

1 **Geochemical characterization of the Oman Crust-Mantle transition zone,**  
2 **OmanDP Holes CM1A and CM2B.**

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25 **Keywords:** Partial melting vs. melt-rock reaction; Dunites and harzburgites serpentinization and  
26 carbonation; Oman ophiolite Crust-Mantle transition; Holes CM1A and CM2B; ICDP Oman  
27 Drilling Project;

28 **Key Points:**

- 29 • The transition from the oceanic crust to the mantle of Oman has been drilled in the CM  
30 Holes during Phase 2 of the ICDP Oman Drilling Project

- 31 • There is large petrological and chemical variability in the dunites and harzburgites from  
32 Holes CM1A and CM2B
- 33 • Partial melting *vs.* melt-rock reaction, and the effects of serpentinization and carbonation  
34 of dunites and harzburgites are investigated

## 35 Abstract

36 The transition from the gabbroic oceanic crust to the residual mantle harzburgites of the Oman  
37 ophiolite has been drilled at Holes CM1A and CM2B (Wadi Tayin massif) during Phase 2 of the  
38 International Continental Scientific Drilling Program (ICDP) Oman Drilling Project (OmanDP)  
39 (Nov. 2017-Jan. 2018). In order to unravel the formation processes of ultramafic rocks in the Wadi  
40 Tayin massif (CM) crust-mantle transition zone and deeper in the mantle sections beneath oceanic  
41 spreading centers, our study focuses on the whole rock major and trace element compositions  
42 (together with CO<sub>2</sub> and H<sub>2</sub>O concentrations) of these ultramafic rocks (56 dunites and 49  
43 harzburgites). Despite extensive serpentinization and some carbonation, most of the trace element  
44 contents (REE, HFSE, Ti, Th, U) record high temperature, magmatic process-related signatures.  
45 Two major trends are observed, with good correlations between (1) Th and U, Nb and LREE on  
46 one hand, and between (2) HREE, Ti and Hf on the other hand. We interpret the first trend as the  
47 signature of late melt/peridotite interactions as LREE are known to be mobilized by such processes  
48 ('lithospheric process'), and the second trend as the signature of the initial mantle partial melting  
49 ('asthenospheric process'), with little or no overprint from melt/rock reaction events.

## 50 Plain Language Summary

51 We focus on the transition from the oceanic crust to the Earth's mantle by studying Holes CM1A  
52 and CM2B, drilled in the Oman ophiolite during Phase 2 of the International Continental Scientific  
53 Drilling Program (ICDP) Oman Drilling Project (OmanDP). Despite extensive serpentinization  
54 and some carbonation, the dunites and harzburgites from the transition zone and the mantle section  
55 show a large variability in their petrological and chemical compositions. Results indicate that most  
56 of the trace element contents (REE, HFSE, Ti, Th, U) record high temperature, magmatic process-  
57 related signatures. Two major trends are observed, with good correlations between (1) Th and U,  
58 Nb and LREE on one hand, and between (2) HREE, Ti and Hf on the other hand. We interpret the  
59 first trend as the signature of late interactions between a percolating melt and the harzburgites  
60 and/or dunites, and the second trend as the signature of the initial mantle partial melting, with little  
61 or no overprint from melt/rock reaction events.

## 62 1 Introduction

63 Melts play a fundamental role in generating lithospheric mantle chemical and mineralogical  
64 heterogeneities, and have a large effect on mantle rheology, viscosity and seismic anisotropy  
65 (Batanova & Savelieva, 2009; Kelemen et al., 1997; Tommasi & Vauchez, 2015). Numerous  
66 studies have been dedicated to melt-rock interaction characterization in both the continental and  
67 the oceanic lithospheric mantle (e.g. Bodinier et al., 1990; Dalton et al., 2017; Dygert et al., 2016;  
68 Godard et al., 2008; Kelemen et al., 1998, 1990; Kelemen & Ghiorso, 1986; Parkinson & Pearce,  
69 1998; Takazawa et al., 1992; Vauchez et al., 2005; Morgan et al., 2008; Navon & Stolper, 1987;  
70 Niu, 1997; Warren et al., 2009; Warren and Shimizu, 2010). Several studies demonstrated that  
71 trace element variations coupled with microstructural, mineralogical and petrological  
72 observations, and trace element numerical modeling, are a pertinent way to evaluate melt transport  
73 and constrain melt-peridotite processes (e.g. Navon & Stolper, 1987; Batanova et al., 1998; Godard  
74 et al., 1995; Kelemen et al., 1995; Kelemen & Ghiorso, 1986; Kourim et al., 2014; Oliveira et al.,  
75 2020). Despite all these studies, the nature of melt and/or fluids involved in the reactional processes  
76 in the oceanic upper mantle below spreading centers remains debated. One of the biggest  
77 challenges to understanding these processes is finding locations where samples that were affected

78 by either incomplete melt extraction or interaction with melt with samples that were not affected  
79 by either of these processes. Abyssal peridotites (e.g. Godard et al., 2008; Johnson et al., 1990;  
80 Niu, 1997; Parkinson & Pearce, 1998) and mantle xenoliths (e.g. Bedini & Bodinier, 1999; Dalton  
81 et al., 2017; Fitzpayne et al., 2018; Grégoire et al., 2001) are good candidates to study mantle  
82 processes in present-day oceanic and continental settings, but their sampling is sparse, lacking  
83 second-order geologic context and is limited to the uppermost oceanic and continental mantle.

84 Oman ophiolite has been instrumental in elucidating the accretion and evolution of oceanic  
85 lithosphere in present-day oceans and exhibits the largest ophiolitic exposures of oceanic  
86 lithosphere worldwide. The mantle section of the Oman ophiolite is mainly composed of depleted  
87 harzburgites and of some dunites, and has been the subject of many petrological, geochemical and  
88 structural studies (e.g. Boudier & Coleman, 1981; Ceuleneer et al., 1988; Dygert et al., 2017;  
89 Kelemen et al., 1995; Godard et al., 2000; Le Mée et al., 2004; Monnier et al., 2006; Nicolas et  
90 al., 2000; Takazawa et al., 2003). The general consensus stands that the dunites, as channels in the  
91 mantle section or massive at the crust-mantle transition, are residues of reaction between a melt  
92 undersaturated in silica at low pressure and mantle harzburgites; this reaction leads to the complete  
93 consumption of orthopyroxene and to the concomitant precipitation of olivine (e.g. Abily &  
94 Ceuleneer, 2013; Boudier & Nicolas, 1995; Braun et al., 2002; Godard et al., 2000; Kelemen et  
95 al., 1995, 1997; Koga et al., 2001; Quick, 1981b; Rabinowicz et al., 1987; Rospabé et al., 2017,  
96 2018a, 2019a). However, the relationship between the harzburgites and the dunites, the nature of  
97 the reactant melt, and the chemical budgets related to the ‘dunitization’ process itself, are still  
98 debated. Oman ophiolite exposes large portions of the mantle and crust-mantle transition zone,  
99 suitable to understanding local to large scale studies of mantle heterogeneities and melt/peridotite  
100 reaction processes. The Oman Drilling Project (OmanDP) enabled sampling of a continuous  
101 section of the crust-mantle transition at Holes CM1A and CM2B (Wadi Tayin Massif, during  
102 Phase 2 of the ICDP OmanDP, Nov. 2017-Jan. 2018), starting from the base of the layered  
103 gabbroic crust and going through the uppermost harzburgitic mantle (Kelemen et al., 2020a,  
104 2020b; Proceedings available at <https://www.omandrilling.ac.uk/>). In this paper, we characterize  
105 the major and trace element contents of the dunites and harzburgites from the Hole CM1A and  
106 CM2B drill cores to better constrain dunitization processes by first, gaining insights into Wadi  
107 Tayin mantle and crust-mantle geochemical characteristics, then, comparing these characteristics  
108 to previously studied Maqсад diapir harzburgites and mantle-crust transition zone, taking  
109 advantage of the continuous and regular, high resolution sampling performed in the Oman Drilling  
110 Project.

## 111 **2 Geological setting and context of CM drill cores**

### 112 **2.1 Geology of the Samail ophiolite**

113 The Samail ophiolite, located in the Sultanate of Oman and the United Arab Emirates (Fig. 1a),  
114 exposes a relatively continuous section of oceanic lithosphere, with, from top-to-depth, a 5-7 km-  
115 thick crust made of pillow basalts, a sheeted dike complex and gabbros, overlying the crust-mantle  
116 transition at the top of the upper mantle peridotites (e.g. Coleman & Hopson, 1981; Glennie et al.,  
117 1974; Searle & Malpas, 1980, Lippard et al., 1986; Nicolas et al., 1988; Nicolas, 2012 and  
118 references therein). According to the ages of pelagic sediments interbedded with basalts and of  
119 zircons in evolved gabbroic and plagiogranitic rocks, the accretion event that led to the formation  
120 of the Oman ophiolite has been estimated at around 94-97 Ma ago (Rioux et., 2012, 2013, 2016;

121 Tilton et al., 1981; Tippit et al., 1981; Warren et al., 2005). The tectonic setting in which the Samail  
122 ophiolite evolved is still debated. The spatial distribution along the ophiolite of the nature of the  
123 (1) mafic dikes cutting-across the mantle section, and (2) lower crustal cumulates, as well as their  
124 geochemical signature, attest that both MORB like and depleted calc-alkaline series coexisted  
125 during the igneous evolution of the ophiolite (e.g. Benoit et al., 1996; 1999; Ceuleneer et al., 1996;  
126 Clénet et al., 2010; Python & Ceuleneer, 2003; Python et al., 2008). The dikes belonging to the  
127 MORB-like volcanic units have been mapped mainly in the SE of the ophiolite and in other more  
128 restricted spots, whereas the depleted calc-alkaline series were observed at a more widespread  
129 scale (Python & Ceuleneer, 2003).

130  
131 Along the Samail ophiolite, spatially constrained vertical flow structures frozen within the  
132 mantle section were interpreted as former asthenospheric diapirs distributed along the oceanic  
133 ridge (Ceuleneer, 1991; Ceuleneer et al., 1988; Jousselein et al., 1998; Nicolas et al., 1988, 2000).  
134 This mantle section is mainly composed of harzburgites (85 to 95%), relatively depleted with a  
135 typical orthopyroxene content of 15-25% and locally grading into lherzolites, and to a lesser extent  
136 of dunites (5-15%) (Boudier & Coleman, 1981; Lippard et al., 1986). The crust-mantle transition  
137 is mainly made of dunites and wehrlites and its thickness varies from a few meters to a few hundred  
138 meters (e.g. Abily & Ceuleneer, 2013; Boudier & Nicolas, 1995; Ceuleneer & Nicolas, 1985;  
139 Jousselein & Nicolas, 2000; Koga et al., 2001; Rospabé et al., 2017, 2018a).

140  
141 Geochemical studies have demonstrated the overprint of partial melting and of  
142 melt/peridotite reaction processes in the mantle harzburgites' signatures (Gerbert-Gaillard, 2002;  
143 Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010; Kanke & Takazawa, 2014; Khedr  
144 et al., 2014; Le Mée et al., 2004; Monnier et al., 2006; Takazawa et al., 2003). In this context,  
145 mantle dunites and dunites from the crust-mantle transition zone (CMTZ) have mostly been  
146 interpreted as replacive in origin, products of melt-harzburgite reaction leading to the complete  
147 consumption of orthopyroxene and concomitant precipitation of olivine (e.g. Abily & Ceuleneer,  
148 2013; Boudier & Nicolas, 1995; Gerbert-Gaillard, 2002; Godard et al., 2000; Kelemen et al., 1995,  
149 1997; Koga et al., 2001; Rabinowicz et al., 1987; Rospabé et al., 2018a). This dunitization process  
150 may have been enhanced by the involvement of a hydrous component in the reaction (Rospabé et  
151 al., 2017, 2018a, 2019a). However, a reaction origin and a cumulate origin are not mutually  
152 exclusive as it has been shown that the uppermost part (~ 20%) of the crust-mantle transition may  
153 have a composition consistent with cumulates while the main lower part (~80%) has a composition  
154 supporting the replacive origin (Abily & Ceuleneer, 2013). Furthermore, as olivine-saturated melt  
155 begins to cool conductively, hybrid processes, termed 'relative crystallization' (Collier & Kelemen  
156 2010) produce reactive characteristics (e.g. Benn et al., 1988; Boudier and Nicolas, 1995; Koga et  
157 al., 2001; Abily and Ceuleneer, 2013; Rospabé et al., 2018a).

## 158 2.2 The crust-mantle transition at Sites CM1 and CM2

159 Samples studied in this paper were drilled in the Wadi Tayin massif in the SE of the ophiolite  
160 during Phase 2 of the ICDP OmanDP (Nov. 2017-Jan. 2018). According to structural and  
161 petrological maps (Gerbert-Gaillard, 2002; Python & Ceuleneer, 2003; Nicolas et al., 2000), this  
162 site is located near the border - or in an intermediate position between the border and the axis of  
163 the frozen paleo-spreading center centered on the Maqsad (Sumail massif) paleo-mantle diapir -  
164 of the MORB segment in this part of the ophiolite. The crust-mantle transition zone (CMTZ) is  
165 relatively well exposed in this area, with a clear transition from harzburgite to the north to dunites

166 then gabbros to the south (Fig. 1b). The two sites CM1 and CM2, separated by about 400 m, have  
167 been drilled twice: with one Hole for core recovery and a second wider Hole for geophysical  
168 logging. At these sites 400 m of core was recovered from CM1A and about 300 m of core was  
169 recovered from CM2B.

170  
171 The geological map produced by the OmanDP group during the preparation of Phases 1 and 2  
172 shows a general tilt of the units by about 30° to the south (Fig. 1c). Considering these petrological  
173 and structural configurations, Hole CM1A borehole was cored 400 m with an inclination of 60°  
174 trending to the north, in order to cut perpendicularly across the mantle-crust transition (Fig. 1d). It  
175 crosses from the gabbroic lower crust (~150 m; the Layered Gabbro “Crustal Sequence”, CS),  
176 through the dunite-rich crust-mantle transition zone (~150 m CMTZ) that includes the Dunite (DS)  
177 and Dunite with Gabbro Sequences (DGS), to the residual upper mantle harzburgites (~100 m,  
178 Mantle Sequence, MS). At Site CM2, the fully cored borehole CM2B is vertical, parallel to the  
179 wider rotary borehole for geophysical logging (Hole CM2A). Hole CM2B starts within the crust-  
180 mantle transition zone (~110 m) and extends deeper in the underlying residual mantle peridotites  
181 than Hole CM1A (~180 m) (Fig. 1d). The main rock types sampled in Holes CM1A and CM2B  
182 are olivine-gabbro, gabbro, dunite, harzburgite and wehrlite, associated with minor gabbro  
183 troctolite, websterite, anorthosite, and chromitite layers (Fig. 1d). The crust-mantle transition zone  
184 sampled in Hole CM1A and CM2B has been divided into two parts according to the rock types  
185 present: the upper half is mainly made of dunites containing rare melt migration features (DS for  
186 Dunite Sequence); in the lower half, the dunites alternate with thin bands containing a higher  
187 proportion of interstitial plagioclase (+/- clinopyroxene), which has been called the Dunite with  
188 Gabbro Sequence (DGS). In the present article we focus on the geochemical compositions of  
189 dunites from the crust sequence (CS), crust-mantle transition zone (CMTZ) and mantle sequence  
190 (MS) and of mantle harzburgites.

### 191 **3 Results**

192 The sample selection strategy and the analytical methods are detailed in [Supporting Information](#)  
193 (see also [Kelemen et al., 2020a, 2020b, 2020c](#)). In summary, one sample was taken every 10 m  
194 along Holes CM1A and CM2B to cover the entire crust-mantle transition and mantle sections.  
195 Additional samples were collected to better characterize some specific levels (e.g. to document  
196 local, minor lithologies). The samples were analyzed for their major (as well as volatile) and trace  
197 element compositions. Sample lithology, macroscopic and microscopic observations and mineral  
198 modes calculated from major elements are reported in [supplementary data table 1](#). Whole rock  
199 major and volatile element compositions are reported in [supplementary data table 2](#). Whole rock  
200 trace element compositions are reported in [supplementary data table 3](#).

#### 201 **3.1. Sample description**

202 The studied samples represent the ultramafic lithologies (harzburgites and dunites) of the mantle  
203 section (46 harzburgites and 12 dunites), the crust-mantle transition zone (45 dunites) and the  
204 crustal Layered Gabbro Sequence (2 dunites) drilled at Holes CM1A (46 samples) and CM2B (59  
205 samples). Lithological classifications were made based on macroscopic and microscopic  
206 observations, and mineral modes calculated from XRF measurements ([supplementary data table](#)  
207 [1](#)). Four rock groups have been defined, harzburgites (36), carbonate-bearing harzburgites (8), pure  
208 dunites (31) and impregnated dunites (29) (Fig. 1e, Fig. SD-1, [supplementary data table 1](#)).

209  
210 The harzburgites and carbonate-bearing harzburgites contain olivine (ol), orthopyroxene (opx) and  
211 minor spinel (sp) as primary minerals (Fig. 2 e, f, g, h, i, j, k and l, Fig. SD-1, supplementary data  
212 table 1), displaying a porphyroclastic texture. Olivine abundance (primary mode) ranges from 72%  
213 to 93% in the harzburgites and from 64% to 75% in carbonate-bearing harzburgites. The olivine  
214 typically show subhedral shape and equant habits. Orthopyroxene abundance ranges from 10 to 28  
215 % with mostly subhedral shapes; some orthopyroxene show sigmoidal crystal-plastic deformation  
216 features (Fig. 2i and j). Grains of spinel are present in all samples (up to 1-3%). Compared to the  
217 dunites, the harzburgites are less altered, the alteration becoming complete only in highly veined  
218 zones and at the bottom of both Holes where carbonate (carb)/serpentine (serp) associations occur  
219 (Fig. 2k and l). The most abundant minerals in the harzburgite background alteration are serpentine  
220 and magnetite (mag). The carbonate alteration occurs at the bottom of both Holes starting at  
221 388.3m depth in CM1A and at 279.5m depth in CM2B. Harzburgite sample CM2B 129Z1 5-10  
222 cm (depth 299.7 m) is the most carbonate vein-rich harzburgite. Secondary amphibole (amph),  
223 chlorite (chl) and hydrogrossular are also present as trace alteration minerals (abundance < 3%,  
224 supplementary data table 1). No patch or deformation features specifically related to the alteration  
225 were observed.

226  
227 The pure dunites are primarily composed of ol >97% and generally contain <1 % sp (Fig. 2a, b, c  
228 and d, supplementary data table1). Mineral modes calculated using bulk rock data (Fig. SD-1,  
229 supplementary data table1) indicate that many dunites have >10% normative pyroxene. The  
230 macroscopic and microscopic descriptions indicate that some dunites contain plagioclase (pl)  
231 and/or pyroxenes. We refer to this group of dunites as ‘impregnated dunites’ in contrast to the pure  
232 dunites containing only ol and sp. The ol in almost all samples has been completely replaced by  
233 serp (Fig. 2a, b, c and d). In many dunites and in impregnated dunites, no relics of porphyroclastic  
234 opx are present, indicating no textural relics of a porphyroclastic texture. The pure and impregnated  
235 dunites’ primary texture was a fine- to medium-grained granular microstructure, characterized by  
236 euhedral ol forming a mosaic of equigranular grain size distribution as preserved by the equant  
237 contacts between serp mesh cores (Fig. 2a, b, c and d). Alteration minerals mainly consist of serp  
238 and mag (Fig. 2a, b, c and d). In addition, brucite (brc) after ol was detected by X-Ray Diffraction  
239 (XRD, performed during the ChikyuOman 2018 Leg 3) in some serpentinized dunites from the  
240 crust-mantle transition zone and in dunites from the mantle sequence (absent in harzburgites).  
241 Where relics of ol are present, in rare cases, they are surrounded by serp and mag. If ol is  
242 completely serpentinized, the mesh cores are mainly composed of serp with minor mag and  
243 accessory grains of sulfides (These sulfides are equant and anhedral, they have grain sizes of 0.1–1  
244 mm, modes of 0.5%, and are very rich in Ni, Kelemen et al., 2020a, 2020b - see also Lorand, 1988 for  
245 an article dedicated the the study of sulfides in Oman peridotites). Serpentine and more abundant  
246 magnetite at mesh rims trace former ol grain and sub-grain boundaries (Fig. 2a, b, c and d).

### 247 3.2. Loss on ignition, CO<sub>2</sub> and H<sub>2</sub>O contents

248 Samples from Holes CM1A and CM2B display high loss on ignition (LOI) values. The LOI varies  
249 from 8.29 to 14.92 wt.% in the harzburgites, from 9.02 to 23 wt.% in the carbonate-bearing  
250 harzburgites, from 10.14 to 15.31 in the pure dunites and from 6.14 to 15.53 wt.% in the  
251 impregnated dunites (Fig. 3 and supplementary table 2). The averaged H<sub>2</sub>O concentration is 12.0  
252 ± 1.7 wt.% in harzburgites, 11.0 ± 5.5 wt.% in carbonate-bearing harzburgites, 14.3 ± 1.2 wt.% in

253 dunites and  $13.7 \pm 1.6$  wt.% in impregnated dunites. The LOI values correlate with measured water  
 254 concentration (H<sub>2</sub>O), which slightly decreases downhole (Fig. 3). The concentrations of CO<sub>2</sub>  
 255 measured in harzburgites, carbonate-bearing harzburgites excluding the carbonate vein-rich  
 256 harzburgite CM2B 129Z1 5-10 cm (depth 299.7 m, CO<sub>2</sub> = 19.54 wt.%), pure dunites and  
 257 impregnated dunites vary from 0.12 to 0.31 wt.%, from 0.36 to 1.30 wt.%, from 0.13 to 0.34 wt.%  
 258 and from 0.03 to 0.41 wt.% respectively. The averaged CaCO<sub>3</sub> concentration excluding the  
 259 carbonate-rich harzburgite mentioned previously (22.77 wt.%), is 0.40 wt.% in harzburgites, 1.57  
 260 wt.% in carbonate-bearing harzburgites, 0.51 wt.% in pure dunites and 0.46 wt.% in impregnated  
 261 dunites. The downhole profile of CaCO<sub>3</sub> shows an increase in concentration in the deepest part of  
 262 Holes CM1A and CM2B, with recovered harzburgites having higher CO<sub>2</sub> concentrations, LOI and  
 263 H<sub>2</sub>O contents, consistent with the particularly high alteration degree and high carbonate content in  
 264 these samples (serp, brc, carb, see Fig. 3 XRD).

### 265 3.3. Whole rock major element compositions

266 Major element analyses were performed on 36 harzburgites, 8 carbonate-bearing harzburgites, 31  
 267 pure dunites, and 29 impregnated dunites sampled along Holes CM1A and CM2B.

#### 268 *The mantle section sequence (MS)*

269 Harzburgite is the most abundant lithology in the MS (36 harzburgites analysed) followed by  
 270 impregnated dunite (9 impregnated dunites analysed) and pure dunite (4 pure dunites analysed).  
 271 Mg# (Mg# =  $100 \times \text{molar Mg}/(\text{Mg} + \text{Fe}_{\text{total}})$ ) in mantle harzburgites is between 90.7 and 92.4. CaO,  
 272 Al<sub>2</sub>O<sub>3</sub>, and TiO<sub>2</sub> contents range from 0.19 to 2.07 wt.%, 0.53 to 0.91 wt.%, and 0.02 to 0.04 wt.%,  
 273 respectively (Fig. 4). Harzburgite with dunite patches contains lower CaO concentrations than the  
 274 average value (0.30 to 0.81 wt.%). Carbonate-bearing harzburgites are characterized by similar  
 275 Mg#, Al<sub>2</sub>O<sub>3</sub>, and TiO<sub>2</sub> contents (90.6, 0.70 wt.% and 0.03 wt.% on average respectively), and very  
 276 high CaO concentrations ranging from 0.68 to 5.42 wt.% compared to carbonate-free harzburgites.  
 277 Dunites from the MS have high Mg#, on average 90.2 for pure dunites and 91.1 for impregnated  
 278 dunites (Fig. 4). The pure dunites have CaO, Al<sub>2</sub>O<sub>3</sub> and TiO<sub>2</sub> contents from 0.30 to 0.94 wt.%,  
 279 0.32 to 0.72 wt.%, and 0.02 to 0.05 wt.% respectively, where the impregnated dunites have CaO  
 280 contents ranging from 0.14 to 1.59 wt.%, Al<sub>2</sub>O<sub>3</sub> from 0.21 to 0.95 wt.% and TiO<sub>2</sub> from 0.02 to  
 281 0.04 wt.% (Fig. 4). Along the MS, TiO<sub>2</sub> and Al<sub>2</sub>O<sub>3</sub> contents do not show any systematic variation  
 282 with depth. On the other hand, the CaO content shows some variations downhole. The vertical  
 283 evolution of the CaO content in harzburgites from Hole CM1A is different from other elements;  
 284 successive trends of increasing and decreasing CaO with depth form a well-defined zigzag pattern  
 285 In detail, it increases from 0.97 to 2.1 wt. % between 311 and 340 m and from 0.54 to 1.8 wt.%  
 286 between 360 and 388 m, and decreases from 2.1 to 0.54 wt. % between 340 and 360 m then from  
 287 1.8 to 0.8 wt.% at most from 388 to around 400 m. CaO contents in dunites and nearby harzburgites  
 288 are correlated. Downhole intervals with the highest CaO contents are characterized by high  
 289 carbonate vein concentrations (Fig. 3 XRD, and Fig. 4).

#### 290 *The crust-mantle transition zone sequence (CMTZ)*

291 The CMTZ is composed mainly of pure and impregnated dunites (Fig. 1d). 27 pure dunites and 18  
 292 impregnated dunites from Holes CM1A and CM2B CMTZ were analyzed. Both pure and  
 293 impregnated dunites from the CMTZ display slightly lower Mg# (89.9 on average, Fig. 4) and



294 CaO concentrations (pure dunites = 0.16 wt.%, and impregnated dunites = 0.28 wt.% on average,  
 295 [Fig. 4](#)) compared to the dunites from the MS. Some dunites from CM1A DGS (between 252.96  
 296 and 271.42 m) display lower Mg# (Mg# = 86.8-88.0) that seems mostly controlled by the increase  
 297 in the FeO content. The TiO<sub>2</sub> contents in pure and impregnated dunites from the CMTZ (0.04 wt.%  
 298 on average) are slightly higher than the TiO<sub>2</sub> contents in the MS dunites (0.03 wt.% on average).  
 299 Al<sub>2</sub>O<sub>3</sub> concentrations are lower in CMTZ pure dunites (0.42 wt.% on average) and higher in  
 300 impregnated dunites (1.12 wt.% on average) compared to pure (0.53 wt.% on average) and  
 301 impregnated (0.51 wt.% on average) dunites respectively from the MS. A similar zigzag pattern to  
 302 the one observed in CaO along the MS is irregularly observed in CMTZ dunites (i.e. in CM1A,  
 303 decreasing from 0.17 to 0.12 wt. % between 170 and 245 m and, increasing from 0.08 to 0.52 wt.  
 304 % between 253 and 310 m, [Fig. 4](#)). This CaO zigzag variation in CM1A CMTZ dunites is  
 305 associated with Mg# zigzag variation (decreasing from 91 to 89 wt. % between 170 and 245 m  
 306 and, increasing from 87 to 91 wt. % between 253 and 310 m, [Fig. 4](#)). The dunites' CaO varies  
 307 over ~ 20 m at the base of the CMTZ following the mantle harzburgites zigzag variation, whereas  
 308 the Mg# varies over ~ 60 m, along with the FeO variations.

### 309 *The crustal sequence (CS)*

310 The two analyzed dunites from the Layered Gabbro Sequence (LGS) (C5707A-51Z-1 W, 31.0-  
 311 39.0 cm, 125.60 m in depth, and C5707A-58Z-2 W, 1.0-6.0 cm, 143.93 m in depth) are  
 312 impregnated ([Fig. 4](#)), they have relatively low Mg# (85.6 and 85.5 respectively), compared to  
 313 dunites in the other sequences. This tendency is similar to the dunites from CM1A DGS. CaO  
 314 (0.04 wt.% and 0.19 wt.% respectively), Al<sub>2</sub>O<sub>3</sub> (0.15 wt.% and 0.93 wt.% respectively) and TiO<sub>2</sub>  
 315 (0.03 wt.%) contents are similar to the impregnated dunites values from the MS and the CMTZ  
 316 ([Fig. 4](#)).

317 CM Holes harzburgites, carbonate-bearing harzburgites, pure dunites and impregnated dunites  
 318 (except CM1A LGS impregnated dunites) show comparable major element compositions to  
 319 previously reported harzburgites, dunites and impregnated dunites from the Oman ophiolite mantle  
 320 and crust-mantle transition zones ([Fig. 5](#)). Most of the pure and impregnated dunites from MS, and  
 321 CMTZ plot above the terrestrial array (Earth differentiation trend, [Jagoutz et al., 1979](#)) in the  
 322 MgO/SiO<sub>2</sub> vs. Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> diagram, similar to previously studied dunites and impregnated dunites  
 323 from the mantle section of Wadi Tayin massif ([Godard et al., 2000](#); [Hanghøj et al., 2010](#)) and the  
 324 CMTZ at the top of Maqсад diapir ([Rospabé et al., 2018a, 2019a](#)) ([Fig. 5a](#)). Whereas the  
 325 harzburgites, carbonate-bearing harzburgites and some pure and impregnated dunites plot below.  
 326 The MgO-FeO variations in most of the CM dunites mimic the stoichiometric variation of the ol  
 327 Mg-Fe composition, similar to other pure/slightly impregnated dunites elsewhere, while only a  
 328 few CM samples fall in the domain of the more highly impregnated dunites. Lower CaO and wider  
 329 range of Al<sub>2</sub>O<sub>3</sub> values characterize all the dunites compared to the harzburgites ([Fig. 5](#)). Similar to  
 330 other previously studied Wadi Tayin massif and Maqсад diapir harzburgites ([Godard et al., 2000](#);  
 331 [Hanghøj et al., 2010](#); [Monnier et al., 2006](#)) and most refractory abyssal peridotites ([Godard et al.,](#)  
 332 [2008](#); [Niu, 1997](#); [Warren et al., 2009](#)), CM harzburgites plot near the most depleted end of the  
 333 mantle fractionation array ([Fig. 5a](#)). They display similar low Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> ratios in comparison to  
 334 other harzburgites from Wadi Tayin and Maqсад harzburgites (0.01-0.02, 0.01-0.04 and 0.01-0.08  
 335 respectively), and a high MgO/SiO<sub>2</sub> ratios typical of refractory peridotites (0.98-1.10, 0.95-1.10  
 336 and 0.96-1.15 respectively). Oman harzburgites are characterized by narrow FeO and MgO  
 337 contents compared to the pure and impregnated dunites ([Fig. 5b](#)), with slightly higher Mg# in CM

338 harzburgites (90.7-92.4) compared to Wadi Tayin and Maqsad harzburgites (89.6-91.5 and 88.4-  
 339 91.1 respectively). Generally,  $\text{Al}_2\text{O}_3$  and CaO show broad positive correlation in harzburgites from  
 340 Wadi Tayin massif and Maqsad diapir (Fig. 5c), Godard et al, (2000) demonstrating that the  
 341 observed  $\text{Al}_2\text{O}_3/\text{CaO}$  ratio variability displayed by Wadi Tayin and Maqsad harzburgites decreases  
 342 with increasing cpx content in the main harzburgites sequence to lower values in the cpx-  
 343 harzburgites at the base of the mantle section.  $\text{Al}_2\text{O}_3/\text{CaO}$  ratio variability is also observed along  
 344 cores CM1A and CM2B, however CaO variability at the bottom of CM Holes is related to  $\text{CO}_2$ -  
 345 bearing fluids interactions with CM harzburgites (see section 4.1.).

346

### 347 **3.4. Whole rock trace element contents**

348 Chondrite-normalized Rare Earth Element (REE) and primitive mantle-normalized trace element  
 349 variations and patterns are shown in Figures 6 and 7 respectively. Similar to other refractory  
 350 peridotites from the Oman ophiolite mantle section and crust-mantle transition, the studied  
 351 harzburgites, carbonate-bearing harzburgites, pure dunite and impregnated dunite whole rock  
 352 concentrations are lower than chondritic (CN) and primitive mantle (PM) values (Figs. 6 and 7).  
 353 Their REEs are characterized by three types of chondrite-normalized patterns: linear or flat linear,  
 354 U- or V-shaped, and concave-upward patterns. CM Holes show vertical trends, continuous over  
 355 tens of meters, in their REE content. These are characterized by ~ 50 m-thick alternations between  
 356 increasing and decreasing concentrations ('zigzag' patterns) (e.g. La, Yb, LREE patterns; Fig. 6).  
 357 These trends along CM Holes are observed in both dunites and harzburgites, independent of the  
 358 lithology, and the changes from one trend to another are commonly correlated with the presence  
 359 of faults described by the structural team during the ChikyuoOman2018 Leg 3 (Kelemen et al.,  
 360 2020a, 2020b).

#### 361 *The mantle section sequence*

362 The mantle harzburgites show two types of chondrite-normalized REE patterns: (1) 19  
 363 harzburgites display linear REE patterns characterized by a progressive depletion from heavy REE  
 364 (HREE) ( $\text{Gd}_{\text{CN}}/\text{Yb}_{\text{CN}} = 0.25 \pm 0.21$ ,  $\text{Yb}_{\text{CN}} = 0.22 \pm 0.11$ ) to middle REE (MREE: Sm, Eu and Gd)  
 365 ( $\text{Sm}_{\text{CN}} = 0.06 \pm 0.05$  and  $\text{Gd}_{\text{CN}} = 0.06 \pm 0.05$ ) and light (LREE) ( $\text{La}_{\text{CN}}/\text{Sm}_{\text{CN}} = 0.98 \pm 0.74$ ,  $\text{La}_{\text{CN}}$   
 366  $= 0.06 \pm 0.05$ ). 6 samples show a positive Eu anomaly ( $(\text{Eu}/\text{Eu}^*)_{\text{CN}} = 2.36 \pm 0.66$ , with  $(\text{Eu}/\text{Eu}^*)_{\text{CN}}$   
 367  $= \text{Eu}_{\text{CN}}/\sqrt{(\text{Sm}_{\text{CN}} \times \text{Gd}_{\text{CN}})}$ ). In detail, 12 harzburgites display linear LREE-depleted patterns (3 from  
 368 CM1A and 9 from CM2B) characterized by a progressive depletion from HREE to LREE, and 7  
 369 harzburgites (5 from CM1A and 2 from CM2B) display flat linear REE patterns characterized by  
 370 slightly lower LREE concentrations compared to HREE concentrations. (2) 15 harzburgites (4  
 371 from CM1A and 11 from CM2B) display U- or V-shaped REE patterns reflecting significant  
 372 MREE depletion relative to LREE ( $\text{La}_{\text{CN}}/\text{Sm}_{\text{CN}} = 3.21 \pm 1.18$ ) and HREE ( $\text{Gd}_{\text{CN}}/\text{Yb}_{\text{CN}} = 0.07 \pm$   
 373  $0.06$ ). 7 samples show a positive Eu anomaly ( $(\text{Eu}/\text{Eu}^*)_{\text{CN}} = 3.36 \pm 1.69$ ).

374 The mantle pure dunites and impregnated dunites show two types of chondrite-normalized REE  
 375 patterns: linear REE patterns (3 pure dunites from CM2B and 6 impregnated dunites (3 from  
 376 CM1A and 3 from CM2B)) and U- or V-shaped REE patterns (5 pure dunites and 1 impregnated  
 377 dunites from CM2B). In detail the linear REE pattern are subdivided to: (1) flat linear REE patterns  
 378 displayed by the three CM1A impregnated dunites, they are characterized by roughly similar

379 LREE ( $\text{La}_{\text{CN}} = 0.14 \pm 0.07$ ), MREE ( $\text{Sm}_{\text{CN}} = 0.10 \pm 0.05$ ), and HREE ( $\text{Yb}_{\text{CN}} = 0.18 \pm 0.05$ )  
 380 concentrations, together with a positive Eu anomaly ( $(\text{Eu}/\text{Eu}^*)_{\text{CN}} = 2.06 \pm 0.65$ ). (2) LREE-  
 381 depleted linear REE patterns displayed by 3 pure dunites and 3 impregnated dunites from CM2B  
 382 show a progressive depletion from HREE ( $\text{Gd}_{\text{CN}}/\text{Yb}_{\text{CN}} = 0.25 \pm 0.16$  and  $0.18 \pm 0.04$ ,  $\text{Yb}_{\text{CN}} = 0.22$   
 383  $\pm 0.05$  and  $0.28 \pm 0.07$  respectively) to MREE ( $\text{Sm}_{\text{CN}} = 0.04 \pm 0.03$  and  $0.04 \pm 0.02$ ,  $\text{Gd}_{\text{CN}} = 0.06$   
 384  $\pm 0.04$  and  $0.05 \pm 0.02$  respectively) to LREE ( $\text{La}_{\text{CN}}/\text{Sm}_{\text{CN}} = 0.71 \pm 0.28$  and  $0.37 \pm 0.23$ ,  $\text{La}_{\text{CN}} =$   
 385  $0.03 \pm 0.02$  and  $0.02 \pm 0.01$  respectively), 1 sample shows a positive Eu anomaly  $(\text{Eu}/\text{Eu}^*)_{\text{CN}} =$   
 386  $1.60$ ). The U- or V-shaped REE patterns (5 pure dunites and 1 impregnated dunites from CM2B)  
 387 are characterized by significant MREE depletion relative to LREE ( $\text{La}_{\text{CN}}/\text{Sm}_{\text{CN}} = 2.66 \pm 1.76$ ) and  
 388 HREE ( $\text{Gd}_{\text{CN}}/\text{Yb}_{\text{CN}} = 0.08 \pm 0.12$ ), 2 pure dunites display a positive Eu anomaly  $(\text{Eu}/\text{Eu}^*)_{\text{CN}} =$   
 389  $2.91$  and  $2.54$ ).

### 390 *The crustal-mantle transition zone sequence (CMTZ)*

391 The pure dunites from the CMTZ show two types of chondrite-normalized REE patterns: (1) U-  
 392 shaped REE pattern displayed by 11 pure dunites (6 from CM1A and 5 from CM2B), with MREE  
 393 depletion relative to LREE ( $\text{La}_{\text{CN}}/\text{Sm}_{\text{CN}} < 3.34$ ) and HREE ( $0.07 < \text{Gd}_{\text{CN}}/\text{Yb}_{\text{CN}} < 0.27$ ). Some  
 394 samples also have positive Eu anomalies (4 samples from CM1A and 3 samples from CM2B,  
 395  $(\text{Eu}/\text{Eu}^*)_{\text{CN}} = 3.62 \pm 2.49$ ). (2) Linear LREE-depleted or slightly concave-upward REE patterns  
 396 displayed by 10 pure dunites from Hole CM1A ( $\text{La}_{\text{CN}}/\text{Yb}_{\text{CN}} = 0.07 \pm 0.05$ ) with similar HREE  
 397 concentrations ( $\text{Yb}_{\text{CN}} = 0.13 \pm 0.02$ ) and lower LREE and MREE variation compared to HREE  
 398 ( $\text{La}_{\text{CN}} = 0.01 \pm 0.01$ ;  $\text{Gd}_{\text{CN}} = 0.02 \pm 0.01$ ). Most samples display a positive Eu anomaly (6 samples,  
 399  $(\text{Eu}/\text{Eu}^*)_{\text{CN}} = 4.43 \pm 2.44$ ).

400 The CMTZ impregnated dunites display three types of chondrite-normalized REE patterns: (1)  
 401 The linear REE patterns displayed by 5 impregnated dunites (2 from CM1A and 3 from CM2B)  
 402 are characterized by a progressive depletion from HREE ( $\text{Gd}_{\text{CN}}/\text{Yb}_{\text{CN}} = 0.38 \pm 0.20$ ,  $\text{Yb}_{\text{CN}} = 0.23$   
 403  $\pm 0.09$ ) to MREE ( $\text{Sm}_{\text{CN}} = 0.07 \pm 0.06$  and  $\text{Gd}_{\text{CN}} = 0.10 \pm 0.09$ ) to LREE ( $\text{La}_{\text{CN}}/\text{Sm}_{\text{CN}} = 0.77 \pm$   
 404  $0.19$ ,  $\text{La}_{\text{CN}} = 0.05 \pm 0.06$ ), 2 samples show a positive and negative Eu anomaly  $(\text{Eu}/\text{Eu}^*)_{\text{CN}} = 2.23$   
 405 and  $0.45$  respectively). (2) 3 samples from CM2B display U-shaped REE pattern, with MREE  
 406 depletion relative to LREE ( $\text{La}_{\text{CN}}/\text{Sm}_{\text{CN}} < 3.93$ ) and HREE ( $0.09 < \text{Gd}_{\text{CN}}/\text{Yb}_{\text{CN}} < 0.25$ ). (3) One  
 407 sample from CM1A and three samples from CM2B display concave-upward patterns characterized  
 408 by a nearly flat slope of the HREE segment ( $\text{Gd}_{\text{CN}}/\text{Yb}_{\text{CN}} = 0.86 \pm 0.39$ ) followed by a progressive  
 409 depletion from MREE to LREE ( $\text{La}_{\text{CN}}/\text{Sm}_{\text{CN}} = 0.13 \pm 0.07$ ). One CM1A impregnated dunite shows  
 410 a negative Eu anomaly  $(\text{Eu}/\text{Eu}^*)_{\text{CN}} = 0.42$ ).

### 411 *The crustal sequence (CS)*

412 The two impregnated dunites from the Hole CM1A CS (C5707A-51Z-1 W, 31-39.0 cm, 125.60 m  
 413 depth, and C5707A-58Z-2 W, 1.0-6.0 cm, 143.94 m depth) display relatively linear (REE) patterns  
 414 characterized by a steady decrease of REE abundances HREE to LREE, as well as by a clear  
 415 positive Eu anomaly  $(\text{Eu}/\text{Eu}^*)_{\text{CN}} = 3.8$ - $5.0$ ).

416 The PM-normalized multi-element patterns of most harzburgites and carbonate-bearing  
 417 harzburgites exhibit strong to moderate enrichments in LILE, Th, U, Nb and Ta relative to LREE  
 418 (e.g. averaged  $\text{Rb}_{\text{PMN}}/\text{La}_{\text{PMN}} = 12.95$  and  $18.24$ ;  $\text{U}_{\text{PMN}}/\text{La}_{\text{PMN}} = 4.64$  and  $4.45$ ;  $\text{Nb}_{\text{PMN}}/\text{La}_{\text{PMN}} = 2.02$   
 419 and  $4.30$  respectively). The carbonate-bearing harzburgites display stronger Pb, Sr and Ti positive

420 anomalies (averaged  $Pb_{PMN}/Ce_{PMN} = 145.23$ ;  $Sr_{PMN}/Nd_{PMN} = 373.50$ ;  $Ti_{PMN}/Gd_{PMN} = 11.09$ )  
 421 compared to the harzburgite (averaged  $Pb_{PMN}/Ce_{PMN} = 25.83$ ;  $Sr_{PMN}/Nd_{PMN} = 30.21$ ,  $Ti_{PMN}/Gd_{PMN}$   
 422  $= 3.52$ ). Most pure dunites show similar enrichments exhibited by the harzburgites and carbonate-  
 423 bearing harzburgites in LILE, Th, U, Nb and Ta relative to LREE (e.g. averaged  $Rb_{PMN}/La_{PMN} =$   
 424  $20.0$ ;  $U_{PMN}/La_{PMN} = 5.12$ ;  $Nb_{PMN}/La_{PMN} = 4.10$ ), but with smaller Pb and Sr positive anomalies  
 425 (averaged  $Pb_{PMN}/Ce_{PMN} = 19.68$ ;  $Sr_{PMN}/Nd_{PMN} = 29.05$ ). The pure dunites display stronger Ti  
 426 positive anomalies compared to the harzburgites and smaller compared to the carbonate-bearing  
 427 harzburgites ( $Ti_{PMN}/Gd_{PMN} = 8.98$ ). The impregnated dunites exhibit moderate LILE, Th, U, Nb  
 428 and Ta enrichments relative to LREE (e.g. averaged  $Rb_{PMN}/La_{PMN} = 3.94$ ;  $U_{PMN}/La_{PMN} = 3.12$ ;  
 429  $Nb_{PMN}/La_{PMN} = 1.05$ ), and the smallest Pb, Sr and Ti positive anomalies compared to all groups of  
 430 rocks (averaged  $Pb_{PMN}/Ce_{PMN} = 9.99$ ;  $Sr_{PMN}/Nd_{PMN} = 14.03$ ;  $Ti_{PMN}/Gd_{PMN} = 1.62$ ).

#### 431 **4 Discussion**

432 Most studies of the ultramafic rocks from the Oman ophiolite consider the dunites to be channels  
 433 in the mantle or massive at the crust-mantle transition, and replacive in origin. In this model, the  
 434 dunites represent residues of reaction between a melt undersaturated in silica at low pressure and  
 435 the host mantle harzburgites during a melt percolation event that led to the complete consumption  
 436 of orthopyroxene and to the concomitant precipitation of olivine (e.g. [Abily & Ceuleneer, 2013](#);  
 437 [Boudier & Nicolas, 1995](#); [Braun et al., 2002](#); [Godard et al., 2000](#); [Kelemen et al., 1995, 1997](#);  
 438 [Koga et al., 2001](#); [Rabinowicz et al., 1987](#); [Rospabé et al., 2017, 2018a, 2019a](#)). The hypothesis  
 439 of the replacive origin of the dunites, also proposed for other ophiolitic sections and for dunites  
 440 associated with abyssal peridotites (e.g. [Dick & Natland, 1996](#); [Kelemen et al., 1990](#); [Godard et](#)  
 441 [al., 2008](#); [Quick, 1981a, 1981b](#)), contrasts with an older cumulative model, in which ol  
 442 crystallization and accumulation created the lowermost part of the oceanic crust (e.g. [Elthon, 1979](#);  
 443 [O'Hara, 1965](#); [Smewing, 1981](#)). Alternatively, it has been proposed that the crust-mantle transition  
 444 of the Oman ophiolite may be of double origin, with the lower 80% as replacive and the upper  
 445 20% as cumulates ([Abily & Ceuleneer, 2013](#)). In the case of the OmanDP CM cores, the  
 446 alternations between dunites and mantle harzburgites at the top of the mantle sequence is consistent  
 447 with a melt/rock reaction origin, a feature observed at crust-mantle transitions of many massifs  
 448 along the Oman ophiolite (e.g. [Boudier & Nicolas, 1995](#)) and that cannot be accounted for by a  
 449 simple fractional crystallization process.

450 Structural and petrological mappings of the Oman ophiolite have revealed contrasting domains  
 451 along the ophiolite. Especially, the spatially varying nature and composition of the dikes cross-  
 452 cutting the mantle section reflect formation involving a MORB-like melt mainly in the south-  
 453 eastern Nakhil, Samail and Wadi Tayin massifs (troctolite and ol-gabbro dikes), contrasting with a  
 454 more widespread depleted, calc-alkaline magma composition elsewhere (mostly gabbro and  
 455 pyroxenite dikes) ([Python & Ceuleneer, 2003](#); [Python et al., 2008](#)). The MOR-like area  
 456 characterizes a NW-SE oriented paleo-spreading segment that seems to have developed within  
 457 older, already accreted lithosphere of depleted calc-alkaline affinity ([Ceuleneer et al., 1988; 1996](#);  
 458 [Gerbert-Gaillard, 2002](#); [Godard et al., 2000](#); [Nicolas et al., 2000](#); [Python & Ceuleneer, 2003](#)). This  
 459 MORB segment is hypothesized to have been centered on, and fed with melts by, the fossil mantle  
 460 diapir of the Maqсад area in the Samail massif ([Rabinowicz et al., 1987](#); [Ceuleneer et al., 1988](#);  
 461 [Jousselin et al., 1998](#)). According to the published structural and petrological maps, the drilling  
 462 site is located within the area where melts were MORB-like, near the NE limit of the paleo  
 463 spreading segment. This delimitation has been defined as the Makhbiyah shear zone in the Wadi

464 Tayin massif, making the contact with the older lithosphere (Nicolas & Boudier, 2008).  
465 Accordingly, it is reasonable to consider that a MORB-like melt played an important role in the  
466 formation of the CM dunites and harzburgites.

467

#### 468 **4.1 Effects of serpentinization and carbonation on the composition of dunites and** 469 **harzburgites**

470 In some cases, alteration, especially serpentinization of ultramafic rocks (sometimes associated  
471 with mineralization related to hydrothermal activity), may significantly modify bulk-rock  
472 chemical composition (e.g. Beinlich et al., 2020; de Obeso & Kelemen, 2018; Gruau et al., 1998;  
473 Hodel et al., 2018; Malvoisin, 2015; Paulick et al., 2006; Snow & Dick, 1995). As described above,  
474 the dunites and harzburgites from the crust-mantle transition and mantle section of Holes CM1A  
475 and CM2B are extensively serpentinized (Fig. 2), up to 100% in many samples. In addition, strong  
476 carbonate-veining affected parts of the cores, especially at the base of Hole CM2B where it is  
477 intensely faulted (carbonate-bearing harzburgites, Fig. 3, CO<sub>2</sub> and CaCO<sub>3</sub> logs). To decipher the  
478 effects of such significant fluid/rock interactions on the composition of OmanDP CM samples is  
479 critical (especially on the trace elements). It particularly concerns the dunites, because of their  
480 more altered character (perhaps related to higher ol mode), and because their trace element budget  
481 was more significantly controlled by primary ol and trace phases such as pyroxenes before  
482 alteration (unlike the harzburgites that contain abundant residual pyroxenes).

483 The LOI content is generally higher in carbonate-bearing harzburgite (averaged LOI = 13.04 ±  
484 9.57 wt.%), pure dunites (averaged LOI = 13.92 ± 1.21 wt.%) and impregnated dunites (averaged  
485 LOI = 13.33 ± 2.02 wt.%) than in harzburgites (averaged LOI = 11.91 ± 1.53 wt.%).  
486 Serpentinization of many of the studied samples appears to have led to enrichment in SiO<sub>2</sub> as  
487 already observed in abyssal peridotites affected by a Si addition, and/or magnesium loss (e.g. de  
488 Obeso & Kelemen, 2018; Paulick et al., 2006; Snow & Dick, 1995). This open system behavior is  
489 confirmed by the plot of some dunites (3 from CM2B) and impregnated dunites (4 from CM1A  
490 and 4 from CM2B) below the mantle fractionation array at the same field as the harzburgites and  
491 the carbonate-bearing harzburgites (Fig. 5a), suggesting MgO loss and/or SiO<sub>2</sub> enrichment as  
492 reported in pervasively serpentinized abyssal peridotites or talc-bearing serpentinites (de Obeso &  
493 Kelemen, 2018; Snow & Dick, 1995; Paulick et al., 2006). This may be the reason why there are  
494 elevated normative pyroxene modes in samples that were classified as dunites based on  
495 macroscopic (hand specimen) and microscopic (thin sections) observations (Fig. SD-1,  
496 supplementary data table 1). Some of these dunites have no pyroxenes or pyroxene pseudomorphs  
497 in thin section. This is supported by XRD analyses performed during ChikyuOman 2018 Phase 2  
498 Leg 3 which revealed the widespread occurrence of brucite associated with other alteration  
499 minerals (Kelemen et al., 2020a, 2020b). It also appears clear that the higher CaCO<sub>3</sub> (averaged  
500 CaCO<sub>3</sub> = 3.93 ± 10.85 wt.% compared to 0.40 ± 0.54 wt.% in harzburgites) together with higher  
501 CO<sub>2</sub> contents (averaged CO<sub>2</sub> = 2.82 ± 9.501 wt.% compared to 0.20 ± 0.25 wt.% in harzburgites)  
502 in the carbonate-bearing harzburgites at the base of CM1A and CM2B (4 from CM1A and 4 from  
503 CM2B) are related to carbonate-veins. However, the covariation of Ni and Co contents with the

504 XMg suggests that the possible precipitation of sulfides (Kelemen et al., 2020a, 2020b) related to  
505 these strong water/rock interactions did not erase the primary compositions in these elements.

506 The plots of the concentration of several trace elements as a function of the LOI (Fig. 8a-c) and of  
507 the CO<sub>2</sub> and CaCO<sub>3</sub> contents show no clear correlation. On the contrary, some reasonably good  
508 covariations are observed between Th and U, Nb and especially La on one hand (Fig. 8 h-j), and  
509 Yb, Ti and Hf on the other hand (Fig. 8 n-o); the Zr content is partially correlated with both Th  
510 and Yb (Fig. 8 f, k). Since Th and Ti are generally considered immobile during alteration processes  
511 (e.g. Kogiso et al., 1997; Niu, 2004; Paulick et al., 2006), these trends probably reflect one or more  
512 overlapping geochemical signatures acquired during high temperature, magmatic processes rather  
513 than during a later serpentinization event. The large ion lithophile elements as well as Li and Pb  
514 do not correlate with the LOI nor with other elements (Fig. 8 b, c, g, l, m), and their compositions  
515 may result from the overprint of several processes having operated over a large range of  
516 temperatures and conditions, from igneous to alteration events. Accordingly, only the  
517 concentrations in REE, HFSE and Th-U, will be used to discuss the igneous processes that led to  
518 the formation of the dunites from the crust-mantle transition zone and mantle sequence sampled  
519 by the CM Holes.

#### 520 4.2. Partial melting vs. melt-rock reaction in the Oman ophiolite mantle section

521 The mantle section of CM Holes (Wadi Tayin massif) is composed of refractory harzburgites with  
522 relatively homogeneous modal and major element compositions (excluding some major elements  
523 e.g. CaO and Na<sub>2</sub>O, Figs. 4 and 5) and more variable trace element contents (Figs. 6, 7a and 7b).  
524 Significant geochemical variability of the mantle peridotites has also been observed all along the  
525 Oman ophiolite. In some cases, previous geochemical studies inferred the overprint of partial  
526 melting and melt/peridotite reaction processes in the mantle harzburgite signatures (e.g. Gerbert-  
527 Gaillard, 2002; Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010; Kanke &  
528 Takazawa, 2014; Khedr et al., 2014; Le Mée et al., 2004; Monnier et al., 2006; Takazawa et al.,  
529 2003). Two different geochemical trends are combined in the trace element contents of all samples  
530 studied here. On one hand, Yb shows a good correlation with Ti and some HFSE (Fig. 8o). Heavy  
531 REE have been demonstrated to be less impacted than MREE and especially LREE during melt-  
532 peridotite reactions such as melt/rock re-equilibration during melt migration or in response to  
533 conversion of harzburgite to dunite (e.g. Godard et al., 1995; Kelemen et al., 1990; Navon &  
534 Stolper, 1987; Prinzhofer & Allègre, 1985; Spiegelman & Kelemen, 2003; Vernières et al., 1997;  
535 Rospabé et al., 2018a). Therefore, these correlations may have been formed during partial melting,  
536 with little or no overprint by subsequent melt-peridotite interactions at shallow depth. On the other  
537 hand, we observe a good correlation between Th, U and La, and partially with the HFSE (e.g. Nb,  
538 Fig. 8h-j), that does not correlate with Yb or Ti concentrations. As the U-/V-shape of the REE  
539 patterns, resulting from the selective enrichment in LREE relative to MREE and HREE in  
540 peridotites, may be attributed to chromatographic fractionation associated with interstitial melt  
541 percolation and/or to the transformation of harzburgite into dunite (e.g. Godard et al., 1995; Navon  
542 & Stolper, 1987; Prinzhofer & Allègre, 1985; Vernières et al., 1997; Rospabé et al., 2018a), we  
543 interpret this second geochemical signature as the overprint of melt-peridotite reaction processes.

##### 544 4.2.1. Geochemical logs

545 Vertical chemical trends are observed along CM Holes, especially in REE (e.g. La, Yb; Fig. 6),  
546 and in some major elements (e.g. FeO, MgO, CaO, Na<sub>2</sub>O; Fig. 4). These chemical trends are

547 continuous over tens of meters, and alternate between increasing and decreasing ('zigzag' patterns)  
 548 with a characteristic thickness of ~ 50 m. Abrupt changes in these trends, particularly in Yb, LREE  
 549 and HREE concentrations, are mostly associated with the presence of faults (Fig. 6). The trends  
 550 observed in CM1A and CM2B dunites and harzburgites do not show any significant correlation as  
 551 a function of LOI, CO<sub>2</sub>, or CaCO<sub>3</sub> contents (Figs. 3 and 8). This, together with the continuity of  
 552 each individual trend, suggests that the trends were imprinted at high temperature and are not  
 553 related to post-magmatic, low temperature events. Furthermore, the trends are observed in both  
 554 dunites and harzburgites in the mantle section, independent of dunite and harzburgite alternations.  
 555 Similar trend changes attributed to the focus of the percolation/migration of melts or HT fluids  
 556 along faults have been observed for the Maqсад CMTZ dunites, with the same characteristic  
 557 thickness of about 50 m (Rospabé et al. 2019a, 2020). Our results seem to confirm the significant  
 558 impact of deep-seated syn-magmatic faults on the development of the crust-mantle transition at  
 559 the expense of the shallower mantle and the recorded whole rock chemical signatures, in addition  
 560 to their impact on the formation of the lower oceanic crust (Abily et al., 2011, 2021; see also Sauter  
 561 et al., 2021). Such structural characters must have developed early and are not just restricted to the  
 562 Maqсад area (e.g. Rospabé (2018) and Abily et al. (2021) for other Oman areas and Sauter et al.  
 563 (2021) for present-day oceans).

#### 564 **4.2.2. The origin of the harzburgite REE pattern shapes and cryptic and modal mantle** 565 **refertilization in Oman ophiolite**

566 CM mantle harzburgites display two types of chondrite-normalized REE patterns: (1) linear to flat  
 567 linear REE patterns characterized by a progressive depletion from HREE to MREE to LREE, some  
 568 samples show similar LREE and HREE concentrations and slightly lower LREE and MREE  
 569 variation compared to HREE; (2) U- or V-shaped REE patterns reflecting significant MREE  
 570 depletion relative to LREE and HREE. CM harzburgites are enriched in LREE relative to MREE  
 571 with 12 harzburgites that are depleted in LREE relative to MREE and HREE; this is not expected  
 572 for mantle residues after near-fractional partial melting (Godard et al., 1995, 2000, 2008; Gruau et  
 573 al., 1998; Johnson & Dick, 1992; Johnson et al., 1990; Kelemen et al., 1997; Navon & Stolper,  
 574 1987; Prinzhofer & Allègre, 1985; Vernières et al., 1997). The linear REE patterns observed in  
 575 CM harzburgites are similar to the main harzburgites mantle section REE patterns from Wadi  
 576 Tayin described by Godard et al. (2000); these harzburgites have relatively homogeneous modal  
 577 and major element compositions. The Maqсад diapir area harzburgites and Samail massif *cpx*-  
 578 harzburgites display concave-upward and 'spoon-shaped' REE patterns respectively, these  
 579 patterns are not observed in CM Wadi Tayin harzburgites. The Maqсад diapir has been interpreted  
 580 as frozen-upwelling mantle that fed a former spreading centre (Ceuleneer et al., 1988). The  
 581 concave-upward REE patterns and the higher Al<sub>2</sub>O<sub>3</sub>/CaO ratios and TiO<sub>2</sub> contents in the diapir  
 582 harzburgites result from the feedback between deformation and melt percolation, and the *cpx*-  
 583 harzburgites REE patterns and chemical characteristics are interpreted as a result of a *cpx* forming  
 584 melt-rock reaction at decreasing melt mass (at near-solidus conditions) along the lithosphere-  
 585 asthenosphere boundary (Godard et al., 2000).

586 Figure 9 shows the modal composition of CM1A, CM2B (harzburgites and carbonate-bearing  
 587 harzburgites) and Nakhil-Samail-Wadi Tayin massif samples (harzburgites, Godard et al., 2000)  
 588 plotted with two published melting models: model 1 represents the Niu (1997) polybaric melting  
 589 model (1a) with and (1b) without excess ol; model 2 represents the Walter et al. (1995) isobaric  
 590 melting at 11 (2a), 16 (2b) and 17 kbar (2c). The figure indicates that CM harzburgites could result

591 from high degrees of partial melting and melt extraction in the range of 15-30 % (e.g. [Asimow et](#)  
592 [al., 2001](#); [Kelemen et al., 1990, 1992, 1995](#); [Niu, 1997](#); [Walter et al., 1995](#)), as suggested by  
593 [Godard et al. \(2000\)](#) for the main harzburgite section of the MORB-like, NW-SE paleo spreading  
594 segment (Nakhl-Samail-Wadi Tayin massifs). This range is higher than melting degrees producing  
595 MORB in present-day oceans (5-10 %; [Langmuir et al., 1992](#)) and high-Ti magmas such as those  
596 forming the sheeted dyke complex and the MORB-like lava sequence in Oman ([Godard et al.,](#)  
597 [2006](#); [Lippard et al., 1986](#)). However, [Niu \(1997\)](#) and [Dick & Natland \(1996\)](#) have reconciled this  
598 inconsistency by considering that abyssal peridotites represent only the shallowest part of the  
599 mantle column affected by partial melting and therefore record the highest melting degrees. In  
600 contrast, MORB are thought to represent integrated, mixed melt fractions from polybaric  
601 decompression melting over 60-100 km at ascent melting column and therefore record average  
602 melting degrees.

603 The linear REE patterns of some studied harzburgites do not show features of strong LREE  
604 depletion, but most of the other CM harzburgites show U-shaped REE patterns, characterized by  
605 LREE enrichment. This, together with the high (primary) proportion of ol (75-90 %) observed in  
606 some samples, point to the fact that near fractional partial melting alone fails to explain the  
607 harzburgite geochemical signatures along the CM cores. [Vernières et al. \(1997\)](#) noted that the  
608 relatively unfractionated REE distribution may simply result from melt transport through the  
609 melting peridotites, as an “open-system melting process”. This process would result in a negative  
610 correlation between LREE/HREE ratios and peridotite fertility, as commonly observed in other  
611 ophiolites (e.g. [Prinzhofer & Allègre, 1985](#)). It results from the competition between the partial  
612 melting continuously depleting the mantle residue on one hand and chromatographic effects  
613 related to the melt extraction that enrich the residue in the most incompatible elements on the other  
614 hand (e.g. [Johnson et al., 1990](#); [Navon & Stolper, 1987](#); [Niu, 2004](#); [Takazawa et al., 1992](#);  
615 [Spiegelman & Kelemen, 2003](#)).

616  
617 CM harzburgites show good correlations among Yb, Ti and Hf concentrations, indicating that the  
618 variability in HREE was more likely controlled by the melting process rather than by the overprint  
619 of melt-rock reaction processes ([Fig. 8 n-o](#)). We compared the linear REE patterns in CM  
620 harzburgites to [Vernières et al. \(1997\)](#) plate model calculations performed by [Godard et al. \(2000\)](#),  
621 who modeled trace elements in peridotites from the Nakhl-Samail-Wadi Tayin massifs to explore  
622 whether the REE variations observed in the studied harzburgites resulted from reactive porous  
623 flow at increasing melt mass, or from partial melting coupled with melt transport ([Fig. 10](#)). The  
624 authors first simulated a standard incremental melting model (experiment a, [Fig. 10a](#), [Johnson et](#)  
625 [al., 1990](#)). Then, they simulated reactive porous flow at increasing melt mass in a second model  
626 (experiment b, [Fig. 10b](#)). Experiment (a) produces strongly LREE-depleted peridotite residues,  
627 quite different in shape from CM and other Oman harzburgites. This experiment does not provide  
628 a better fit to the data when the presence of the trapped melt in the residue is included ([Fig. 10c](#)).  
629 Experiment (b) produces peridotite residues moderately depleted in LREE and with a small amount  
630 of trapped melt (0.5-1 %) ([Fig. 10d](#)). The linear REE patterns of harzburgites from Holes CM1A  
631 and CM2B are similar to Nakhl-Samail-Wadi Tayin harzburgites of [Godard et al. \(2000\)](#) and well  
632 reproduced by experiment (b). The model suggests that part of the studied harzburgites were  
633 pervasively percolated by diffuse melt flow which affected their geochemical signature. However,  
634 the presence of many dunite intervals at the top of the mantle section requires an orthopyroxene-  
635 consuming reaction between the residual peridotites and infiltrated melts.

636



637 The CMTZ pure dunites from CM holes (Wadi Tayin massif) display two types of chondrite-  
638 normalized REE patterns: (1) U-shaped REE patterns; (2) Linear LREE-depleted or slightly  
639 concave-upward REE patterns with similar HREE, MREE and LREE concentrations, with slightly  
640 lower LREE and MREE concentrations compared to HREE in some samples. The CMTZ  
641 impregnated dunites display three types of chondrite-normalized REE patterns: (1) linear REE  
642 patterns; (2) U-shaped REE patterns; (3) concave-upward patterns characterized by a nearly flat  
643 slope of the HREE segment followed by a progressive depletion from MREE to LREE. The  
644 Maqsad mantle-crust transition zone pure dunites trace elements are characterized by U-shaped to  
645 concave-upward REE patterns (Godard et al., 2000; Rospabé et al., 2018a, 2019a), similar to  
646 Maqsad diapir harzburgites (Godard et al., 2000) and to CM mantle dunites and harzburgites trace  
647 element patterns but with larger range of LREE variations in Maqsad MTZ dunites compared to  
648 Wadi Tayin CMTZ (CM cores). Maqsad impregnated dunites described by Rospabé et al. (2018a,  
649 2019a) are characterized by similar trace element patterns to CM impregnated dunites, varying  
650 between linear LREE-depleted to variably concave-upward trace element patterns. Maqsad MTZ  
651 pure and impregnated dunites have been interpreted as end-members that recorded different stage  
652 of an initially shared same igneous process (Rospabé et al., 2018a, 2019a). Boudier & Nicolas  
653 (1995) and Godard et al (2000) attest that Maqsad MTZ dunites are diapir harzburgites that were  
654 strongly modified by ol-forming melt-rock reactions at high melt/rock ratios. Furthermore,  
655 Rospabé et al. (2018a) argue that the pure dunites are residues left after extraction of a percolating  
656 melt, whereas the impregnated dunites correspond to a frozen stage before complete melt  
657 extraction.

658 Relatively good covariations are observed in CM dunites and harzburgites between Th and U, Nb  
659 and especially La (Fig. 8 h, j), whereas their concentrations are not correlated with the HREE.  
660 Following many previous works, we interpret these correlations as the result of melt/peridotite  
661 reaction contemporaneously with, and/or subsequent to, the partial melting event discussed above.  
662 The crust-mantle transition zone pure and impregnated dunites from CM Holes display similar  
663 REE patterns to the dunites and harzburgites in the mantle section, U-/V-shaped REE patterns  
664 (displayed by 11 pure dunites and 3 impregnated dunites from the CM, CMTZ) cannot be  
665 explained by a pure cumulate origin (Fig. 7a-b). Several studies argued that the LREE enrichment  
666 relative to MREE cannot be explained by REE partition coefficients between ol and melt (Frey et  
667 al., 1978; Hauri & Hart, 1995; Kelemen et al., 1993; Lee et al., 2007; McKenzie & O’Nions, 1991;  
668 Sun & Liang, 2014) and may better be explained by peridotite metasomatism as a result of melt-  
669 peridotite reactions (e.g. Agranier & Lee, 2007; Godard et al., 1995; Navon & Stolper, 1987;  
670 Vernières et al., 1997). Most CM harzburgites display U-shaped REE patterns (15 harzburgites  
671 and 8 carbonate-bearing harzburgites) characterized by a LREE-enrichment compared to the linear  
672 REE patterns (12 harzburgites), with strong LREE enrichments indicating extensive interaction  
673 with a pervasive melt (Fig. 7a-b; e.g. Gerbert-Gaillard, 2002; Godard et al., 2000; Monnier et al.,  
674 2006). CM samples, in particular in Hole CM2B, show downhole variations that indicate a  
675 decreasing degree of melting with increasing depth (see Section 4.2.1. below geochemical logs).  
676 Two intervals from 170 m to 260 m, and 230 m to 300 m depth, are particularly good examples of  
677 this (Fig. 6, CM2B e.g. LREE, La, U). The correlations observed between LREE enrichment and  
678 an increasing fraction of trapped melt - calculated from experiment (b and d) in the most residual  
679 peridotites from ‘Plate model’ of Vernières et al. (1997) applied by Godard et al. (2000)  
680 (experiment b and d, see Fig.10) - in CM samples (e.g. La, Fig. 6), suggest that CM samples have

681 experienced extensive interaction with a pervasive melt or fluid (for fluid-melt-rock interactions  
682 at Maqsad CMTZ see also [Rospabé et al. 2017, 2018a, 2019a](#)). In the mantle section sampled by  
683 the CM cores, crosscutting dunites are widespread and represent end products of the opx-  
684 consuming reaction. The harzburgites experienced a more extensive melt flow at a shallow level  
685 that contributed to their ol enrichment ([Fig. 1e, logs](#)). The dunites were probably individualized  
686 by channeled percolation at the top of the melting column.  
687

#### 688 **4.3. Petrological and geochemical overview of the mantle-crust transition in the SE of the** 689 **Oman ophiolite**

690 The CM ultramafic rocks studied in this paper (Wadi Tayin) compared to previously studied  
691 Maqsad MTZ dunites and diapir harzburgites show that: (1) the dunites have a replacive origin,  
692 they are the products of melt-harzburgite reaction leading to the complete consumption of  
693 orthopyroxene and concomitant precipitation of ol (e.g. [Abily & Ceuleneer, 2013; Boudier &](#)  
694 [Nicolas, 1995; Gerbert-Gaillard, 2002; Godard et al., 2000; Kelemen et al., 1995, 1997; Koga et](#)  
695 [al., 2001; Rabinowicz et al., 1987; Rospabé et al., 2018a, and references therein](#)). (2) The  
696 alternations between dunites and mantle harzburgites observed at the top of the mantle sequence  
697 of CM Holes were also observed at the base of the crust-mantle transition in many other massifs  
698 along the Oman ophiolite (e.g. [Boudier & Nicolas, 1995](#)), recording a snapshot of melt/harzburgite  
699 reaction frozen at the time of the uppermost mantle dunitization. (3) The vertical chemical trend  
700 changes related to the focus of the percolation/migration along faults observed at CM Holes  
701 samples have been previously observed at the Maqsad CMTZ dunites, with the same characteristic  
702 thickness of about 50 m ([Rospabé et al. 2019a, 2020](#)), confirming the control of synmagmatic  
703 faulting on melt/peridotite ( $\pm$  fluids) reactions and the petrological and geochemical structuration  
704 of the CMTZ.

705 The two main differences with the crust-mantle transition and mantle section of the neighboring  
706 Maqsad area, that have been extensively studied due to their diapir-related features ([Ceuleneer et](#)  
707 [al., 1988](#)), are that (1) if the two sites (Maqsad and OmanDP CM) are both more or less related to  
708 the rise of the mantle diapir, the CM site may have had less pronounced magmatic activity due to  
709 its distance to the axis of the diapir, and (2) the CM site may have been contaminated by possible  
710 remelting of the base of the old lithosphere during the development of the MORB segment, as  
711 evidenced in the Sumail massif at the borders of the area influenced by the diapir ([Amri et al.,](#)  
712 [1996; Benoit et al., 1999; Clénet et al., 2010](#)). In cores from Holes CM1A and CM2B, core  
713 description revealed the widespread presence of magmatic impregnations in the dunites (i.e.  
714 minerals that crystallized interstitially between ol grains during melt migration; e.g. [Benn et al.,](#)  
715 [1988; Dick, 1989](#)), and of magmatic segregation or more intruding dikes ([Kelemen et al., 2020a,](#)  
716 [2020b](#)). This evidence for melt migration was described within both the crust-mantle transition  
717 and the mantle sequence. The impregnations in dunites are composed of plagioclase and  
718 clinopyroxene, and the magmatic segregations and dikes are mostly troctolitic and gabbroic,  
719 consistent with a MORB-like parent melt produced by decompression melting in the mantle. On  
720 the other hand, more exotic websterite and anorthositic gabbros to anorthosites were also observed  
721 in the mantle sequence and the dunite/harzburgite alternations.

722 CM ultramafic rocks studied in this paper (Wadi Tayin) show thus multiple differences compared  
723 to the previously studied Maqsad CMTZ dunites and diapir harzburgites: (1) the CM samples have

724 been more intensively affected by low temperature alteration features (serpentinization up to 100%  
725 in many samples). (2) The crust-mantle transition zone is thinner along CM drilled Holes (~150  
726 m, Kelemen et al., 2020a, 2020b) compared to the Maqсад CMTZ (300-400 m, Abily and  
727 Ceuleneer, 2013, Boudier and Nicolas, 1995; Jouselin and Nicolas, 2000; Rabinowicz et al., 1987;  
728 Rospabé, 2018; Rospabé et al., 2019a). (3) CM pure dunites display lower LREE enrichments  
729 compared to the Maqсад ones. (4) Two major trends are observed in the trace element signatures  
730 of CM dunites and harzburgites, with good correlations between the Th and U, Nb and LREE on  
731 one hand, and between the HREE, Ti and Hf on the other hand, that are not observed at Maqсад.  
732 (5) The CM site CMTZ is mostly composed of pure dunites at the top and impregnated dunites at  
733 the bottom, whereas the typical structuration of the Maqсад CMTZ is generally composed by  
734 impregnated dunites at the top and pure dunites at the bottom (Rospabé, 2018; Rospabé et al.  
735 2019a). (6) The dunites' impregnation characteristic appears to be different in Maqсад (e.g., mostly  
736 plagioclase and clinopyroxene (Boudier and Nicolas, 1995; Koga et al., 2001; Abily and  
737 Ceuleneer, 2013) but also widespread opx and amphibole impregnations in the higher level of the  
738 transition zone (Rospabé et al., 2017, 2018a, 2019a) compared to CM Hole dunites (plagioclase  
739 and clinopyroxene impregnations only). Similarly, exotic silicate inclusions (opx, amph, mica for  
740 the more abundant) enclosed in disseminated chromite grains in the dunites from the Maqсад area  
741 as well as in associated chromitite ore bodies (Lorand and Ceuleneer, 1989; Leblanc and  
742 Ceuleneer, 1991; Schiano et al., 1997; Borisova et al., 2012; Rollinson et al., 2018; Zagrtednov et  
743 al., 2018; Rospabé et al., 2019b; 2020, 2021), suggest the involvement of a fluid or fluid-rich melt  
744 in the melt/rock reactions, which were not investigated in existing work on cores CM1A and  
745 CM2B - a few chromite schlierens have been sampled along the cores but not studied in details  
746 yet. Further investigation of CM ultramafic rock mineral chemistry is need to evaluate these  
747 discrepancies between the two CM and Maqсад sites.

748 All the above similarities and differences between CM Wadi Tayin and Maqсад ultramafic rocks  
749 point to a lighter imprint of melt-rock reaction at CM compared to Maqсад (e.g., thinner transition  
750 zone, lighter LREE enrichment in dunites, two distinct geochemical trends in trace elements, that  
751 perhaps are totally overprinted by stronger melt/rock reaction at Maqсад). This could be the  
752 consequence of one or the several factors: It could be related to the structural position of Wadi  
753 Tayin (at the periphery of the diapir) and Maqсад (centered on the diapir) leading to different  
754 melt/rock ratios (i.e., the geological context), to the timing of the occurrence of the melt-rock  
755 interaction, and/or to the nature of the percolating magma/fluid involved in the melt-rock reactions.  
756 Further detailed mineral chemistry, trace element geochemical modelling, structural and  
757 microstructural studies will be of use for addressing the above questions.

## 758 **Conclusions**

759 Continuous sampling of the Oman crust-mantle transition zone at Holes CM1A and CM2B,  
760 recovered by the Oman Drilling Project allows the study of the large range of petrological and  
761 geochemical variations in Oman ultramafic rocks with an unprecedented high resolution. Volatile  
762 (H<sub>2</sub>O and CO<sub>2</sub>), and major and trace elements of 56 dunites and 49 harzburgites from Holes CM1A  
763 and CM2B have been analyzed. CM1A and CM2B volatile element contents reflect extensive  
764 serpentinization (+/- carbonation) linked to the late-stage interaction with H<sub>2</sub>O- and/or CO<sub>2</sub>-  
765 bearing fluids. The refractory samples are characterized by relatively homogeneous modal and  
766 major element compositions, whereas other samples show primary cryptic and modal  
767 refertilization. Bulk rock Mg, Si and Al systematics and normative mineral modes suggest that

768 open system behavior during alteration (Mg loss and/or Si gain) affected many samples. However,  
769 the trace element concentrations are interpreted as reflecting magmatic processes and exhibit  
770 significant variations: the refractory harzburgites characterized by linear REE patterns are  
771 interpreted as mantle residues after  $\geq 15$  % melt extraction. The REE signatures in these samples  
772 can be explained by melt transport associated with partial melting. Other harzburgites displaying  
773 U-/V-shaped REE patterns are interpreted as the result of interstitial melt percolation. The pure  
774 and impregnated dunites from the mantle and the crust-mantle transition zones are characterized  
775 by similar trace element patterns as the mantle harzburgites, they are interpreted as reflecting  
776 different stages of ol-forming melt/rock reactions at high melt/rock ratio. The plagioclase and  
777 pyroxene minerals present in impregnated dunites indicate that the impregnated dunites represent  
778 the stage before complete extraction of the melt from the dunites, whereas the pure dunites  
779 represent the melt/rock reaction end-product after complete melt extraction (i.e. compaction of the  
780 ol matrix and system closure before/without interstitial plagioclase and clinopyroxene  
781 crystallization).

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800

## 801 **Data Availability Statement**

802 The dataset is available on the PANGAEA data archiving platform  
803 (<https://doi.org/10.1594/PANGAEA.942402>).

804

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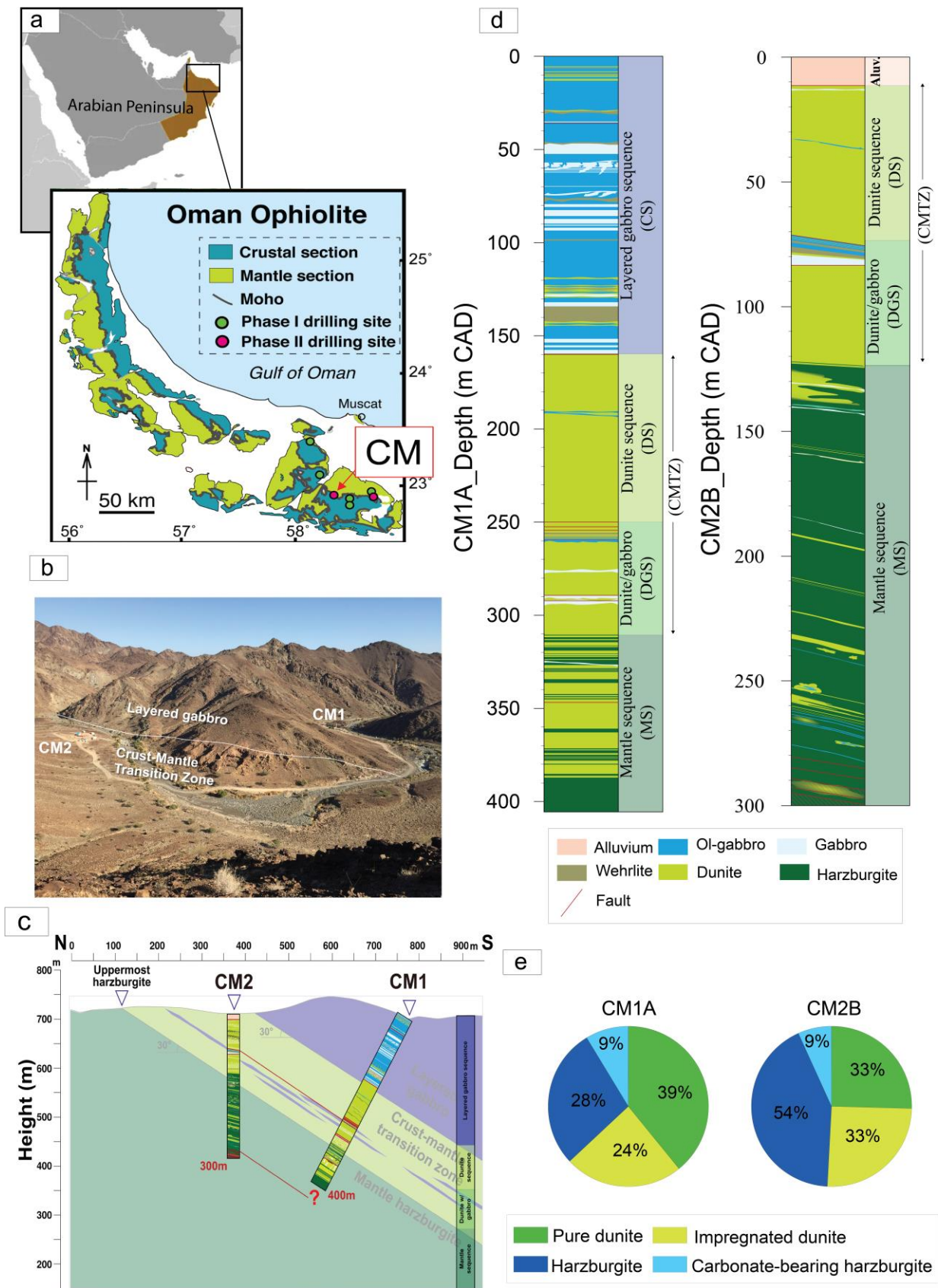
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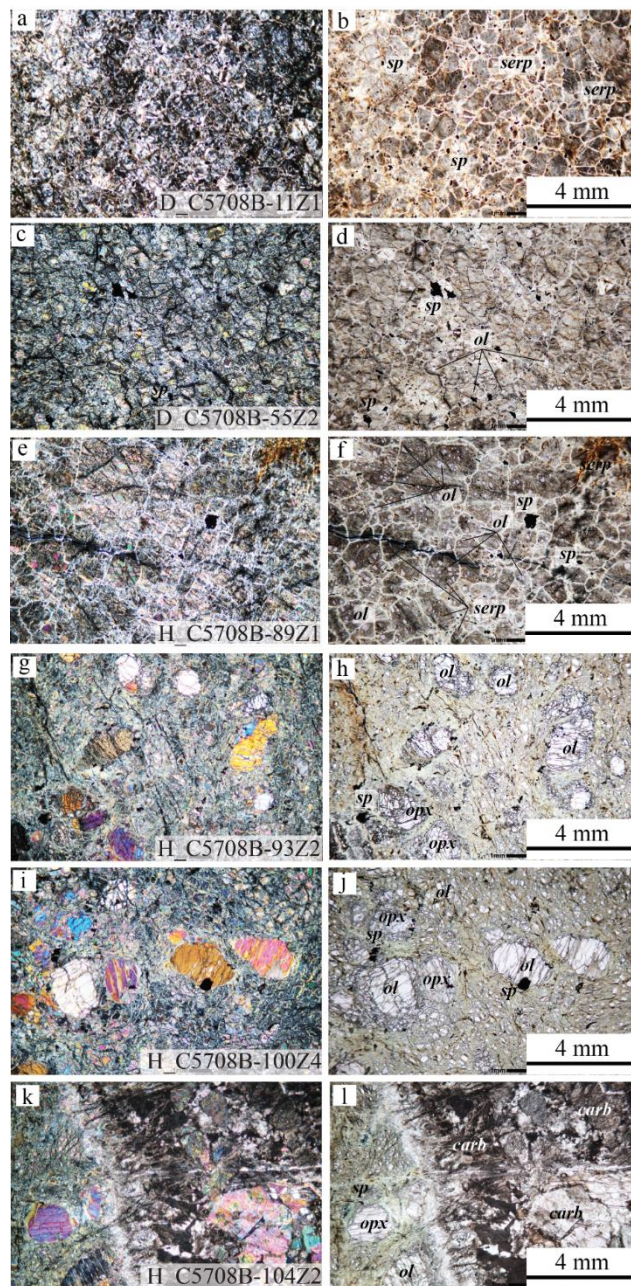
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1240 **Figure captions**





1242 **Figure 1** Geological context, **a** Simplified geological map showing the location of Holes CM1A  
 1243 and CM2B in the context of the regional and local geology (after **Nicolas & Boudier, 1995**). **b**  
 1244 Field photographs showing the location of Holes CM1A and CM2B. **c** Simplified N-S cross-  
 1245 section showing the different sampled lithologies and the correlation in depth between Holes  
 1246 CM1A and CM2B. **d** Downhole stratigraphy of CM1A and CM2B showing the distribution of the  
 1247 lithologies in the crust, crust-mantle and mantle sequences. **e** pie charts showing the studied  
 1248 samples lithology count in Hole CM1A and CM2B (see Section 3 in the Supporting Information).  
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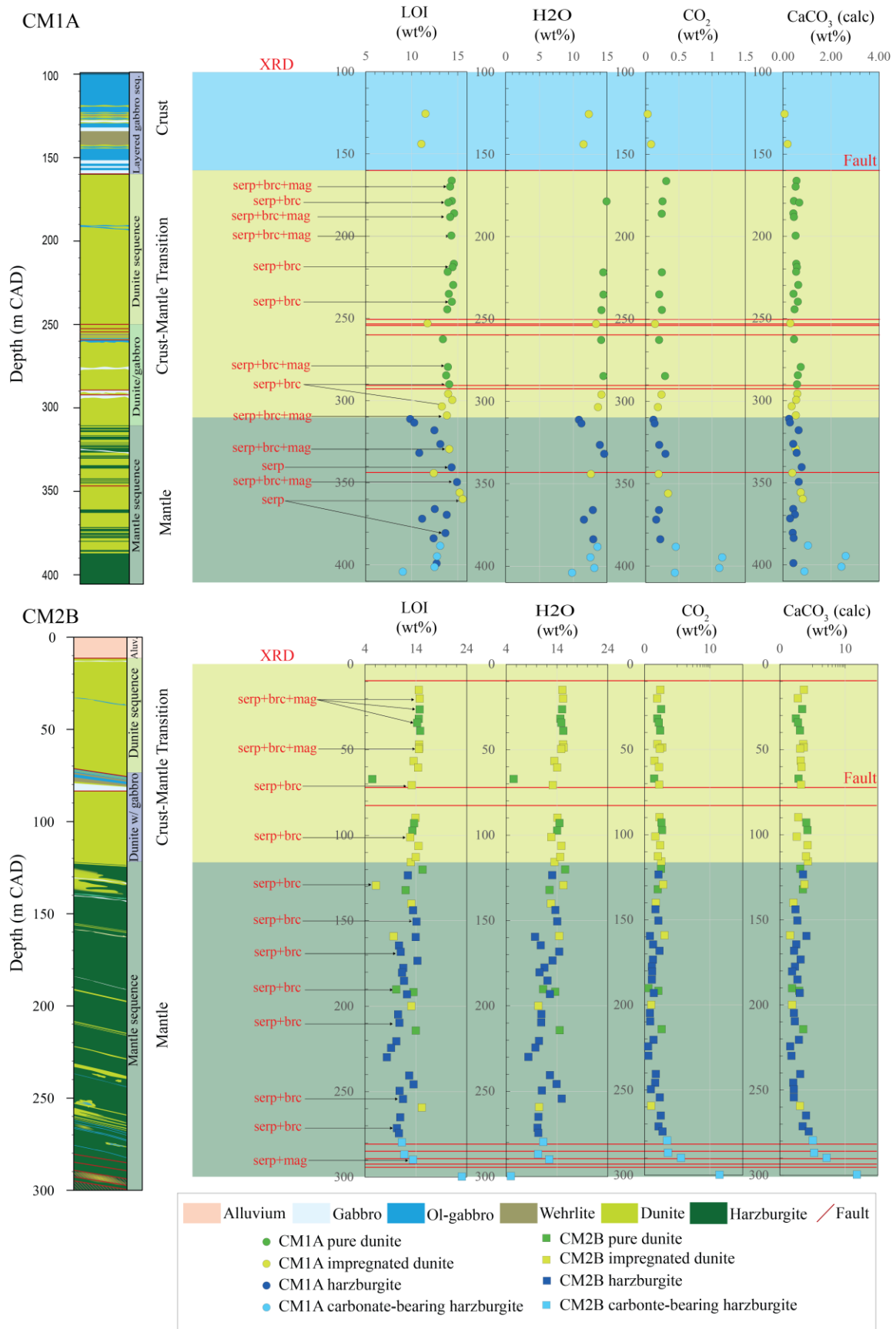
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1252 **Figure 2** microphotographs of petrographic details of a selection of CMTZ dunites (**a, b, c** and **d**)  
 1253 and harzburgites (**e, f, g, h, i, j, k** and **l**) from Holes CM1A and CM2B. The dunites are  
 1254 characterized by fine- to medium grained granular texture and the harzburgites by porphyroclastic

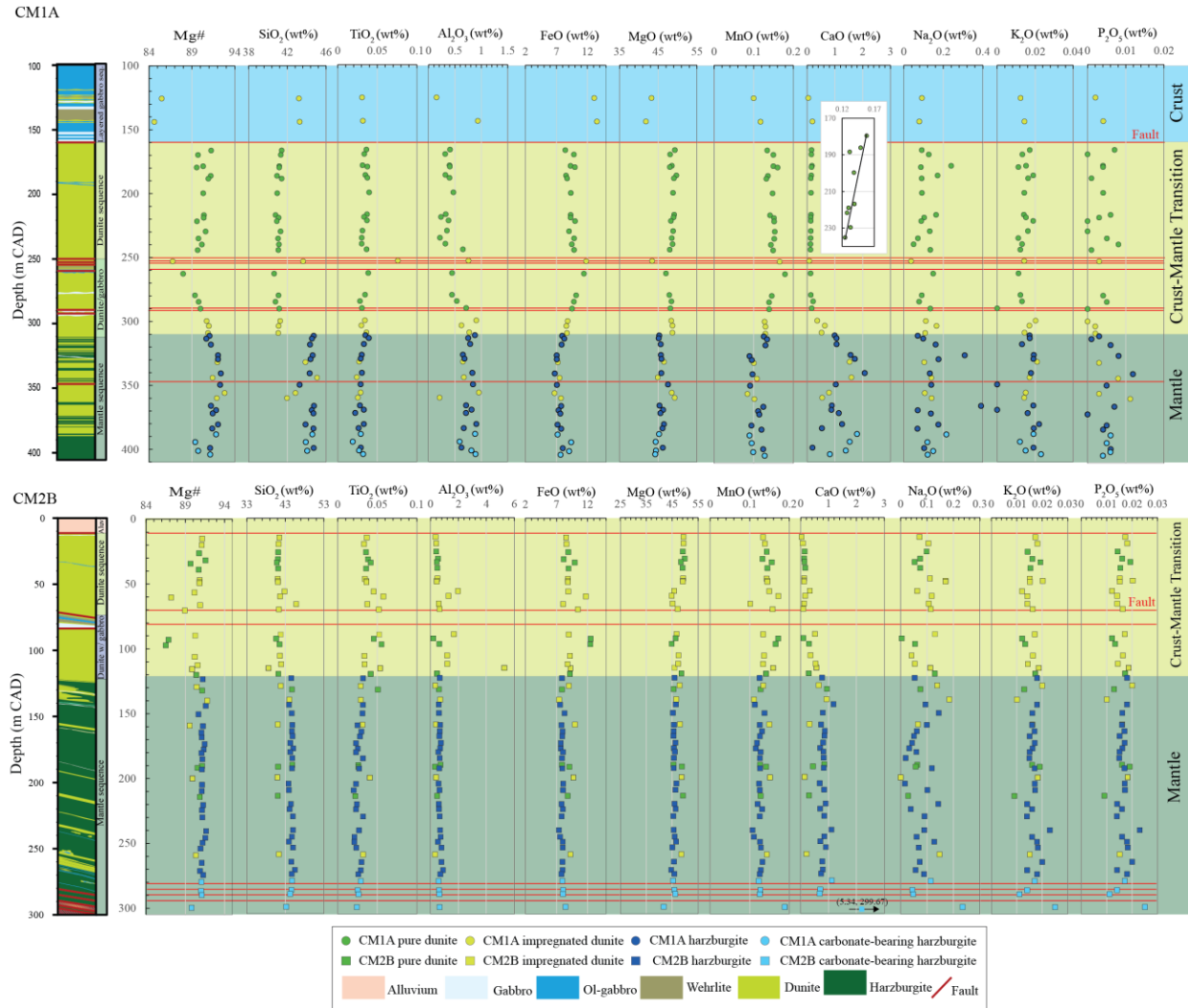
1255 texture. The dunites and harzburgites have generally preserved their primary high temperature  
1256 texture (granular or porphyroclastic textures) after complete (**a**, and **b**), or partial (**c**, **d**, **e**, **f**, **g**, **h**, **i**,  
1257 **j**, **k** and **l**) replacement of olivine crystals by serpentine. **i** and **j** illustrate microphotographs of a  
1258 plastically deformed orthopyroxene surrounded by neoblasts. **k** and **l** show harzburgite CM2B-  
1259 104Z2 crosscut by low-temperature veins (Carbonate).

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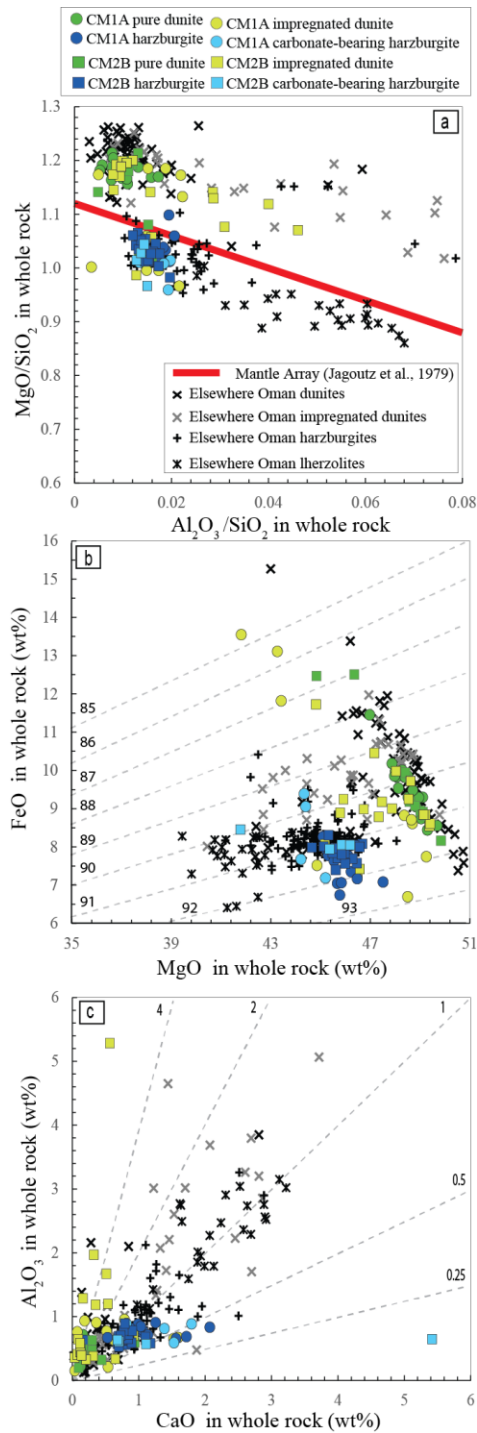


1263 **Figure 3** Downhole plots of (from left to right) the lithology, LOI (wt.%), H<sub>2</sub>O (wt.%), CO<sub>2</sub> (wt.%)  
 1264 and calculated CaCO<sub>3</sub> (wt.%) contents in pure dunites, impregnated dunites, harzburgites and  
 1265 carbonate-bearing harzburgites recovered samples at Holes CM1A and CM2B. The thick solid red  
 1266 lines indicate the faults. XRD: X-Ray Diffraction, serp: serpentine, brc: brucite, mag: magnetite.  
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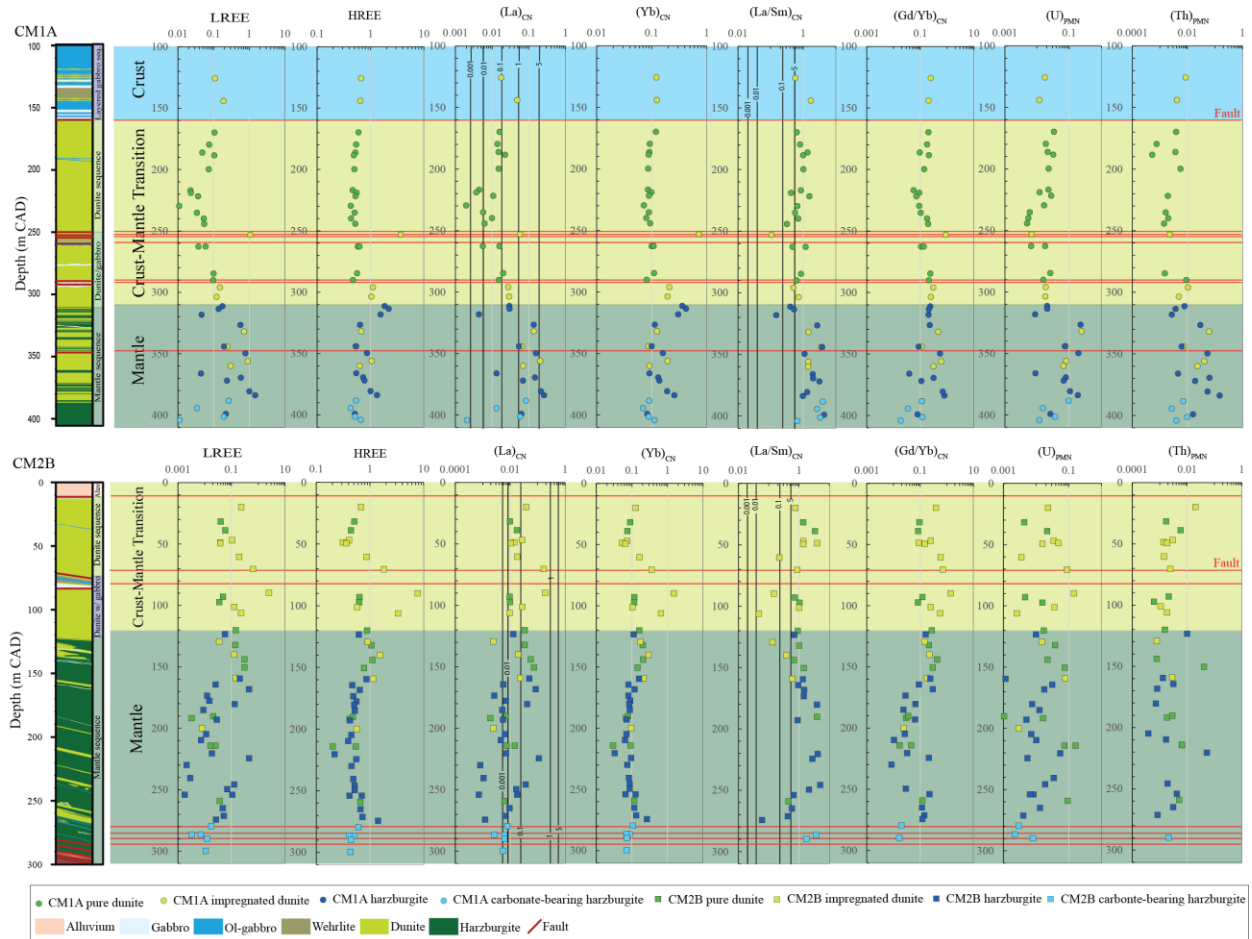
1269 **Figure 4** Downhole plots of (from left to right) the lithology, Mg# (cationic 100 x Mg/(Mg+Fe<sub>total</sub>);  
 1270 calculated assuming all Fe as FeO) (mol%), SiO<sub>2</sub> (wt.%), TiO<sub>2</sub> (wt.%), Al<sub>2</sub>O<sub>3</sub> (wt.%), FeO (wt.%),  
 1271 MgO (wt.%), MnO (wt.%), CaO (wt.%), Na<sub>2</sub>O (wt.%), K<sub>2</sub>O (wt.%), and P<sub>2</sub>O<sub>5</sub> (wt.%) in whole  
 1272 rock samples recovered at Holes CM1A and CM2B. The thick solid red lines indicate the faults.  
 1273



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1275 **Figure 5** Whole rock major compositions of samples recovered at Holes CM1A and CM2B  
 1276 compared to other crust-mantle transition dunites and mantle harzburgites/lherzolites from the  
 1277 Oman ophiolite (Gerbert-Gaillard, 2002; Godard et al., 2000; Hanghøj et al., 2010; Khedr et al.,  
 1278 2014; Monnier et al., 2006; Nicolle et al., 2016; Rospabé et al., 2018a, 2019a; Takazawa et al.,  
 1279 2003). (a)  $\text{MgO}/\text{SiO}_2$  versus  $\text{Al}_2\text{O}_3/\text{SiO}_2$ , (b) total iron as FeO versus MgO, and (c)  $\text{Al}_2\text{O}_3$  versus  
 1280 CaO. Compositions are recalculated on a volatile-free basis. Red bar in panel (a) represents the

1281 silicate Earth differentiation trend (or “terrestrial array”) (Jagoutz et al., 1979). Dashed grey lines  
 1282 in panels (b) and (c) represent constant Mg# and Al<sub>2</sub>O<sub>3</sub>/CaO ratios respectively.

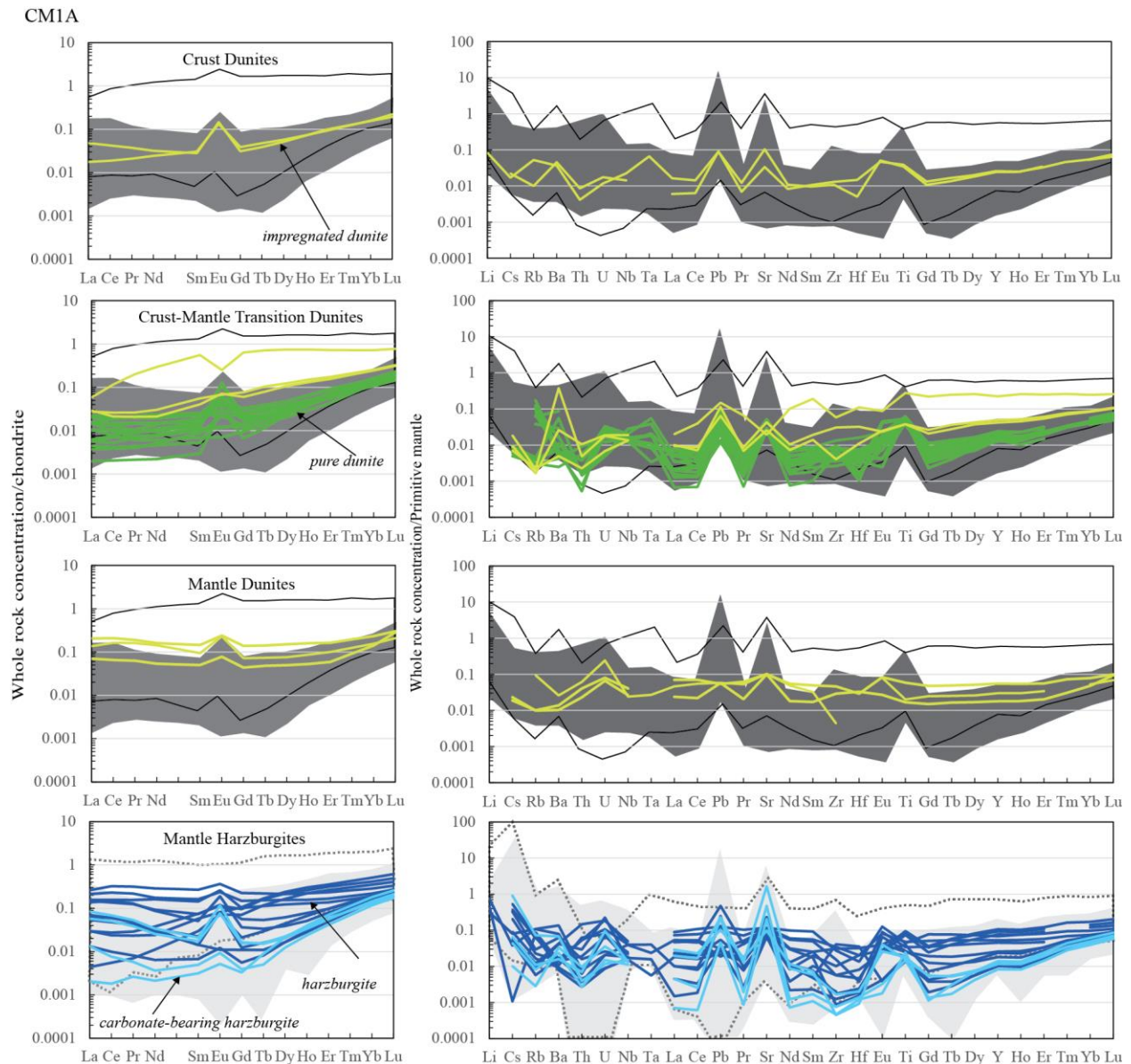


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1284 **Figure 6** Downhole plots of (from left to right) (LREE: (La+Ce+Pr+Nd)<sub>CN</sub>), (HREE:  
 1285 (Ho+Er+Tm+Yb+Lu)<sub>CN</sub>), (La)<sub>CN</sub>, (Yb)<sub>CN</sub>, (La/Sm)<sub>CN</sub>, (Gd/Yb)<sub>CN</sub>, (U)<sub>PMN</sub> and (Th)<sub>PMN</sub> in whole  
 1286 rock samples recovered at Hole CM1A and CM2B (CN: chondrite-normalized; PMN: primitive  
 1287 mantle-normalized). The thick solid red lines indicate the faults, and the thicker solid black lines  
 1288 indicate the (La)<sub>CN</sub> and (La/Sm)<sub>CN</sub> concentrations in the most residual peridotites from ‘Plate  
 1289 model’ of Vernières et al. (1997) applied by Godard et al. (2000) (experiment b and d, see Fig.9),  
 1290 the numbers on the black lines indicate the proportions of trapped melt (in percentage) issued from  
 1291 the models. Normalizing chondrite and Primitive Mantle values are from Barrat et al. (2012) and  
 1292 Sun and McDonough (1989) respectively.

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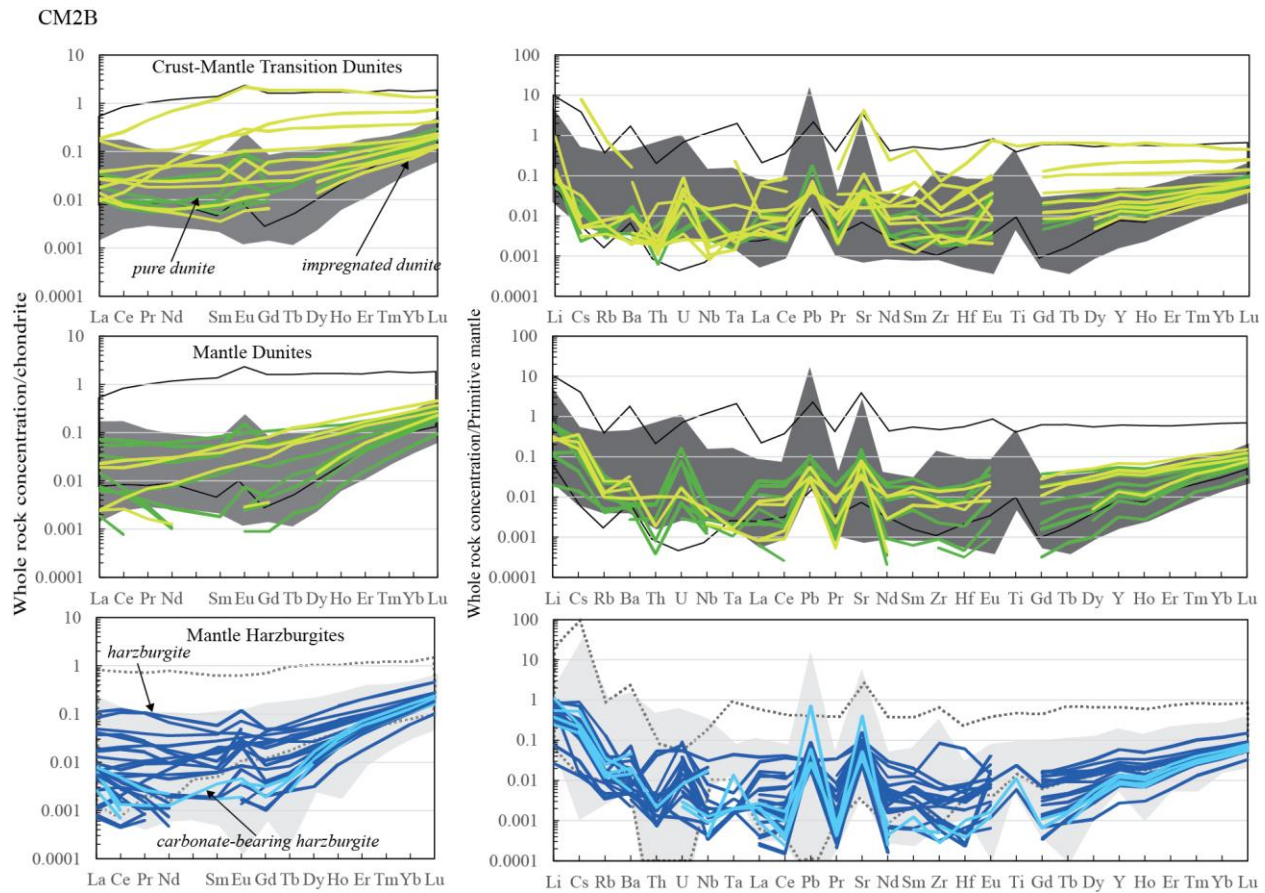
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1296 **Figure 7a** Chondrite-normalized REE and Primitive Mantle-normalized multi-element patterns of  
 1297 pure dunites, impregnated dunites, harzburgites and carbonate-bearing harzburgites from the crust,  
 1298 crust-mantle transition zone and mantle sections recovered at Hole CM1A. Other pure dunites  
 1299 (dark gray field formed) and impregnated dunites (field formed by black line) from the crust-  
 1300 mantle transition, and harzburgites (light gray field) and lherzolites (field formed by dashed line)  
 1301 patterns from the mantle section of the whole Oman ophiolite are reported for comparison  
 1302 (Gerbert-Gaillard, 2002; Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010; Khedr  
 1303 et al., 2014; Lippard et al., 1986; Monnier et al., 2006; Nicolle et al., 2016; Rospabé et al., 2018a,  
 1304 2019a; Takazawa et al., 2003). Normalizing chondrite and Primitive Mantle values are from Barrat  
 1305 et al. (2012) and Sun and McDonough (1989) respectively.

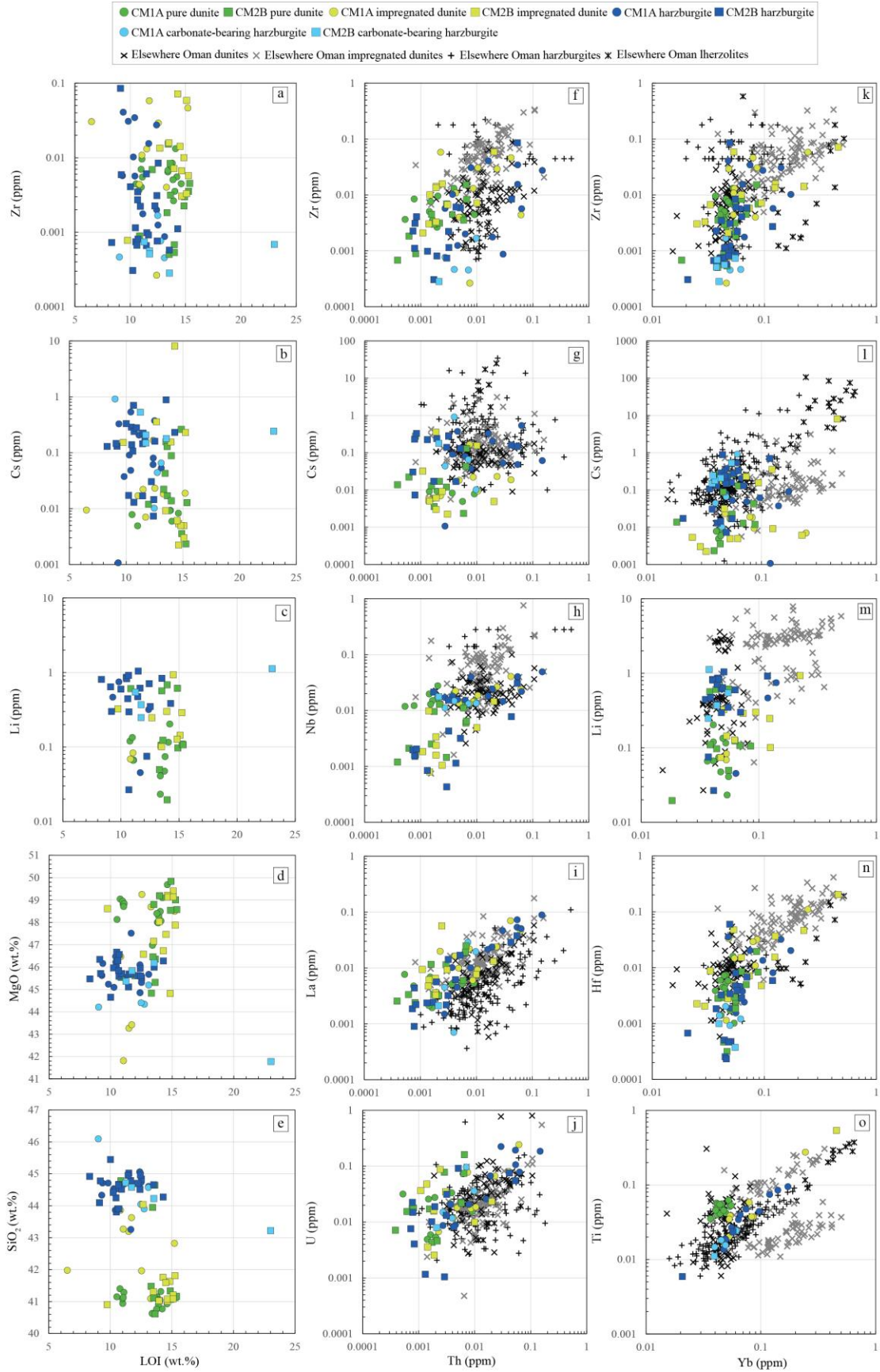




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1307 **Figure 7b** Chondrite-normalized REE and Primitive Mantle-normalized multi-element patterns of  
 1308 pure dunites, impregnated dunites, harzburgites and carbonate-bearing harzburgites from the crust-  
 1309 mantle transition zone and mantle sections recovered at Hole CM2B. Other pure dunites (dark gray  
 1310 field formed) and impregnated dunites (field formed by black line) from the crust-mantle  
 1311 transition, and harzburgites (light gray field) and lherzolites (field formed by dashed line) patterns  
 1312 from the mantle section of the whole Oman ophiolite are reported for comparison (Gerbert-  
 1313 Gaillard, 2002; Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010; Khedr et al., 2014;  
 1314 Lippard et al., 1986; Monnier et al., 2006; Nicolle et al., 2016, Rospabé et al., 2018a, 2019a;  
 1315 Takazawa et al., 2003). Normalizing chondrite and Primitive Mantle values are from Barrat et al.  
 1316 (2012) and Sun and McDonough (1989) respectively.

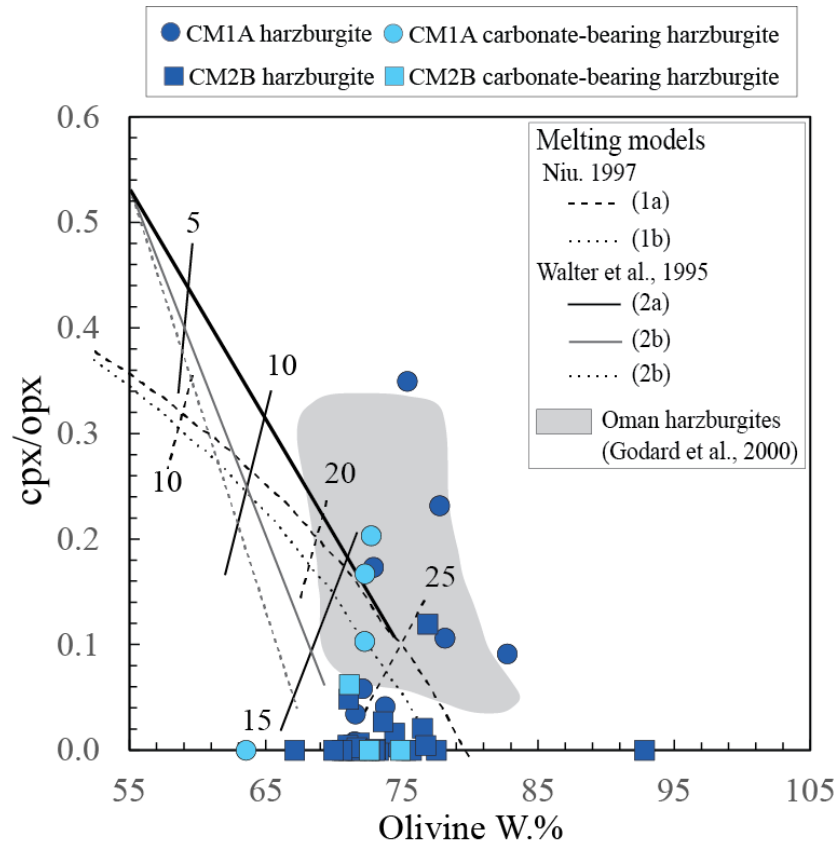
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1320 **Figure 8** Plots of LOI (wt.%) vs. Zr (ppm), Cs (ppm), Li (ppm), MgO (wt.%), and SiO<sub>2</sub> (wt.%); Th (ppm) vs. Zr  
 1321 (ppm), Cs (ppm), Nb (ppm), La (ppm), and U (ppm); Yb (ppm) vs. Zr (ppm), Cs (ppm), Li (ppm), Hf (ppm), and Ti  
 1322 (ppm) in dunite and harzburgites recovered at Hole CM1A and CM2B. Other pure dunites, impregnated dunites,  
 1323 harzburgites and lherzolites compositions from the crust-mantle and the mantle section of the whole Oman ophiolite  
 1324 are reported for comparison (Gerbert-Gaillard, 2002; Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010;  
 1325 Khedr et al., 2014; Lippard et al., 1986; Monnier et al., 2006; Nicolle et al., 2016, Rospabé et al., 2018a, 2019a;  
 1326 Takazawa et al., 2003).

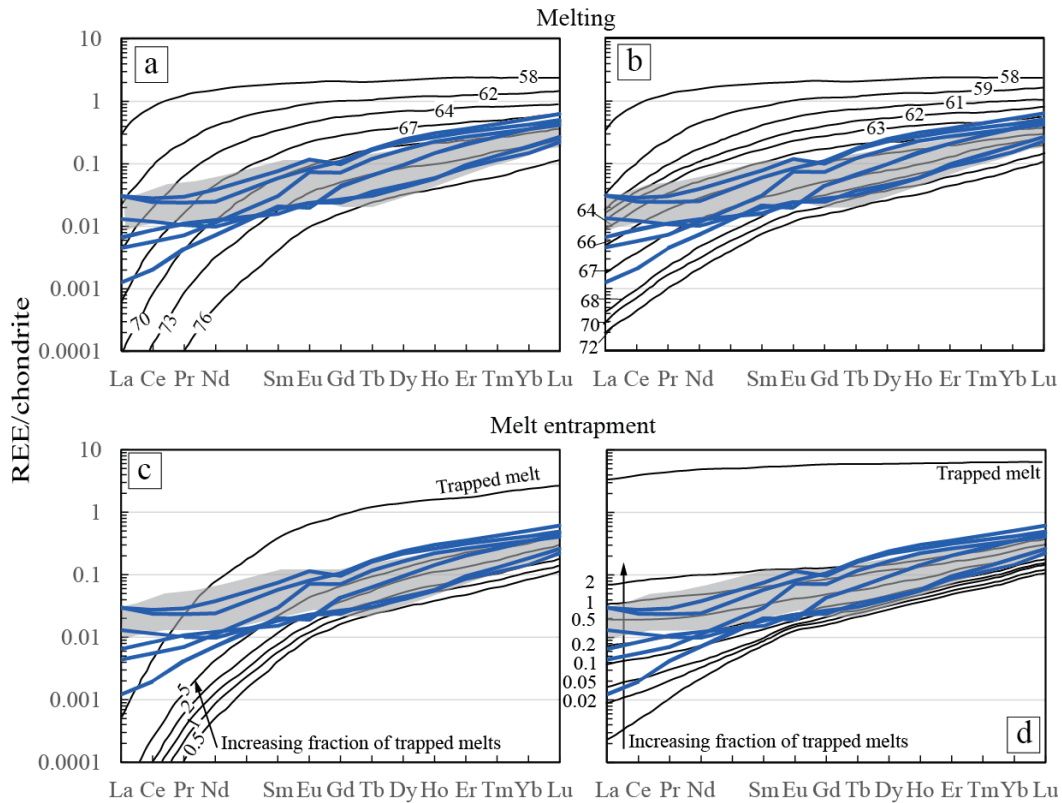
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1329 **Figure 9** Modal compositions of the analyzed CM harzburgites plotted on a cpx/opx vs. olivine  
 1330 diagram. The field defined by the mantle harzburgites studied in Godard et al., 2000 is reported  
 1331 for comparison. Published melting models are also shown for comparison: model 1 represents the  
 1332 polybaric melting model after Niu (1997), (1a) with and (1b) without excess olivine; model 2  
 1333 represents the isobaric melting after Walter et al. (1995) at 11 (2a), 16 (2b) and 17 kbar (2c). The  
 1334 initial modal composition is given by Niu (1997) for polybaric melting and was fixed for isobaric  
 1335 melting as: 55% ol, 28% opx, 15% cpx and 2% sp. Numbers refer to percent melting degrees.  
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1339 **Figure 10** CM1A and CM2B REE linear flat shaped REE patterns (Thicker blue lines) compared  
 1340 to the ‘Plate model’ of Vernières et al. (1997) applied by Godard et al. (2000) to simulate REE  
 1341 variations in a peridotite affected by partial melting with (a) or without (b) melt infiltration. The  
 1342 chondrite-normalized REE patterns of the Oman harzburgites from Godard et al. (2000) (the gray  
 1343 fields, main harzburgite section) are also shown for comparison. The authors simulate standard  
 1344 incremental melting in model (a) and the percolation of fixed N-MORB melt through molten  
 1345 peridotites in model (b). The peridotite initial modal composition was (spinel free): 57% ol, 28%  
 1346 opx and 15% cpx. The melting reaction was taken from Walter et al. (1995). Mineral/melt partition  
 1347 coefficients are the same as those selected by Bedini & Bodinier (1999). Numbers on the chondrite-  
 1348 normalized REE patterns indicate olivine proportion (in percentage) in residual peridotites.  
 1349 Thicker black lines indicate the REE patterns of the less residual peridotites. In model (a), the most  
 1350 residual peridotite (76% olivine) is produced after 21.1% melt extraction. In model (b), the ratio  
 1351 of infiltrated melt to peridotite varies from 0.02 to 0.19. (Bottom) Modifications of the REE  
 1352 patterns of residual peridotites due to the presence of equilibrium, trapped melt. Models (c) and  
 1353 (d) show the effect of trapped melt on the most residual peridotites of models (a) and (b),  
 1354 respectively (thicker solid lines). Numbers on the REE patterns indicate the proportions of trapped  
 1355 melt (in percentage). Normalizing chondrite values are from Barrat et al. (2012).

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