

Compact nano-void spectrometer based on a stable engineered scattering system

QI SUN,^{1,*} PRZEMYSŁAW FALAK,^{1**} TOM VETTENBURG,² TIMOTHY LEE,¹ DAVID B. PHILLIPS,³ GILBERTO BRAMBILLA,¹ MARTYNAS BERESNA,¹

¹*Optoelectronics Research Centre, University of Southampton, Southampton, SO17 1BJ, United Kingdom*

²*University of Dundee, Nethergate, Dundee, DD1 4HN, United Kingdom*

³*University of Exeter, Exeter, EX4 4QL, United Kingdom*

**qs3g15@soton.ac.uk*

***plf1n15@soton.ac.uk*

Abstract: Random scattering of light in disordered media can be used for highly-sensitive speckle-based wavemeters and spectrometers. However, the multiple scattering events that fold long optical paths within a compact space also make such devices exceedingly sensitive to vibrations and small disturbances to the disordered media. Here, we show how scattering can be engineered so that it can be used for a compact computational spectrometer that is largely insensitive to environmental factors. We designed and fabricated a three-dimensional pseudo-random nano-void pattern with 62% scattering efficiency. The controlled amount of multiple scattering ensured a sufficiently long optical path for the target resolution of 100 pm, with optimal long-term stability. The 200 μm -thick scattering silica substrate was integrated in a compact assembly with a low-cost camera sensor. The target resolution was achieved for full spectrum measurements while single wavelengths could be determined with 50 pm resolution. Such tailored scattering systems can improve the trade-off between cost, size, stability, and spectral resolution in computational spectrometers.

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1. Introduction

Accurate wavelength measurement is central for analysis and characterization in various disciplines, including biology, chemistry, material analysis and astronomy. The conventional approach for extracting spectral information relies on a diffraction grating to separate the spectral components. A key parameter for such a system is the optical path length between the dispersive element and the detector. Fine spectral resolution requires a large optical path length and correspondingly large instrument footprint.

An alternative method is to exploit the dispersion inherent in multiple cascaded light scattering events to implement spatial mapping of spectral components [1, 2]. Multiple scattering would increase the equivalent optical path of light, beyond the actual physical size of the system. Therefore, a small shift in the wavelength of input light can result in a large change in its generated speckle pattern. Many approaches for implementing scattering media have been exploited: multimode fibre, integrating spheres, alumina powder and 2D scattering chip. These have all provided a high spectral resolution, yet a key challenge remained unsolved: the mechanical stability of the scattering sample itself. Here we propose a 3D scattering chip made of material with intrinsic environmental (thermal and mechanical) stability and inscribed with densely-packed laser-written nano-voids. These arrays of scattering centers create a high degree of spectral dispersion within a small volume, effectively folding the optical path. As a result, multiple scattering events produce intricate speckles that are highly wavelength dependent [3]. Despite their random appearance, such patterns are deterministic provided that the scattering element is static in nature.

46 Scattering spectrometers have been demonstrated using several different scattering elements.
47 A multimode fiber-based system reaching a fine resolution of 1 pm was demonstrated using
48 100 m of fiber waveguide [4], where the speckle pattern is formed by interference of many higher
49 order modes. The resolution in this arrangement is proportional to the length of the fiber. To
50 achieve a 0.1 nm resolution, a length of at least 1 meter of fiber (NA = 0.22) is required [5].
51 Unfortunately, the longer waveguide also makes it sensitive to small environmental perturbations
52 and vibrations, thus rendering it impractical for many applications. Such systems require regular
53 re-calibration and must remain unchanged after the calibration for each test [6]. Wavemeters
54 based on an integrating sphere have been used to demonstrate sub-femtometer resolution [7, 8],
55 yet size and stability requirements limit applications to specialist and laboratory settings. Since
56 2013, a compact spectrometer chip based on a silicon-on-insulator scattering medium has been
57 utilised to increase the optical path length and enable better spectral resolution within the same
58 physical space [9, 10]. The first version of such a device was designed as a two-dimensional
59 random air holes etched into the silicon plane [3]. It had a resolution of 0.75 nm at a wavelength
60 of 1500 nm in a 25-mm-radius structure. However, as the scattering medium was designed
61 in a single plane, the chip could only operate with limited detectors thus limited wavelength
62 channels in 2 dimensions. Soon after, a wavelength meter using alumina (Al_2O_3) powder
63 (3D approach) was demonstrated [11]. By analyzing the intensity pattern created by multiple
64 scattering within a dried drop of alumina, the illumination wavelength could be determined with
65 13 pm accuracy. However, the instability of alumina powder renders performance extremely
66 sensitive to vibrations and alignment drift. Because of this, interest was refocused on chip-based
67 systems. As a result, the chip strategy evolved, either by miniaturisation, increasing number
68 of output ports or expanding the structure from 2D to 3D (multiple planes approach) [12–14].
69 Transition into 3D design enabled key advantages: applying 3D detector arrays (more detection
70 channels) and generating more sophisticated speckle patterns (cascaded scattering events), which
71 increased optical resolution of such devices. Nowadays the chip-based spectrometers can achieve
72 single-photon sensitivity [15], yet still the main challenge persists: the stability of both the
73 scattering medium and the device itself. More generally, while such highly scattering dispersive
74 systems offer greater potential resolution, the practical resolution will be stability-limited by
75 its higher environmental sensitivity [16]. It is therefore important to maximize stability while
76 adapting the level of scattering to optimally match the required spectral resolution for the
77 application.

78 Here, we demonstrate a highly stable, low-cost, compact and high-resolution spectrometer.
79 To minimize the influence of vibrations and thermal gradients, our design is based on a single
80 monolithic block of silica. We first describe how femtosecond laser direct writing can be used to
81 create a pseudo-random distribution of nano-voids which scatter light in a wavelength dependent
82 manner, tunable by design. The following sections discuss the calibration and characterization of
83 its high-stability operation.

84 **2. Pseudo-random scattering spectrometer**

85 *2.1. Nano-void scattering chip*

86 We designed a planar pattern of pseudo-random scattering nano-voids that can be readily produced
87 using direct laser writing. Femtosecond lasers can generate nano-voids with sub-micro joule
88 pulse energies [17]. The highly non-linear interaction of femtosecond pulses with the transparent
89 substrate enables the formation of densely packed spherical cavities without collateral damage.
90 We exploit this to produce a pseudo-random geometry with an average void-separation of $1\ \mu\text{m}$.

91 Fig. 1a shows a schematic of the optical system used for writing the custom scattering chip.
92 Micro-cavities are formed using a femtosecond laser (Pharos, Light Conversion Ltd., Lithuania)
93 with central wavelength of $\lambda = 1.03\ \mu\text{m}$, pulse duration of $\tau = 200\ \text{fs}$, and repetition rate of
94 $f = 200\ \text{kHz}$. To enhance the writing resolution, we used the second harmonic $\lambda = 515\ \text{nm}$ for

95 laser writing. A computer controlled translation stage was used to translate the silica substrate
96 with respect to the beam in the horizontal directions x and y , while the height of the oil-immersion
97 objective (NA 1.25) was adjusted in the vertical z direction. The overhead CMOS camera allowed
98 in-situ visualization of the writing process and the initial quality inspection of the imprinted
99 structures. Each void was produced by a single pulse of energy 220 nJ.

100 The scattering chip matrix consists of multiple planes of pseudo-randomly placed nano-voids
101 on both sides of a 1 mm thick (z) substrate. The use of multiple planes increases the fraction of
102 scattered light to induce long optical path length differences. A high purity fused silica glass
103 (UVFS C7980 0F) with low thermal expansion coefficient ($0.57 \times 10^{-6} K^{-1}$) and OH content
104 of 800 – 1000 ppm was used as a substrate (Fig. 1c). To achieve a scattering efficiency of
105 $62 \pm 2\%$ (measured with a superluminescent diode at 1080 nm, SLD-1080-30-YY-100, Innolume
106 GmbH), we defined 40 scattering planes with a $p_z = 5 \mu\text{m}$ separation and 20 μm below the
107 sample surface. The nano-voids in each plane were uniformly randomly distributed ± 400 nm
108 from the regular $p_x \times p_y = 1 \times 1 \mu\text{m}^2$ grid. To facilitate rapid laser writing, the randomization
109 was alternated in the x and y direction for consecutive planes. Each individual dot in Fig. 1d is a
110 microexplosion void induced by a single femtosecond pulse that changes the effective refractive
111 index by up to 0.45 [18]. Notice there are few factors we were considered during the chip design:
112 pulse energy, plane number, plane separation, average void distance and fabrication time. Refer
113 to our experiment, multiple planes and stronger pulse energy (under sub-micro joule) gives
114 stronger scattering. However, there is a trade-off between the pulse energy and plane number.
115 Our experiment shows that with 500 nJ laser pulse writing, when the plane number increases
116 to 20 planes, the substrate starts to have visually crack. Indeed, we could increase the plane
117 separation to release the stress inside the substrate so that we can reach a higher number of planes.
118 However, limited by the objective focal depth, the deepest plane for 1.25 NA oil immersion lens
119 is around 400 μm below the surface. We tried to increase the plane separation to 10 μm which
120 gives the similar scattering. Overall, the final scattering chip with 40 scattering planes, 5 μm
121 plane separation, 1 μm average void distance and pseudo-randomized structure is an optimised
122 chip version with the trade-off among fabrication time, scattering efficiency – given by the ratio
123 of scattered light to incident light intensity – and fabrication system limitation. The resulting
124 scattering chip displays a collective scattering efficiency of 62%. This bottom-up design and
125 fabrication approach enables excellent control of the scattering properties by varying the number
126 of scattering planes, their separation, and the nano-void distribution.

127 2.2. Packaging of the spectrometer optical system

128 Performance of the scattering spectrometer is highly dependent on the mechanical stability of
129 the whole system. For the enclosure and packaging, a 3D printed monocoque construction was
130 chosen to deliver several key functions: (1) to accommodate key device components in confined
131 space (compact device for spectral analysis); (2) to ensure stable measurements by shielding the
132 interior from fluctuating environmental conditions (including temperature, humidity and ambient
133 light) and (3) to provide mechanical protection for delicate electronic parts [5].

134 Ambient temperature and humidity can potentially affect the scattering process via thermal
135 expansion or swelling (Sec. 3.4). To monitor the influence of the environment, a temperature
136 and humidity sensor was added to the system (SHTC3, Adafruit). The spectrometer enclosure
137 was 3D printed from Tough PLA (Ultimaker), with external dimensions of $57 \times 35 \times 35$ mm and
138 four parts: the monocoque body providing mounting space and protection for all components;
139 camera retainer fixing CMOS sensor to the main frame and preventing its displacement during
140 operation; scattering chip holder with an adjustable (± 3.5 mm in width and height) mounting
141 space for optical alignment and top cover enclosing whole system. The spectrometer was fully
142 assembled by mounting the following components: F110SMA-1064 collimator (Thorlabs), Pi
143 NoIR Camera V2 with Sony IMX219 photodetector (Raspberry Pi Foundation) and SHTC3

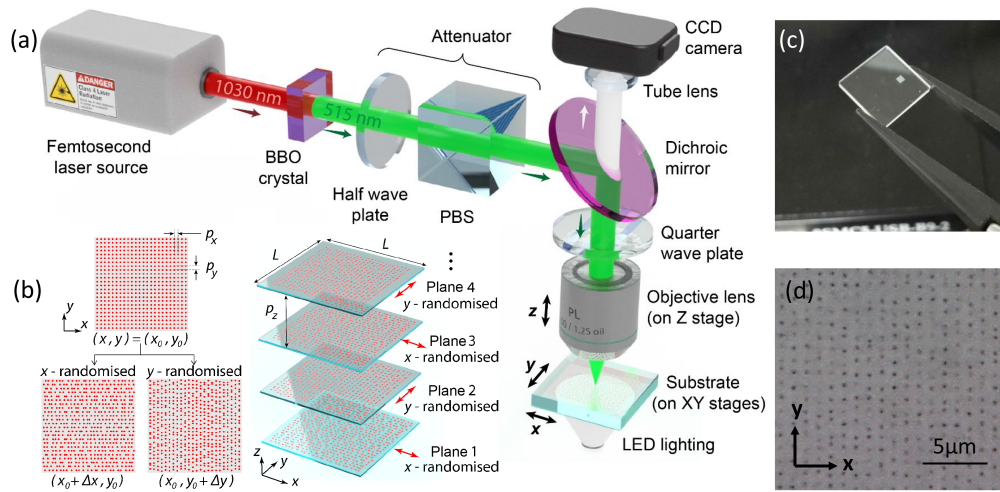


Fig. 1. Production of pseudo-random scattering chips. (a) Femtosecond laser writing experimental setup for scattering medium. (b) Design of scattering planes: a regular grid of scattering voids (red dots) is randomized in either the x or y direction, by adding uniformly distributed random offsets ($\pm 0.4 \mu\text{m}$) to either the x or y coordinates for alternating planes. Mean transverse pitch $p_x = p_y = 1 \mu\text{m}$; plane spacing $p_z = 5 \mu\text{m}$. (c) Photograph of $10 \times 10 \times 1 \text{ mm}^3$ silica substrate and $1 \times 1 \text{ mm}$ scattering pattern. (d) Microscope image of one plane of a y -axis randomized scattering pattern.

144 Temperature/Humidity sensor (Adafruit).

145 3. Calibration and characterization

146 3.1. Speckle pattern registration

147 By calibrating the relation between individual wavelengths and their speckle pattern, we can
 148 reconstruct the spectrum with our spectrometer. A tunable laser source (Thorlabs TLK-L1050M)
 149 with 40 pm linewidth, 1 mW output power and 40 dB optical signal-to-noise (SNR) ratio was used
 150 to calibrate speckle patterns (Fig. 2b) with a 0.05 nm wavelength tuning step in the 1035-1070 nm
 151 range, from which the calibration matrix was formed. The speckle patterns were captured by a
 152 CMOS camera with near-infrared sensitivity (Raspberry Pi NoIR Camera V2). The tunable laser
 153 source (TLS) output was connected to a 99:1 polarisation maintaining splitter coupler where the
 154 99% port output was fed to the collimator inside the spectrometer box, which directs the light
 155 through the scattering chip, and the 1% splitter output port was connected to an optical spectral
 156 analyser (OSA) (Yokogawa AQ6370D) to determine the calibration wavelength. Note the camera
 157 has 3 color channels, but we only use the green channel to reduce processing time.

158 3.2. Single wavelength measurement

159 We first demonstrate the operation of our system to measure a single wavelength [19]. We
 160 collected a test set of speckle patterns for a number of wavelengths and compared each to all of
 161 the speckle patterns in the calibration set. Before the comparison, the pixel values in both the
 162 calibration and test set images were first normalized to the range from -1 to 1 (by subtracting
 163 mean intensity and then normalized) and applied to the Euclidean norm. The comparison is then
 164 done by calculating the sum of the point-wise product of the test measurement pixel values with
 165 those in the calibration set. The calibration wavelength corresponding to the largest product sum
 166 is taken to be the measured wavelength.

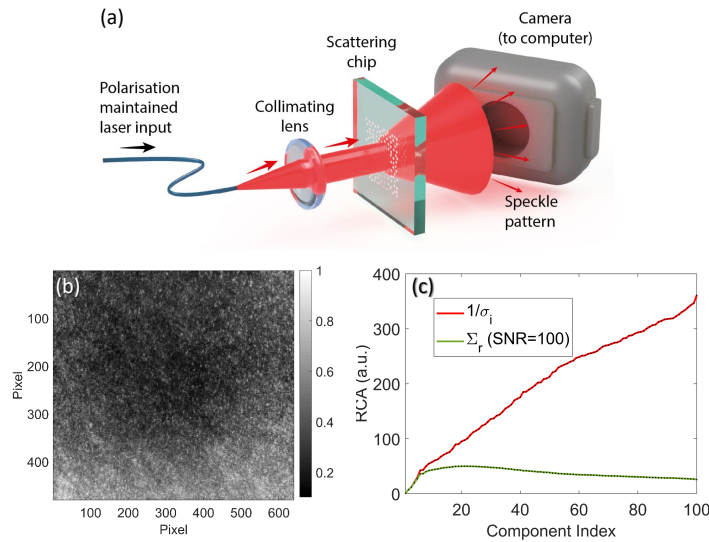


Fig. 2. (a) Scattering spectrometer setup. (b) Example of speckle intensity pattern captured with camera. (c) Inverse of normalized singular values before (red) and after (green) applying Wiener filter. The signal-to-noise (SNR) ratio of the system is set to 100. RCA is the reciprocal component amplitude.

167 3.3. Spectral reconstruction

168 Spectral reconstruction is a less well-posed problem. While wavemeters expect a single
 169 wavelength input, a scattering spectrometer must determine a complete spectrum from an
 170 incoherent superposition of speckle patterns [20]. When taking into account measurement noise,
 171 monochromatic illumination at one wavelength might produce a speckle pattern that is too similar
 172 to that of a neighboring wavelength. The resolution of a speckle spectrometer will be limited by
 173 the spectral memory effect of the scattering system (chip) [21–27]. Highly scattering materials
 174 can enable high resolution provided that the signal-to-noise ratio is sufficient to distinguish all
 175 independent speckle patterns across the working spectrum. Optimal performance can thus be
 176 achieved by engineering the scattering system so that the optimal memory effect matches the
 177 target resolution.

178 The spectral speckle decomposition is undertaken by noting that the measured intensity at
 179 the sensor, $I_s(x, y)$ is a linear combination of the individual speckle patterns that make up the
 180 calibration matrix, $C_s(x, y, \lambda)$, weighted by the *a priori* unknown spectrum, $S(\lambda)$, and an error
 181 term, $\epsilon(x, y, \lambda_i)$, to account for detection noise [28]:

$$I_s(x, y) = \sum_{i=1}^N C_s(x, y, \lambda_i) S(\lambda_i) + \epsilon(x, y, \lambda_i), \quad (1)$$

182 where N is the number of wavelengths distinguishable by the spectrometer. This linear set of
 183 equations can be written more succinctly as the matrix equation $I_s = C_s S + \epsilon$. Ideal, noise-free
 184 measurements, would enable us to calculate the spectrum, S , by considering that $S = C_s^{-1} I_s$ and
 185 calculating the inverse from the singular value decomposition $C_s = U \Sigma V^*$ as $C_s^{-1} = V \Sigma^{-1} U^*$ [29].
 186 Here, U and V are unitary matrices with Hermitian conjugate denoted by $*$, Σ is a diagonal
 187 matrix with the singular values, σ_i , of C_s , and Σ^{-1} is a diagonal matrix with the reciprocal of its
 188 singular values, $1/\sigma_i$. The singular values relates specific combinations of wavelengths (columns
 189 of V) to measurable patterns (columns of U). Each singular value thus represents an independent

190 relation between the wavelengths and their effect on the measured speckle pattern. The relation's
 191 contribution is proportional to the singular value.

192 However, in practice, the matrix C_s may not be full-rank and division by near-zero σ_i -values
 193 would amplify any measurement error, ϵ , that may be present. Error amplification can be avoided
 194 by discarding components with low signal-to-noise [5]. Choosing the cut-off is critical. Instead
 195 of outright discarding components beyond an arbitrary cut-off, we weigh each component as
 196 $\hat{S} = C_r M \approx S$ using the Wiener filter $C_r = V \Sigma_r U^T$, where diagonal matrix Σ_r has as elements

$$\Sigma_{r,ii} = \frac{\sigma_i}{\sigma_i^2 + |n/s|^2}, \quad (2)$$

197 and $n/s = 0.01\sigma_0$, where σ_0 is the first and largest singular value, corresponding to a conservative
 198 signal-to-noise ratio of 20 dB, is an estimate of the noise-over-signal ratio to regularize the
 199 inversion. Fig. 2c shows the inverse singular values before and after applying a Wiener filter with
 200 an example calibration set. It shows that the Wiener filter efficiently suppressed the noise after
 201 setting the SNR level. Note that σ_i in Eqn. 2 has been normalized to the largest number, so the
 202 SNR level is 100 in Fig. 2c. The Wiener filter is the linear operation that minimizes the spectral
 203 error for a given signal-to-noise model. Although more accurate estimates of the spectrum, S ,
 204 may be obtained using non-linear methods, these tend to require resource-intensive iterative
 205 algorithms [30]. The Wiener filter efficiently minimizes the average value of $\|C_s \hat{S} - M\|^2$ in a
 206 single step algorithm.

207 It can be noted that large singular values are approximately inverted, while singular values
 208 below $|n/s|$ are suppressed. The number of independent components with values above the noise
 209 level gives an indication of the number of degrees of freedom that can be quantified, and thus
 210 the number of wavelengths that can be reliably distinguished by the scattering spectrometer.
 211 This places a direct limit on the ratio between the spectrometer's bandwidth and its resolution.
 212 To maximize its bandwidth, the scattering system should thus be designed to have the spectral
 213 memory effect match the target resolution of the spectrometer.

214 3.4. Stability

215 Strongly scattering systems, whilst providing fine spectral resolution, are sensitive to small
 216 environmental fluctuations. Temporal noise or environmental fluctuations in temperature or
 217 humidity manifest as changes to the speckle pattern and their lateral displacement, and such
 218 instabilities ultimately reduce spectral reconstruction accuracy by invalidating the calibration
 219 dataset. To understand their respective importance, we investigate these effects individually.

220 To experimentally evaluate the speckle pattern stability, both long-term and short-term fixed
 221 wavelength experiments were performed. In the first long-term experiment, speckle pattern
 222 images were captured every 10 minutes over 7 days (168 hr), while the spectrometer was connected
 223 via single mode fiber to the tunable laser source at one fixed wavelength. While for the short-term
 224 experiment, the fixed wavelength switches every 20 hours among 5 different wavelength steps,
 225 the total time for short-term experiment is 180 hours. The time, humidity and temperature were
 226 also recorded. To quantify the stability of the captured speckles and reconstruction, three metrics
 227 were proposed: root-mean-square (RMS), transformation matrix [31], and wavelength shift.

228 4. Results and discussion

229 4.1. System stability

230 Based on the long-term stability experimental data, post-collection analysis was achieved by
 231 computing the speckle difference RMS and displacement using the first captured image as the
 232 reference. Results are presented in Fig. 3a-b. The temporal stability analysis revealed fluctuations
 233 in both coordinate displacements (Fig. 3a) and RMS (black line in Fig. 3b).

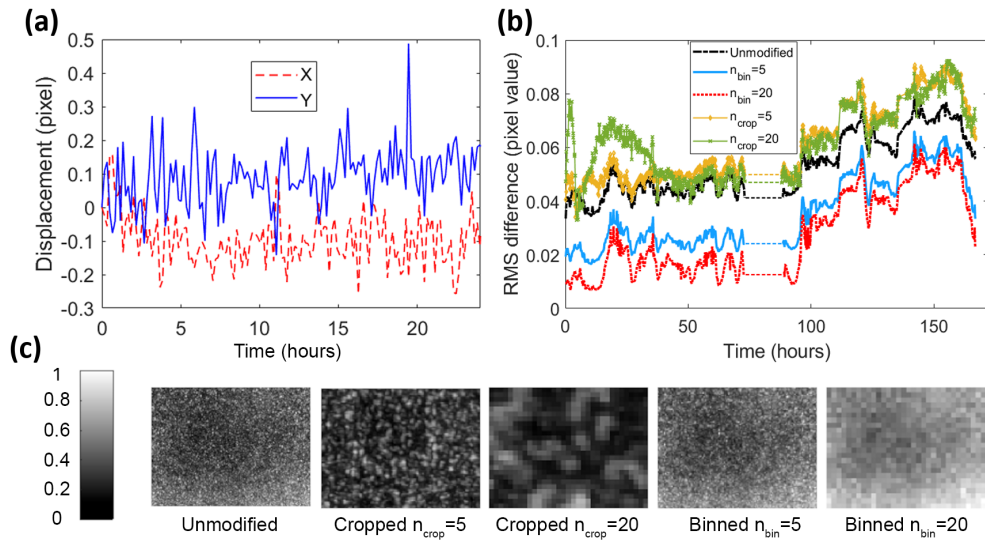


Fig. 3. Cropping and binning effect on speckle stability: (a) Unmodified speckle time-wise displacement in x and y . Note: for clarity, only first 24 hours of displacement are plotted. The behaviour of a fluctuation of x and y coordinates remained the same at all time, (b) Speckle temporal stability (RMS difference) for unmodified, binned and cropped speckle patterns. Note: dashed flat line indicates period when laser was turned off, (c) Example speckle patterns with and without cropping and binning. Pixel peak intensity value normalized to unity.

234 To elucidate on the origin of these speckle differences, the displacement analysis showed
 235 that both x (horizontal) and y (vertical) axis displacements vary in both positive and negative
 236 directions over a total range of 0.5 pixels. It was noted that the angle of rotation ranged between
 237 0 and 0.06° which comes from the registration error and mathematical limit of approximating the
 238 transformation matrix. Hence, its impact on the stability was negligible and transformations can
 239 be treated as purely translational. The RMS analysis shows a low (within 0.08 pixels) difference,
 240 confirming high stability of the scattering chip.

241 In terms of environmental influence, there was no relation between temperature/humidity and
 242 displacements or RMS over the recorded temperature ranges of $22.7\text{--}23.8^\circ\text{C}$ and humidity range
 243 $39.5\text{--}41\%$, implying these day-to-day thermal and humidity variations had minimal systematic
 244 impact on stability.

245 Even if the spectrometer is environmentally stable, its crucial characteristic is the reconstructive
 246 stability (shown in Sec. 3.4). Fig. 4a demonstrates the wavelength reconstruction over 1 week
 247 (168 hr) with the long-term stability experiment data. The blue line is the reconstructed
 248 wavelength, the red dashed line shows the reference wavelength monitored by an OSA. The
 249 fluctuation of the reference wavelength is due to the instability of motor and error from the OSA.
 250 Notice that the reconstructed wavelength only fluctuates within 0.05 nm, which is the finest
 251 tuning step of the laser source.

252 To verify whether device stability improved, it has been cross-compared to another scattering
 253 system of the same resolution. For this purpose, a 50 cm section of straight, jacketed MMF
 254 (FG105LCA, Thorlabs) has been selected. The results (Fig. 4b) confirm the MMF-based system
 255 has lower stability. Whilst the chip retains stability until at least 60 hr and its fluctuation is
 256 within a single laser tuning step, the MMF stability lasts only 3.5 hr and system deviates more
 257 from the reference level (3-8 nm). Moreover, the MMF setup has a greater footprint (fixing the

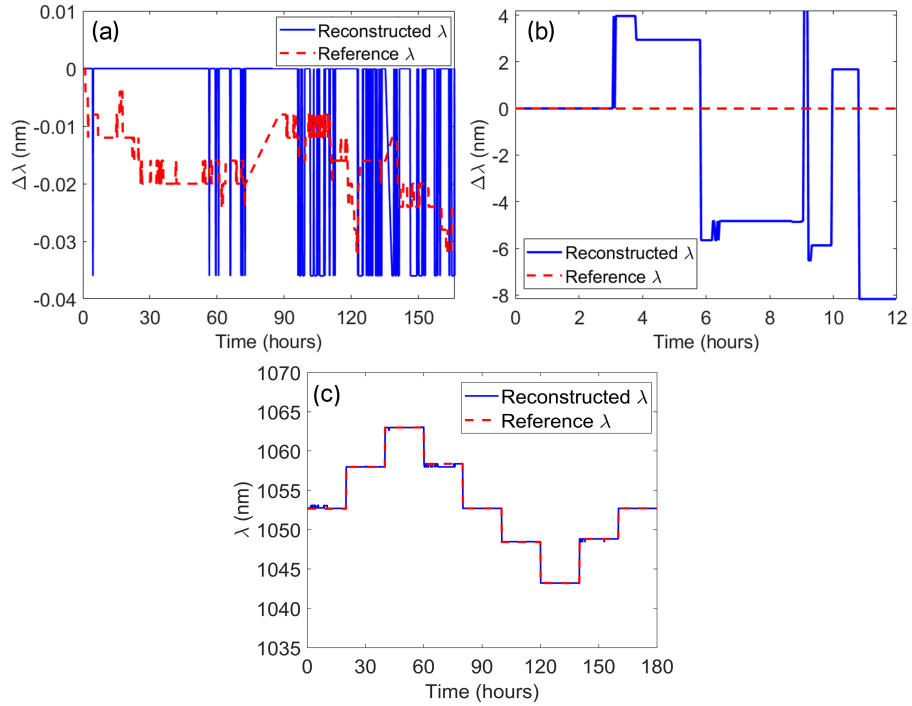


Fig. 4. Wavelength reconstruction stability test for (a) long-term fixed wavelength over 168 hours (chip), (b) long-term fixed wavelength over 12 hours (50 cm MMF) and (c) short-term wavelength steps over a total of 180 hours, comparing the reconstructed and OSA-measured reference wavelengths (chip).

258 MMF required a 50 cm guiding rail whilst the chip-based system can be confined in a 10 times
 259 smaller box). Therefore, the chip-based device has been investigated further by checking the
 260 short-term step wavelengths stability over the full range (1035-1070 nm) of the tunable source
 261 (Fig. 4c). This also shows a stable performance over 180 hours. It is important to note that in
 262 both long-term and short-term experiments, the error did not increase with time, indicating that
 263 the optical properties of the scattering chip do not diverge over the time and are repeatable. This
 264 indicates that a stable performance can be achieved, making a scattering spectrometer a viable
 265 solution for practical applications.

266 4.2. Spectrometer operation

267 To evaluate how well the device can resolve neighbouring wavelengths, we simulated the
 268 superposed speckle pattern for a spectrum consisting of two wavelength lines of equal brightness.
 269 The speckle patterns are selected from the calibration data group with the minimum corresponding
 270 wavelength separation of 0.05 nm. By testing different spectral separations between these
 271 wavelengths and reconstructing the resulting spectrum, the spectral resolution of the scattering
 272 wavemeter can be estimated. Note that we can treat the light fields with a wavelengths in range
 273 1050-1060 nm with separation of 0.05 nm as effectively incoherent since their phase difference
 274 averages out thousands of times during the integration time of 60 ms that was used in our
 275 experiments.

276 Fig. 5a shows the closest distinguishable probe wavelengths reconstructed with our algorithm.
 277 We obtained a 0.1 nm resolution with 0.05 nm calibration wavelength step. Note the speckle
 278 patterns were normalized and applied with a 70% filter where the pixel intensity below 0.7 is set

279 to zero. The resolution can be further improved by using finer tuning steps.

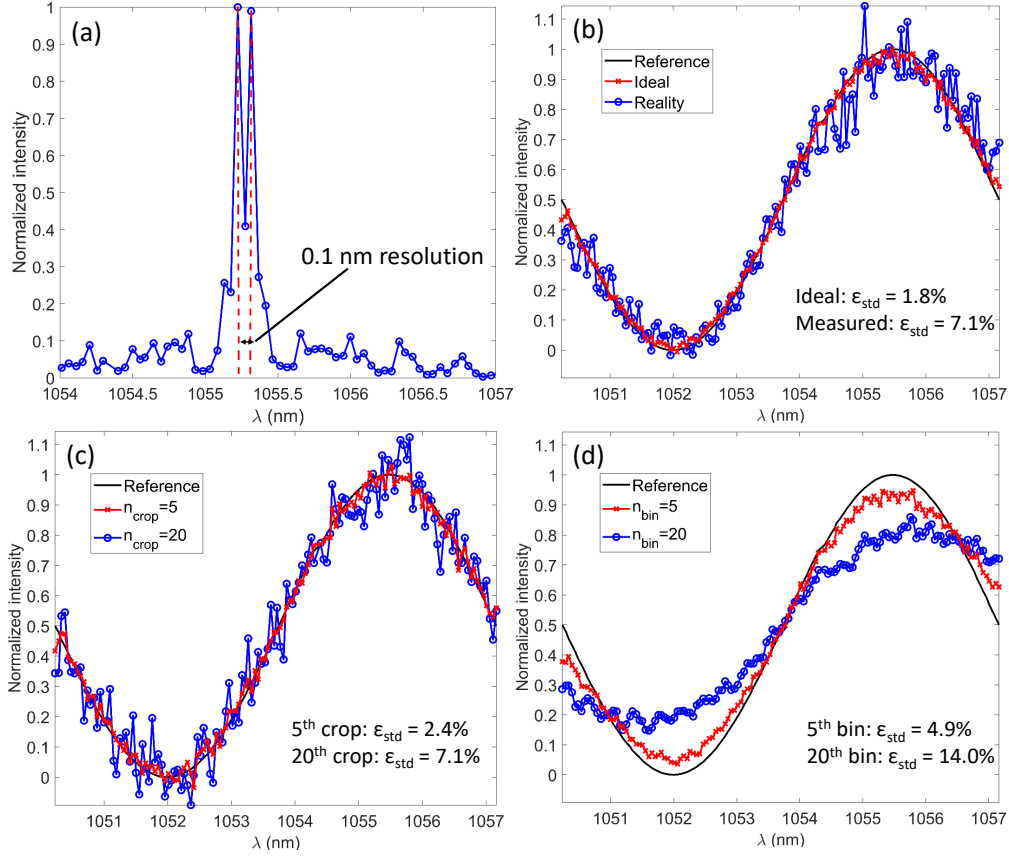


Fig. 5. (a) Reconstruction of a spectrum with two wavelengths separated by 0.1 nm. (b) Spectrum with sinusoidal shape (black), its ideal reconstruction from the calibration data (red), and its reconstruction from test speckle patterns (blue). (c) Impact of binning on reconstructed spectrum, showing an increase in standard deviation. (d) Impact of binning on reconstructed spectrum, showing reduction in spectral contrast. ϵ_{std} is the standard error of the obtained spectra.

280 For the spectrum reconstruction test, we simulate arbitrary test spectra by applying an intensity
 281 modulation $Y(N)$ to each speckle pattern $I(x, y)$ in the calibration group, where N is the number
 282 of wavelengths:

$$I_{mix}(x, y) = \frac{1}{N} \sum_{i=1}^N I_i(x, y) \cdot Y(i) \quad (3)$$

283 By averaging all the modulated images, we obtain a mixed speckle pattern $I_{mix}(x, y)$. The
 284 reconstruction results are shown in Fig. 5b, for a sinusoidal modulation. The black reference
 285 line is the intensity modulation applied to the speckle patterns. Notice here we have two
 286 different reconstructed spectra, the ‘ideal’ and ‘real’ scenarios. In the ‘ideal’ case (Fig. 5b),
 287 both speckle sets were captured at the same motor position of the tunable laser, so they have the
 288 same wavelengths. As a result, the error from wavelength reconstruction is contributed by the
 289 correlation with neighbouring wavelength’s speckle patterns and the accuracy of the algorithm.
 290 This reconstructed spectra intensity showed 1.8% std. error, in the ‘ideal’ case. However, in

291 the ‘real’ case, when capturing the test group images, the motor cannot move to the exact same
292 position as for calibration, so the wavelength reconstruction has its intrinsic error which results a
293 higher error (7.1%) than the ‘ideal’ case.

294 4.3. Optimization of speckle pattern imaging

295 The spectrometer performance depends on how the speckle pattern images are captured by the
296 camera, in particular the sensor size (number of pixels) and pixel pitch. Here, we analyse how an
297 appropriate choice of these parameters can improve the system stability and computational speed.

298 A smaller size sensor allows for a more compact and cost-effective design, and so with the
299 cropping and binning methods we modelled the spectrometer performance for different sensor
300 sizes and pixel pitches as a guidance for sensor selection. Cropped speckle images contain only
301 part of the pattern with a similar noise level as the unmodified image. Such an approach requires
302 careful selection of the cropping window, since the reconstruction is based on the probability of
303 finding characteristic patterns in this area. Therefore, with a reduced image size, the cropping
304 method would give a lower similarity (higher RMS difference), compared to the unmodified
305 dataset (Fig. 3b).

306 To investigate different pixel pitches, $n_{\text{bin}} \times n_{\text{bin}}$ groups of neighboring pixels were combined
307 into super-pixels by summing their intensity values (Fig. 3c). Although the loss of fine detail
308 may be of concern as it increases the similarity of the calibration speckle patterns and may thus
309 hamper faithful spectral reconstruction (and lower RMS difference than unmodified speckle
310 shown in Fig. 3b), this has the additional advantage of alleviating data-processing and memory
311 requirements.

312 From the same data set, we investigated the reconstruction accuracy for binning and cropping
313 in the wavelength range $\lambda = 1035 - 1065$ nm (Fig. 6) with an average laser wavelength tuning
314 step of 0.05 nm. A stronger contrast on the diagonal line signal in Fig. 6a-c, e-f means better
315 correlation between test and calibration wavelength. Fig. 6g shows that cropped images allow
316 reconstruction of spectra with an average standard error of 0.2 nm, and a higher cropping order
317 (n_{crop}) gives higher error. Notice the error is defined as the standard deviation of the difference
318 between reconstructed and calibrated wavelengths. However, the exact error value depends on
319 the cropping order and the position of the cropping region (in Fig. 6g, the red bars indicate
320 the range over which the error varies for different cropping window positions along the main
321 diagonal of the speckle image). It can be noted that for $n_{\text{crop}} \leq 13$ (corresponding cropped image
322 size greater than 49×37 px), the error is relatively small (0.175–0.3 nm); however, once this
323 threshold is exceeded, the error increases because the cropped regions become too similar. It
324 can be explained from the interpretation of the spectrum matrix in Fig. 6e and Fig. 6f, which
325 shows the correlation between the reference and reconstructed wavelengths, where the increasing
326 cropping order amplifies the background noise. These cropped speckle image reconstruction
327 results indicate that our spectrometer can achieve similar wavelength reconstruction resolution
328 with more than 13 times smaller sensor size, which has the potential to be integrated into a
329 smaller system.

330 Next, for single wavelength reconstruction with binned speckle patterns, images of native
331 resolution 640×480 can be binned down to 32×24 pixels ($n_{\text{bin}} = 20$) and still maintain
332 full reconstructive ability within the whole spectral range with a reconstruction accuracy of
333 approximately 0.07 nm; three times better compared to cropping. Furthermore, beyond $n_{\text{bin}} = 5$
334 the image noise is sufficiently reduced so that further binning up to $n_{\text{bin}} = 20$ does not significantly
335 affect the reconstructive error. The rationale for such behavior can be seen from the spectral
336 matrices (Fig. 6b and Fig. 6c) — increasing binning order do not amplify noise, but rather
337 decreases the diagonal, causing it to ‘blur’ into neighboring wavelengths. Initially, it reduces the
338 reconstruction error because the background noise and pixel displacement decrease (each binned
339 pixel was $1/20^{\text{th}}$ of the original one) what caused RMS similarity to decrease from 0.025 to 0.023

340 (Fig. 3b). After reaching a minimum for $n_{\text{bin}} = 5$, the reconstruction error remains relatively low
 341 for binning orders up to $n_{\text{bin}} = 17$. Beyond that, the characteristic speckle patterns become less
 342 distinguishable and the error increases more rapidly. Even for $n_{\text{bin}} = 20$, the spectrometer can
 343 distinguish wavelengths with a reconstruction error as low as 0.07 nm. The binning reconstruction
 344 results indicate better wavelength reconstruction could be achieved with the larger pixel pitch
 345 sensor.

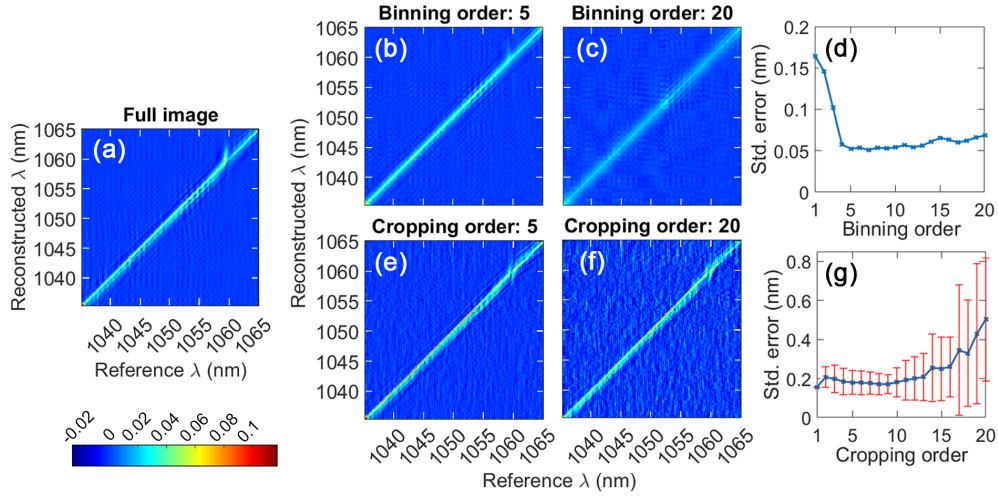


Fig. 6. Impact of cropping and binning on spectral reconstruction. Spectrum matrices for (a) the full image size, (b) $n_{\text{bin}} = 5$ and (c) $n_{\text{bin}} = 20$. (d) Standard reconstruction error vs binning order n_{bin} . (e) Spectrum matrices for $n_{\text{crop}} = 5$ and (f) $n_{\text{crop}} = 20$. (g) Standard reconstruction error vs cropping order n_{crop} , red bars indicate range of standard errors for different positions of cropping window over the main diagonal. Blue line indicates average from all cropping area positions for given order. Spectrum matrices in (a-f) are normalized to unity; better contrast on the diagonal line signal means better correlation between test and calibration wavelength.

346 On the other hand, for the spectrum reconstruction, increasing cropping (Fig. 5c) clearly shows
 347 increased distortion in reconstructed spectrum, which was expected due to less similar speckles
 348 (Fig. 3c and Fig. 6f). The trade-off between sensor size and accuracy need to be considered.
 349 Similarly, increased binning order (Fig. 5d) maintains the spectrum shape, but its extremes
 350 (minima and maxima) are flattened - which is understandable from the reduced correlation for
 351 high spectral intensities (Fig. 6c). Therefore, in the ideal case, the algorithm can reconstruct the
 352 spectrum with a standard error of only 1.8%. After applying the data reduction technique, with
 353 only a 32×24 pixel image after cropping, the algorithm can still reconstruct the spectrum with the
 354 standard error of $\sim 7\%$. However, binning is not an ideal technique for spectrum reconstruction
 355 because a high binning order makes the reconstructed speckle patterns solved more similar to
 356 each other thus results in a flattened spectrum.

357 5. Conclusion

358 Whilst multiple scattering reconstructive systems are extremely sensitive to incoming wavelengths
 359 and can thus detect slightest change in light frequency, their instability limits their use in practice.
 360 We demonstrated a highly stable, compact, low cost spectrometer, based on a fs-laser-written
 361 tailored scattering chip medium.

362 The pseudo-random nano-void pattern scatters 62% of the incoming light and allows for the
 363 resolution of two wavelengths separated by 100 pm. We found that variations in temperature

364 and humidity (22.7-23.8 °C, 39.5-41%) had minimal systematic impact on the spectrometer.
365 Also, during 180 hours stability test, the reconstruction of single wavelength has confirmed with
366 <0.05 nm error. A systematic analysis of the sensor size and effective pixel area showed that
367 adequate binning can improve the robustness to mechanical instabilities. The density and size
368 of the pseudo-random pattern could be adapted to balance the trade-off between resolution and
369 stability for a given particular application.

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376 Author contributions statement

377 M.B., D.B.P., T.V. and G.B. directed the project, Q.S. and P.F. carried out the wavemeter,
378 spectrometer and stability experiments, P.F. designed the spectrometer case, T.V., T.L., Q.S. and
379 P.F. developed the analysis algorithm. Q.S. and P.F. wrote the manuscript. All authors reviewed
380 the manuscript.

381 Disclosures

382 The authors declare no competing interest.

383 Data availability statement

384 Data underlying the results presented in this paper are available in [32].

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