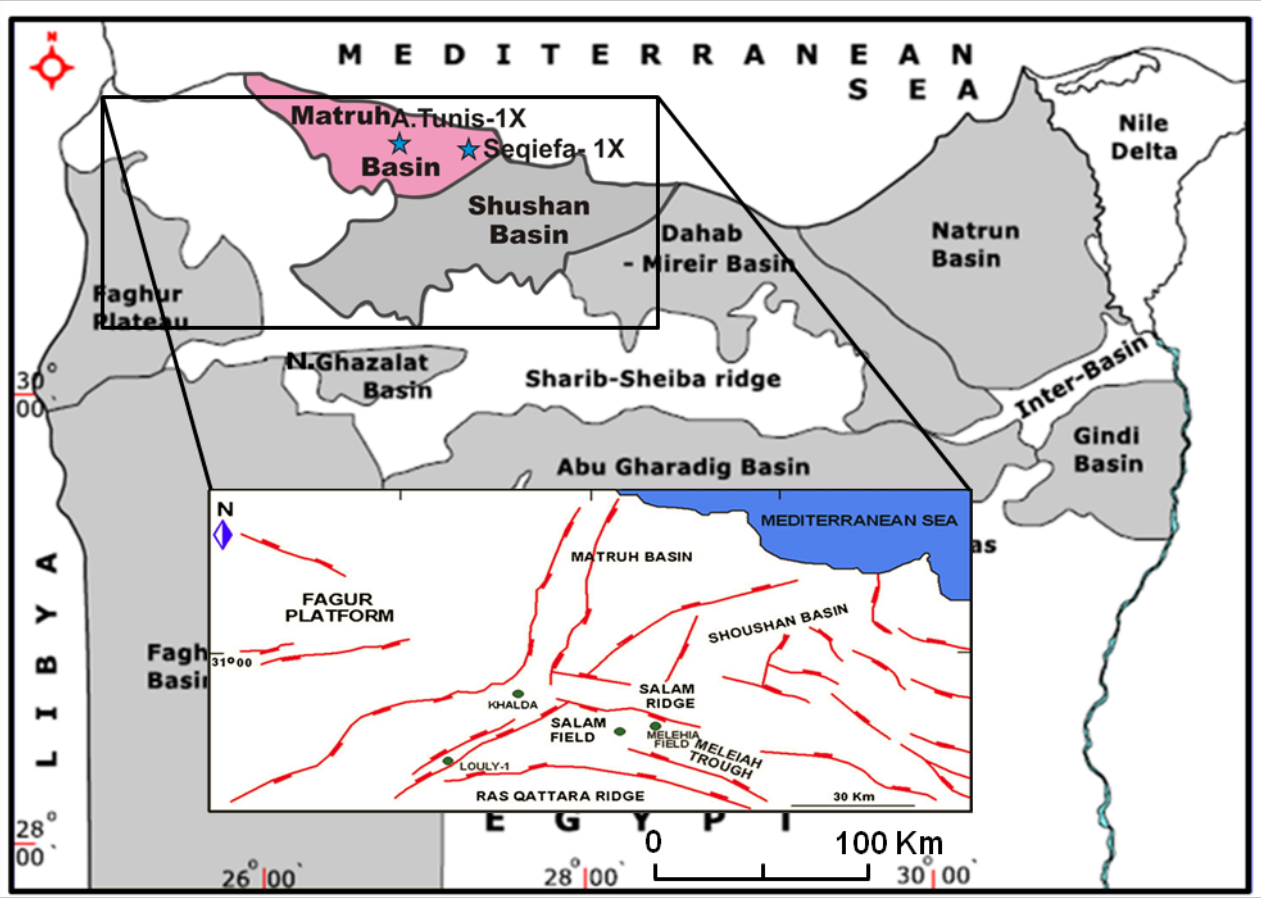
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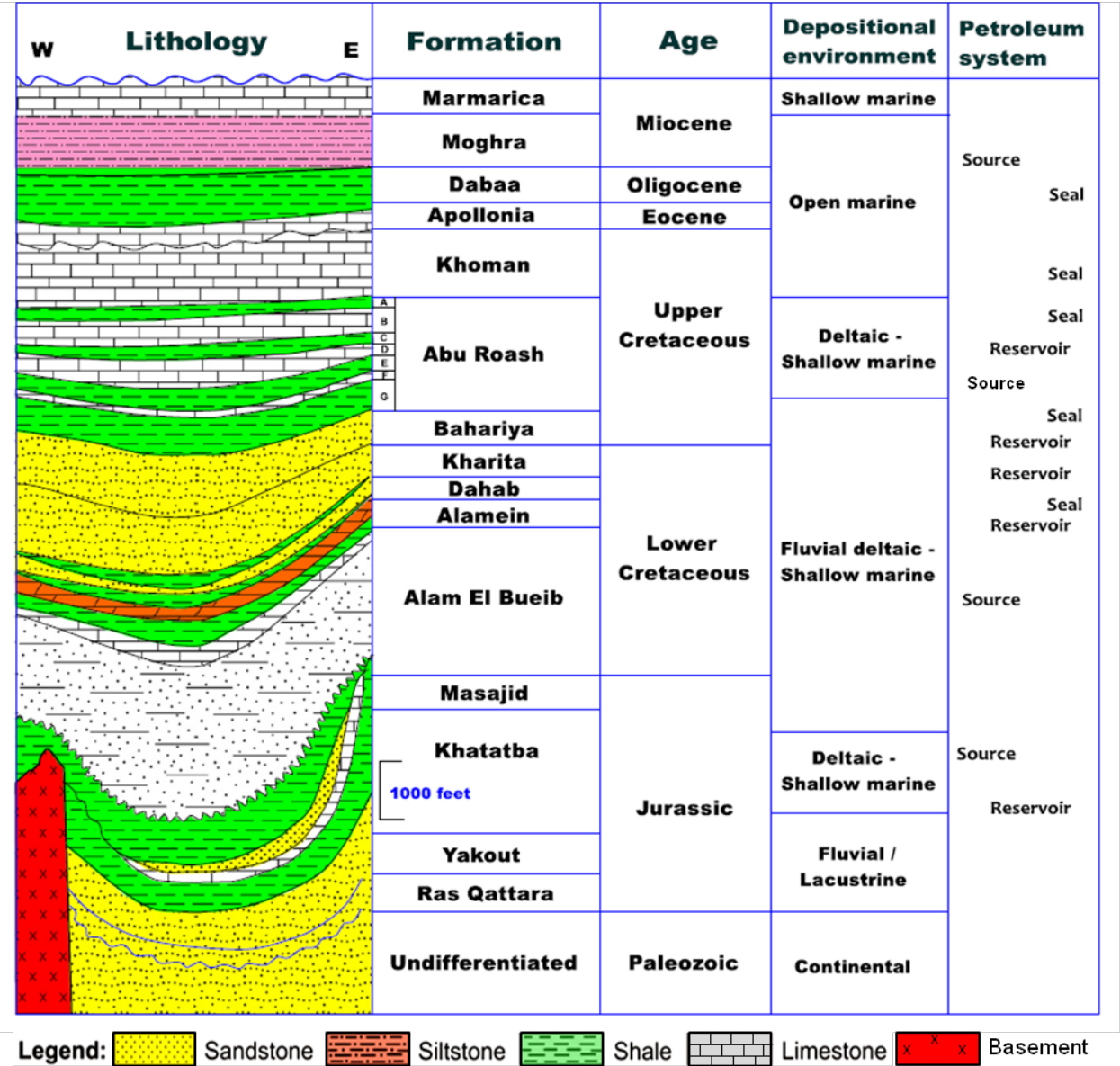
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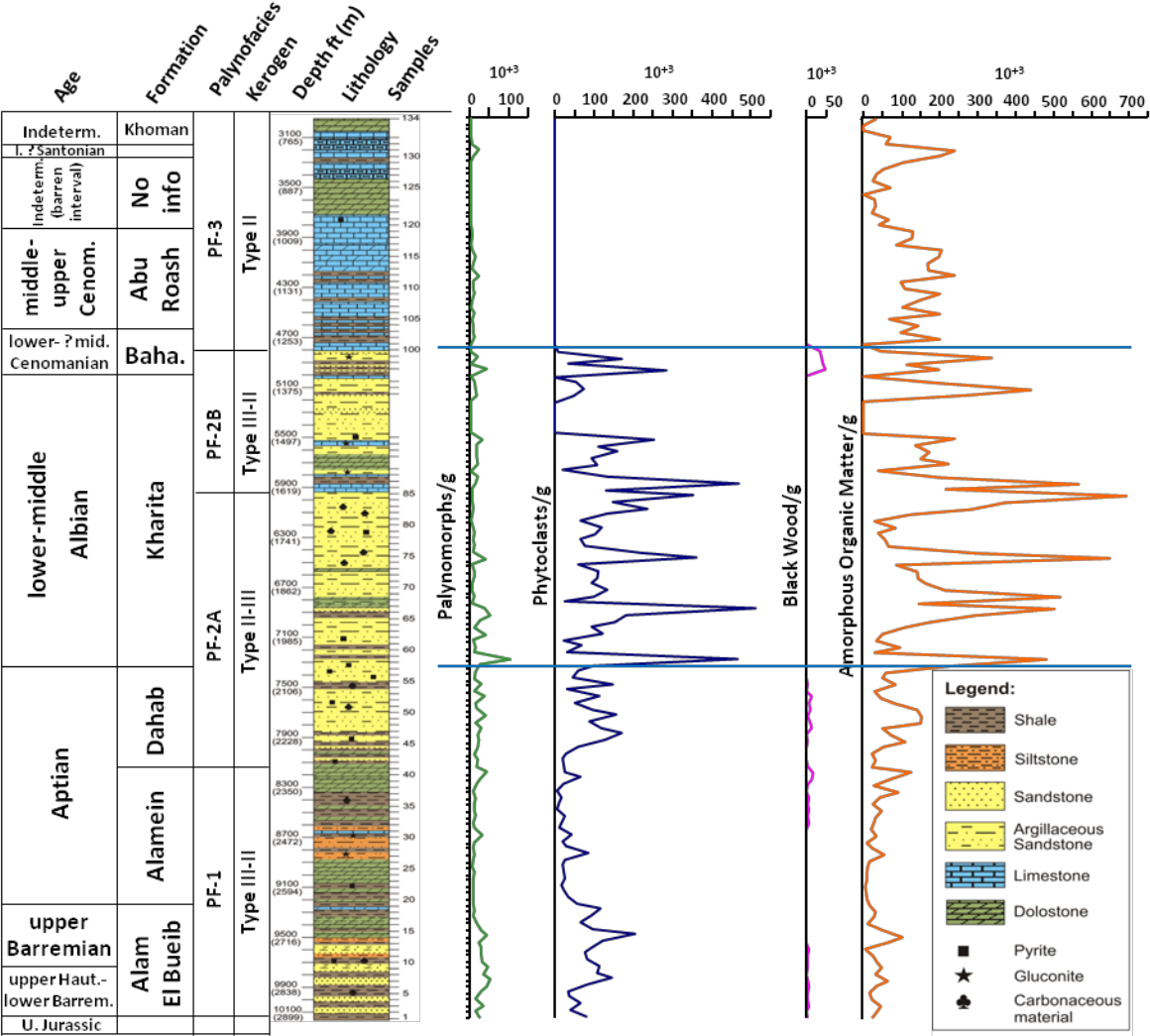
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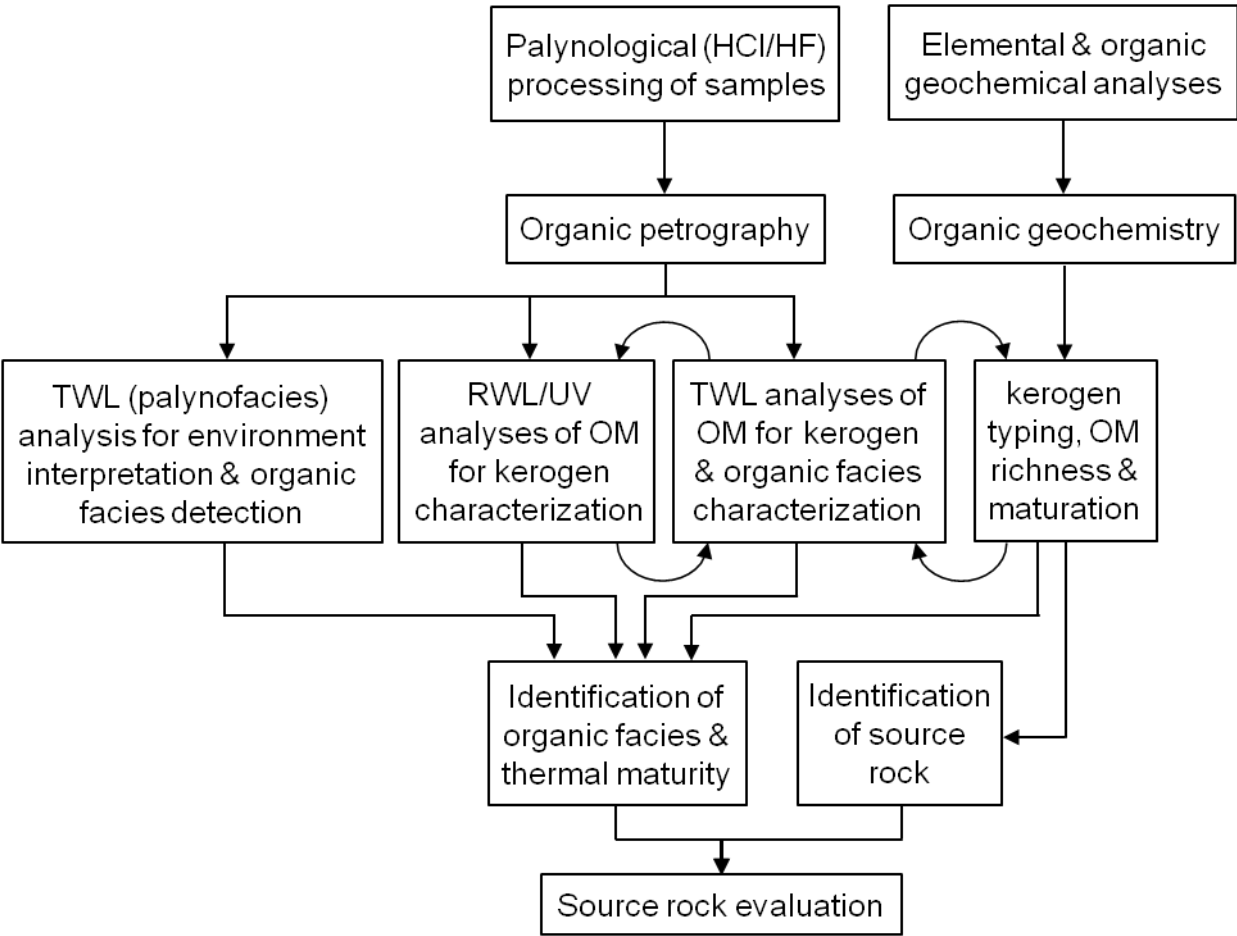
1227 **Figure 1**

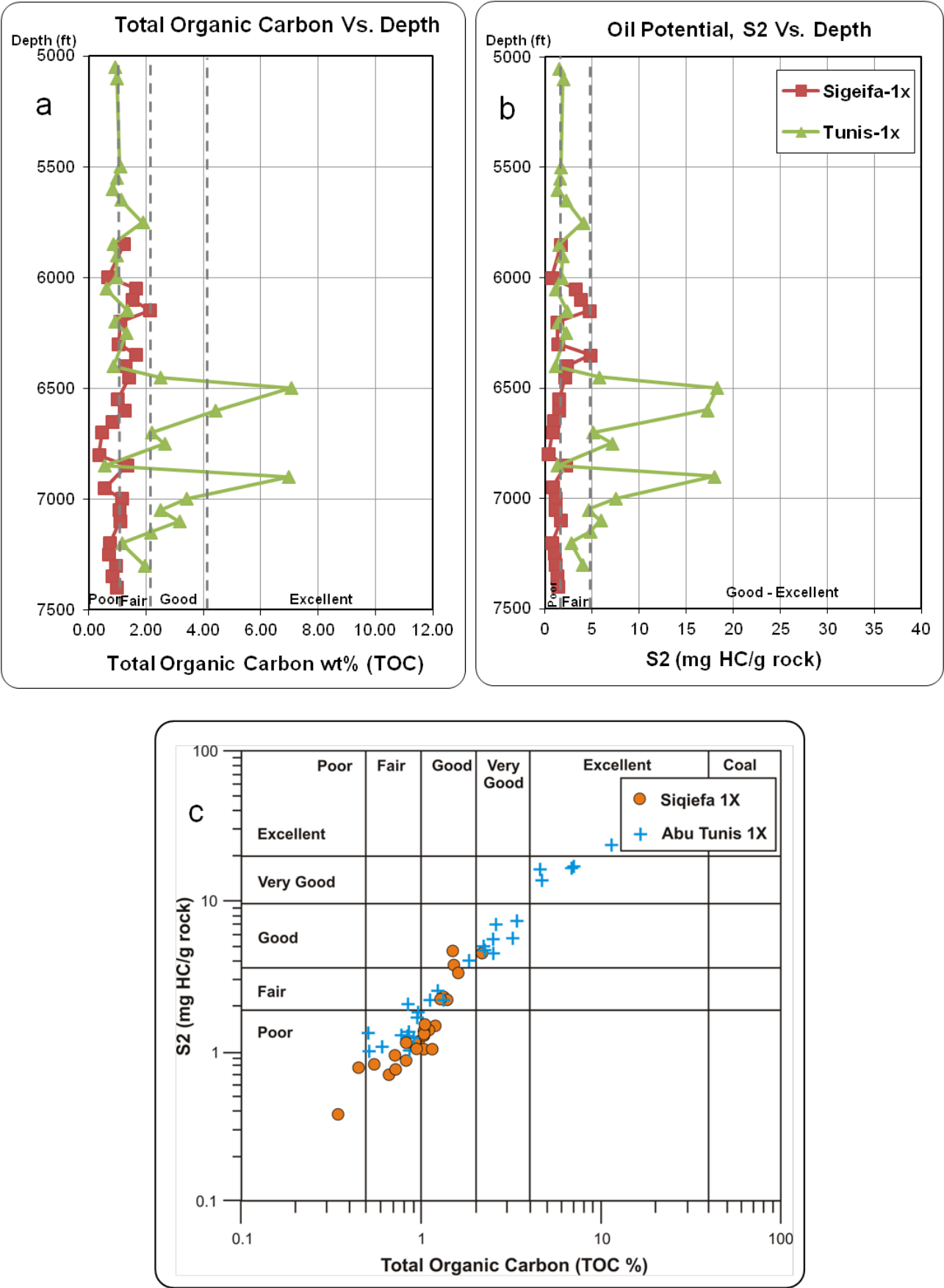


1228 **Figure 2**

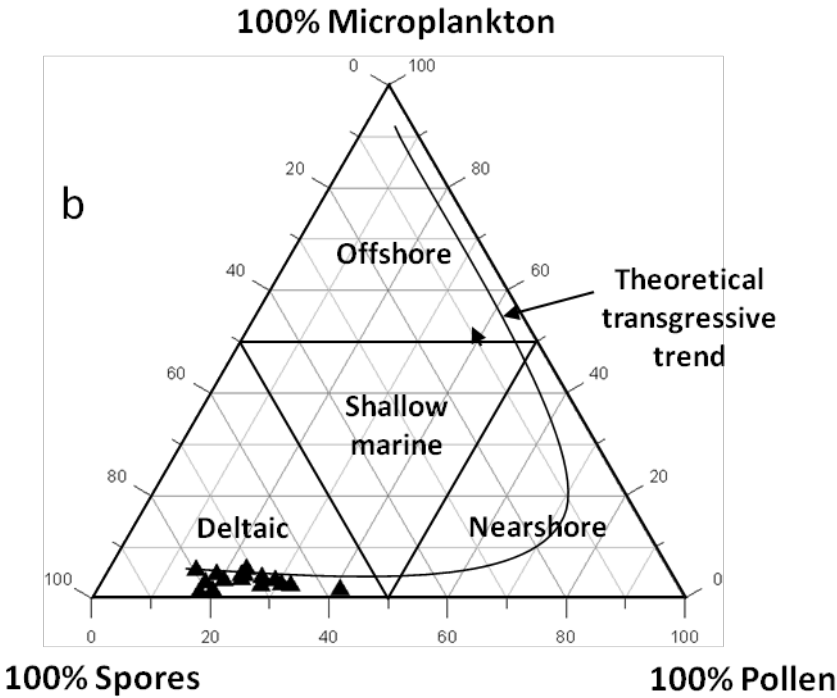
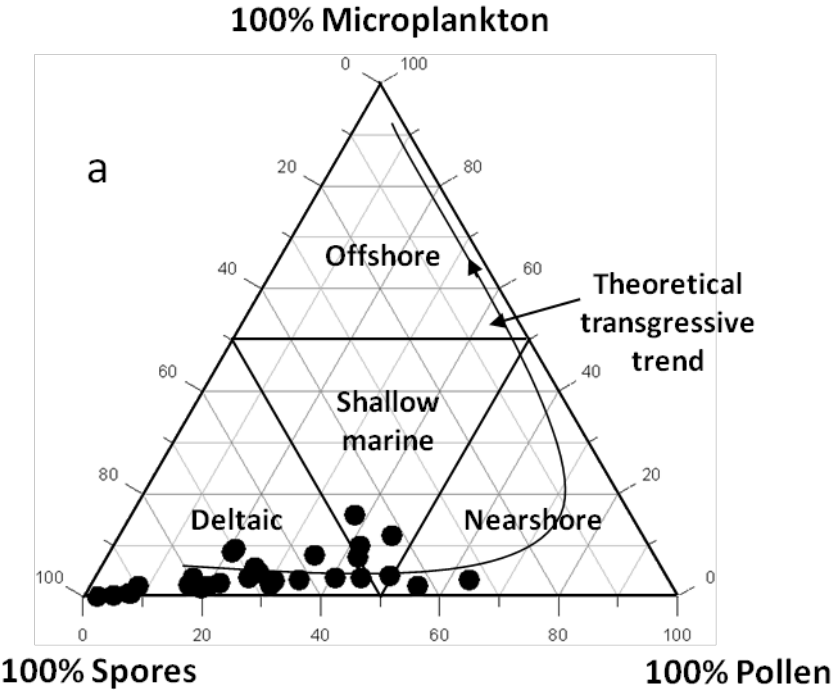








1232 **Figure 6**



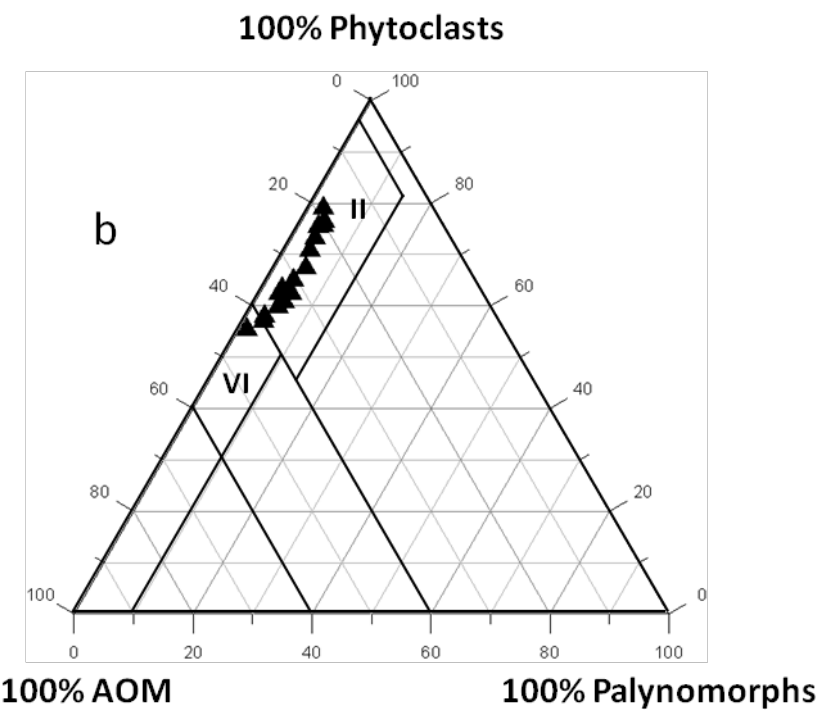
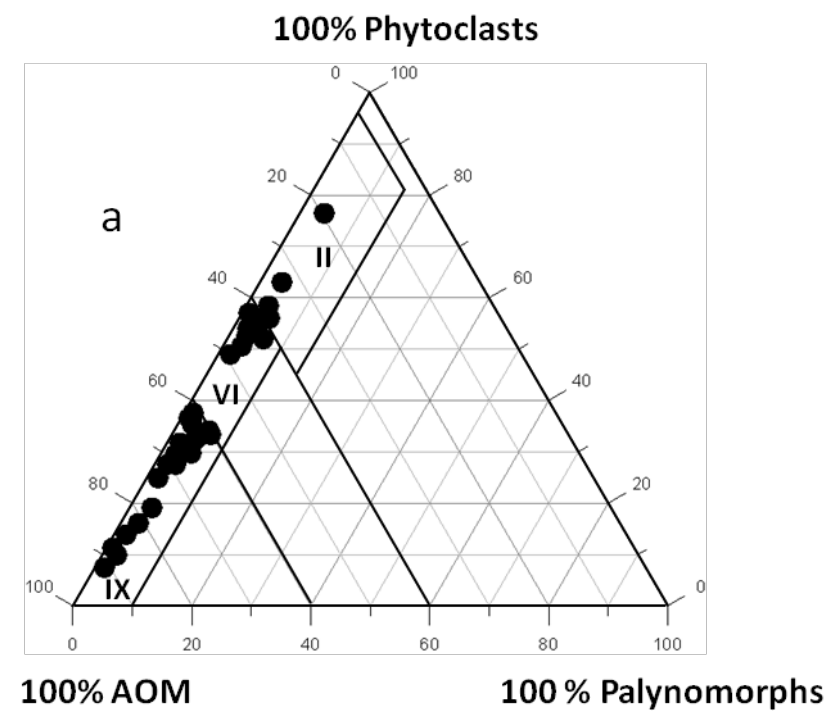


Figure 8

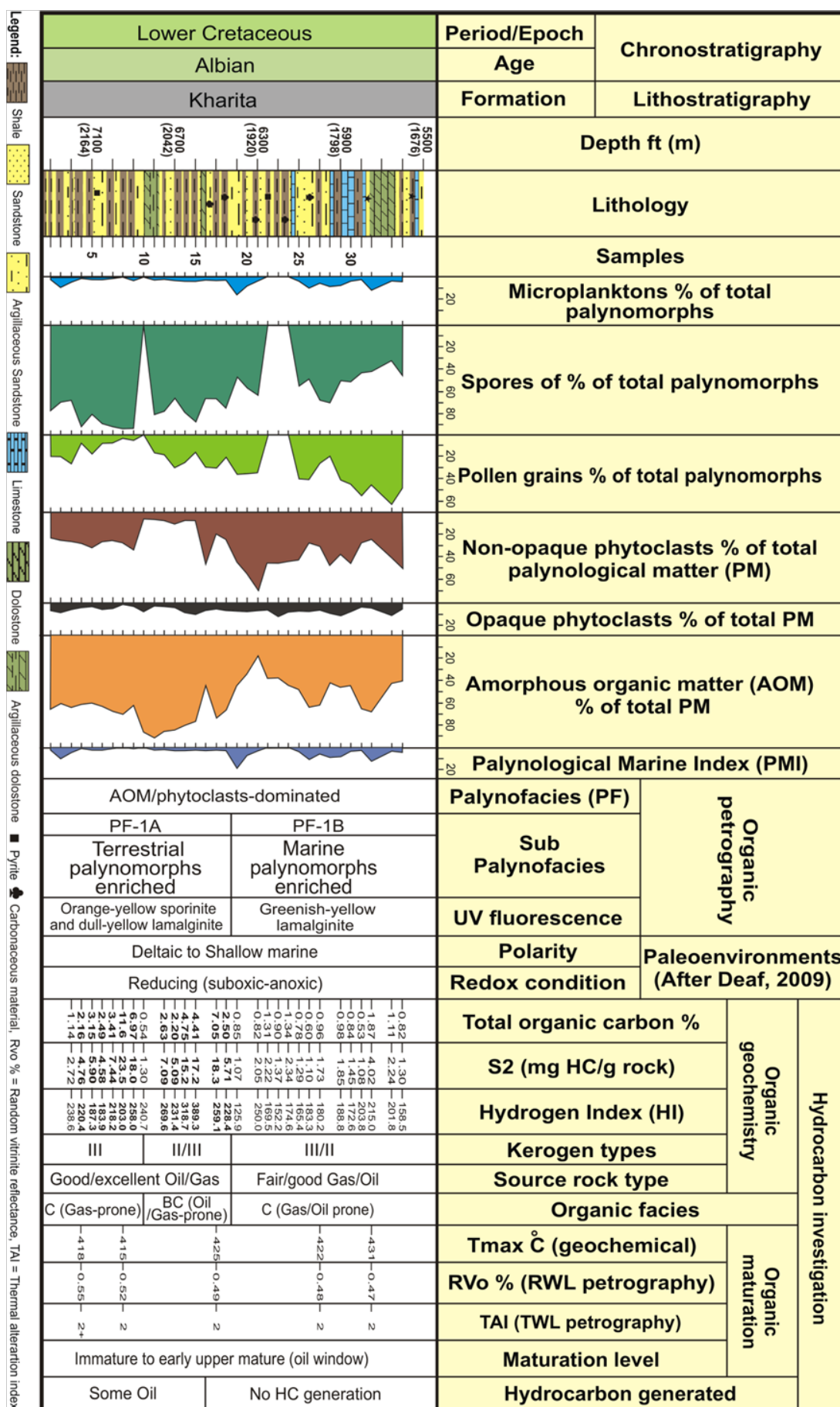
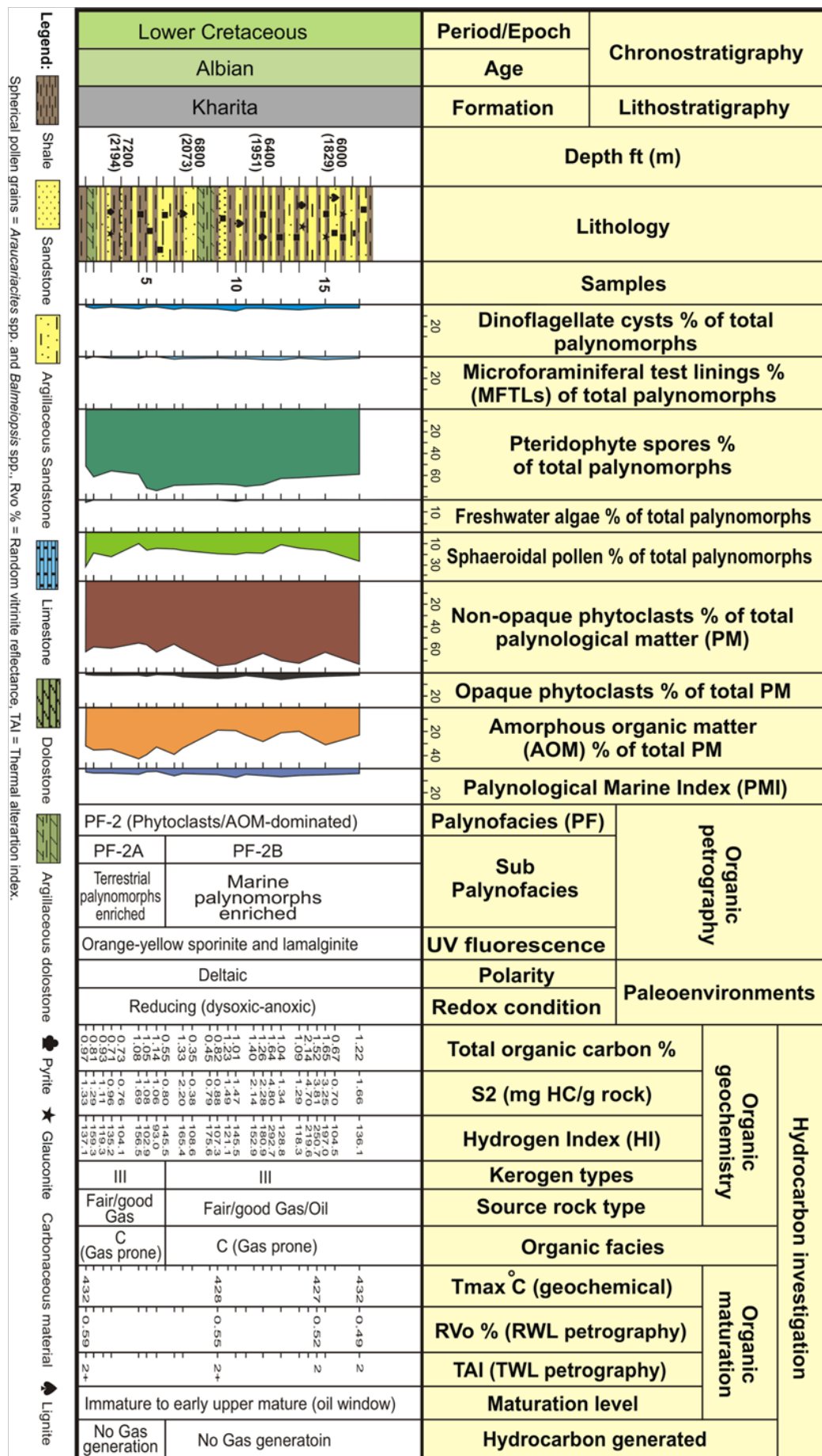
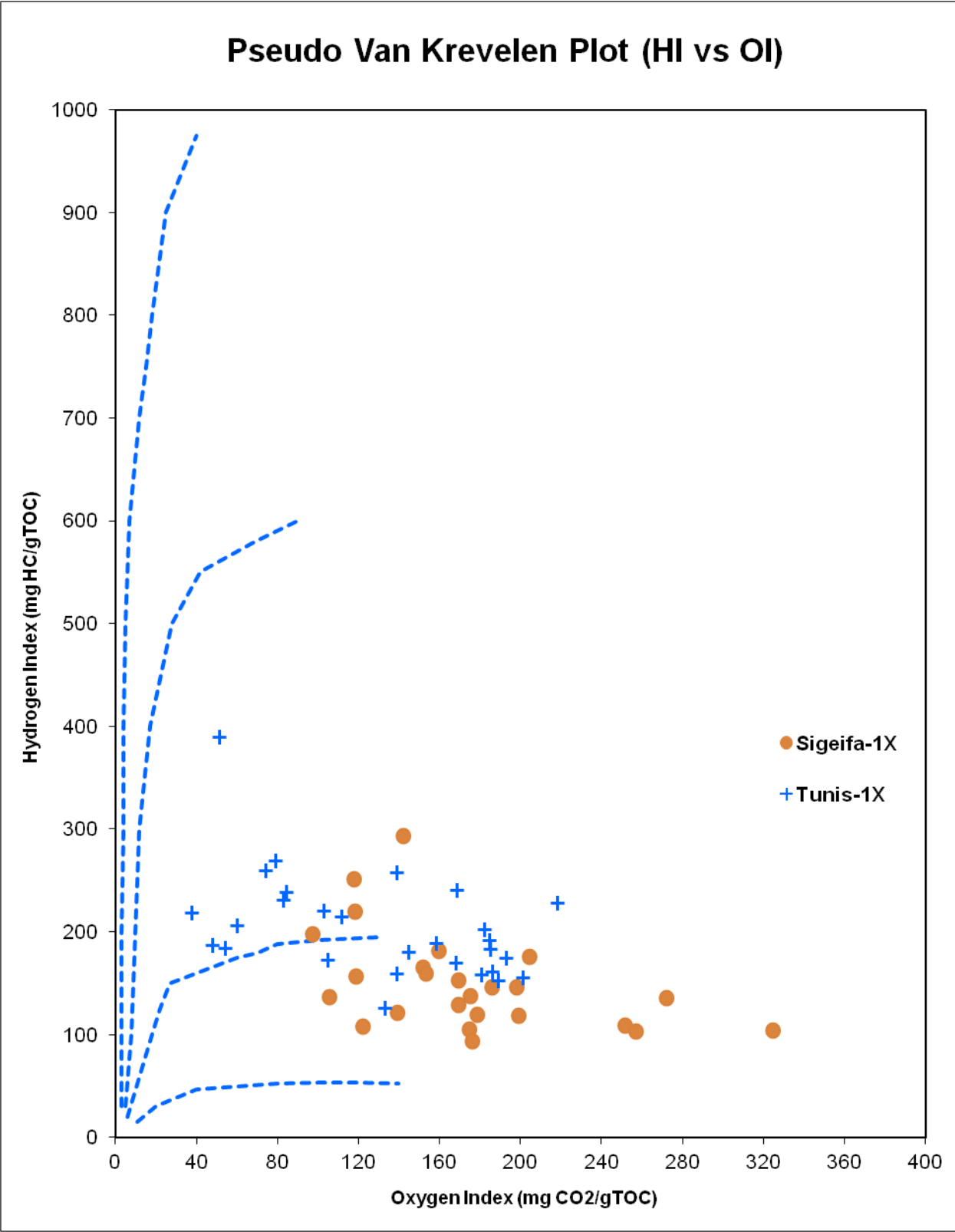
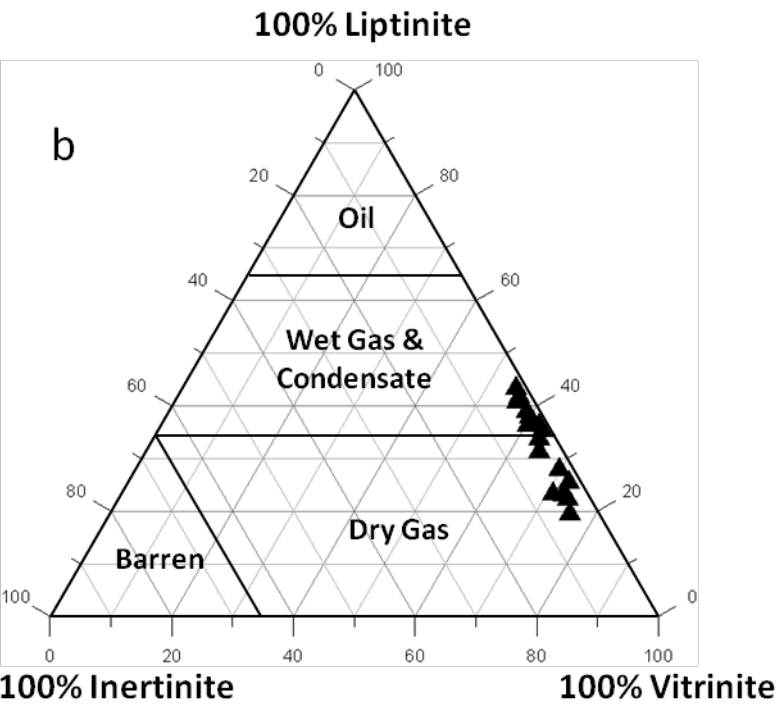
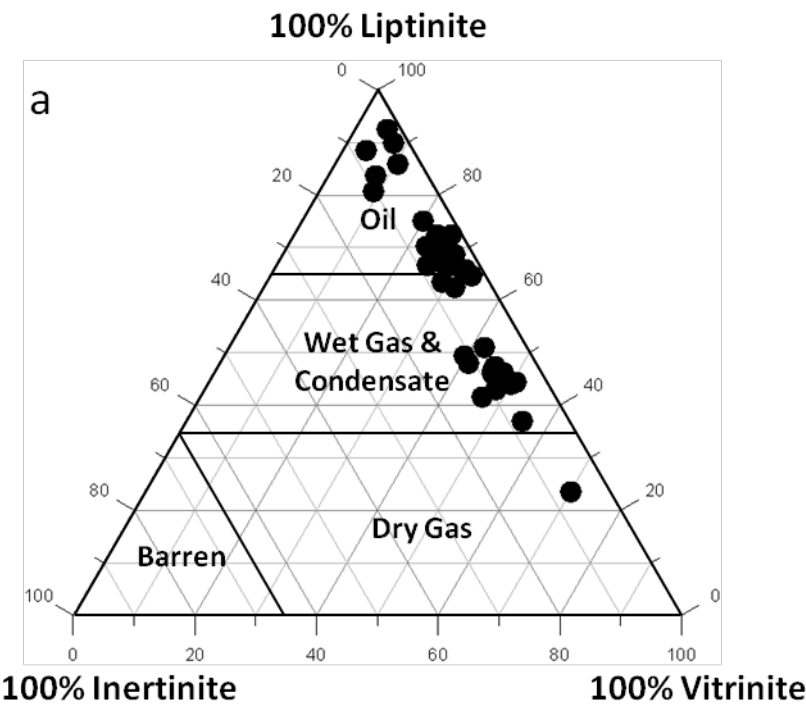


Figure 9







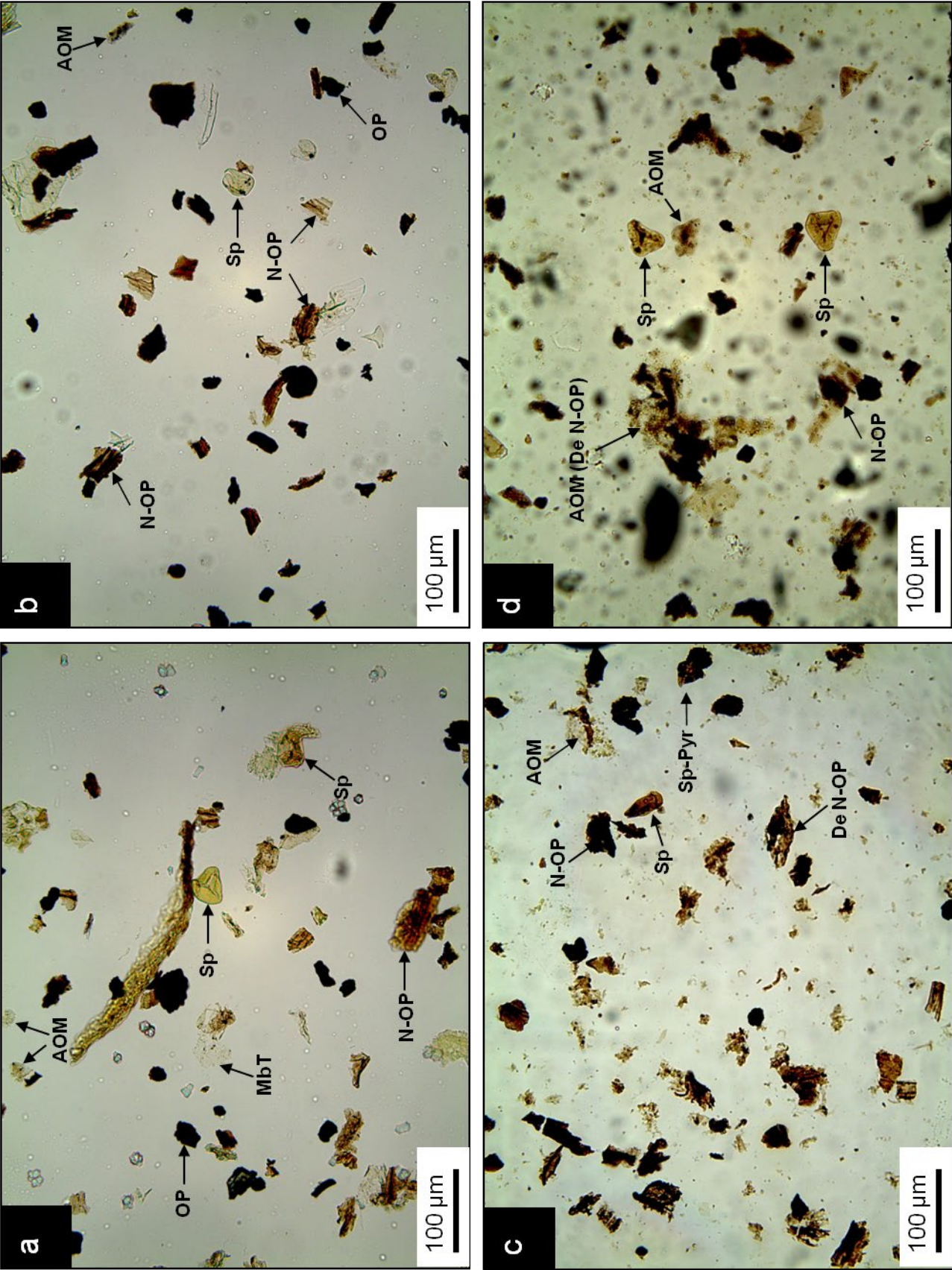
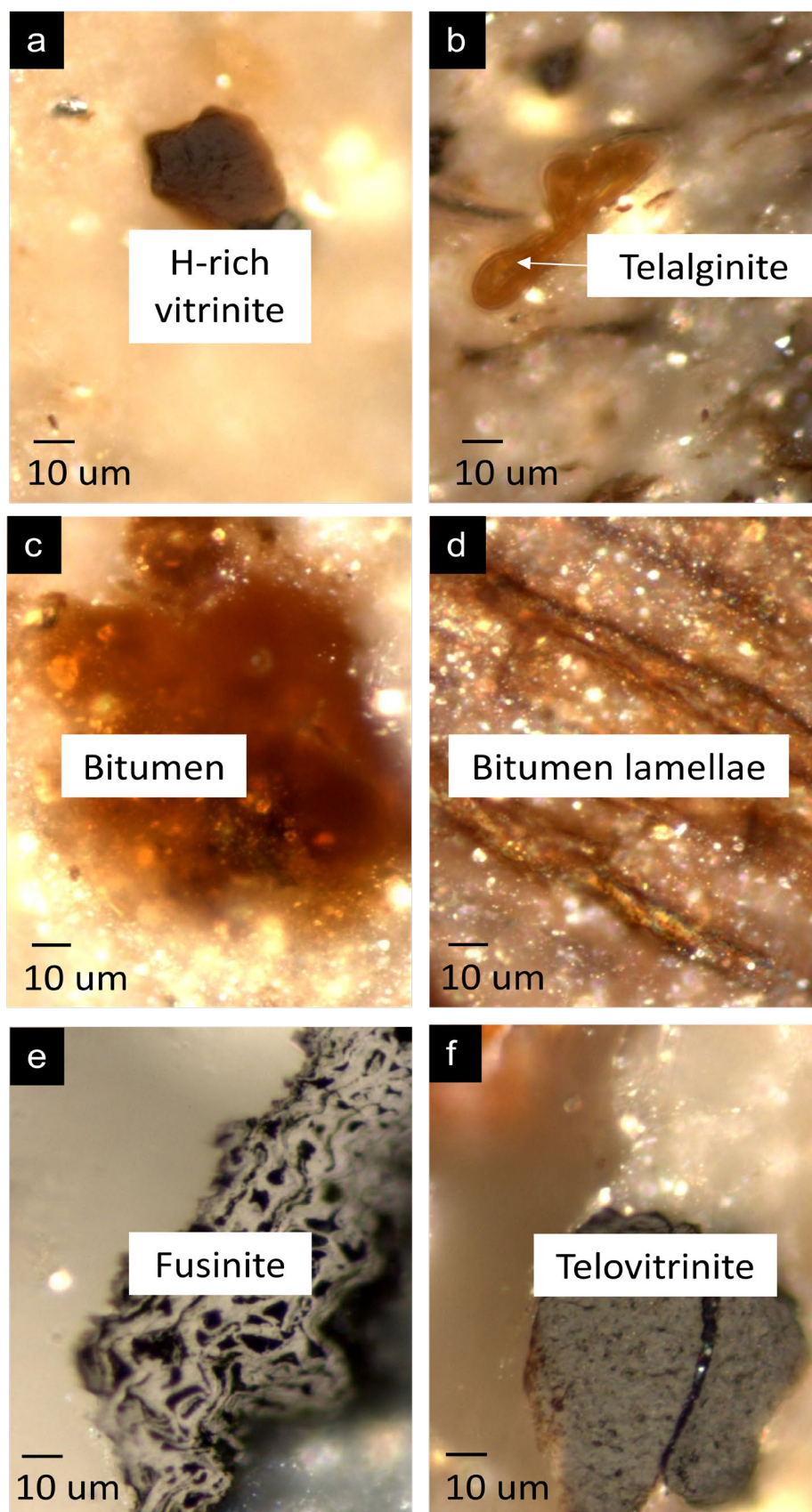
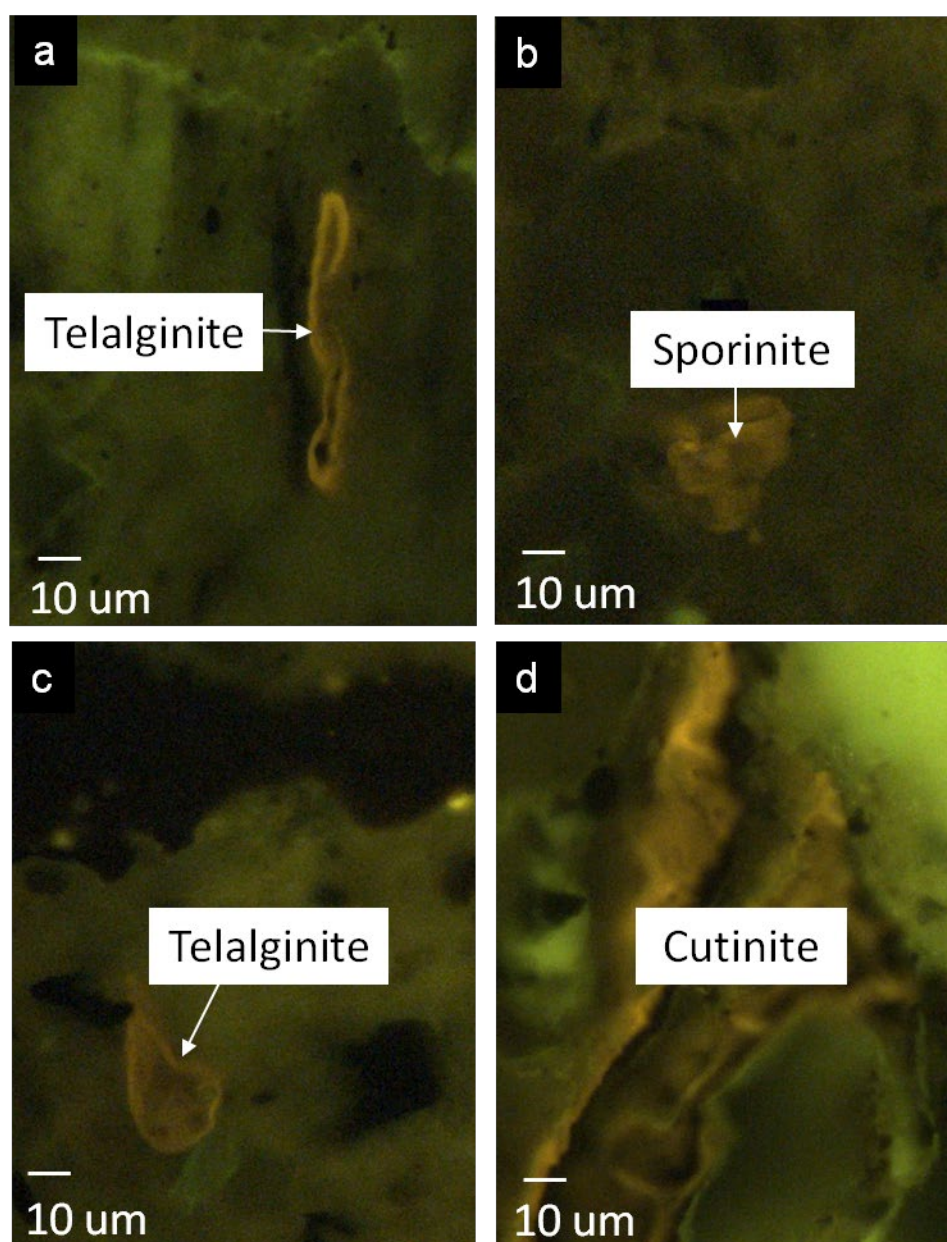
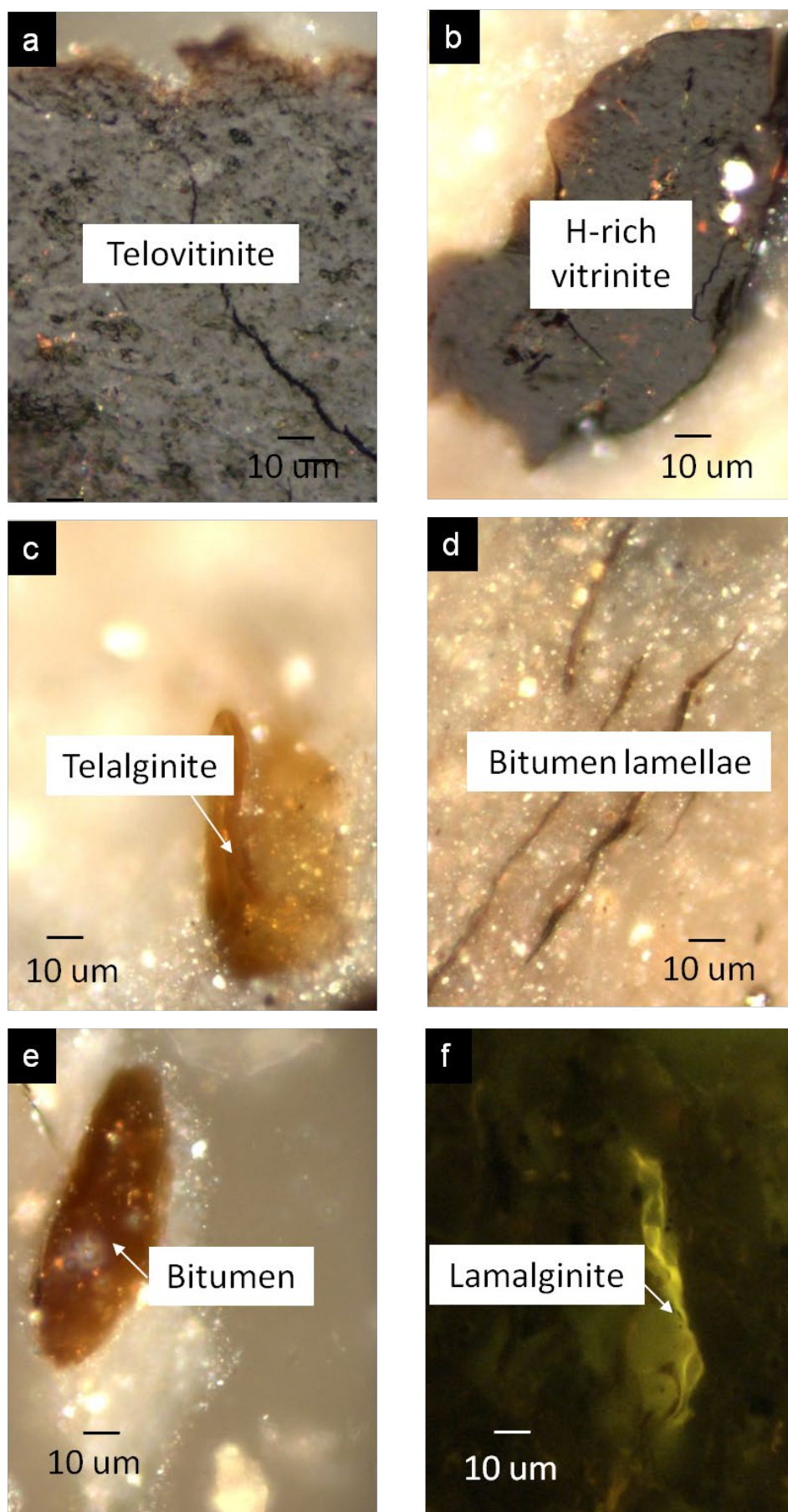
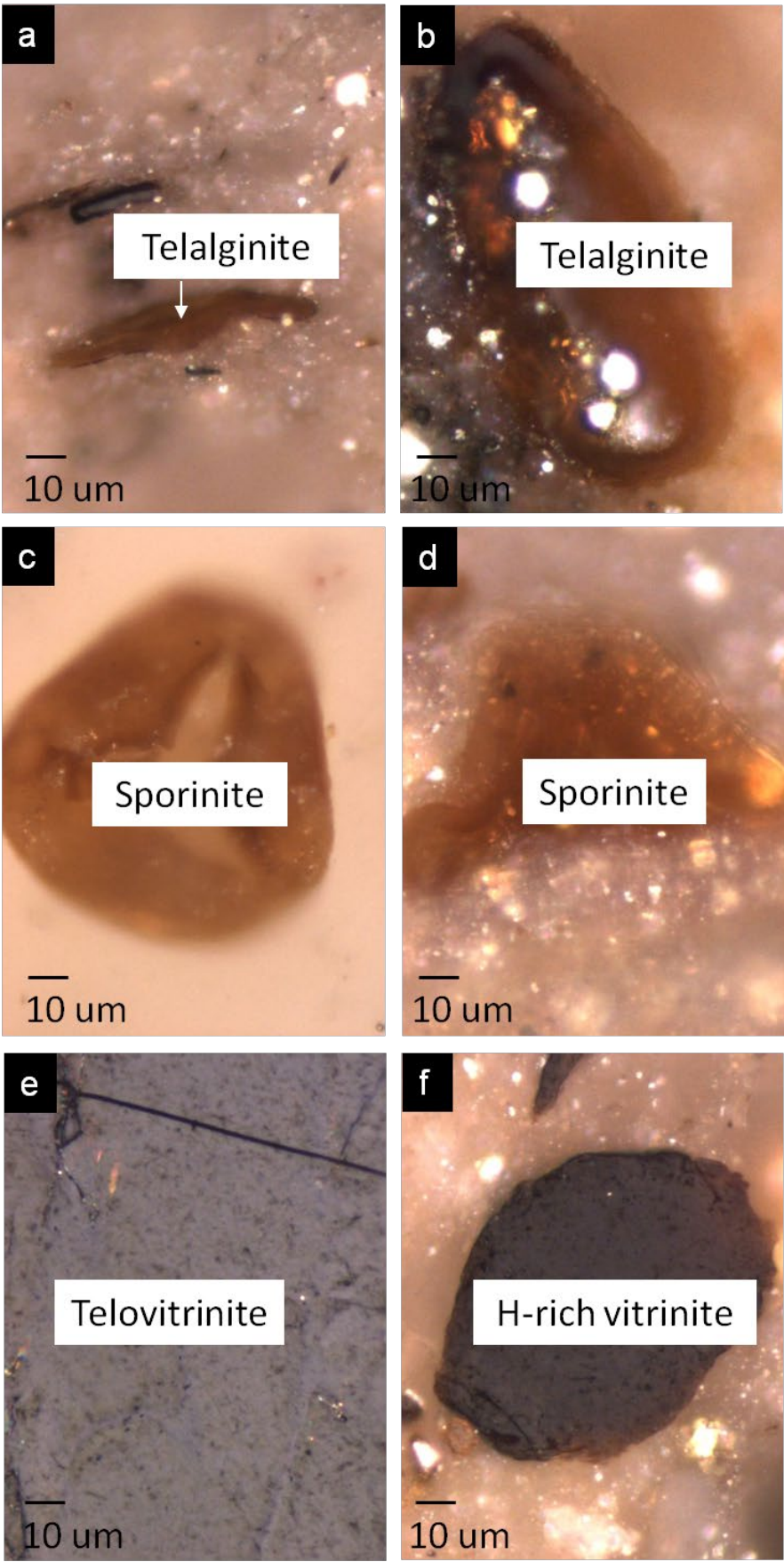
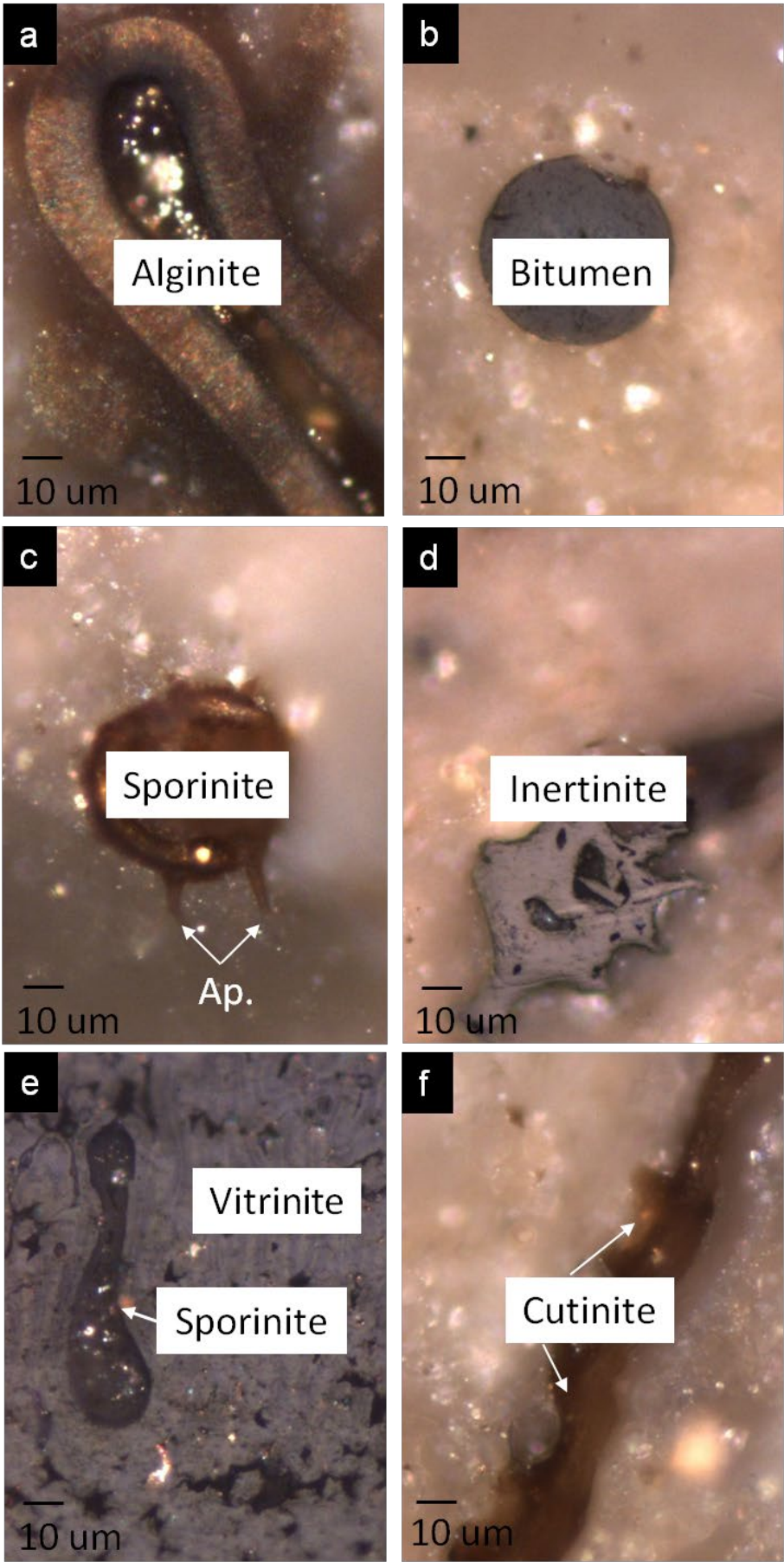


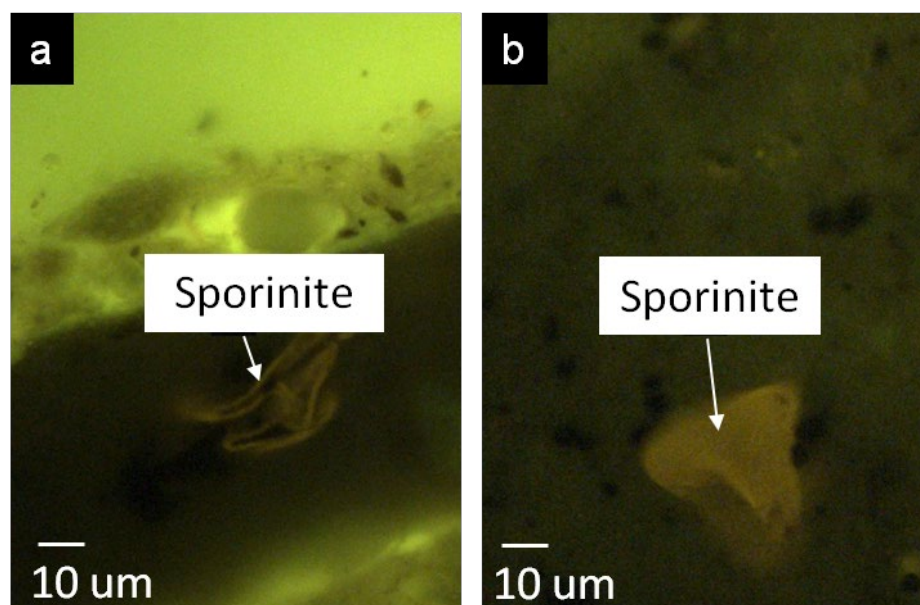
Figure 13

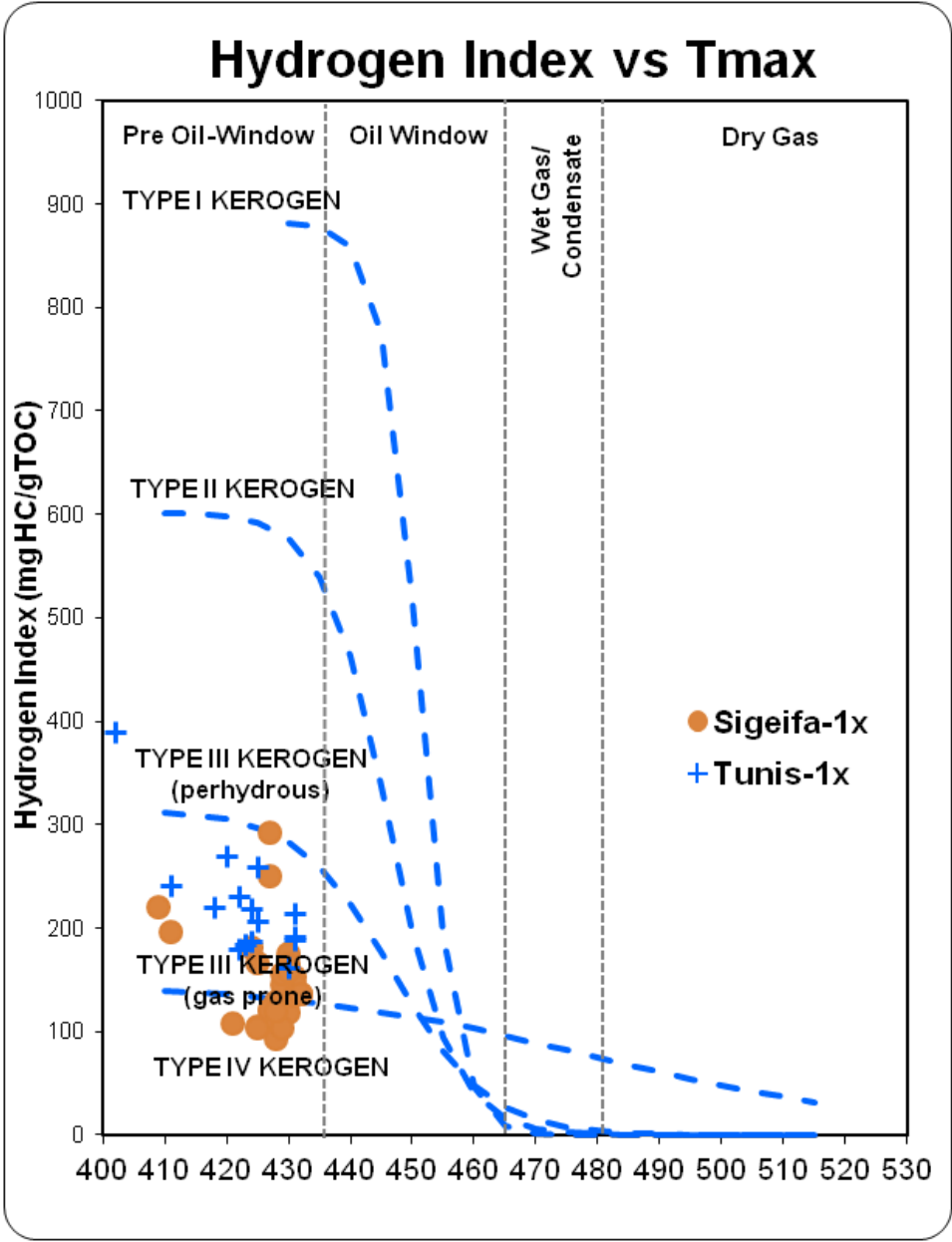


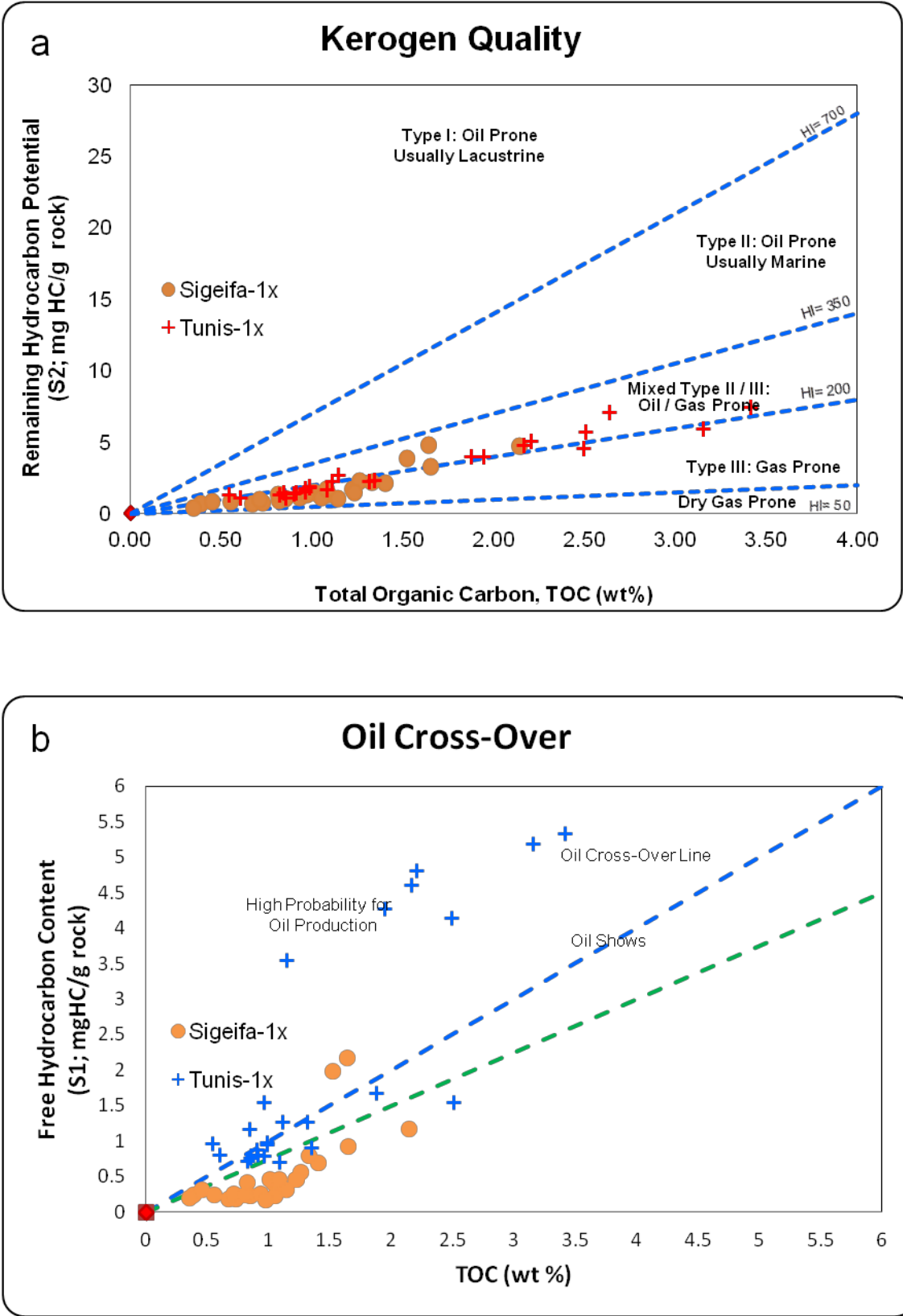












1247 **Figure 21**

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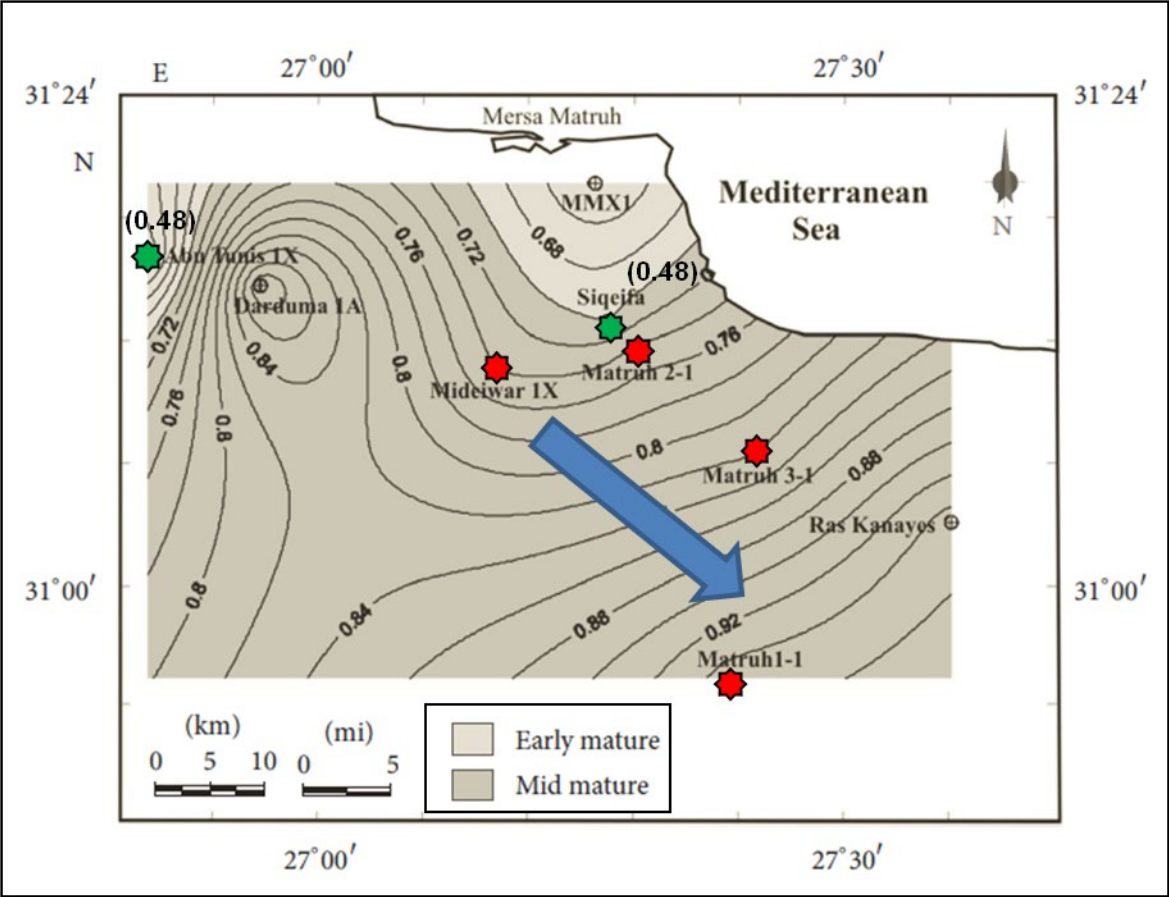


Table 1

Samples No.	Depth ft	Depth m	TOC wt %	S1 mg HC/g rock	S2 mg HC/g rock	S3 mg CO2/g rock	S3 mg HC/g rock	HI mg HC/g TOC	OI mg HC/g TOC	PI mg HC/g rock	OSI mg HC/g TOC	Oil in Rock bbl oil/ac-ft	Tmax °C	VRo (%)
28	5600	1706.9	0.82	0.72	1.3	1.48	0.28	158.54	180.49	0.36	87.80	15.76	298	1250 0.46
27	5650	1722.1	1.11	1.27	2.24	2.02	0.43	201.80	181.98	0.36	114.41	27.80	309	0.47
26	5750	1752.6	1.87	1.69	4.02	2.08	0.74	214.97	111.23	0.30	90.37	36.99	431	1252 0.47
25	5800	1767.8	0.53	0.77	1.08	0.81	0.22	203.77	152.83	0.42	145.28	16.86	428	
24	5850	1783.1	0.84	1.18	1.45	0.88	0.11	172.62	104.76	0.45	140.48	25.83	292	1253 0.47
23	5900	1798.3	0.98	0.98	1.85	1.55	0.41	188.78	158.16	0.35	100.00	21.45	431	0.49
22	6000	1828.8	0.96	1.56	1.73	1.39	0.39	180.21	144.79	0.47	162.50	34.15	422	1254 0.48
21	6050	1844.0	0.60	0.81	1.10	1.11	0.42	183.33	185.00	0.42	135.00	17.73	312	0.49
20	6100	1859.3	0.78	0.89	1.29	1.55	0.40	165.38	198.72	0.41	114.10	19.48	426	
19	6150	1874.5	1.34	0.92	2.34	2.58	0.95	174.63	192.54	0.28	68.66	20.14	318	1255 0.52
18	6200	1889.8	0.90	0.88	1.37	1.70	0.53	152.22	188.89	0.39	97.78	19.26	318	0.49
17	6250	1905.0	1.31	1.27	2.22	2.20	1.04	169.47	167.94	0.36	96.95	27.80	321	1256 0.48
16	6300	1920.2	0.82	1.04	2.05	1.65	0.42	250.00	201.22	0.34	126.83	22.76	423	
15	6400	1950.7	0.85	0.77	1.07	1.13	0.31	125.88	132.94	0.42	90.59	16.86	295	1257 0.51
Avg			0.98	1.05	1.79	1.58	0.47	181.54	164.39	0.38	112.20	23.06	358.9	
14	6450	1966.0	2.50	1.55	5.71	5.45	1.92	228.40	218.00	0.21	62.00	33.93	305	1258 0.49
13	6500	1981.2	7.05	3.77	18.27	5.20	2.46	259.15	73.76	0.17	53.48	82.53	425	
12	6600	2011.7	4.41	7.18	17.17	2.26	0.79	389.34	51.25	0.29	162.81	157.17	402	1259 0.49
11	6650	2026.9	4.75	7.69	15.14	13.92	0.80	318.74	293.05	0.34	161.89	168.33	416	
10	6700	2042.2	2.20	4.82	5.09	1.82	0.39	231.36	82.73	0.49	219.09	105.51	422	1260 0.49
9	6750	2057.4	2.63	6.04	7.09	2.08	0.55	269.58	79.09	0.46	229.66	132.22	420	0.49
8	6850	2087.9	0.54	0.97	1.30	0.91	0.25	240.74	168.52	0.43	179.63	21.23	411	0.50
7	6900	2103.1	6.97	7.70	17.98	9.67	4.58	257.96	138.74	0.30	110.47	168.55	312	1261 0.52
6	6950	2118.4	11.59	11.48	23.53	4.43	1.88	203.02	38.22	0.33	99.05	251.30	415	0.52
5	7000	2133.6	3.41	5.34	7.44	1.28	0.52	218.18	37.54	0.42	156.60	116.89	424	1262 0.57
4	7050	2148.8	2.49	4.15	4.58	1.34	0.56	183.94	53.82	0.48	166.67	90.84	423	
3	7100	2164.1	3.15	5.20	5.90	1.51	0.34	187.30	47.94	0.47	165.08	113.83	424	1263 0.51
2	7150	2179.3	2.16	4.62	4.76	2.22	0.62	220.37	102.78	0.49	213.89	101.13	418	
1	7200	2194.6	1.14	3.56	2.72	0.96	0.22	238.60	84.21	0.57	312.28	77.93	301	1264 0.55
Avg			3.93	5.29	9.76	3.78	1.13	246.19	104.97	0.39	163.76	115.81	394.1	

1266

Table 2

Samples No.	Depth ft	Depth m	TOC wt %	S1 mg HC/g rock	S2 mg HC/g rock	S3 mg CO2/g rock	S3 mg CO/g rock	HI mg HC/g TOC	OI mg HC/g TOC	PI mg HC/g rock	OSI mg HC/g TOC	Oil in Rock bbl oil/ac-ft	Tmax °C	VRo (%)
25	5850	1783.1	1.22	0.47	1.66	1.29	0.26	136.07	105.74	0.22	38.52	10.29	432	0.49
24	6000	1828.8	0.67	0.19	0.7	1.17	0.22	104.48	174.63	0.21	28.36	4.16	425	0.51
23	6050	1844.0	1.65	0.93	3.25	1.61	0.47	196.97	97.58	0.22	56.36	20.36	411	0.49
22	6100	1859.3	1.52	1.99	3.81	1.79	0.35	250.66	117.76	0.34	130.92	43.56	427	0.52
21	6150	1874.5	2.14	1.18	4.7	2.53	0.47	219.63	118.22	0.20	55.14	25.83	409	0.52
20	6200	1889.8	1.09	0.34	1.29	2.17	0.19	118.35	199.08	0.21	31.19	7.44	430	0.55
19	6300	1920.2	1.04	0.4	1.34	1.76	0.21	128.85	169.23	0.23	38.46	8.76	428	0.50
18	6350	1935.5	1.64	2.18	4.8	2.33	0.46	292.68	142.07	0.31	132.93	47.72	427	0.51
17	6400	1950.7	1.26	0.57	2.28	2.01	0.26	180.95	159.52	0.20	45.24	12.48	424	0.53
16	6450	1966.0	1.40	0.7	2.14	2.37	0.39	152.86	169.29	0.25	50.00	15.32	431	0.56
15	6550	1996.4	1.01	0.47	1.47	1.88	0.25	145.54	186.14	0.24	46.53	10.29	431	0.55
14	6600	2011.7	1.23	0.47	1.49	1.71	0.36	121.14	139.02	0.24	38.21	10.29	427	0.53
13	6650	2026.9	0.82	0.42	0.88	1	0.23	107.32	121.95	0.32	51.22	9.19	428	0.53
12	6700	2042.2	0.45	0.32	0.79	0.92	0.08	175.56	204.44	0.29	71.11	7.00	430	0.54
11	6800	2072.6	0.35	0.21	0.38	0.88	0.06	108.57	251.43	0.36	60.00	4.60	421	0.55
10	6850	2087.9	1.33	0.8	2.2	2.02	0.44	165.41	151.88	0.27	60.15	17.51	430	0.54
9	6950	2118.4	0.55	0.25	0.8	1.09	0.09	145.45	198.18	0.24	45.45	5.47	429	0.56
Avg			1.14	0.69	1.99	1.67	0.28	161.79	159.19	0.26	57.64	15.31	425.88	0.54
8	7000	2133.6	1.14	0.32	1.06	2.01	0.46	92.98	176.32	0.23	28.07	7.00	428	0.58
7	7050	2148.8	1.05	0.23	1.08	2.7	0.19	102.86	257.14	0.18	21.90	5.03	425	0.56
6	7100	2164.1	1.08	0.46	1.69	1.28	0.26	156.48	118.52	0.21	42.59	10.07	429	0.55
5	7200	2194.6	0.73	0.19	0.76	2.37	0.14	104.11	324.66	0.20	26.03	4.16	429	0.56
4	7250	2209.8	0.71	0.26	0.96	1.93	0.17	135.21	271.83	0.21	36.62	5.69	431	0.55
3	7300	2225.0	0.93	0.26	1.11	1.66	0.22	119.35	178.49	0.19	27.96	5.69	428	0.56
2	7350	2240.3	0.81	0.24	1.29	1.24	0.27	159.26	153.09	0.16	29.63	5.25	430	0.57
1	7400	2255.5	0.97	0.18	1.33	1.7	0.2	137.11	175.26	0.12	18.56	3.94	432	0.59
Avg			0.93	0.26	1.16	1.86	0.23	129.91	201.61	0.20	32.11	6.91	429	0.57

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Table 3

Sample no.	Depth (ft)	Depth (m)	Spores	Microplankton	pollen	Rt (Sporomorphs)	Rm (Marine palynomorphs)	PMI	Spores %	Microplankton %	Pollen %
34	5600	1706.9	139	12	149.0	288.0	12	4.2	46.3	4.0	49.7
33	5650	1722.1	100	10	190.0	290	10	3.4	33.3	3.3	63.3
32	5750	1,753	126	36	138	264	36	13.6	42.0	12.0	46.0
31	5800	1,768	128	6	166	294	6	2.0	42.7	2.0	55.3
30	5850	1,783	154	11	135	289	11	3.8	51.3	3.7	45.0
29	5900	1,798	150	23	127	277	23	8.3	50.0	7.7	42.3
28	5950	1,814	212	26	62	274	26	9.5	70.7	8.7	20.7
27	6000	1,829	205	17	78	283	17	6.0	68.3	5.7	26.0
26	6050	1,844	145	30	125	270	30	11.1	48.3	10.0	41.7
25	6100	1,859	167	11	122	289	11	3.8	55.7	3.7	40.7
24	6150	1,875	0	0	0	0	0	0	0	0	0
23	6200	1,890	0	0	0	0	0	0	0	0	0
22	6250	1,905	0	0	0	0	0	0	0	0	0
21	6300	1,920	186	10	104	290	10	3.4	62.0	3.3	34.7
20	6350	1,935	171	24	105	276	24	8.7	57.0	8.0	35.0
19	6400	1,951	139	48	113	252	48	19.0	46.3	16.0	37.7
Avg						279.7	20.3		51.8	6.8	41.4
18	6450	1,966	227	8	65	292	8	2.7	75.7	2.7	21.7
17	6500	1,981	199	9	92	291	9	3.1	66.3	3.0	30.7
16	6550	1,996	202	7	91	293	7	2.4	67.3	2.3	30.3
15	6600	2,012	239	11	50	289	11	3.8	79.7	3.7	16.7
14	6650	2,027	211	11	78	289	11	3.8	70.3	3.7	26.0
13	6690	2,039	200	9	91	291	9	3.1	66.7	3.0	30.3
12	6750	2,057	238	5	57	295	5	1.7	79.3	1.7	19.0
11	6800	2,073	243	7	50	293	7	2.4	81.0	2.3	16.7
10	6850	2,088	0	0	0	0	0	0	0	0	0
9	6900	2,103	284	1	15	299	1	0.3	94.7	0.3	5.0
8	6950	2,118	293	0	7	300	0	0.0	97.7	0.0	2.3
7	7000	2,134	275	2	23	298	2	0.7	91.7	0.7	7.7
6	7050	2,149	269	6	25	294	6	2.0	89.7	2.0	8.3
5	7100	2,164	242	6	52	294	6	2.0	80.7	2.0	17.3
4	7150	2,179	276	2	22	298	2	0.7	92.0	0.7	7.3
3	7200	2,195	204	14	82	286	14	4.9	68.0	4.7	27.3
2	7250	2,210	209	28	63	272	28	10.3	69.7	9.3	21.0
1	7300	2,225	234	6	60	294	6	2.0	78.0	2.0	20.0
Avg						288.8	7.3		79.3	2.6	18.1

1284 **Table 4**

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Sample no.	Depth (ft)	Depth (m)	Spores	Microplankton	pollen	Rt (Sporomorphs)	Rm (Marine palynomorphs)	PMI	Spores %	Microplankton %	Pollen %
16	5850	1783.08	199	10	91	290	10	3.4	66.3	3.3	30.87
15	6050	1844.04	207	14	79	286	14	4.9	69.0	4.7	26.3
14	6200	1889.76	215	15	70	285	15	5.2	71.7	5.0	23.3
13	6300	1920.24	238	18	44	282	18	6.4	79.3	6.0	14.78
12	6400	1950.72	217	15	68	285	15	5.2	72.3	5.0	22.7
11	6500	1981.2	227	12	61	288	12	4.2	75.7	4.0	20.89
10	6550	1996.44	212	19	69	281	19	6.7	70.7	6.3	23.0
9	6650	2026.92	218	13	69	287	13	4.5	72.7	4.3	23.0
8	6850	2087.88	237	11	52	289	11	3.8	79.0	3.7	12.90
7	6900	2103.12	229	16	55	284	16	5.6	76.3	5.3	18.3
Avg			219	14.3	65.	285.7	14.3	5.0	73.3	4.8	21.9
6	7000	2133.6	243	5	52	295	5	1.7	81.0	1.7	17.31
5	7050	2148.84	236	6	58	294	6	2.0	78.7	2.0	19.3
4	7100	2164.08	201	12	87	288	12	4.2	67.0	4.0	29.07
3	7250	2209.8	195	9	96	291	9	3.1	65.0	3.0	32.0
2	7350	2240.28	210	9	81	291	9	3.1	70.0	3.0	27.0
1	7400	2255.5	171	7	122	293	7	2.4	57.0	2.3	40.93
Avg			209.33	8.2	74.8	292	8	2.7	69.8	2.7	27.6

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1295 **Table 5**

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Samples No.	Depth ft	Depth m	Liptinite Group		Vitrinite (Non-opaque) Group %	Inertinite (Opaque) Group %
			(Palynomorphs) %	(AOM) %		
34	5600	1706.9	4	40.4	50.6	5
33	5650	1722.1	6	42	41	11
32	5750	1752.6	3.4	69	23.6	4
31	5800	1767.8	3.0	65.6	28.4	3.0
30	5850	1783.1	3.6	43	45.6	7.8
29	5900	1798.3	3	46.4	39.6	11
28	5950	1813.6	1	42	48	9
27	6000	1828.8	1.4	61	31.6	6
26	6050	1844.0	1.2	62.2	29	7.6
25	6100	1859.3	2	49	42	7
24	6150	1874.5	2.4	43.6	46	8
23	6200	1889.8	3.6	38	46.4	12
22	6250	1905.0	5	39	50	6
21	6300	1920.2	4	19.6	70	6.4
20	6350	1935.5	3.6	33.4	55.4	7.6
19	6400	1950.7	3.0	44.4	45.6	7.0
Avg.			3.1	46.2	43.3	7.4
18	6450	1966.0	2	66.2	25.4	6.4
17	6500	1981.2	1.8	73.2	20	5
16	6550	1996.4	2.8	43.4	47.6	6.2
15	6600	2011.7	3.6	77.2	9	10.2
14	6650	2026.9	3	80.8	7.8	8.4
13	6700	2042.2	2	84	10.4	3.6
12	6750	2057.4	2.4	87.6	7.6	2.4
11	6800	2072.6	1.6	91	5.4	2
10	6850	2087.9	1	87.6	3.8	7.6
9	6900	2103.1	2.4	62.2	33.2	2.2
8	6950	2118.4	2	70.4	26	1.6
7	7000	2133.6	2.4	68	25.2	4.4
6	7050	2148.8	4.4	62.2	27.8	5.6
5	7100	2164.1	5.8	60	31.6	2.6
4	7150	2179.3	5.4	61	29.8	3.8
3	7200	2194.6	4.4	63	27	5.6
2	7250	2209.8	6.4	60.2	25	8.4
1	7300	2225.0	5	65.2	23	6.8
Avg.			3.2	70.2	21.4	5.2

1319 **Table 6**

Palynofacies field and environment		Comments	Spores: Bisaccate pollen	Microplankton	Kerogen type
II	Marginal dysoxic-anoxic basin	AOM diluted by high phytoclast input, but AOM preservation moderates to good. Amount of marine TOC dependent on basin redox state.	High	Very low	III (gas prone)
VI	Proximal suboxic-anoxic shelf.	High AOM preservation due to reducing basin conditions. Absolute phytoclast content may be moderate to high due to turbiditic input and/or general proximity to source.	Variable low to moderate	Low to common dinocysts dominant	II (oil prone)
IX	Distal suboxic- anoxic basin.	AOM-dominant assemblages. Low abundances of palynomorphs partly due to masking. Frequently alginate-rich. Deep basin or stratified shelf sea deposits, especially sediments starved basins.	Low	Generally low, prasinophyte often dominant	II ≥ I (highly oil prone)

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Samples No.	Depth ft	Depth m	Liptinite Group		Vitrinite (Non-opaque) Group %	Inertinite (Opaque) Group %
			(Palynomorphs) %	(AOM) %		
16	5850	1783.08	3.60	22.6	72.20	1.60
15	6050	1844.04	4	30.4	63.2	2.4
14	6200	1889.76	3.8	20	72.6	3.6
13	6300	1920.24	3	21	70.6	5.4
12	6400	1950.72	5	27	64.4	3.6
11	6500	1981.2	4	24.6	69.4	2.342
10	6550	1996.44	3.6	19.4	73.6	3.4
9	6650	2026.92	2	18.2	75.4	4.4
8	6850	2087.88	3	34	60	3
7	6900	2103.12	3	39.4	56	1.6
Avg.			3.50	25.7	67.7	3.1644
6	7000	2133.6	3	33	63	1
5	7050	2148.84	2.8	38.6	56	2.645
4	7100	2164.08	1	43	54.6	1.4
3	7250	2209.8	4.6	34	59.6	1.8
2	7350	2240.28	4	35.6	58.6	1.8
1	7400	2255.5	5	32	62	1
Avg.			3.4	36	58.9	1.647

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Table 8

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Sample no.	Depth (ft)	Depth (m)	Dinocyst %	MFTLs %	Pteridophyte spores %	Freshwater Algae %	Spherical pollen grains %	PMI	Non-opaque wood %	Opaque wood %	1353 AOM %
16	5850	1783.08	2.7	0.7	59.0	0.0	26.3	3.4	72.20	1.60	22.6
15	6050	1844.04	2.7	2.0	62.3	0.0	16.3	4.9	63.2	2.4	30.5 1354
14	6200	1889.76	4.3	0.7	64.7	0.0	14.3	5.2	72.6	3.6	20
13	6300	1920.24	3.7	2.3	66.7	0.0	11.0	6.4	70.6	5.4	21.5 1355
12	6400	1950.72	3.0	2.0	68.0	0.0	17.7	5.2	64.4	3.6	27
11	6500	1981.2	3.0	1.0	70.0	0.0	16.7	4.2	69.4	2	24.6 1356
10	6550	1996.44	5.3	1.0	67.7	0.3	20.3	6.7	73.6	3.4	19.4
9	6650	2026.92	3.7	0.7	66.7	0.0	19.3	4.5	75.4	4.4	18.2
8	6850	2087.88	2.7	1.0	69.3	0.0	16.0	3.8	60	3	13.5 1357
7	6900	2103.12	4.0	1.3	67.0	0.0	15.0	5.6	56	1.6	39.4
Avg			3.5	1.3	66.1	0.03	17.3	5.0	67.7	3.1	25.7 1358
6	7000	2133.6	1.7	0.0	74.0	0.0	14.3	1.7	63	1	33
5	7050	2148.84	2.0	0.0	71.0	0.0	16.0	2.0	56	2.6	38.6 1359
4	7100	2164.08	3.3	0.7	58.3	0.0	10.0	4.2	54.6	1.4	43
3	7250	2209.8	1.7	0.7	56.7	0.0	24.7	3.1	59.6	1.8	34
2	7350	2240.28	3.0	0.0	63.3	0.0	18.3	3.1	58.6	1.8	35.6 1360
1	7400	2255.5	1.3	1.0	52.0	1.0	31.7	2.4	62	1	32
Avg			2.2	0.4	62.6	0.2	26.3	2.7	58.9	1.6	38.6 1361

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