

1 Integrated geophysical characterization of crustal domains in the Eastern 2 Black Sea

3 **Vanessa Monteleone^{1,2}, Tim A. Minshull¹, and Hector Marin-Moreno²**

4 *¹School of Ocean and Earth Science, University of Southampton, National Oceanography Centre*
5 *Southampton, Waterfront Campus, European Way, Southampton SO14 3ZH, UK*
6 *(v.monteleone@soton.ac.uk; tmin@noc.soton.ac.uk).*

7 *²National Oceanography Centre, University of Southampton Waterfront Campus, European*
8 *Way, Southampton SO14 3ZH, UK (hector.marin.moreno@noc.ac.uk).*

9 10 **ABSTRACT**

11 Rifting may lead ultimately to continental breakup, but the identification and
12 characterization of the resulting crustal distribution remains challenging. Also, spatial and
13 temporal changes in breakup magmatism may affect the geophysical character of the newly
14 formed oceanic crust, resulting in contrasting interpretations of crustal composition and
15 distribution. In the Eastern Black Sea Basin (EBSB), the evolution from rifting to breakup has
16 been long debated, with several interpretations for the distribution of stretched continental and
17 oceanic crust. We interpret basement morphological variations from long-offset seismic
18 reflection profiles, highlighting a NW-SE transition from faulted and tilted continental blocks, to
19 a rough and then smoother basement. We model magnetic anomalies to constrain further the
20 various basement domains, and infer the presence of a weakly-magnetized, stretched continental
21 crust in the NW, and a 0.4-2.8 A/m layer coinciding with the smooth basement in the central and
22 SE area. We conclude that the EBSB oceanic crust extends further to the NW than was suggested

23 previously from an abrupt change in crustal thickness and lower-crustal velocity. The apparent
24 discrepancy between these different types of geophysical evidence may result from changes in
25 magma supply during breakup, affecting thickness and velocity structure of the resulting oceanic
26 crust.

27 **1. INTRODUCTION and METHOD**

28 Observations of the structure and composition of the basement underlying rifted margins can
29 unravel the geodynamic processes driving their formation. High-quality, 2D/3D regional seismic
30 datasets provide insights into basement structures and sediment filling within the distal part of
31 the margin, thus helping geodynamic reconstruction (e.g., Tucholke et al., 2007; Hauptert et al.,
32 2016). In deep-water and thick sedimentary infill settings, and when the basement is neither
33 exposed nor sampled by well data, our ability to interpret rifting and breakup features and
34 basement structures is limited. Particularly in these cases, refraction seismic (e.g., Dunn and
35 Martinez, 2011) and potential field data (e.g., Ball et al., 2013) can be used to investigate the
36 nature of the crust, its thickness and velocity structure, and the amount of magmatism
37 contributing at its formation (e.g., Franke, 2013). Only a few studies integrate these approaches
38 (e.g., Prada et al., 2014; Tugend et al., 2015).

39 In the Eastern Black Sea Basin (EBSB), interpretations of different datasets commonly disagree.
40 Crustal thickness and lower-crustal velocities from wide-angle seismic data define a small area
41 of oceanic crust to the SE of the basin (Shillington et al., 2009). Evidence for hummocky
42 basement structures along seismic reflection profiles suggest a narrow and elongated oceanic
43 crust extending further NW (Nikishin et al., 2015). Gravity modeling has also been used to
44 define crustal boundaries, with different results (e.g., Starostenko et al., 2004; Graham et al.,
45 2013) (Fig. DR1 in GSA Data Repository¹). Therefore, although it is generally agreed that the

46 EBSB formed as a Late Cretaceous-early Cenozoic back-arc basin (e.g., Zonenshain and Le
47 Pichon, 1986; Finetti et al., 1988; Okay et al., 1994), its crustal distribution is still debated.
48 New long-offset seismic reflection profiles, acquired in 2011 by Geology Without Limits (GWL)
49 and ION GXT (Table DR1), allowed Monteleone et al. (2019) to define the spatial and temporal
50 distribution of extension and breakup processes in the EBSB. Here, we build on that analysis
51 using additional data to untangle the lack of consensus on the EBSB crustal distribution. We first
52 differentiate crustal domains based on basement morphological changes and the interaction
53 between basement structures, stratigraphic and structural elements visible from long-offset
54 seismic reflection profiles. We then apply magnetic anomaly modeling constrained by our
55 seismic interpretation to investigate the magnetization character of the seismically identified
56 domains, using the Earth Magnetic Anomaly Grid (EMAG2-v3) over the EBSB region (Meyer et
57 al., 2017) (Fig. 1; Fig. DR2-8) (see GSA Data Repository). Magnetic data have been previously
58 used in the Black Sea to define the age of basin opening (Kazmin et al., 2007), crustal scale
59 structures and faulting (Starostenko et al., 2015), and the thermal structure of the lithosphere
60 (Starostenko et al., 2014), but no modeling has been attempted yet to define crustal composition.
61 Our results show a NW-SE change in morphological and magnetic character of the EBSB
62 basement. Our inferred oceanic domain extends into a region where the presence of thinned
63 continental crust was previously inferred from the seismic velocity structure. We relate our new
64 interpretation to along-axis changes in breakup kinematics and magma supply during spreading.

65 **2. RESULTS and DISCUSSION**

66 **2.1. Seismic Interpretation and Magnetic Anomaly Modeling**

67 We identify the EBSB basement as the most continuous seismic reflection underlying the
68 sedimentary infill, showing a variable morphological and structural character (Fig. 2). Over the
69 structural highs (Shatsky Ridge and MBSH) and within the narrow NW rift, the basement is
70 affected by normal faults generating half-graben structures bounded by faulted and tilted crustal
71 blocks overlain by clear wedge-shaped syn-rift deposits (Fig. 2B). These elements are typically
72 associated with stretched continental basement (Domain I) (Fig. 2A, 2B). In this area, the
73 observed magnetic anomaly is best fit by a weakly-magnetized layer (< 0.4 A/m), also
74 suggesting a continental nature for Domain I, bounded by two highly magnetized layers
75 corresponding to mafic magmatism over the Shatsky Ridge and MBSH (e.g., Starostenko et al.,
76 2004; Nikishin et al., 2015) (Fig. DR2-7). Previous results from wide-angle seismic data support
77 this interpretation (Fig. 2A).

78 Seaward of Domain I, the basement has relief of similar magnitude but with no evidence for
79 extensional faulting, nor clear syn-rift deposits (Fig. 2C). This basement extends 70-80 km
80 along-axis, and from its distinctive morphological character we interpret it as a separate crustal
81 domain (Domain II) (Fig. 2A, 2C). This rough basement was interpreted by Nikishin et al.
82 (2015) as having oceanic affinity. Rough basement may result from slow seafloor spreading
83 (e.g., Malinverno, 1991), or mantle exhumation processes (e.g., Sauter et al., 2018). The
84 observed magnetic anomaly along Domain II is best fit by a weakly-magnetized layer,
85 suggesting a continental rather than oceanic nature (Fig. 2C; Fig. DR4-5). Although we cannot
86 rule out the presence of exhumed mantle, which also may have rather weak magnetization (e.g.,

87 Sibuet et al., 2007), seismic velocities suggest the presence of highly stretched continental crust
88 in this area (Shillington et al., 2009) (Fig. 2A).

89 The rough basement of Domain II becomes smoother towards the central and SE part of the
90 basin where well-layered, post-rift deposits are predominant (Domain III) (Fig. 2D, 2E; Fig.
91 DR4-7). No extensional faults are visible from seismic data, although high-angle NE-SW-
92 trending transform faults are present (Fig. 2A, 2E). Along some profiles, volcano-like structures
93 have been interpreted as evidence for enhanced magmatism in this area (Nikishin et al., 2015)
94 (Fig. 2A; Fig. DR7). Smooth basement with extrusive volcanic material may relate to
95 anomalously thick and/or fast spreading oceanic crust (e.g., Small, 1994; Searle et al., 2010),
96 exhumed mantle (e.g., Pickup et al., 1996), or thinned continental crust overlain by basaltic flows
97 (e.g., Zhao et al., 2016). At Domain III, magnetic anomaly data is best fit with negatively
98 magnetized layers with intensity bands between 0.4 and 2.8 A/m, in contrast to the weakly
99 magnetized Domain II (< 0.4 A/m) (Fig. 3; Fig. DR2), thus providing evidence against
100 continental crust or exhumed mantle. Serpentinized exhumed mantle may be significantly
101 magnetized (e.g., Sibuet et al., 2007), but its presence is excluded based on crustal seismic
102 velocities (Shillington et al., 2009).

103 Seismic data and magnetic anomaly modeling results support the presence of smooth and
104 magnetized oceanic crust along Domain III (Fig. 2-3; Fig. DR2-7). However, wide-angle data
105 indicate the presence of a thin crust (7-9 km) with low lower-crustal velocities (6.4-6.6 km/s) in
106 the NW, and of a thick crust (11-13 km) with high lower-crustal velocities (6.8-7.2 km /s) in the
107 SE. This change in crustal structure occurs close to the Ordu-Pitsunda transform fault (OPf) (Tari
108 et al., 2018), within Domain III (Fig. 2A, 2E). Shillington et al. (2009) interpreted this change as
109 a transition from highly stretched continental (NW) to thick oceanic (SE) crust (Fig. 2A), an

110 interpretation inconsistent with the uniform magnetization character of Domain III. Our results
111 show that the best-fitting model along profile bs-110, over crust interpreted as continental by
112 Shillington et al. (2009), is obtained with magnetizations of 2.4-3.2 A/m over the Shatsky Ridge
113 and 0.6 A/m towards the Pontides magmatic arc (Fig. 3; Fig. DR7). In the central basin, the
114 observed anomaly is best fit by a ~ 125 km wide, negatively magnetized layer, with intensity
115 bands between 0.4 and 2.4 A/m (Fig. 3; Fig. DR7). Similar results are obtained along profile bs-
116 100 at the NW end of Domain III, and profile bs-120 at the SE end of Domain III where the
117 velocity structure is clearly oceanic (Fig. 1, 2A; Fig. DR2-6). Although it is possible that
118 magnetic anomalies NW of the OPf are caused by syn-rift magmatic intrusions through thinned
119 continental crust (Starostenko et al., 2004), their intensity, overall continuity, and similarity to
120 anomalies in the oceanic domain to the SE, support their formation by seafloor spreading. Thus,
121 Domain III has a smooth and highly magnetized basement resembling that of oceanic crust,
122 extending the EBSB oceanic domain further NW than previously inferred from wide-angle data
123 (Fig. 2A). Gravity modeling also supports a more extended oceanic domain, with boundaries
124 comparable to our results (Graham et al., 2013) (Fig. DR1B).
125 Crustal thickness and velocity structure are usually reliable parameters to discriminate crustal
126 composition in extensional settings (e.g., Prada et al., 2014). However, observations from wide-
127 angle data and magnetic anomaly modeling suggest contradicting results on crustal distribution
128 in the EBSB. Below we explain this apparent contradiction based on the tectono-magmatic
129 evolution of the EBSB.

130 **2.2. Tectono-Magmatic Evolution**

131 The EBSB rotational opening resulted in a V-shaped basin with increased extension to the SE
132 (Okay et al., 1994; Shillington et al., 2008). Analytical and numerical models of V-shaped basins

133 show that, as rifting propagates, stretching increases with distance from the pole of rotation and
134 breakup initiates away from the pole, propagating towards it and into stretched continental
135 lithosphere (e.g., Vink, 1982; Le Pourhiet et al., 2018). As spreading progresses, spatial and
136 temporal variations in magmatic crustal accretion driven by mantle temperature, strain
137 localization, and spreading rates, can influence the emplaced crustal structure and its geophysical
138 character (Franke, 2013). Changes in crustal structure within Domain III may reflect along-axis
139 changes in magmatism (e.g., Muller et al., 1999; Hooft et al., 2000), and/or variations in melt
140 supply over time (e.g., Tucholke et al., 1997; Cannat et al., 2003). Such changes would mainly
141 cause variations in oceanic Layer 3 thickness (e.g., Mutter and Mutter, 1993; Grevemeyer et al.,
142 2018). Oceanic Layer 2, characterized by lower seismic velocities (e.g., Mutter and Mutter,
143 1993) and considered to be the main source of oceanic magnetic anomalies (e.g., Talwani et al.,
144 1971), would not be much affected. An enhanced melt supply, driven by the advection of warm
145 asthenosphere during lithospheric thinning or by the presence of a hot mantle plume (e.g.,
146 Hopper et al., 2003), can form atypically thick oceanic crust with a well-developed, high-
147 velocity Layer 3. Lower melt supply, due to an initially cool lithosphere or conductive cooling at
148 slow extension rates suppressing partial melting (Bown and White, 1995), will result in a thinner
149 oceanic crust mostly characterized by Layer 2 velocities. Models of rotational basin opening
150 suggest that a decrease in extension rates toward the rotational pole results in an along-axis
151 reduction in the volumes of decompression melt (Franke, 2013), but do not explain the abrupt
152 change in crustal structure at the OPf (Fig. 2A, yellow line, Fig. 4).

153 Discontinuities and/or transform faults can control magma supply along spreading segments
154 (e.g., Fox and Gallo, 1984; Georgen and Lin, 2003), and thermal perturbation by transforms
155 often results in a reduction in crustal thickness near the transforms (e.g., Hooft et al., 2000),

156 resulting in anomalous oceanic crustal structure (e.g., Dunn and Martinez, 2011). Numerical
157 models show that transfer/transform faults can act as rift propagation barriers delaying the
158 opening of consecutive rift segments (Koopmann et al., 2014). Rift delay can affect both mantle
159 and melt flow between segments, limiting the propagation of low-viscosity mantle across the
160 segment boundary, and favoring melt accumulation at the opening segment (Koopmann et al.,
161 2014). The model predicts rift-parallel flow and concomitant magmatic peaks near the
162 propagation barrier caused by the lateral pressure gradient between sequentially opening
163 segments. This flow may result in abrupt changes in crustal thickness at segment boundaries
164 matching observations from the EBSB. Thus, the OPf may have acted as a rift propagation
165 barrier between SE and NW segments, resulting in pooling of melt to the SE. Such a model
166 would explain the abrupt transition from a thick oceanic crust SE of the OPf, with a well-
167 developed Layer 3, and thin oceanic crust NW of the OPf, with a thin/absent Layer 3, velocities
168 similar to those of thinned continental crust, but still with a distinctive magnetic character (Fig.
169 4).

170 Based on these considerations, we propose that the observed variations in crustal thickness and
171 lower-crustal velocities in Domain III result from changes in oceanic crustal accretion rather than
172 a change in crustal type. Variations in magma supply driven by the geometry of the opening rift,
173 and the effect of segmentation, were major controls on the basin architecture and the crustal
174 structure at the newly emplaced EBSB crust.

175 **CONCLUSIONS**

176 Based on basement morphology and magnetization, we have drawn new boundaries for the
177 oceanic crust in the EBSB. Changes in crustal thickness and velocity structure from wide-angle
178 seismic data within this oceanic domain, previously interpreted to show oceanic crust only to the
179 very SE, are interpreted instead as evidence for changes in magma supply during spreading. Rift
180 segmentation and transform faults may have played an important role in controlling the along-
181 axis flow of mantle and melt. Initial spreading focused in the wider SE rift segment, where
182 enhanced decompression melting and melt pooling caused the emplacement of thick oceanic
183 crust with a well-developed Layer 3. Melt propagation to the NW was limited by transform
184 faults, particularly the OPf, and/or by rift narrowing suppressing partial melting, resulting in the
185 emplacement of a thinner oceanic crust lacking a well-developed Layer 3.

186 This study highlights the ambiguity in the geophysical characterization of crustal type and
187 distribution in rifted margin settings. Although large amount of data are often available, different
188 and perhaps contradicting interpretations are possible, especially where geophysical evidence
189 deviates from what is “normally” expected for continental and oceanic crusts. In such cases, to
190 avoid contradicting interpretations, data comparison and integration is essential.

191

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336

337 **FIGURE CAPTIONS**

338 Figure 1. Reduced-to-pole magnetic anomaly map (EMAG2-v3) for the EBSB region. Black
339 lines mark GWL seismic profiles. Prominent positive anomalies correspond to the Shatsky Ridge
340 (SR), the Mid Black Sea High (MBSH), the Eastern Pontides (EP) magmatic arc, and contrast
341 with the negative anomaly of the central basin.

342

343 Figure 2. Summary of the basement morphology and distribution across the EBSB based on
344 seismic interpretation. (A) Map of the identified crustal domains (Domain I, II and III). Overlain
345 on map, the location of the GWL profiles, of sections shown in (B)-(E), and of wide-angle
346 seismic profiles (Shillington et al., 2009). The color code for the wide-angle lines shows the
347 crustal distribution interpreted by Shillington et al. (2009). Volcanics (red stars) are identified
348 within Domain III (Nikishin et al., 2015). Transform faults interpreted in the SE of the basin are
349 named after Tari et al. (2018). To the SE, we have insufficient data to constrain the seismic
350 character in Domain III(?). (B) Domain I - tilted and faulted crustal blocks. (C) Domain II -
351 rough basement with no extensional faulting, nor syn-rift deposits. (D) Domain III - smooth
352 basement with no extensional faulting. (E) Domain III - smooth basement affected by high-angle
353 transform faults; the thick sub-vertical dashed line shows the location of the Ordu-Pitsuda fault
354 (OPf).

355

356 Figure 3. Magnetic anomaly modeling along profile bs-110 (location shown in Figs.1 and 2A).
357 Top panel displays the comparison between calculated (red line) and observed (black line)
358 anomalies. The root-mean-squared (RMS) error, expressed in nT, between observed and
359 calculated anomaly is shown at the bottom left corner. The grey band represents the anomaly

360 variations 15 km either side of the profile. Colored backgrounds show the lateral extent and
361 intensity (dark - high intensity; pale - low intensity) of the magnetized layers: pink - continental
362 domain; blue - oceanic crust. Bottom panel shows the model generating the calculated anomaly
363 and the magnetization for each layer.

364

365 Figure 4. Conceptual 3D and 2D models for oceanic crustal accretion in the EBSB. After back-
366 arc extension, breakup focused in the wider SE rift where higher amount of decompression melt
367 and melt focusing favored the emplacement of a thick oceanic crust with a well-developed Layer
368 3 (L3). Some volcanics are also visible in this area, probably extruded along fracture zones
369 separating different spreading segments. Transform faults, particularly the OPf, acted as barriers
370 limiting melt propagation towards the NW. In this area, a reduced supply of decompression melt
371 and a limited melt migration from the SE segment, caused the emplacement of a thinner oceanic
372 crust characterized mostly by the upper oceanic Layer 2 (L2). The main geophysical difference
373 between the SE and the NW oceanic domains is the change in oceanic crustal thickness and
374 velocity structure, related to the absence/near absence of L3 along the NW segment.

375

376 ¹GSA Data Repository item 2019xxx, additional information on seismic data (Table DR1),
377 magnetic anomaly data and methodology used in this study, together with a comparison between
378 our and previous results on crustal distribution in the basin (Figure DR1), and additional
379 magnetic anomaly modeling results (Figures DR2-8), is available online at
380 <http://www.geosociety.org/datarepository/2019/>, or on request from editing@geosociety.org or
381 Documents Secretary, GSA, P.O. Box 9140, Boulder, CO 80301, USA.

Figure 1

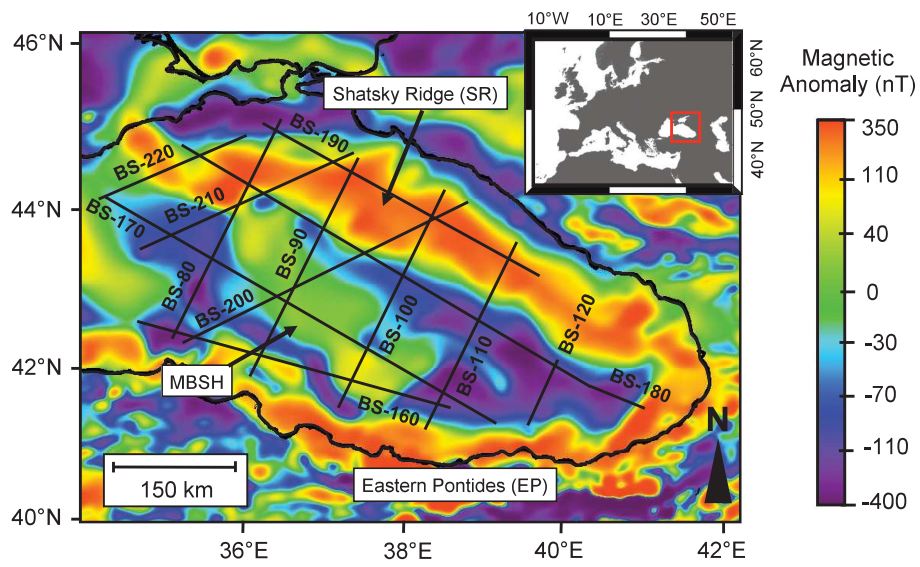


Figure 2

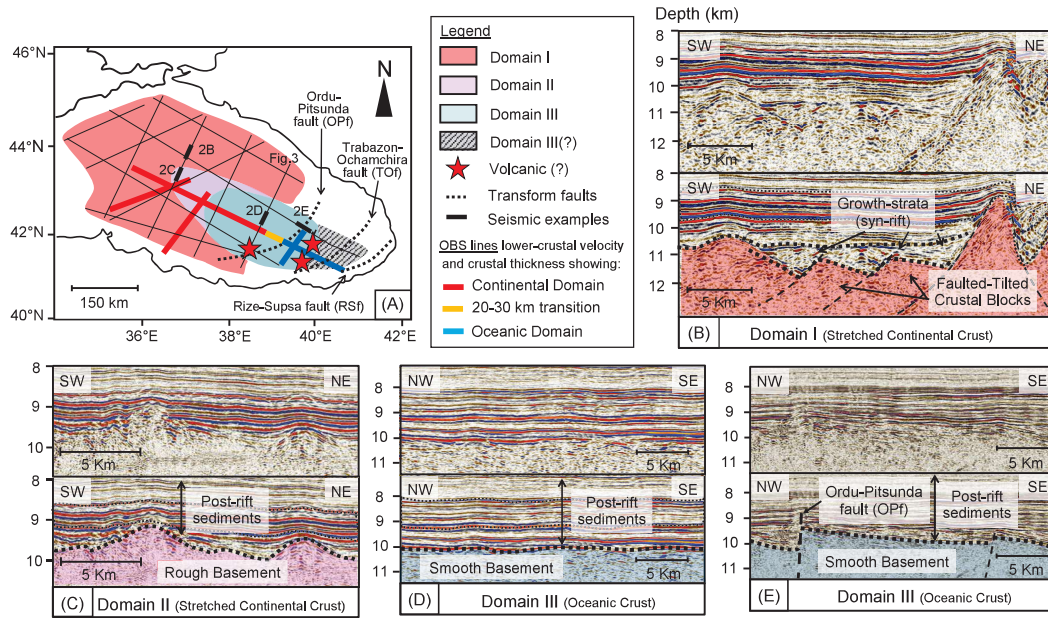


Figure 3

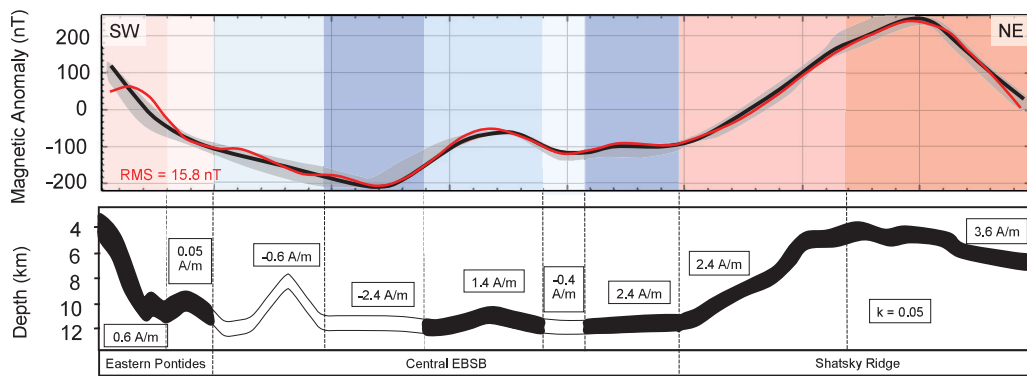


Figure 4

