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3 **Supporting Information for**

4 A marine record of Patagonian ice sheet changes over the past
5 140,000 years.

6

7 Julia R. Hagemann^{a,b*}, Frank Lamy^{a*}, Helge W. Arz^c, Lester Lembke-Jene^a, Alexandra
8 Auderset^{b,d}, Naomi Harada^e, Sze Ling Ho^f, Shinya Iwasaki^g, Jérôme Kaiser^c, Carina B.
9 Lange^{h,i,j}, Masafumi Murayama^{k,l}, Kana Nagashima^m, Norbert Nowaczykⁿ, Alfredo Mar-
10 tínez-García^b & Ralf Tiedemann^a

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12 ^aMarine Geology Section, Alfred Wegener Institute Helmholtz Centre for Polar and Marine
13 Research, 27568 Bremerhaven, Germany; ^bClimate Geochemistry, Max Planck Institute for
14 Chemistry, 55128 Mainz, Germany; ^cDepartment of Marine Geology, Leibniz Institute for
15 Baltic Sea Research, 18119 Rostock-Warnemünde, Germany; ^dSchool of Ocean and Earth
16 Science, University of Southampton, Southampton SO14 3ZH, UK; ^eAtmospheric and
17 Ocean Research Institute, The University of Tokyo, 277-8564 Kashiwa, Japan; ^fInstitute
18 of Oceanography, National Taiwan University, 106 Taipei, Taiwan; ^gMARUM-Center for
19 Marine Environmental Sciences, University of Bremen, Leobener Straße 8, 28359 Bremen,

20 Germany; ^hDepartamento de Oceanografía & Centro Oceanográfico COPAS Coastal, Uni-
21 versidad de Concepción, 4030000 Concepción, Chile; ⁱCentro de Investigación Dinámica
22 de Ecosistemas Marinos de Altas Latitudes (IDEAL), Universidad Austral de Chile, Val-
23 divia, Chile; ^jScripps Institution of Oceanography, University of California San Diego, La
24 Jolla, United States; ^kFaculty of Agriculture and Marine Science, Kochi University, 783-
25 8502 Kochi, Japan; ^lCenter for Advanced Marine Core Science, Kochi University, 783-
26 8502 Kochi, Japan; ^mResearch Institute for Global Change, Japan Agency for Marine-Earth
27 Science and Technology, 237-0061 Yokosuka, Japan; ⁿClimate Dynamics and Landscape
28 Evolution, Helmholtz Centre Potsdam German Research Centre for Geosciences, 14473
29 Potsdam, Germany

30

31 *Julia R. Hagemann and Frank Lamy.

32 **Email:** Julia.Hagemann@awi.de; Frank.Lamy@awi.de

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35 **This PDF file includes:**

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37 Supporting text

38 Figures S1 to S9

39 Tables S1 to S4

40 SI References

41

42

43 ***SI Age Model***

44 At orbital timescales, we fitted the $\delta^{18}\text{O}$ record of the intermediate-dwelling foraminifera
45 *Globorotalia truncatulinoides* to the $\delta^{18}\text{O}$ intermediate Pacific stack (IP; **Fig. S2, Table**
46 **S1; 1**). We preferred planktic deep dwellers, which reflect intermediate water mass char-
47 acteristics (2, 3) over benthic foraminifers because the latter might be influenced by high,
48 but largely unconstrained, ages of Pacific Deep Water masses prevalent at the study site
49 (4). In particular, during glacials and glacial terminations, aged Pacific Deep Water (5, 6)
50 may well lead to unanticipated temporal offsets when aligned with reference stack records,
51 which may not fully compensate for such bias (1, 7, 8). In contrast, intermediate waters in
52 the study area are not known to exhibit similar aging behavior (9) and were thus used here.
53 Based on the good agreement between the EDML (10) oxygen isotope reference record
54 and the XRF-based Rubidium (Rb) record from our site, thought to reflect fine-grained
55 terrigenous input as a clay mineral component, we added four more tuning points, mainly
56 around MIS 4 (**Fig. S2**), which are all within the principal error ranges of the existing ox-
57 ygen isotope-based age model. We checked our age model for consistency with $\delta^{18}\text{O}$ of *G.*
58 *bulloides* and alkenone-derived SSTs on EDML, but did carry not out further adjustments.

59 The age model of the younger section of our core (<42 ka) is based on ^{14}C dating (*G.*
60 *bulloides*). We calibrated our samples with MARINE20 (11) and an ΔR of 400 years (12)
61 using the program Calib 8.2 (**Table S2**). We choose all dating points between the TIPs
62 (**Fig. S2**) to avoid the major impact of glacial meltwater. We also associated all TIPs with
63 Antarctic stadials, considering the 2-sigma range of the radiocarbon calibration-derived

64 calendar ages. An association of all TIPs with Antarctic interstadials instead was not possible for TIPs 2c, 2b, and 3b. In addition, we also used the onset and termination of the geomagnetic Laschamps excursion centered at 41 ka, which matches well with two ^{14}C ages (Fig. S3; Table S2). We conclude that the millennial-scale variations during MIS 3 and MIS 2 are related to Antarctic timing, indicating PIS advances occurring during stadials.

70

71 ***SI Methods***

72 **Sediment Core.** Piston core MR16-09 PC03 was retrieved from the Southeast Pacific, ~150 km west of the Chilean coast (46° 24.32' S, 77° 19.45' W; 3082 m water depth) during RV Mirai Cruise Leg 2 of the Trans South Pacific Project Expedition (13). The 17.53 m long sediment core consists mainly of glacial gray to olive gray silty clay with diatoms and nannofossils. Lighter, more carbonate-rich intervals are restricted to the thin Holocene sequence at the top and the last interglacial sediments towards the bottom of the core, accompanied by increased biogenic components. In addition, a ~15 cm thick dark brownish tephra layer in core depth ~9.1 m was identified as a turbidite and removed from the core sequence.

81

82 **Biomarkers.** A total of 224 samples, taken every 4.3 cm in the interglacial intervals and 8.6 cm in the glacial intervals, were analyzed for *n*-alkanes, alkenones, and isoprenoid (not used in this study), and branched glycerol dialkyl glycerol tetraethers (GDGTs). Sample

85 preparation, extraction, and measurement were done at the Max Planck Institute for Chem-
86 istry in Mainz, according to the method described in Auderset *et al.* (14). In this method,
87 the biomarker extraction was done by an Accelerated Solvent Extraction (DIONEX ASE
88 350; Thermo Scientific). During the extraction, the samples were separated into two frac-
89 tions (fraction I containing *n*-alkanes and alkenones; fraction II containing GDGTs)
90 through using an ASE. The ASE cells were filled with 16 g silica gel (deactivated by adding
91 5 % milli-Q water) and 2 g of freeze-dried and homogenized sediment samples. Subse-
92 quently, the cells were filled with solvent and heated up to 100° C for 5 min. The extraction
93 was performed using *n*-hexane (15 ml; 1-time), DCM (22 ml; 3-times), and DCM:MeOH
94 (1:1, v/v; 22 ml; 3-times). Fractions I (*n*-Hexane and DCM) and II (DCM:MeOH) were
95 collected separately. After ASE extraction, 80 µl of 2-nonadecanone and hexatriacontane
96 were added to fraction I and 60 µl of C₄₆-GDGT (15, supplied by Pandion Labs) to fraction
97 II as internal standards. In a Rocket Evaporator (Genevac – SP Scientific), both fractions
98 were dried under vacuum and converted into 1.5 ml vials. For the measurement, fraction I
99 was dissolved in 30 – 80 µl isooctane, and 5 µl were injected into a GC-FID, a Gas Chro-
100 matography – Flame Ionization Detector (GC-FID) 7890B GC System from Agilent Tech-
101 nologies with helium as the carrier gas, a VF200ms column (60 m x 250 µm x 0.25 µm),
102 and a PTV injector. The oven temperature increased from 44 to 300° C at 20° C/ min and
103 held at 300° C for 14 minutes. Fraction II was diluted in 400 µl *n*-hexane:isopropanol
104 (1.5 %) and 20 µl were injected in a HPLC-MS, a High-Performance Liquid Chromatog-
105 raphy (Agilent, 1260 Infinity) with a single quadrupole mass spectrometer detector (Ag-

106 ilent 6130) with the method as described in Hopmans *et al.* (16). We measured the follow-
107 ing masses: m/z 744 (C_{46} standard), m/z 1302.3 (GDGT-0), m/z 1300.3 (GDGT-1), m/z
108 1298.3 (GDGT-2), m/z 1296.3 (GDGT-3), m/z 1292.3 (crenarchaeol and crenarchaeol iso-
109 mer), m/z 1022 (GDGT-Ia), m/z 1020 (GDGT-Ib), m/z 1018 (GDGT-Ic), m/z 1036 (GDGT-
110 IIa and IIa'), m/z 1034 (GDGT-IIb and IIb'), m/z 1032 (GDGT-IIc and IIc'), m/z 1050
111 (GDGT-IIIa and IIIa'), m/z 1048 (GDGT-IIIb and IIIb'), m/z 1046 (GDGT-IIIc and IIIc').
112 We used the GC Analysis Software ChemStation by Agilent Technologies to quantify *n*-
113 alkanes and alkenones and Agilent MassHunter Quantitative Analysis Version B.07.01 for
114 isoprenoid- and branched GDGTs.

115

116 **Mass Accumulation Rate (MAR).** *n*-alkanes, brGDGT and titanium (Ti), considered here
117 as terrestrial proxies (e.g., 17, 18-22), are presented as mass accumulation rates (MAR).
118 We determined the accumulation rate as follows (23, 24):

119

$$120 \quad \text{MAR} = \text{SR} * \text{terrigenous Input} * \text{DBD}$$

121

122 MAR – Mass Accumulation Rate [$\text{ng/ka} \cdot \text{cm}^2$], SR - Sedimentation rate [cm/ka], DBD -
123 Dry Bulk Density [g/cm^3], and terrigenous input based on the concentrations of *n*-alkanes
124 and branched GDGTs [ng/g Sed.]. For Ti, counts per second (cps) were used to determine
125 MAR instead of concentrations. Concentrations, counts per second, and MAR of the *n*-
126 alkanes, brGDGTs, and Ti can be found in **Fig. S5**.

127

128 ***n*-alkanes and branched GDGTs.** Long-chain ($C_{27} - C_{33}$) *n*-alkanes are leaf waxes from
129 terrestrial plants that can be used to infer changes in the transport of terrestrial material to
130 the ocean (17). We used the carbon preference index (CPI), which compares even and odd
131 carbon numbered *n*-alkanes, to assess potential changes in *n*-alkanes sources through time
132 (**Fig. S5B**). CPI values of 1 are typical of mature samples, while leaf waxes are character-
133 ized by high CPI values (25, 26). In the absence of petrogenic markers, low CPI values
134 have been attributed to increased microbial reworking of the odd carbon numbered *n*-al-
135 kanes (22, 27). At our site the CPI varies between ~5-6 during warm phases and ~3 during
136 millennial scale cold periods when the terrestrial input is the highest. These observations
137 suggest a change in *n*-alkane sources during these events. The higher CPI values observed
138 during the warmer intervals point towards a dominant contribution from unaltered higher
139 plant material that may have been transported to our core site by runoff or winds. In con-
140 trast, the higher proportion of microbially altered *n*-alkanes during these events is con-
141 sistent with a more important contribution of older *n*-alkane deposits transported to the
142 sediment by glacial erosion associated with the advance of the PIS.

143 The average chain length (ACL; **Fig. S5C**), which traces changes in the distribution of
144 odd *n*-alkanes, can provide information on the source vegetation, with higher values (i.e.
145 30.66 ± 0.83) typically associated with C4 grasses and lower values (i.e. 29.00 ± 0.83) asso-
146 ciated with C3 trees, shrubs and grasses (26). At our site ACL values range between ~30
147 during warmer intervals and 29.7 during millennial-scale cold periods. These observations
148 are consistent with wetter conditions during these periods favoring the development of a
149 larger proportion of C3 vegetation.

150 Branched GDGTs (brGDGTs) are produced mostly in soils (28). The branched vs iso-
151 prenyl index (BIT; **Fig. S5E**) provides information about the terrigenous fraction of all
152 GDGTs in the sediment, with higher values indicating a dominance of soil GDGTs (19).
153 However, it has also been shown that in some open ocean location the BIT can be affected by
154 the production and/or preferential degradation of isoprenoid GDGTs, and high BIT values
155 can be obtained despite of lower brGDGTs inputs (20). Thus, it is recommended to com-
156 pare BIT values with brGDGTs concentrations and MARs (20). Our BIT index shows
157 higher values during the TIPS, together with higher concentrations and MAR of brGDGTs
158 indicating a higher proportion of soil-derived organic matter during these events.

159

160 **Sea surface temperature (SST).** We used the Unsaturation Ketone index (U^{K}_{37} – index;
161 $U^{K}_{37} = [C_{37:2}] - [C_{37:4}] / ([C_{37:2}] + [C_{37:2}] + [C_{37:4}])$) with a chain length of 37 carbon atoms, with
162 the di-, tri- and tetra-unsaturated alkenones $C_{37:2}$, $C_{37:3}$ and $C_{37:4}$ (29) and the modified $U^{K'}_{37}$
163 – index ($U^{K'}_{37} = [C_{37:2}] / ([C_{37:2}] + [C_{37:2}])$) after Prahl and Wakeham (30) excluding the tetra-
164 unsaturated alkenone $C_{37:4}$. The SSTs based on both indices were calculated with $SST =$
165 $(U^{K}_{37} + 0.104) / 0.04$ (31) and $SST' = (U^{K'}_{37} - 0.044) / 0.033$ (32).

166

167 **XRF-Scan.** Scanning X-ray fluorescence spectroscopy was performed by an ITRAX mi-
168 cro-XRF scanner at Kochi Core Center, Kochi University, Japan, using a Mo X-ray tube
169 with settings of 30 kV, 55 mA, down-core step sizes of 0.5 cm and a counting time of 15
170 seconds per step.

171

172 **$\delta^{18}\text{O}$.** The stable oxygen isotope ratio ($\delta^{18}\text{O}$) of planktic foraminifera were measured in the
173 Marine Geology stable isotope laboratory facility at the Alfred-Wegener-Institute in Brem-
174 erhaven, Germany. 121 samples of *G. bulloides* (size range 250–500 μm), in addition to
175 the dataset of Iwasaki *et al.* (4), and a total of 171 samples of *G. truncatulinoides* (size
176 range 250–500 μm) were picked to determine $\delta^{18}\text{O}$. Analyses were carried out on a Thermo
177 MAT253 mass spectrometer and a Thermo MAT253Plus, both connected to Kiel IV
178 CARBO units. All values are reported as ‰ vs. V-PDB. Calibration was obtained with
179 NIST 19 and an internal carbonate standard of Solnhofen limestone.

180

181 **Magnetostratigraphy.** The core was sampled every ~23 mm with cubic plastic boxes
182 (6 cm³). Remanence Measurements were performed with a 2G Enterprises superconduct-
183 ing long-core cryogenic magnetometer with in-line 3-axis alternating field demagnetizer.
184 All samples were stepwise demagnetized and demagnetization results were subjected to
185 principal component analysis in order to determine the characteristic remanent magnetiza-
186 tion (ChRM). An anhysteretic remanent magnetization (ARM) was imparted as a proxy for
187 the concentration of magnetic particles and subsequently also stepwise demagnetized. The
188 slope of NRM versus ARM of common demagnetization levels was then used to determine
189 the relative paleointensity (RPI). Documented geomagnetic excursion (Laschamps 41 ka),
190 as well as the obtained RPI curve, correlated to the reference curve provided by Liu *et al.*
191 (33), provided age tie points.

192

193 ***SI Calibration index and C_{37:4}***

194 The two indices, $U^{K_{37}}$ and $U^{K'_{37}}$, differ in alkenone $C_{37:4}$, which occurs mainly in cold
195 regions (e.g., 34) and was therefore removed from the global $U^{K'_{37}}$ index (30). The choice
196 of the index at this site is critical for a qualitative statement about the effect of temperature
197 on ice sheet dynamics, as they are inconsistent during TIPs. The alkenone $C_{37:4}$ is particu-
198 larly conspicuous because it is present in relatively large amounts (10 – 25 %) during TIPs
199 (**Fig. S9A** – gray area). Otherwise, $\%C_{37:4}$ is ≤ 10 , with only minor differences between the
200 indexes (**Fig. S9A**). This large amount of $\%C_{37:4}$ during TIPs is reflected in a linear rela-
201 tionship (**Fig. S9B**) between $U^{K_{37}}$ and $U^{K'_{37}}$ (when slope = 1, then $U^{K_{37}} = U^{K'_{37}}$), with a
202 slope of ~ 0.9 ($\%C_{37:4} < 5$), ~ 0.8 ($\%C_{37:4}$ between 5 – 10), and ~ 0.2 ($\%C_{37:4} > 10$). Studies
203 have demonstrated a correlation of $C_{37:4}$ with the temperature only when $\%C_{37:4} \leq 5$ (35-
204 37), implying that the $U^{K_{37}}$ – based temperature calibration at our site is biased during TIPs.

205 In addition to cold temperatures, $C_{37:4}$ also appears to correlate with salinity negatively
206 (34, 36, 37). The correlation between increasing $\%C_{37:4}$ and decreasing salinity could be
207 due to (1) an adjustment in the biosynthesis of the ocean-dominant haptophyte *Emiliana*
208 *huxleyi* or (2) increased occurrence of other alkenone-forming haptophyte species. How-
209 ever, adjustment of *E. huxleyi* biosynthesis results in only a small increase (≤ 10 $\%C_{37:4}$)
210 (35), suggesting that salinity changes during TIPs are not the only reason for the high
211 $\%C_{37:4}$ values at our site. High (> 10) $\%C_{37:4}$ values could be attributed in some studies to
212 other alkenone-forming haptophyte species that exhibit a different biochemical response
213 to growth temperature than *E. huxleyi* (e.g., 35, 38). *E. huxleyi* and *Gephyrocapsa oceanica*
214 are the predominant alkenone-producing species in the ocean (e.g., 39, 40, 41) but may be

215 locally complemented by other haptophyte species, e.g., in regions of sea ice formation
216 (42). They also occur in considerable genetic variation in lakes, where they have higher
217 levels of unsaturation (e.g., 38).

218 High %C_{37:4} values were also observed in previous studies. For example, Bendle *et al.*
219 (34) found high values (40 – 77 %) of %C_{37:4} in polar waters, intermediate values (0 –
220 25 %) in Arctic and Norwegian coastal waters, and low values (0 – 3 %) in Atlantic waters.
221 This high occurrence of %C_{37:4} in polar and coastal waters supports the assumption that
222 other haptophyte species may have influenced sites in regions with sea ice or increased
223 terrigenous input. %C_{37:4} data from core GeoB3327-5 (Fig. S9C; 43), located ~425 km
224 further offshore, shows no %C_{37:4} values >10 for ~500 ka, implying that freshwater input
225 never reaches this site. Other %C_{37:4} offshore data from site PS75/034-2 (Fig. S9D; 43)
226 further south at 54° S instead show C_{37:4} values up to 20 % during glacial phases. In this
227 case, the elevated %C_{37:4} values could be related to northward extending sea ice during the
228 glacials (44, 45). The sea ice most likely did not reach the site but the increased northward
229 expansion of Antarctic water masses (46, 47) may have transported alkenones from the
230 phylogenetic Group of 2i Isochrysidales which are associated with sea ice (42) to site
231 PS75/034-2.

232 The high %C_{37:4} values during TIPs at our site suggest that alkenones deposited in the
233 sediment are not exclusively from *E. huxleyi*. Other alkenone-forming haptophyte species
234 could have been deposited autochthonous as well as allochthonous. The sharp drop in δ¹⁸O
235 of the surface-dwelling foraminifer *Globigerina bulloides* during the TIPs supports a large
236 freshwater influx that may have caused a shift in the haptophyte community. On the other

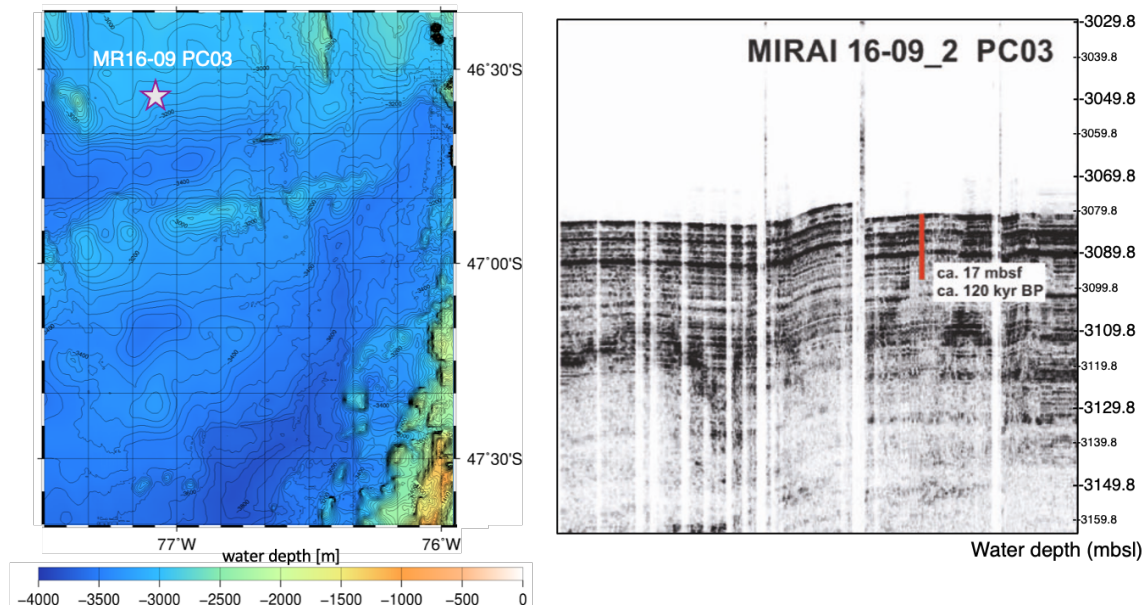
237 hand, the high %C_{37:4} values are consistent with a high terrigenous input (e.g., *n*-alkanes,
238 Ti, brGDGTs), which does not exclude alkenones produced in lakes. Furthermore, the dif-
239 ferential biochemical response of other haptophyte species to growth temperature most
240 likely also affects the U^{K'}₃₇ – index. Therefore, Alkenone SSTs during TIPs cannot be un-
241 ambiguously determined at our site due to the potential influence of other haptophyte spe-
242 cies. Instead, we use here %C_{37:4} as an additional proxy for freshwater supply. The terri-
243 genous influence during the TIPs on the SSTs is illustrated by the absence of the charac-
244 teristic orbital-scale fluctuations, like MIS 4 or the last glacial maximum (LGM), which
245 are not detectable in U^{K'}₃₇ – derived SSTs and only hinted at in U^K₃₇ – derived SSTs. Stud-
246 ies north at ~41° S (**Fig. S8B**; 48, 49) and south at ~53° S (**Fig. S8E**; 50) of our site show
247 a very prominent MIS 4 and LGM, instead.

248 In summary, both indices are most likely influenced by other haptophyte species at our
249 site. However, the U^K₃₇ – index shows a consistent pattern of cooler SSTs during TIPs,
250 broadly consistent with Antarctic stadials and coolings in the more northern site ODP 1233
251 and southern MD07-3128. Haptophytes living in lakes have a higher degree of unsaturation
252 but still, show a correlation with temperature, so U^K₃₇ – based calibrations are applied (38).
253 With such an influence from freshwater-induced haptophyte community or lake-derived
254 alkenones, C_{37:4} would still correlate with temperature, albeit in a different biogeochemical
255 dependence. Such a correlation of the alkenone C_{37:4} with temperature is consistent with
256 studies from the North Atlantic and Nordic Seas with a high number of samples north of
257 70° N, an area of sea ice formation, showing a stronger correlation of the U^K₃₇ – index than
258 the U^{K'}₃₇ – index with SST (51, 52) and no correlation of the U^{K'}₃₇ – index with SST at all

259 (36). It is possible, although highly speculative, that U^K_{37} – index is also the better choice
260 at our site and reflects the general temperature changes, albeit with a larger amplitude than
261 expected.

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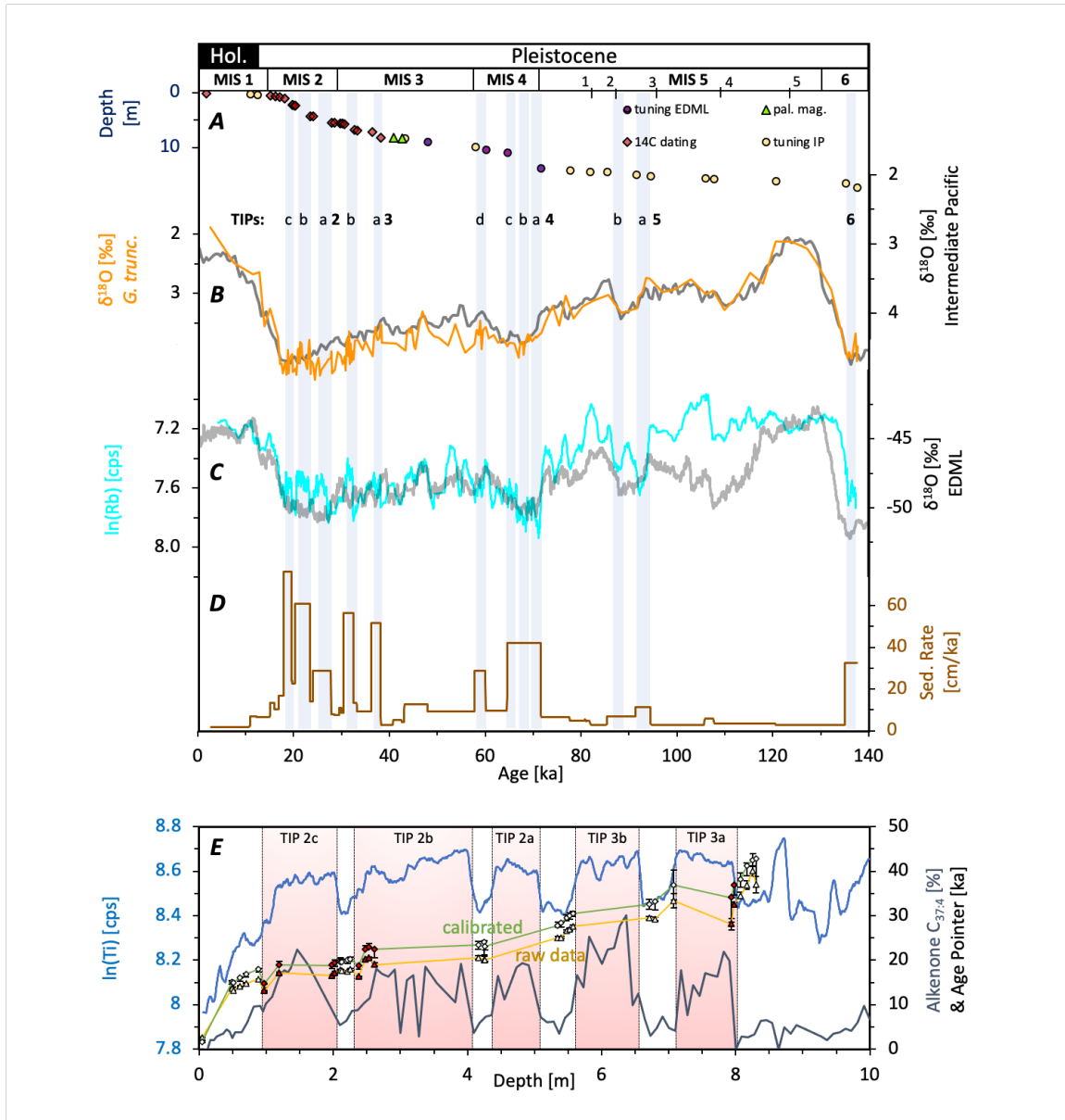
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266 **Fig. S1.** Site location of MR16-09 PC03 with a bathymetric map (left) and 3.5 kHz acoustic
267 sub-bottom profiler (right).

268

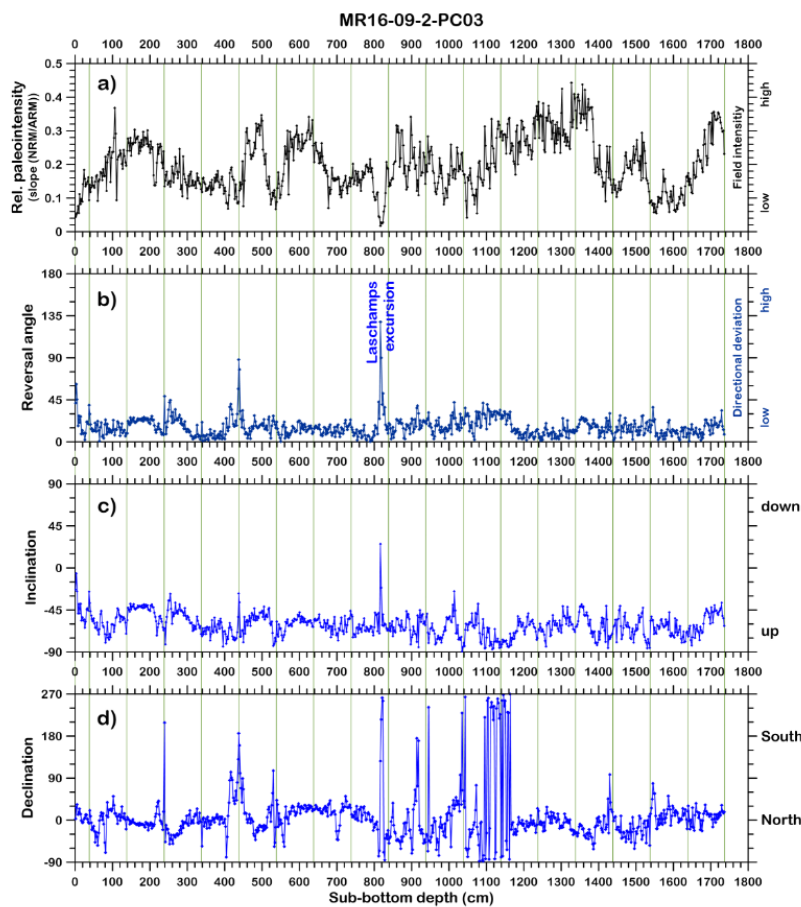


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271 **Fig. S2.** Age model of core MR16-09 PC03. Gray shadings and numbers at the top mark
 272 the times of the Terrigenous Input Phase (TIP). (A) ^{14}C and tuning points. (B) $\delta^{18}\text{O}$ values
 273 of Intermediate Pacific waters (IP; 1) with $\delta^{18}\text{O}$ values of the deep-dwelling (200 - 500 m;
 274 53, 54) foraminifer *G. truncatulinoides*. (C) $\delta^{18}\text{O}$ values of the Antarctic ice core EDML

275 (10) with XRF-scan of the element composition of rubidium (Rb). (D) Sedimentation rate
276 of the Age model. (E) Depth vs. age plot of surface-dwelling foraminifer *Globigerina*
277 *bulloides* ¹⁴C raw data (yellow) with the error of measurement and ¹⁴C calibrated data
278 (green) with MARINE20 (11), an ΔR of 400 years (12), and an error bar showing the 2-
279 sigma range. In the background are shown XRF-derived ln(Ti) and the %C_{37.4} alkenone to
280 evaluate the sections of terrigenous input. Only samples from sections with little terri-
281 genous input were chosen for the age model to avoid a bias of glacier meltwater. Red-
282 shaded stipes show sections of high terrigenous input. The age model did not consider red-
283 marked ¹⁴C age pointers (raw and calibrated data).

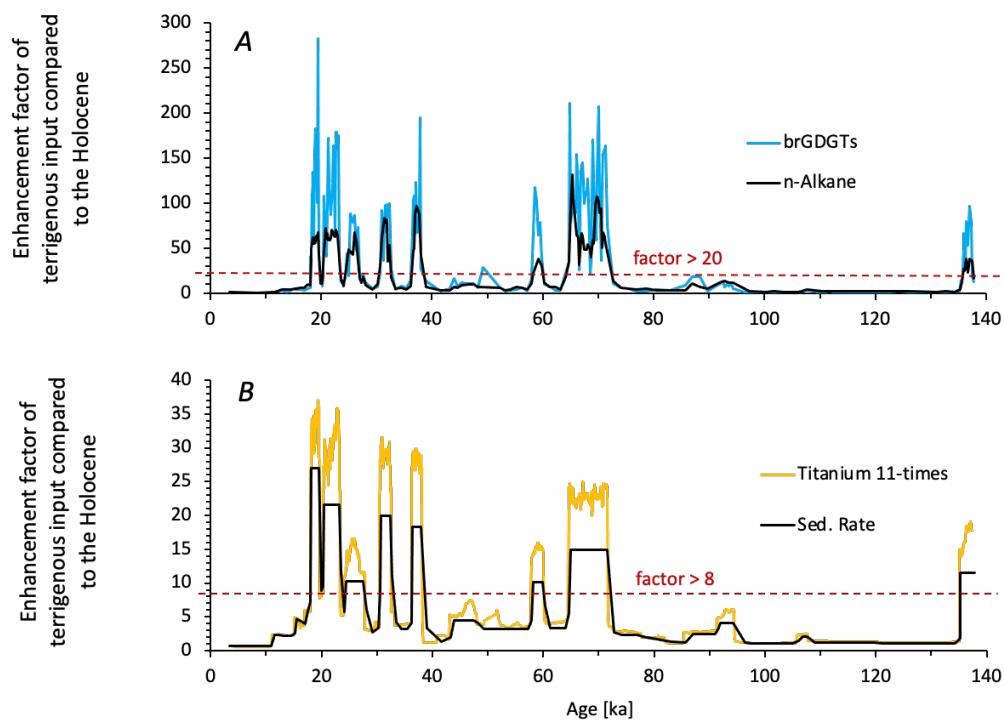
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286 **Fig. S3.** Magnetostratigraphic results from Mirai core PC03: (a) relative paleointensity es-
 287 timated by the slope of NRM (natural remanent magnetization) versus ARM (anhysteretic
 288 remanent magnetization) of common demagnetization steps, (b) reversal angle, defined as
 289 the angle along a great circle between the expected dipole direction and the measured di-
 290 rection, shown as Inclination in (c) and declination in (d). Directions pointing up and north-
 291 ward (down and southward) are associated with normal (reversed) polarity. The geomag-
 292 netic Laschamps excursion at 41 ka (55) is also associated with low (relative) paleointen-
 293 sities.

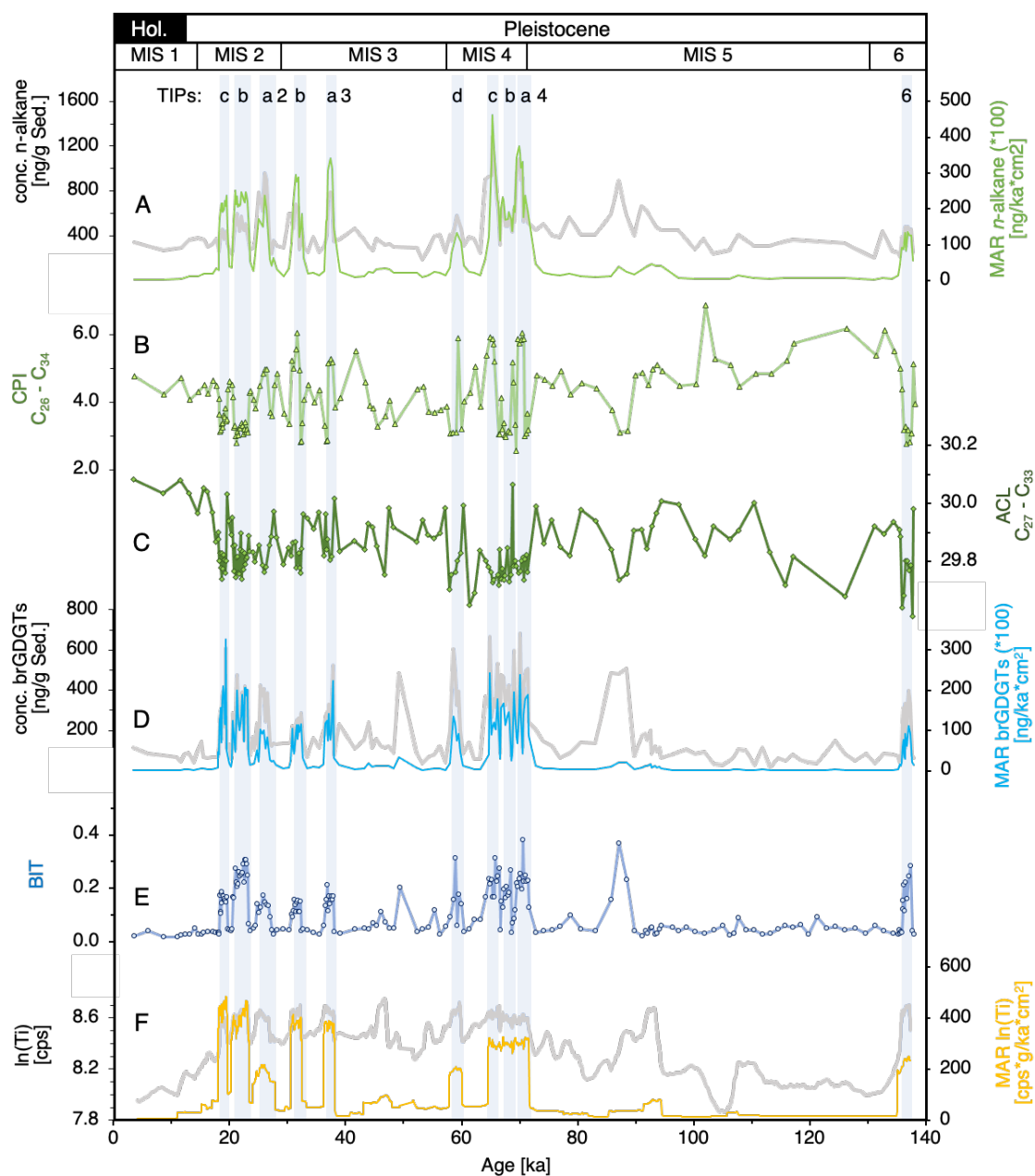
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297 **Fig. S4.** Biomarker (A) and sedimentation rate (B) normalized to the averaged Holocene
 298 background sedimentation.

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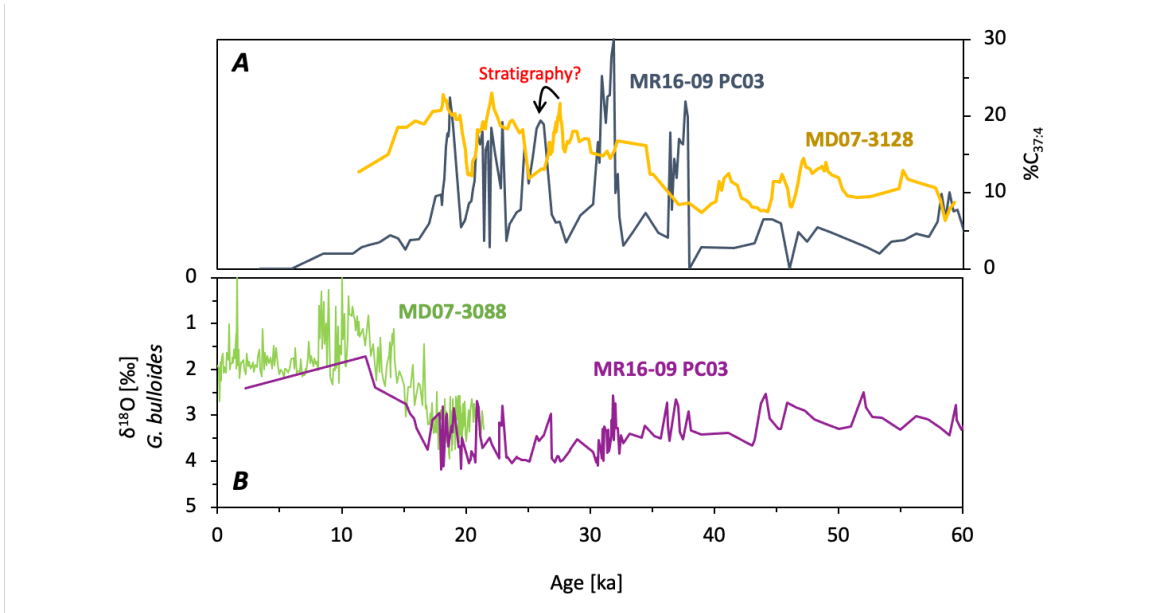


300

301 **Fig. S5.** Concentrations (left axis) and Mass Accumulation Rates (MAR; right axis) of *n*-
 302 alkanes (A). Carbon preference index (CPI) for the *n*-alkanes C₂₆ – C₃₄ (B). Average chain
 303 length (ACL) for the *n*-alkanes C₂₇ – C₃₃ (C). Concentrations (left axis) and Mass Accu-

304 mulation Rates (MAR; right axis) of branched GDGTs (brGDGTs) (**D**). Branched vs iso-
305 prenoic index (BIT) to determine the terrigenous fraction in the sediment (**E**). Concentra-
306 tions (left axis) and Mass Accumulation Rates (MAR; right axis) of Titanium ($\ln(\text{Ti})$),
307 whereby the MAR of Ti is based on their cps and not on their concentration. Furthermore,
308 for Ti, an 11-point running mean was chosen (**F**). Gray shadings and numbers at the top
309 mark Terrigenous Input Phases (TIP).

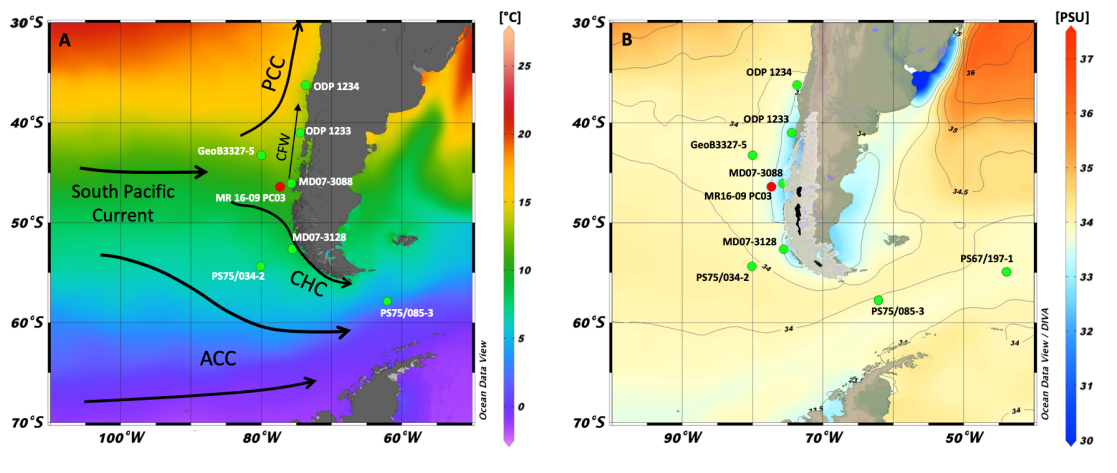
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311

312 **Fig. S6.** Comparison of the alkenone %C_{37:4} of the southern core MD07-3128 (yellow line;
 313 50) with MR16-09 PC03 (gray line; this study). The data of core MD07-3128 were cleaned
 314 from outliers and shown here as a 5-point moving average (A). Comparison of the δ¹⁸O
 315 values of the foraminifer *G. bulloides* of the nearby core MD07-3088 (green line; 56, 57)
 316 with MR16-09 PC03 (violet line; this study) (B).

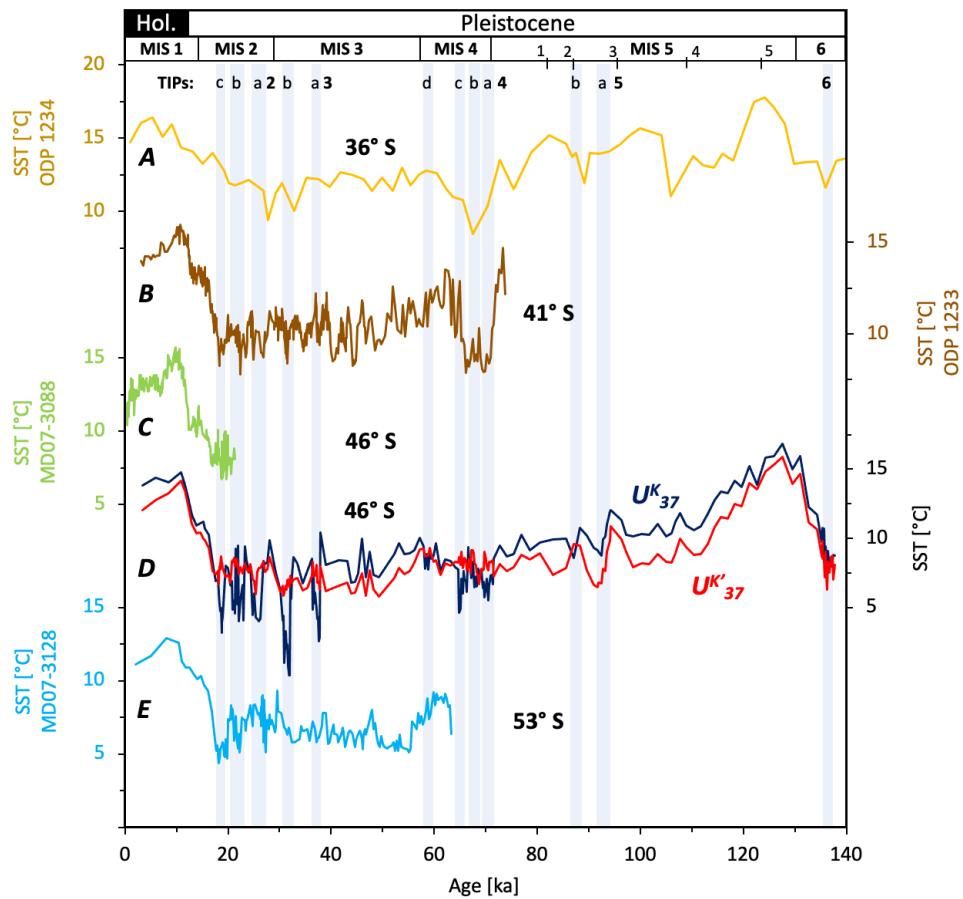
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318

319 **Fig. S7.** World Ocean Atlas (WOA13) derived (A) sea surface temperature (SST; 58) and
 320 (B) sea surface salinity (SSS; 59) maps of the study area with for this paper relevant cores
 321 sites (GeoB3327-5, PS75/034-2: 43, PS75/085-3, PS67/197-1 47, ODP 1233: 48, MD07-
 322 3128: 50, MD07-3088: 56, ODP 1234: 60). ACC: Antarctic Circumpolar Current; PCC:
 323 Peru-Chile Current; CHC: Cape Horn Current; CFW: Chilean Fjord Water; Semi-transparent
 324 pale shading: Patagonian Icesheet (PIS) during its maximum extension. Black poly-
 325 gons: Recent icefields in Patagonia (61).

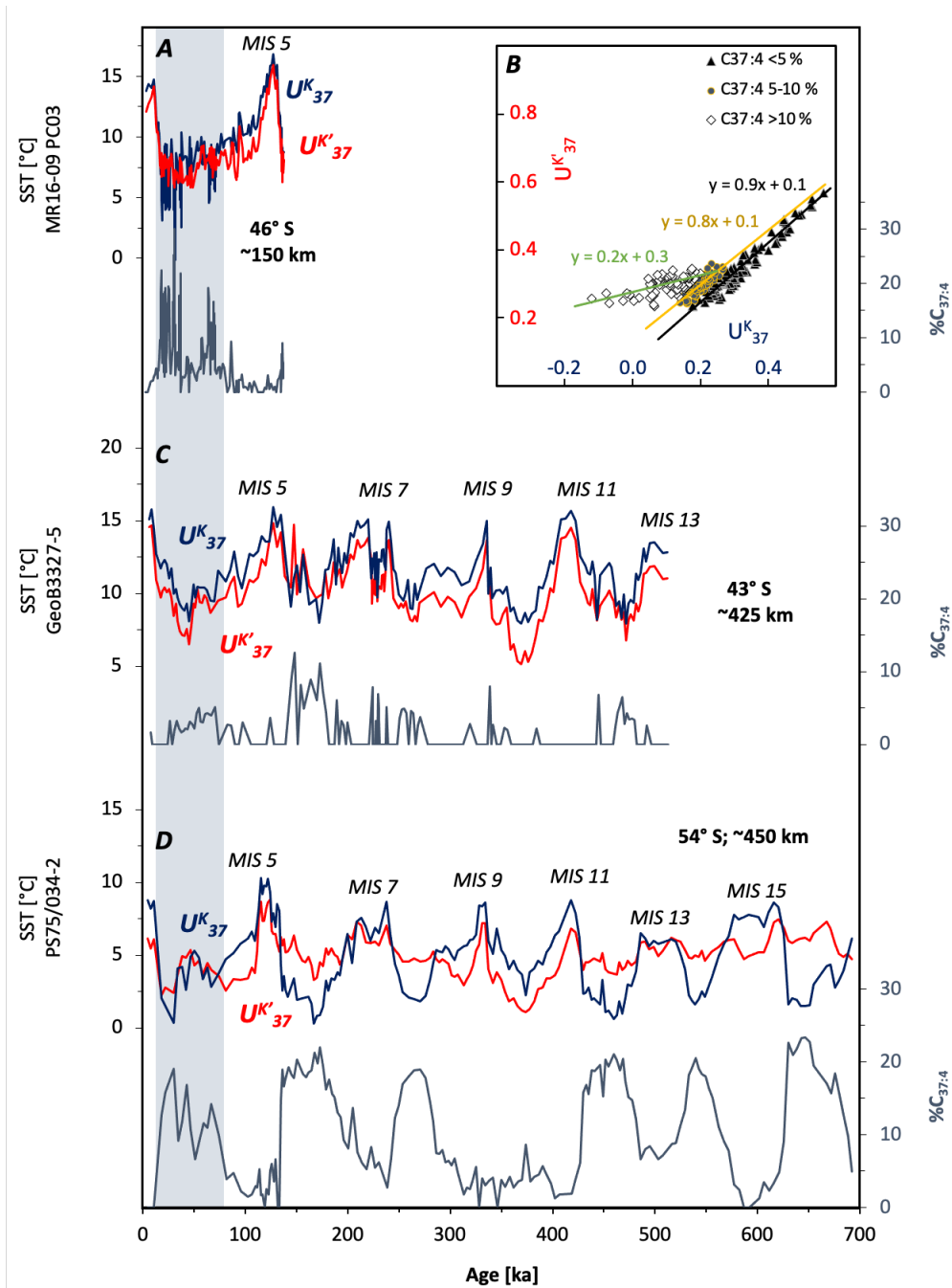
326



327

328 **Fig. S8.** SSTs ($U^{K'}_{37}$) of additional cores along the Chilean continental margin from north
 329 to south. Gray shadings and numbers at the top mark Terrigenous Input Phases (TIP). (A)
 330 ODP site 1234 (60). (B) ODP Site 1233 (48, 49). (C) MD07-3088 (56, 57). (D) MR16-09
 331 PC03 (this study). (E) MD07-3128 (50).

332



333

334 **Fig. S9.** U_{37}^K (gray; 31), U'_{37}^K (red; 32) based SSTs and $\%C_{37:4}$ of cores (A) MR16-09
335 PC03 (this study), (C) GeoB3327-5 and (D) PS75/034-2 from Ho *et al.* (43; $\%C_{37:4}$

336 published in this study). Gray marked area shows the last glacial with high %C_{37:4} values
337 of this study. (B) Relationship of U^K₃₇ – index, and U^{K'}₃₇ – index of core MR16-09 PC03.
338

339 **Table S1.** Age pointers of the age model.

Depth [m]	Age [ka]		Depth [m]	Age [ka]	
0.04	1.69	¹⁴ C dating	8.14	40.68	Laschamps
0.21	10.9	δ ¹⁸ O tuning	8.23	42.47	Laschamps
0.3	12.28	δ ¹⁸ O tuning	8.26	43.06	δ ¹⁸ O tuning
0.48	15.08	¹⁴ C dating	8.88	47.97	δ ¹⁸ O tuning
0.61	16	¹⁴ C dating	9.78	57.79	δ ¹⁸ O tuning
0.7	16.9	¹⁴ C dating	10.43	60.05	δ ¹⁸ O tuning
0.88	18	¹⁴ C dating	10.87	64.55	δ ¹⁸ O tuning
2.11	19.6	¹⁴ C dating	13.84	71.59	δ ¹⁸ O tuning
2.12	19.65	¹⁴ C dating	14.22	77.5	δ ¹⁸ O tuning
2.21	20.05	¹⁴ C dating	14.43	81.79	δ ¹⁸ O tuning
2.26	20.24	¹⁴ C dating	14.53	85.34	δ ¹⁸ O tuning
4.16	23.37	¹⁴ C dating	14.93	91.31	δ ¹⁸ O tuning
4.25	24	¹⁴ C dating	15.29	94.41	δ ¹⁸ O tuning
5.35	27.80	¹⁴ C dating	15.61	105.71	δ ¹⁸ O tuning
5.40	28.43	¹⁴ C dating	15.73	107.6	δ ¹⁸ O tuning
5.48	29.52	¹⁴ C dating	16.16	120.43	δ ¹⁸ O tuning
5.53	29.96	¹⁴ C dating	16.58	135.09	δ ¹⁸ O tuning
5.57	30.47	¹⁴ C dating	17.36	137.48	δ ¹⁸ O tuning
6.71	32.49	¹⁴ C dating			
6.80	33.17	¹⁴ C dating			
7.08	36.2	¹⁴ C dating			
8.07	38.12	¹⁴ C dating			

340

341

342 **Table S2.** All ^{14}C data based on *G. bulloides* for the age model of core MR16-09 PC03.
343 The calibration is based on MARINE20 (11) with a reservoir age of 400 years (12) in the
344 Program Calib. 8.2. Samples were excluded if they were located directly in a TIP event;
345 shaded samples were considered in the age model. Two (1875, 2209; red) samples could
346 not be included in the age model, even considering the maximum 2-sigma range and under
347 the assumptions described in the text above. A third sample (7373) was removed because
348 it did not fit the following $\delta^{18}\text{O}$ tuning. Samples 8315 and 8318 corresponded to the onset
349 and termination of the paleomagnetic marker Laschamps.

Internal AWI	Core	^{14}C raw	\pm Error	1-Sigma	1-Sigma	2-Sigma	2-Sigma	Used
Lab-Nr.	Depth [cm]	Age [ka]	Age [ka]	Age [ka]	Age [ka]	Age [ka]	Age [ka]	Age [ka]
2206.1.1	4.3	2.657	0.074	1.572	1.795	1.469	1.915	1.686
2207.1.1	48.4	13.632	0.123	14.886	15.27	14.631	15.488	15.08
1875.1.1	50.6	13.251	0.140	14.216	14.758	14.004	14.972	
7360.1.1	60.8	14.104	0.143	15.486	15.923	15.265	16.122	16
7361.1.1	70	14.819	0.152	16.403	16.849	16.184	17.036	16.9
7363.1.1	88.4	15.674	0.135	17.452	17.876	17.277	18.079	18
7364.1.1	97.2	13.051	0.126	13.882	14.344	13.739	14.664	
1876.1.2	119.2	17.142	0.132	19.102	19.485	18.899	19.681	
8333.1.1	197.9	16.518	0.165	18.379	18.798	18.197	18.987	
8320.1.1	202.6	17.144	0.286	18.927	19.636	18.665	20.04	
1877.1.1	210.8	17.821	0.167	19.907	20.362	19.634	20.564	19.6
7365.1.1	212	17.594	0.156	19.616	20.072	19.427	20.298	19.65
8313.1.1	221.1	17.462	0.180	19.448	19.943	19.196	20.195	20.05
8332.1.1	225.5	17.915	0.179	20.006	20.482	19.764	20.739	20.244
1878.1.1	238	16.303	0.153	18.216	18.586	18.025	18.755	
7366.1.1	247.7	20.099	0.184	22.552	22.995	22.345	23.24	
2208.1.1	253.4	20.490	0.191	22.975	23.472	22.819	23.726	
8334.1.1	261.4	19.019	0.385	21.115	22.071	20.613	22.456	
8331.1.1	416.1	20.613	0.411	22.911	23.804	22.419	24.278	23.369
8327.1.1	425	20.127	0.674	22.152	23.653	21.228	24.385	24
2209.1.1	426.1	19.987	0.185	22.466	22.891	22.248	23.094	
7367.1.1	534.6	25.067	0.265	27.694	28.29	27.443	28.604	27.795

7368.1.1	539.6	25.093	0.270	27.713	28.318	27.461	28.635	28.426
7369.1.1	547.9	26.559	0.305	29.213	29.843	28.887	30.102	29.524
7370.1.1	552.6	26.989	0.315	29.62	30.301	29.238	30.646	29.957
7371.1.1	557.1	27.530	0.320	30.167	30.785	29.912	31.033	30.472
8314.1.1	670.8	29.465	0.453	31.865	33.068	31.382	33.644	32.49
8306.1.1	679.9	29.227	0.376	31.689	32.729	31.254	33.168	33.168
8317.1.1	707.6	33.409	1.625	35.204	38.759	33.526	40.329	36.866
8325.1.1	793.5	28.128	1.299	29.81	32.644	28.552	34.027	
8307.1.1	798	32.528	0.514	35.207	36.266	34.634	36.843	
8308.1.1	806.9	34.555	0.746	37.293	39.085	36.34	39.679	38.234
8315.1.1	816.3	36.914	0.924	39.727	41.214	38.934	41.896	41.21
8318.1.1	825.6	40.054	1.041	41.835	43.089	41.246	44.017	42.489
7373.1.1	830.2	37.009	1.869	38.904	41.99	36.429	42.747	

350

351

352 **Table S3.** Terrigenous Input Phases (TIP). The age was read from the terrigenous entries,
 353 at the beginning and end of each TIP. DO: Dansgaard-Oeschger

TIP	Depth [m]		Age [ka]		Duration [ka]	Tentative to DO
6	17.38	16.79	137.5	135.7	1.8	-
4a	13.91	12.55	72.7	68.5	4.2	DO18?
4b	12.55	1.73	68.5	66.6	1.9	-
4c	11.73	10.73	66.6	63.1	3.5	DO17?
4d	10.45	9.73	60.3	57.2	3.1	DO16
3a	8	7.1	38.0	36.2	1.8	DO8
3b	6.64	5.55	32.4	30.2	2.2	DO5?
2a	5.19	4.27	27.3	24.1	3.2	-
2b	4.09	2.37	23.3	20.4	2.8	DO2?
2c	2.1	1	19.6	18.2	1.4	-

354

355

356 **Table S4.** Selection of previously published land-based advances mainly from Patagonian
357 glaciers of moraines and out-washed sediments and used for Figures 2 and Figure 4. ¹Ra-
358 diocarbon dating, recalibrated with SHCal13 by Davies *et al.* (61; file “Radiocarbon_ages”,
359 tab “Terrestrial”, column “W”). The difference to the originally used IntCal13 calibration
360 of Moreno *et al.* (62) are negligible. We used an arithmetic average calculation and the
361 standard error (1σ) of several dating points for each advancing event. ²Cosmogenic dating,
362 recalibrated with Lm scaling and an erosion rate of 0 mm/ka by Davies *et al.* (61; file “¹⁰Ba
363 and ²⁶Al ages August 2019”, tab “0 erosion Kaplan et al.” column “Y”). We used an arith-
364 metic average calculation and the standard error (1σ) of several dating’s for each advancing
365 event. For single advancing events, we used the external standard deviation (given in col-
366 umn “AA”). ³Data, based on Lm scaling, originally taken from the reference.

Reference	Advances (ka)					
North Patagonia: 38 – 46° S, west side of the Andes						
¹ Moreno <i>et al.</i> (62)	17.8 ± 0.6	25.8 ± 1.4	26.7 ± 0.3	30.9 ± 0.8	33.3 ± 0.5	
³ García <i>et al.</i> (63)	26.0 ± 2.9	57.8 ± 4.7				
Central Patagonia: 46 – 50° S, east side of the Andes						
² Hein <i>et al.</i> (64)	17.4 ± 1.7	19.2 ± 1.8	20.9 ± 1.2	24.9 ± 0.5	28.4 ± 1.0	
³ Mendelova <i>et al.</i> (65)	24.8 ± 1.0	75.2 ± 2.8				
² Glasser <i>et al.</i> (66)	15.3 ± 1.7	23.5 ± 2.2	35.2 ± 3.2	(60.0 ± 5.9)	93.6 ± 3.8	
South Patagonia: 50 – 53° S, east side of the Andes						
² García <i>et al.</i> (67), (68)	14.1 ± 0.5	21.7 ± 2.0	35.5 ± 4.1	40.0 ± 4.1	48.3 ± 1.8	
² Kaplan <i>et al.</i> (69)	20.1 ± 0.9	21.4 ± 2.1	23.3 ± 1.3	28.7 ± 1.2		
³ Peltier <i>et al.</i> (70)	18.1 ± 0.6	19.1 ± 0.7	23.9 ± 0.8	25.7 ± 0.8	27.4 ± 0.8	65.4 ± 2.0 67.5 ± 2.1
New Zealand						
³ Strand <i>et al.</i> (71)	18.0 ± 0.4	20.0 ± 0.5	26.7 ± 0.7	36.5 ± 0.9	41.8 ± 1.1	44.0 ± 1.0
³ Schaefer <i>et al.</i> (72)	65.1 ± 2.7					

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